





Axion Production and Detection using NMR

DESY Theory Workshop
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collaborate with Jeff Dror and Qiushi Wei at the University of Florida [in preparation]

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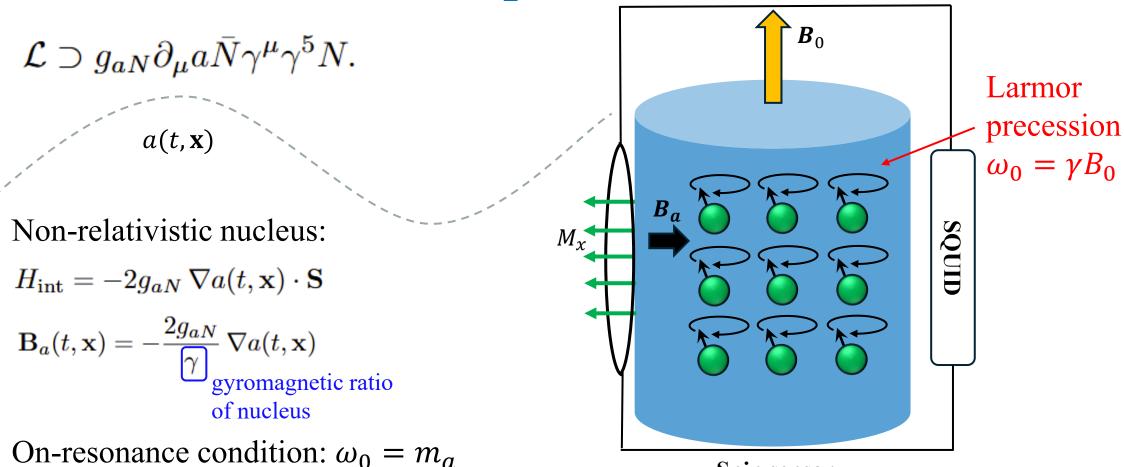
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Introduction

- Axion is a well-motivated hypothetical pseudoscalar in the extension of the Standard Model.
- Axion dark matter searches through axion-nucleon coupling g_{aN} using NMR techniques, e.g., CASPEr. [Budker et al., 2014]
- Non-relic axion search: production and detection in the lab.
 - Light-shinning-through-wall experiments [Bibber et al., 1987;...] Resonant cavity
 - ARIADNE [Arvanitaki & Geraci, 2014; Geraci et al., 2018]

mechanical oscillators • Ferroaxionic force [Arvanitaki et al., 2024] or rotators + NMR

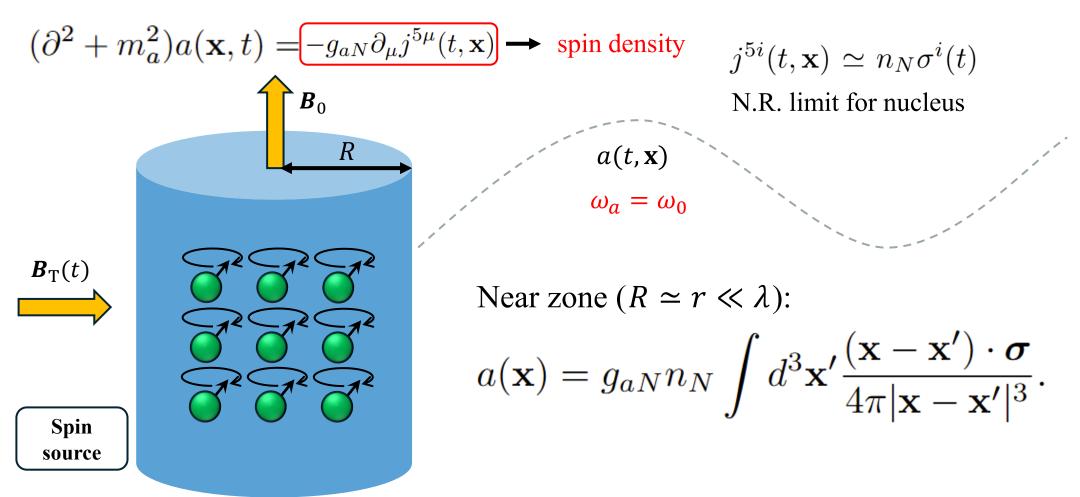
Axion Precession Experiment



Spin sensor

Ambient axion DM field induced AC magnetic field drives the spin precession and produces detectable (transverse) magnetization.

Axion Production



AC magnetic field drives the spin precession and produces axion fields.

Maximum axion mass

• Full-form axion field profile:

$$a(\mathbf{x}) = g_{aN} n_N \int d^3 x' \frac{(\mathbf{x} - \mathbf{x}') \cdot \boldsymbol{\sigma}}{4\pi |\mathbf{x} - \mathbf{x}'|^3} \left(1 - ik_0 |\mathbf{x} - \mathbf{x}'|\right) e^{ik_0 |\mathbf{x} - \mathbf{x}'|}$$

$$k_0 \equiv \sqrt{\omega_0^2 - m_a^2}$$
 (if $\omega_0 \ge m_a$) or $i\kappa_0$ (if $\omega_0 < m_a$)

When the axion mass is too large,

$$e^{-\kappa_0|\mathbf{x}-\mathbf{x}'|} \rightarrow e^{-m_a r} \ll 1$$
 when $m_a R \gg 1$

Near-zone profile

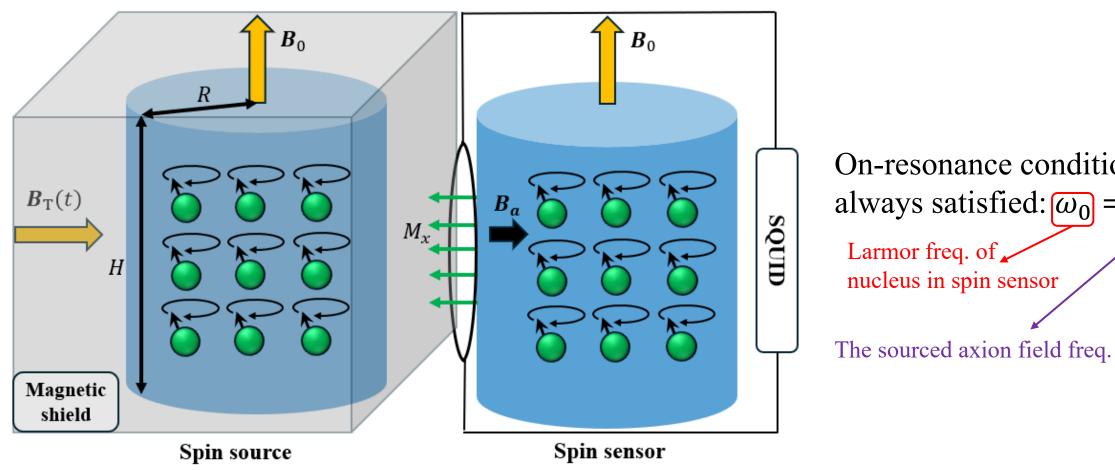
• Axion field generated by spin source:

$$|\partial_x a/a| \sim 1/R$$

• Axion DM field:

$$\frac{\nabla a}{a} \sim m_a v_0 \sim 10^{-3} m_a \ll R^{-1}$$

Experimental setup



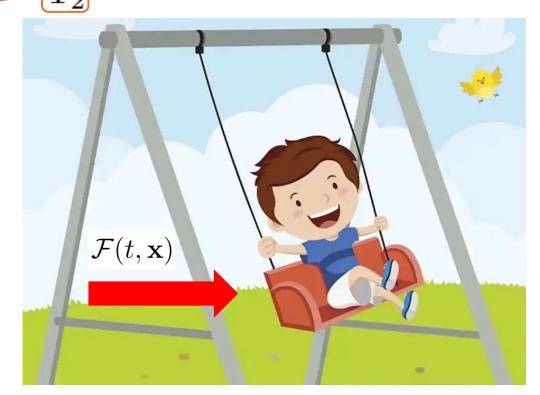
On-resonance condition is always satisfied: $\omega_0 = \omega_0$ nucleus in spin sensor

Growth of transverse magnetization

• Driven harmonics of the transverse magnetization M_{χ} :

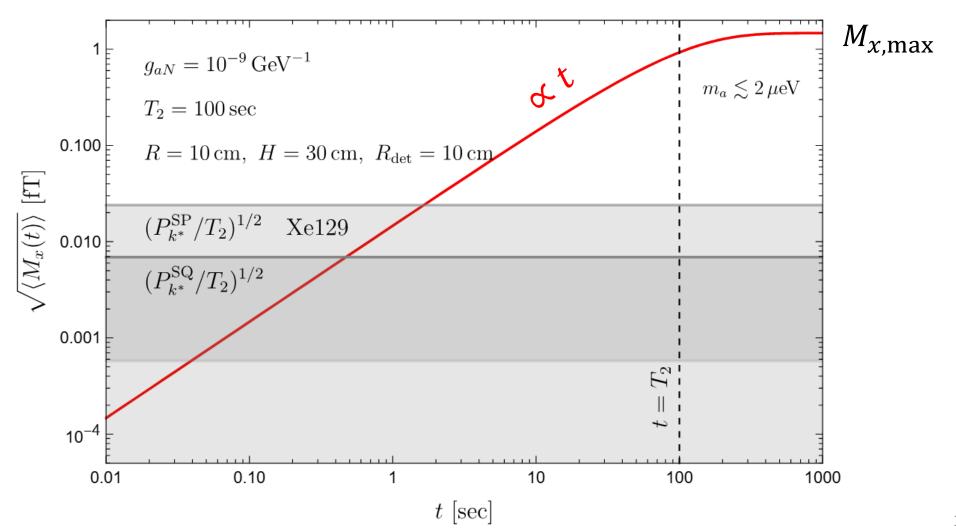
$$\ddot{M}_x + \frac{2}{T_2}\dot{M}_x + \omega_0^2 M_x \simeq \boxed{\gamma M_0(\omega_0 B_{ax} - \dot{B}_{ay})} \qquad \mathcal{F}(t, \mathbf{x})$$

Transverse relaxation time



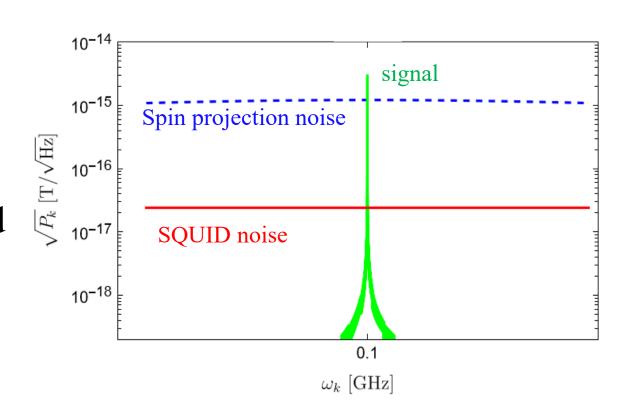
driving force

Transverse Magnetization

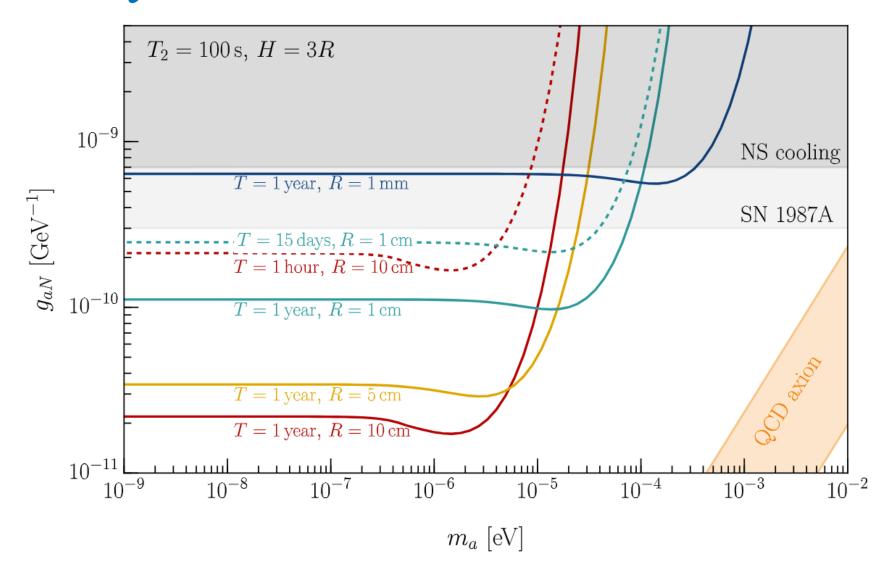


Noise source

- Thermal noise → cool down the apparatus
- SQUID noise → high-precision magnetometer
- Spin projection noise → limited and fundamental
- Transverse AC magnetic field → magnetic shielding



Sensitivity



Scaling of model parameters

• In the plateau region

$$g_{aN}^{(\mathrm{UL})} \simeq 2 \times 10^{-11} \,\mathrm{GeV}^{-1} \left(\frac{n_N}{1.3 \times 10^{22} \,\mathrm{cm}^{-3}}\right)^{-3/4}$$

$$\times \left(\frac{T_2}{100 \,\mathrm{s}}\right)^{-1/4} \left(\frac{T}{1 \,\mathrm{yr}}\right)^{-1/4} \left(\frac{V}{0.01 \,\mathrm{m}^3}\right)^{-1/4}$$

Advantage

- A broad-band search with large axion mass $m_a \sim R^{-1}$ beyond the Larmor frequency ω_0 of the NMR device.
- The sourced axion field is coherent, with a known frequency ω_0 .
 - Can always set the detector intrinsic frequency to satisfy on-resonance condition
 - Can measure for a long observation time
- The sensitivity of a continuous experiment for T_{obs} is equivalent to that of N multiple experiments with a time T_{obs}/N .
 - the experiment is more applicable and reliable since one can calibrate the apparatus daily or hourly.
 - conduct multiple experiments simultaneously with multiple sets of devices.

Conclusion

• 1. Axion production and detection through pure g_{aN} is studied.

• 2. The dual NMR setup can probe the non-relic axion up to $m_a \le R^{-1}$.

• 3. The sensitivity can surpass the astrophysical constraint within days' observation time and centimeter-scale NMR device.

Backup slides

Near-zone profile

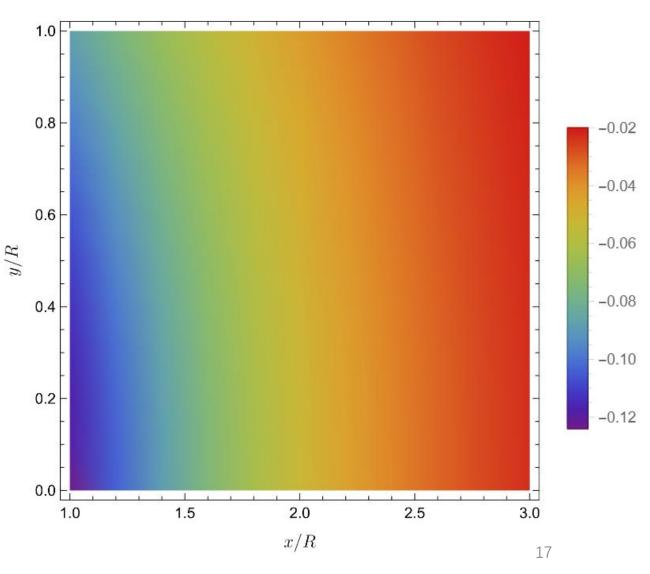
• The profile of ∇a in near zone:

$$a(\mathbf{x} = x\hat{x}) \simeq 2.4 \times 10^{-7} \text{eV} \left(\frac{x}{R}\right)^{-1.6}$$

$$\times \left(\frac{g_{aN}}{10^{-10} \text{GeV}^{-1}}\right) \left(\frac{n_N}{1.3 \times 10^{22} \text{cm}^{-3}}\right) \left(\frac{R}{1 \text{cm}}\right) \lesssim$$

$$\partial_x a(\mathbf{x} = x\hat{x}) \simeq -7.6 \times 10^{-12} \text{eV}^2 \left(\frac{x}{R}\right)^{-2.6}$$

$$\times \left(\frac{g_{aN}}{10^{-10} \text{GeV}^{-1}}\right) \left(\frac{n_N}{1.3 \times 10^{22} \text{cm}^{-3}}\right).$$



Axion Production

• *Far zone: $e^{ik_0|\mathbf{x}-\mathbf{x}'|}$ oscillating

