

Graphing the Dark Sector: Anomalous Shower Detection with GNNs in the ATLAS Calorimeters

Scientific Computing Seminar
20.06.2025, DESY

Lukas Bauckhage^{1,2}

¹Deutsches Elektronen-Synchrotron DESY

²Physikalisches Institut Universität Bonn

HELMHOLTZ

UNIVERSITÄT **BONN**



Previous Talk


- Talk at last year's SciComp Workshop

FH SciComp Workshop 2024

 Jul 1, 2024, 2:00 PM → Jul 2, 2024, 2:00 PM Europe/Berlin

 Seminar Room 4a/b (DESY Campus Hamburg)

8. CNNs and GNNs for tagging anomalous showers with ATLAS

 Lukas Bauckhage (ATLAS (ATLAS Upgrade))

 7/2/24, 11:00 AM

Scientific Computing III

This talk will discuss recent work towards developing a ML (CNN/GNN) tagger to distinguish anomalous showers caused by the decays of long-lived particles from QCD jets with the ATLAS detector

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⇒ Loads of updates!

Previous Talk

(in this Seminar)

(indico)

CaloClouds: Ultra-Fast Geometry-Independent Highly-Granular Calorimeter Simulation

Scientific Computing Seminar
DESY, 23.05.2025

E. Buhmann², H. Day-Hall¹, T. Buss², F. Gaede¹, G. Kasieczka²,
W. Korcar², **A. Korol^{1,*}**, K. Krüger¹, P. McKeown^{1,3}

¹ Deutsches Elektronen-Synchrotron, DESY

² University of Hamburg, UHH

³ CERN

*anatolii.korol@desy.de

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Introduction

HEP Data Pipeline

Standard Tools and Methods

Simulation

Matrix-element
Calculation

Parton-shower /
Hadronization

Numerical
Integration

Detector
Simulation

Markov Chain
Monte Carlo

Digitization

Topological
clustering

Reconstruction

Topoclusters
& Spacepoints

Track Finding
& Fitting

Kalman Filtering
& Fitting

Jet Tagging
& Vertexing

Particle ID
& Particle Flow

Conformal Fits
& Hough
Transform

Analysis

Calibration

Likelihood
Fitting

Unfolding

Statistical
Techniques,
Bayesian
Inference

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ML?

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Generative Models:
GANs, VAEs,
Normalizing Flows
and Diffusion

Metric Learning,
Object
Condensation

Deep Full Event
Reconstruction

CNNs, Graph
Neural Networks &
Transformers

Symmetric ML
& Equivariance

Autoencoders
& Anomaly Detection

Omnifold and
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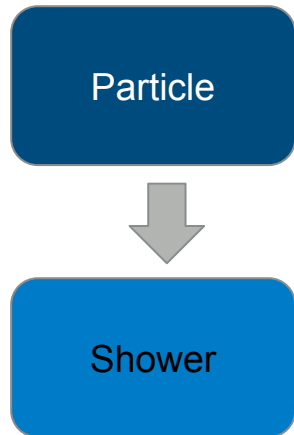
*anatolii.korol@desy.de

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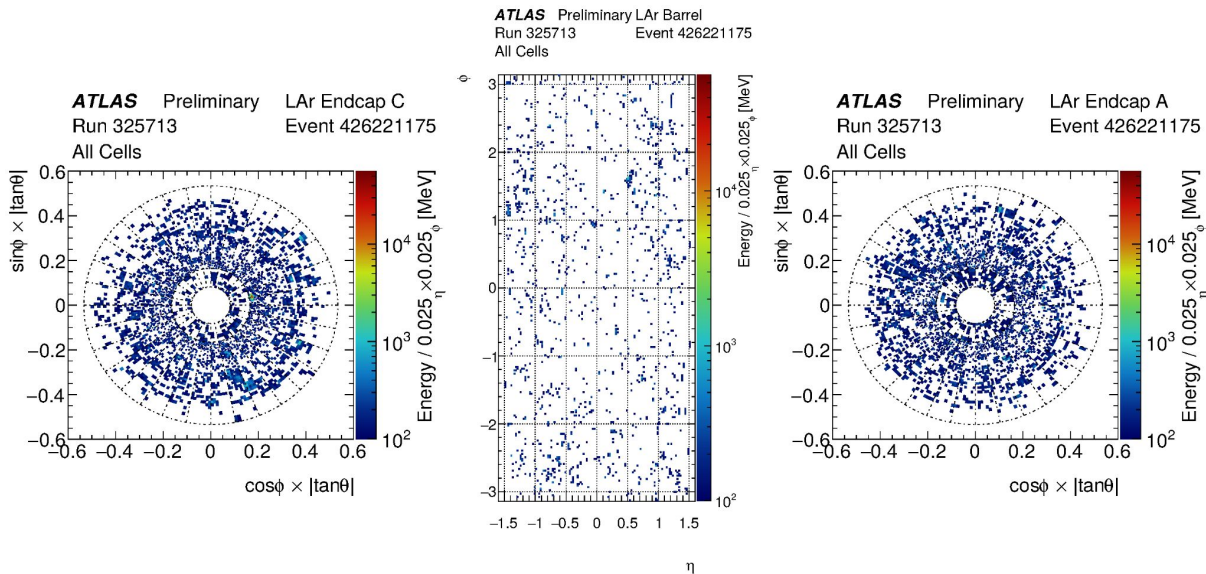


TODAY!

Shower Reconstruction

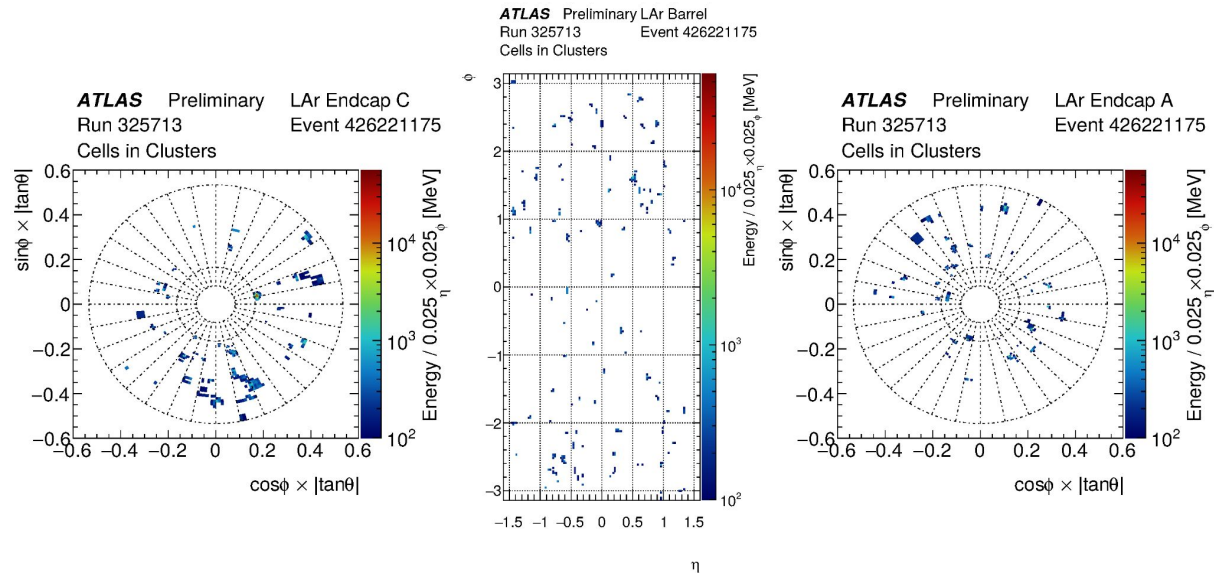
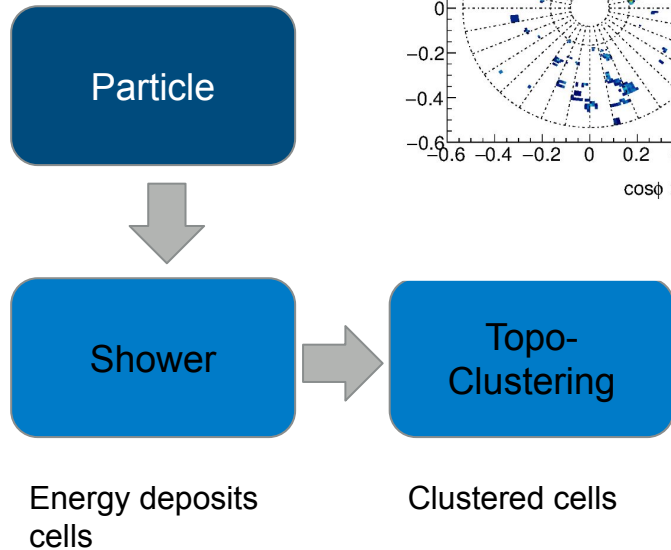


Energy deposits
cells



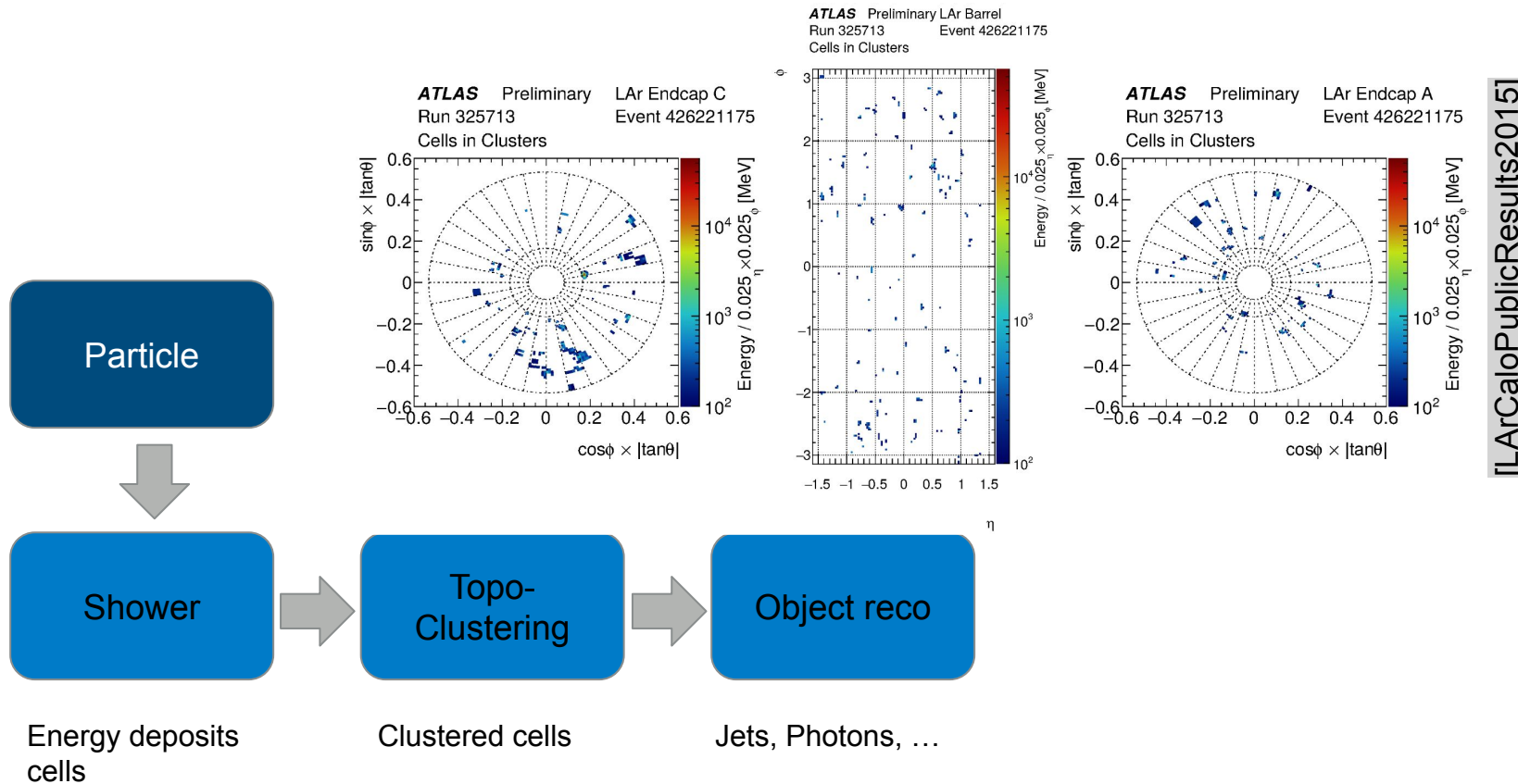
[LArCaloPublicResults2015]

Shower Reconstruction

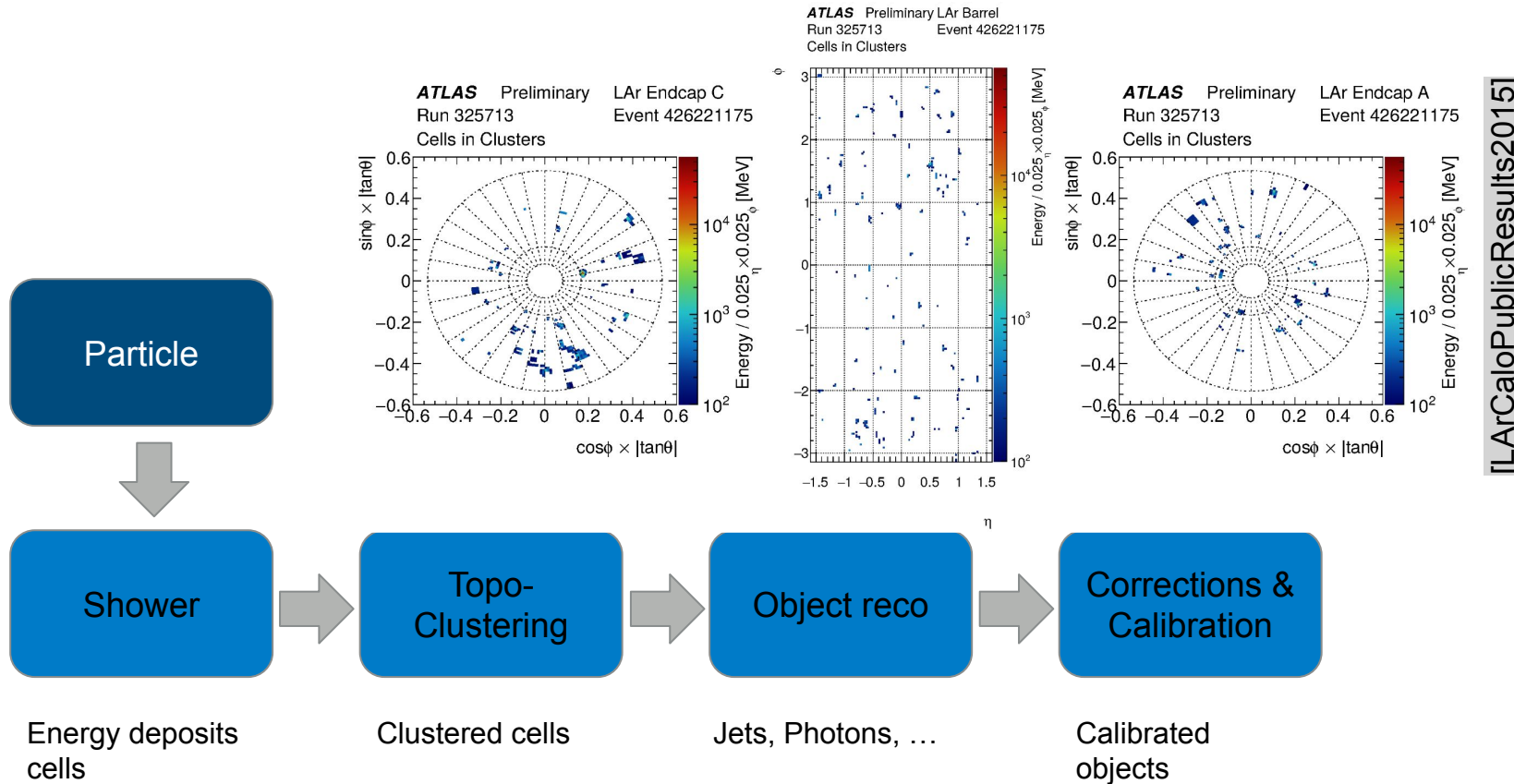


[LArCaloPublicResults2015]

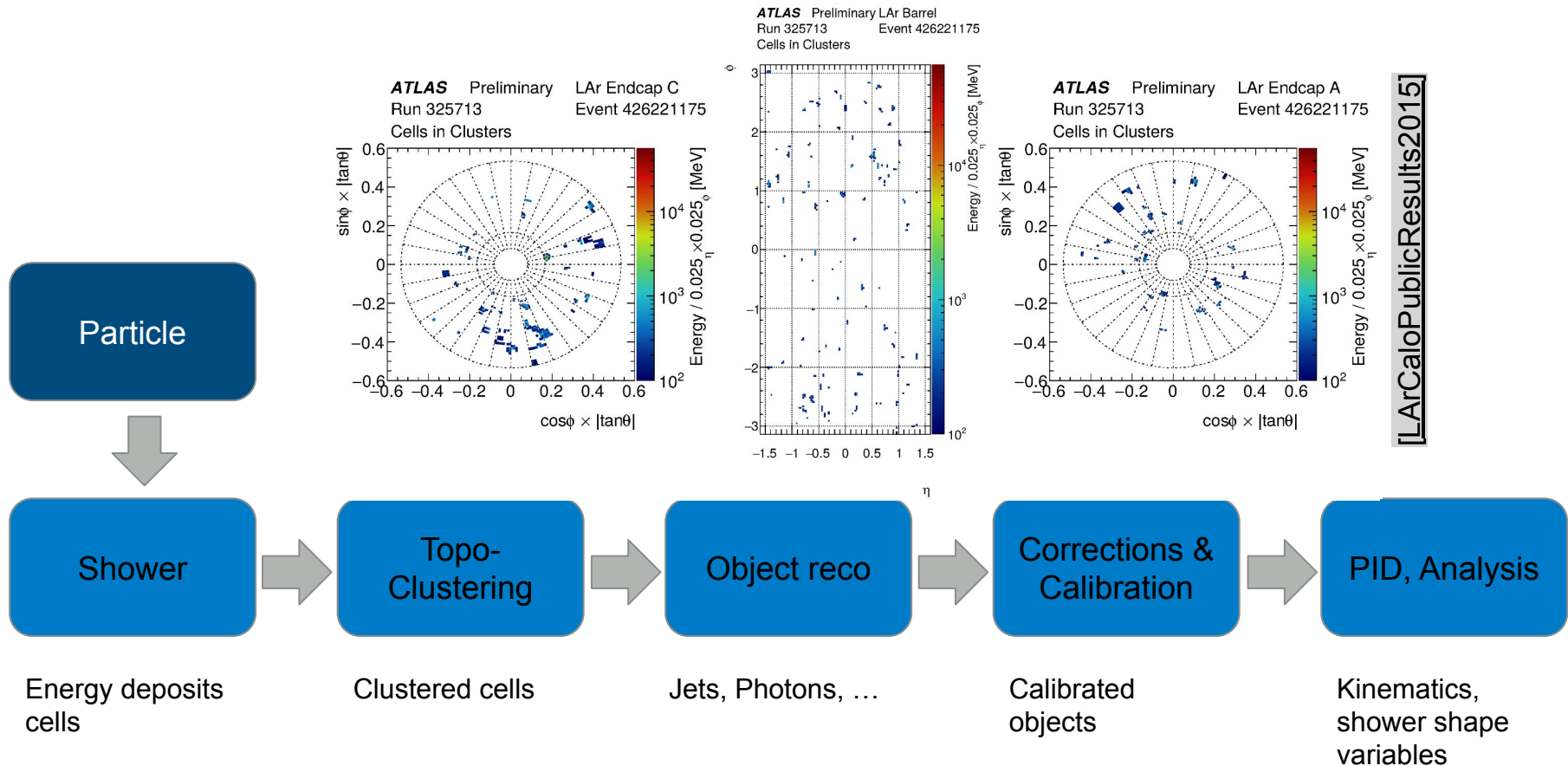
Shower Reconstruction



Shower Reconstruction



Shower Reconstruction



Shower Reconstruction

BSM Physics,
LLPs, ...

Particle

Shower

Energy deposits
cells

Topo-
Clustering

Clustered cells

~~Object reco~~

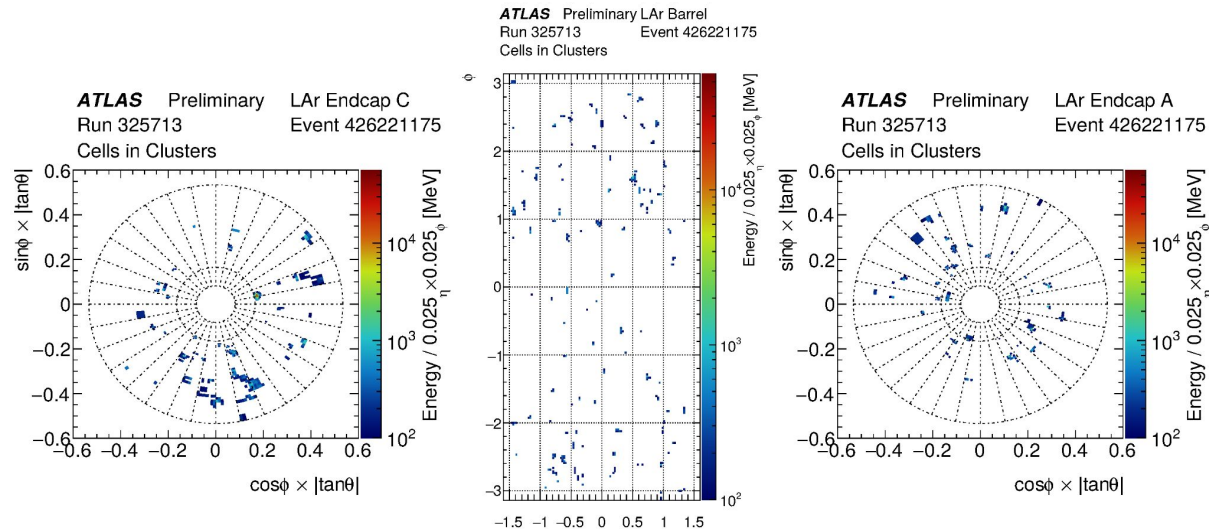
~~Jets, Photons, ...~~

Corrections &
Calibration

Calibrated
objects

PID, Analysis

Kinematics,
shower shape
variables



[LArCaloPublicResults2015]

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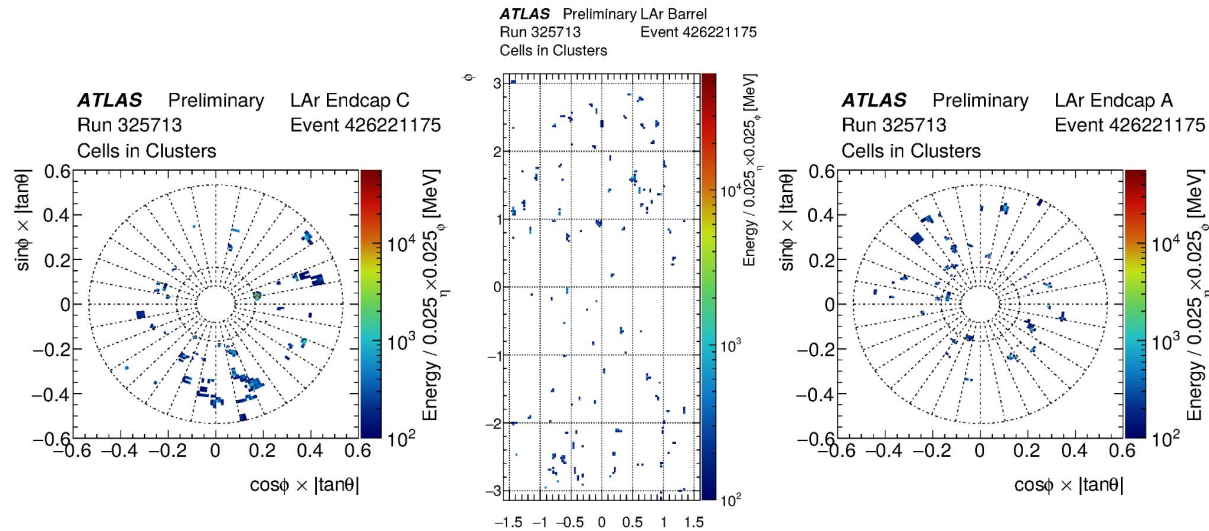
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Can we use something more “low-level”?



[LArCaloPublicResults2015]

Shower Reconstruction

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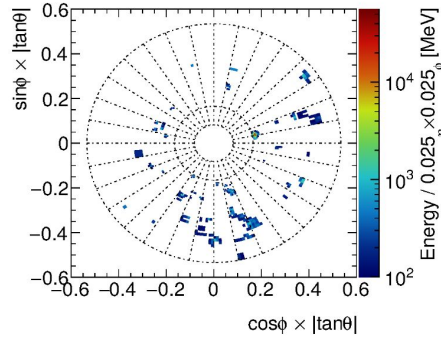
Particle

Shower

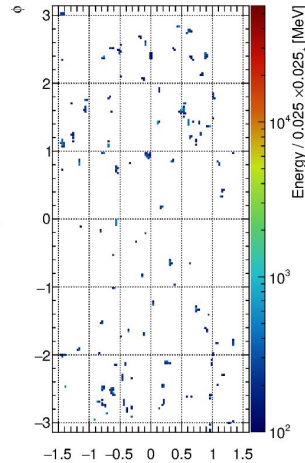
Energy deposits
cells

too expensive

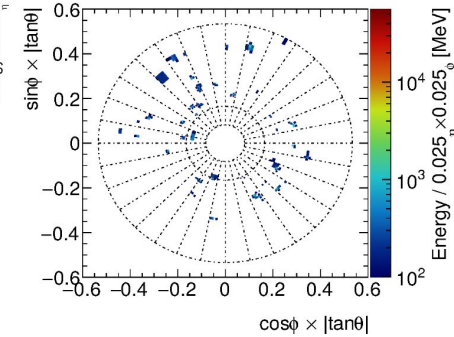
ATLAS Preliminary LAr Endcap C
Run 325713 Event 426221175
Cells in Clusters



ATLAS Preliminary LAr Barrel
Run 325713 Event 426221175
Cells in Clusters



ATLAS Preliminary LAr Endcap A
Run 325713 Event 426221175
Cells in Clusters



[LArCaloPublicResults2015]

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Clustering

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Shower Reconstruction

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Shower

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too expensive

Topo-
Clustering

Clustered cells

(ML) Tagger based on
clusters?

~~Object reco~~

Jets, Photons, ...

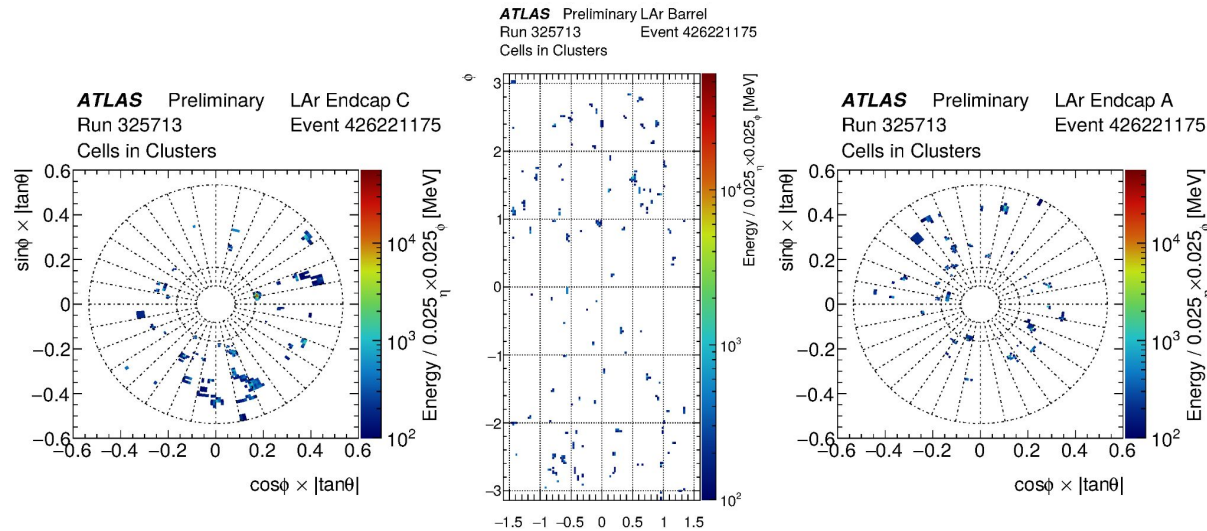
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[LArCaloPublicResults2015]

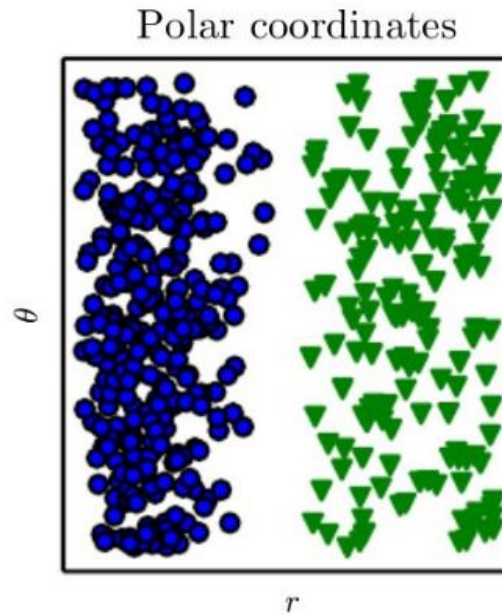
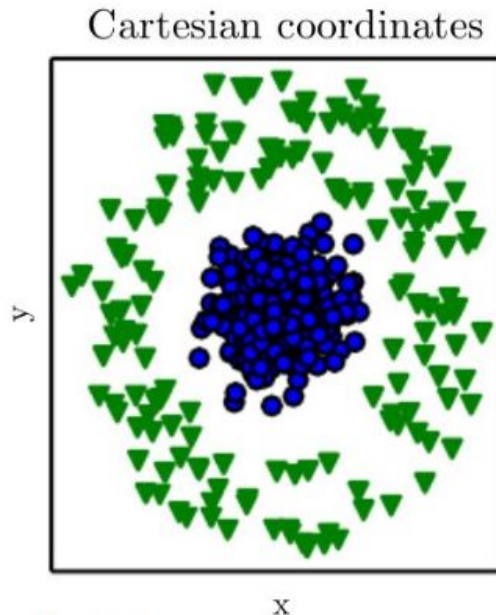
Clusters and Representation of Showers

How to represent calorimeter showers in data?

(indico)

The Right Data Representation Can Turn an Impossible Problem into an Easy One

impossible task for
linear model



easy to solve with
vertical line

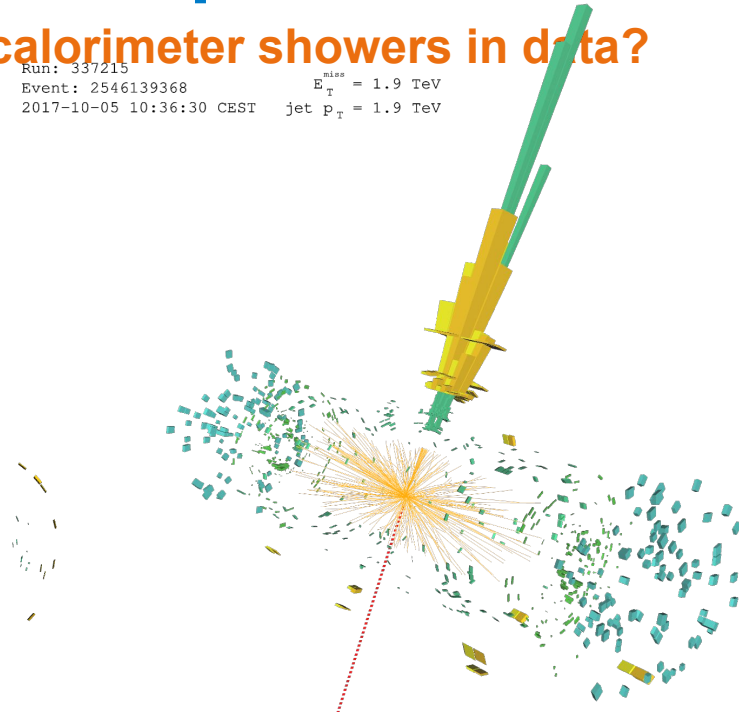
Figure from:

Deep Learning, Ian Goodfellow and Yoshua Bengio and Aaron Courville,
<https://www.deeplearningbook.org/>

Clusters and Representation of Showers

How to represent calorimeter showers in data?

Run: 337215
Event: 2546139368 $E_T^{\text{miss}} = 1.9 \text{ TeV}$
2017-10-05 10:36:30 CEST jet $p_T = 1.9 \text{ TeV}$

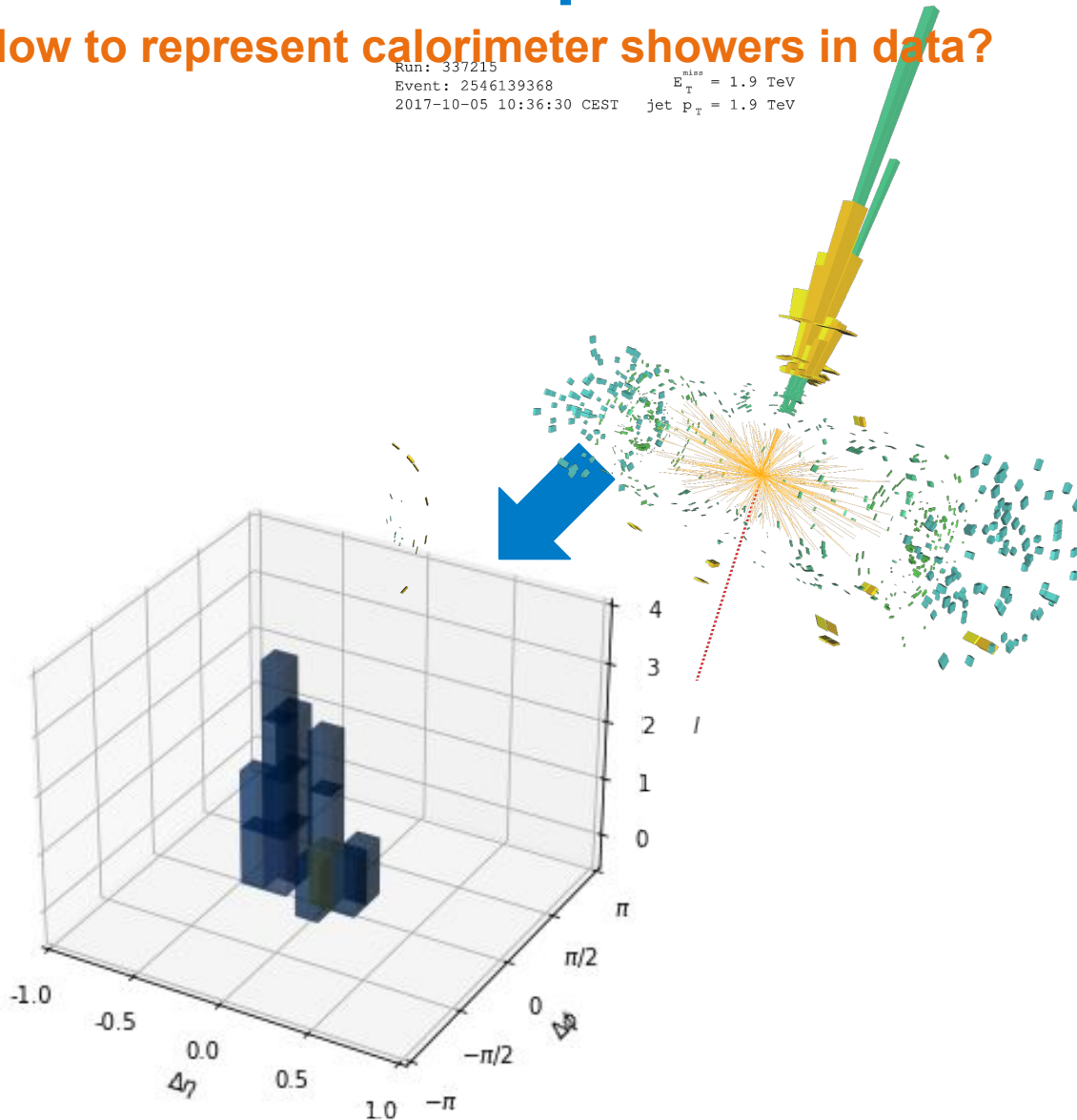


[EXOT-2018-06]

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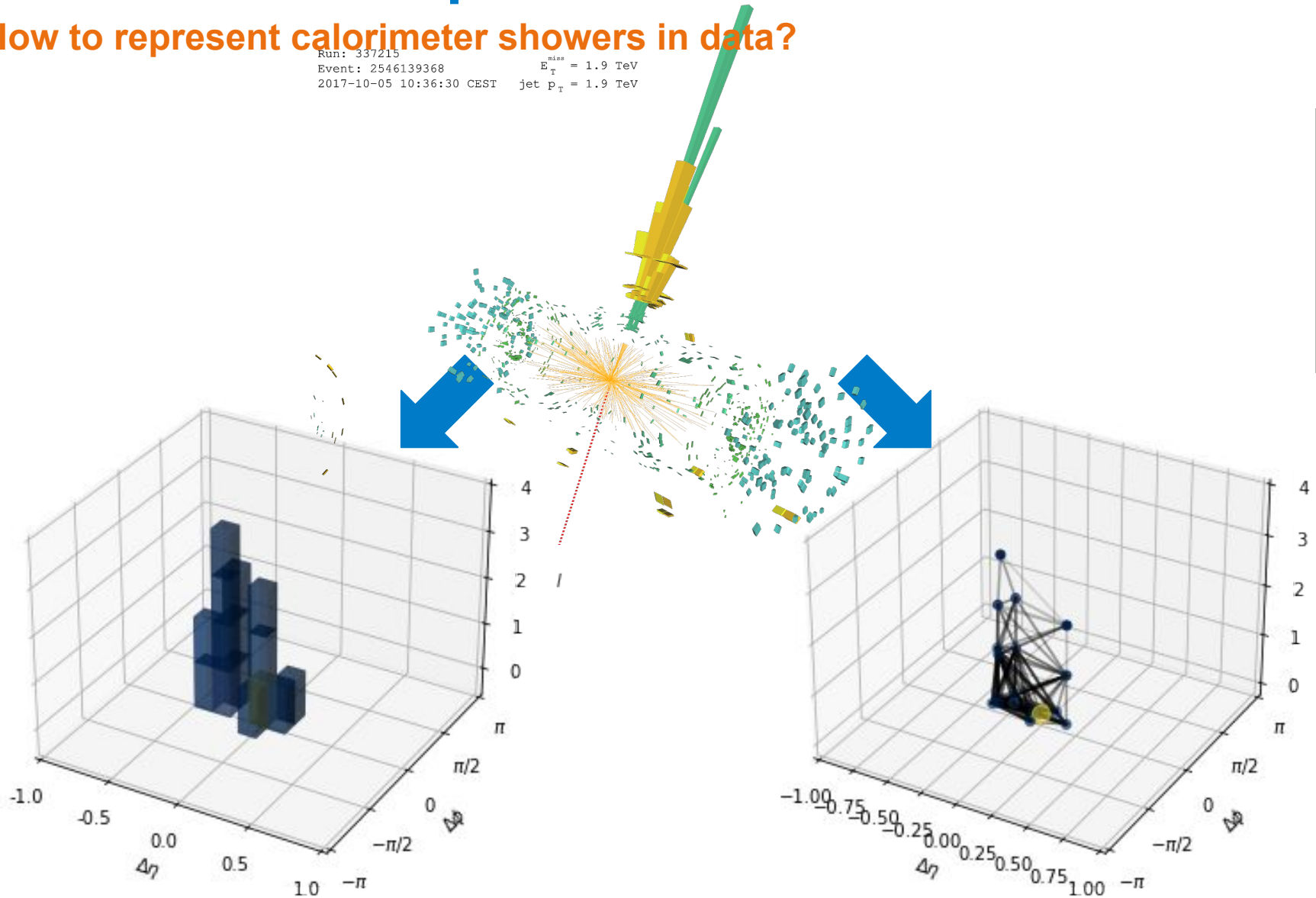


[EXOT-2018-06]

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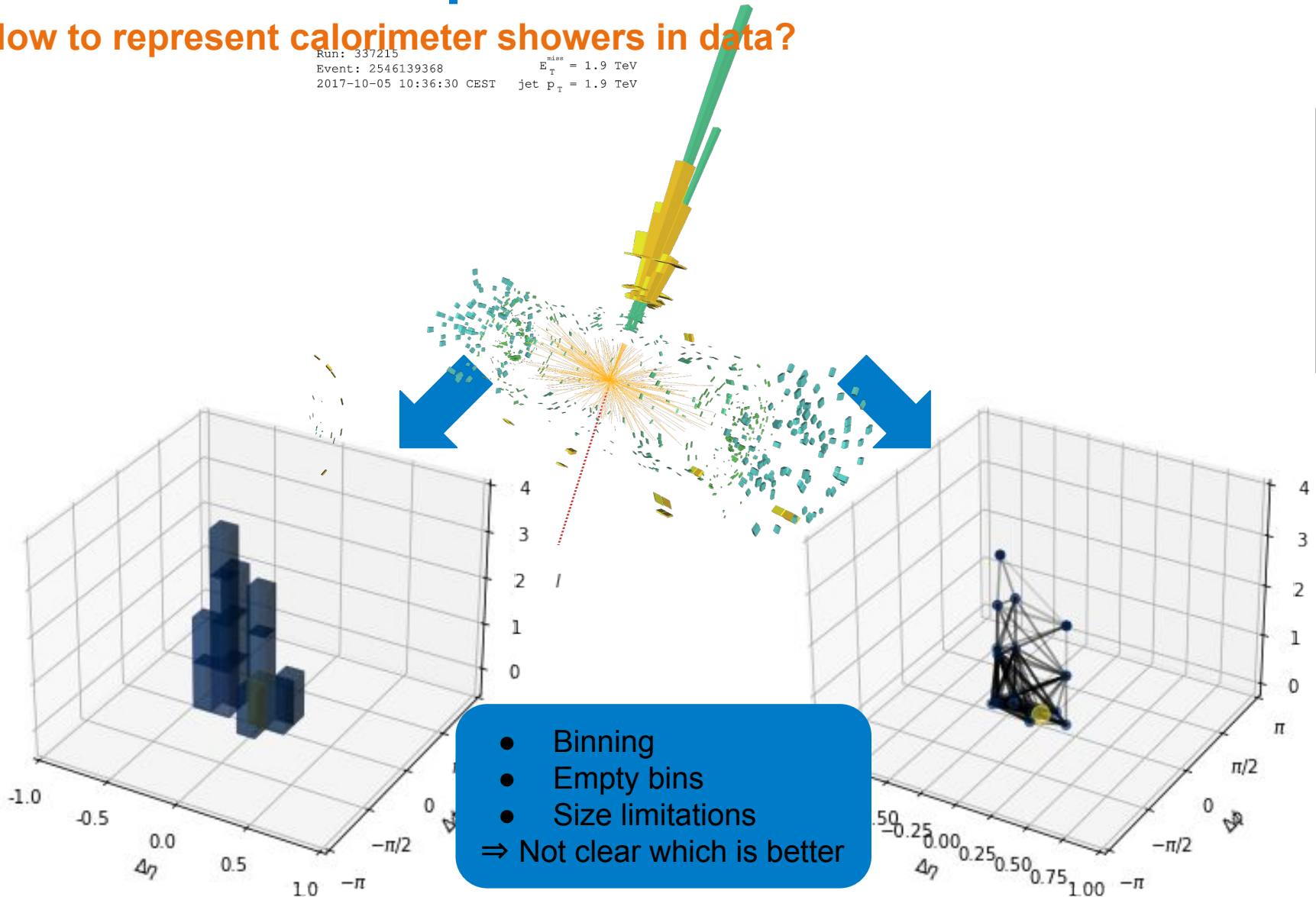


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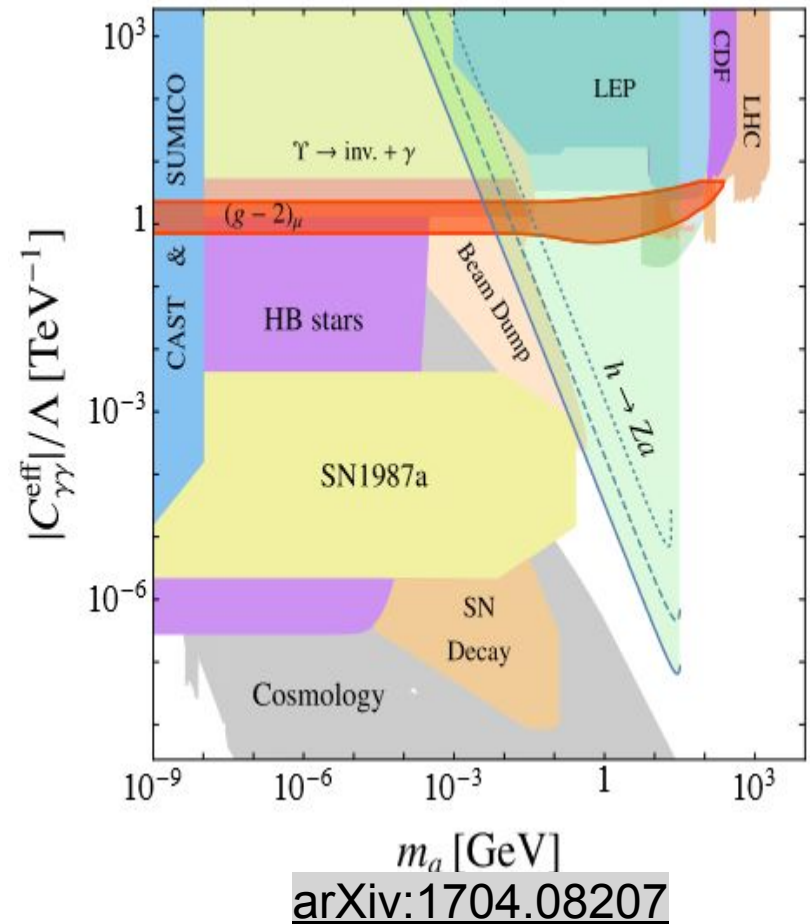


[EXOT-2018-06]

Selected Benchmark Study

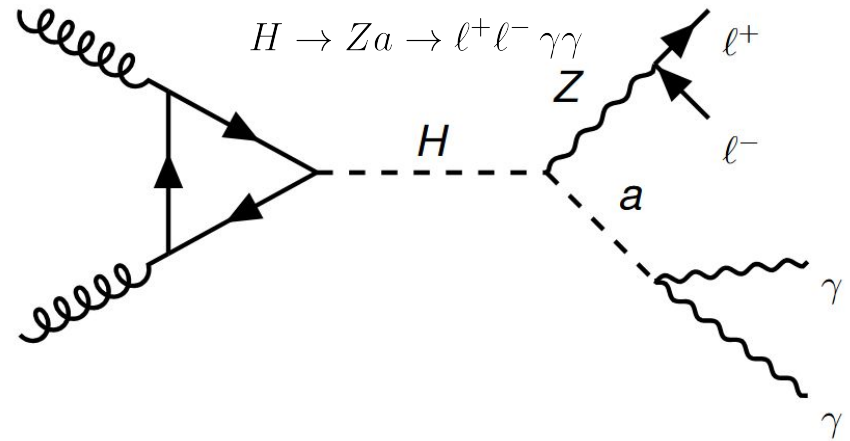
Axion-Like Particles (ALPs)

- $g_{a\gamma\gamma}, m_a$ free parameters
- Dark Matter candidate:
non-thermally in early universe
- Mediator in Dark Sector theories
- Produced in H decays
(Higgs portal theories)



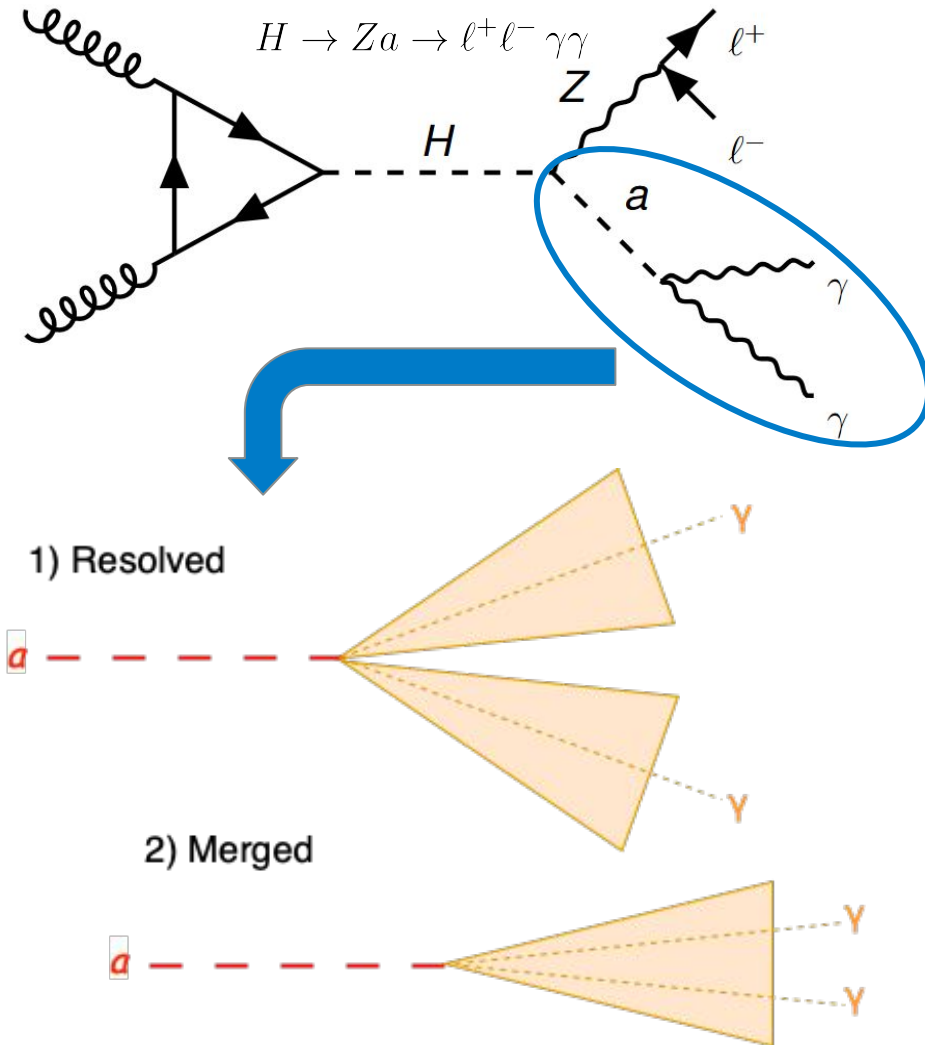
ALPs in Higgs Decays

- $Z \rightarrow \ell\ell$: simple and clean signature
- Main background:
 $Z + \text{jets}$ or $Z + \text{photons}$



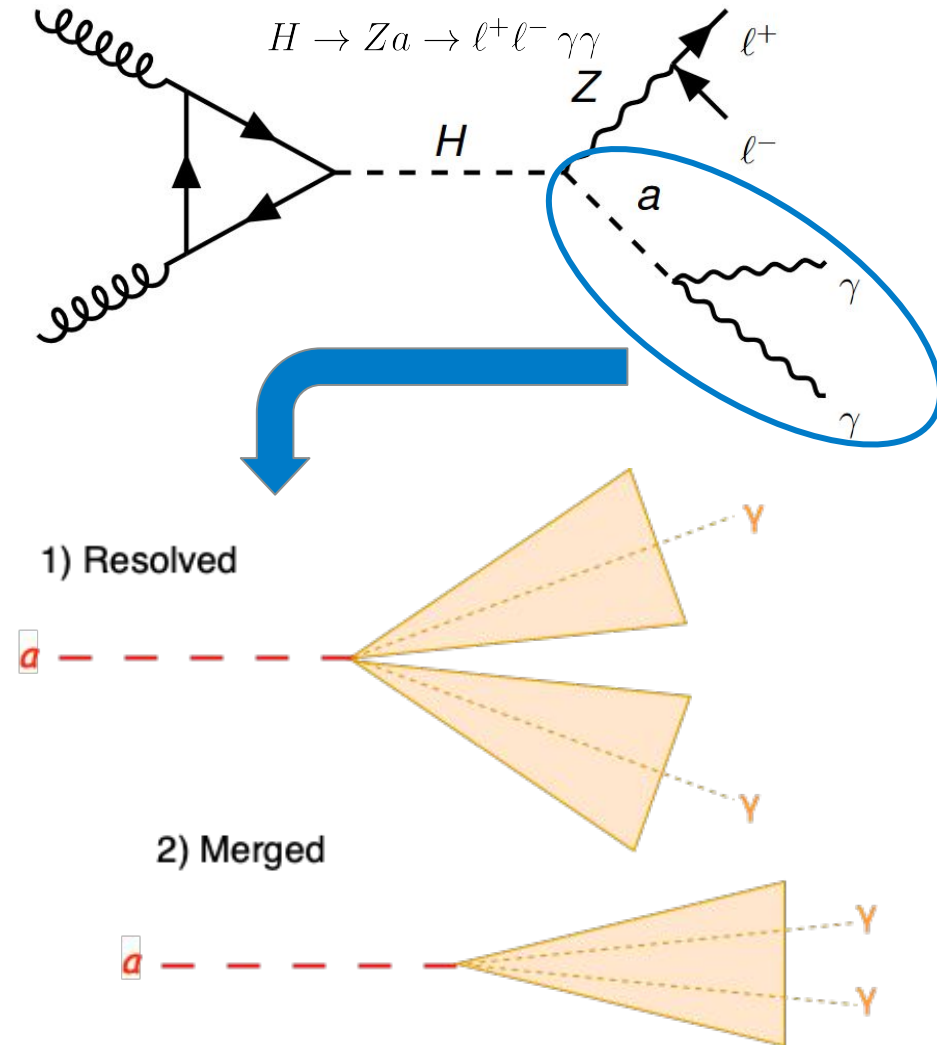
ALPs in Higgs Decays

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- 2 categories:
 - “Resolved”/Separated:
high ALP mass
 - “Merged”/Collimated:
low ALP mass



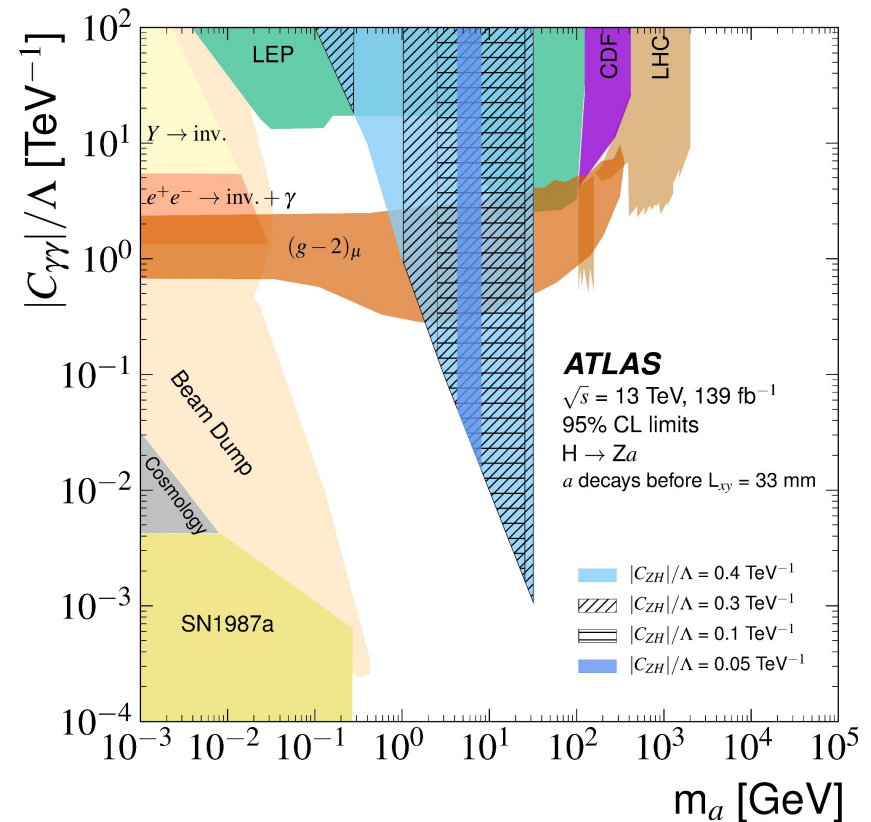
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- $H \rightarrow 2a \rightarrow 4\gamma$ at ATLAS:
[arXiv:2312.03306](https://arxiv.org/abs/2312.03306)



Prompt Decay Analysis

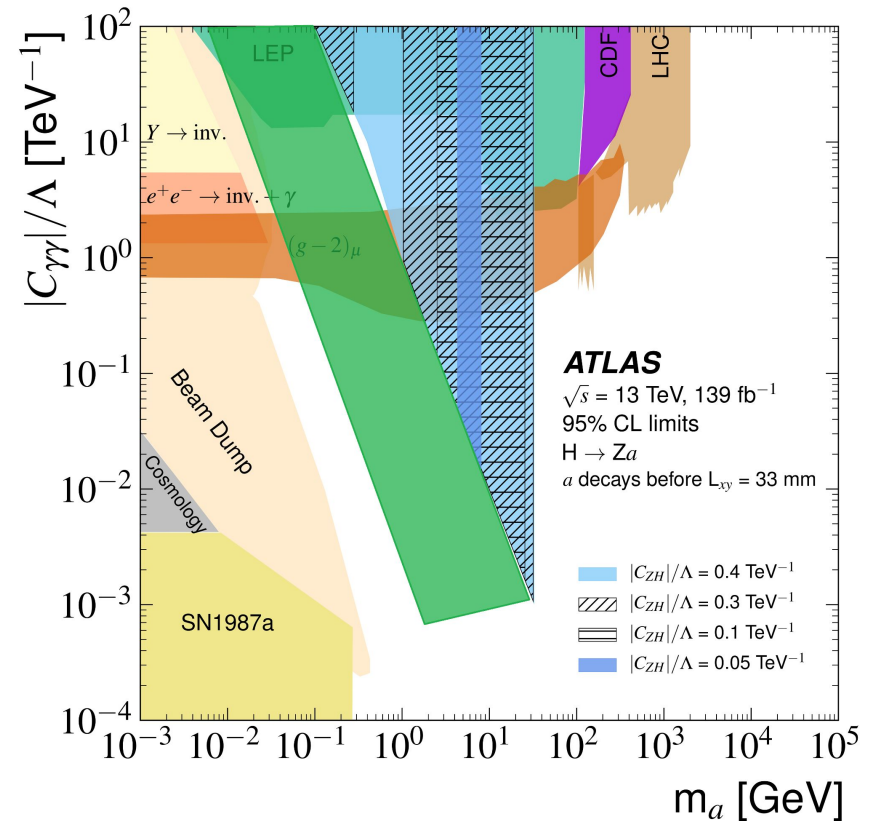
- Search for $H \rightarrow Za$ in Run 2 data
- Promptly decaying ALP
- Trigger on Z decay leptons
- $m_a \in [0.1, 31] \text{ GeV}$



arXiv:2312.01942

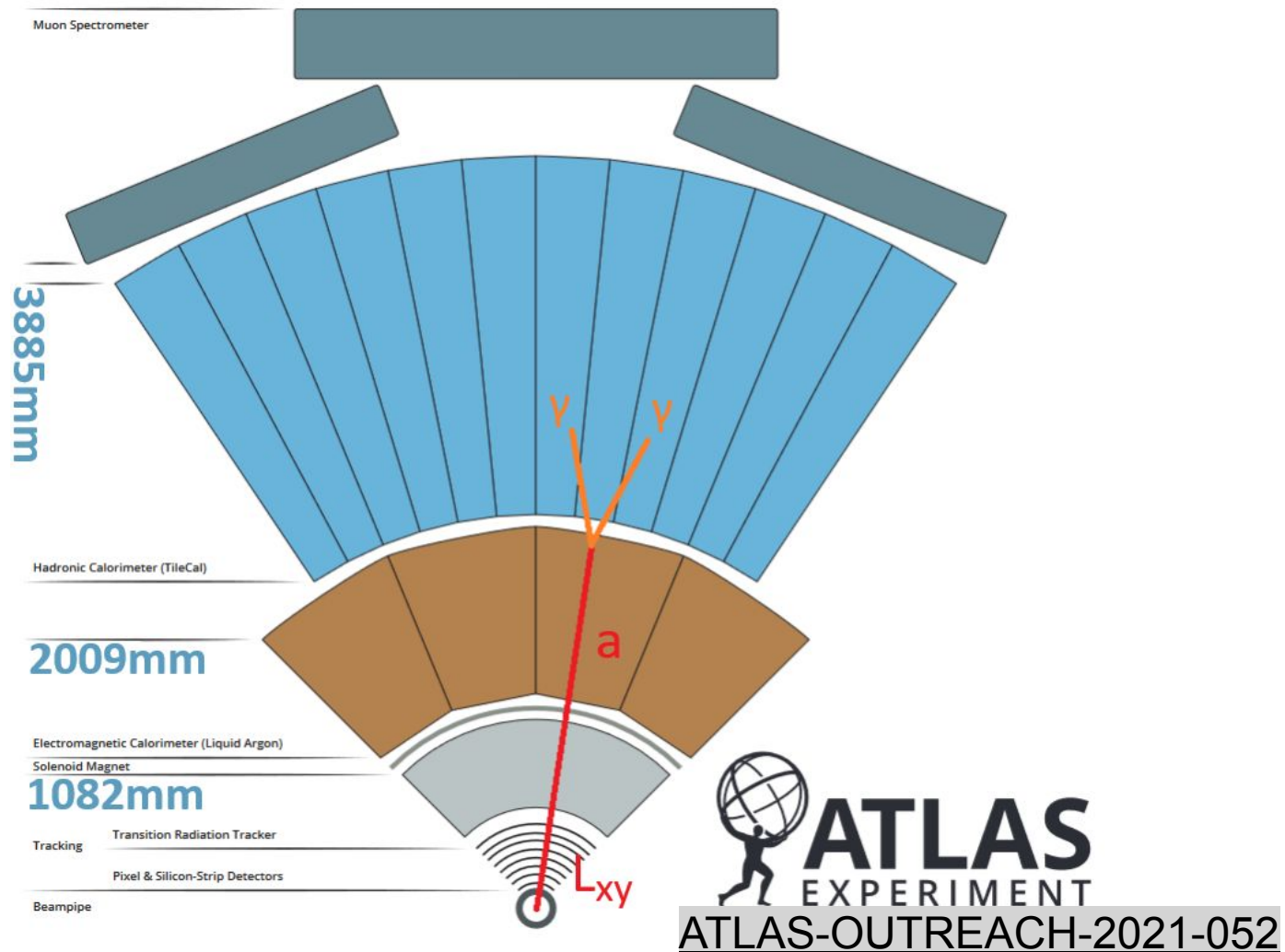
Prompt Decay Analysis

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- $m_a/C_{a\gamma\gamma}$ also allow for long-lived ALPs

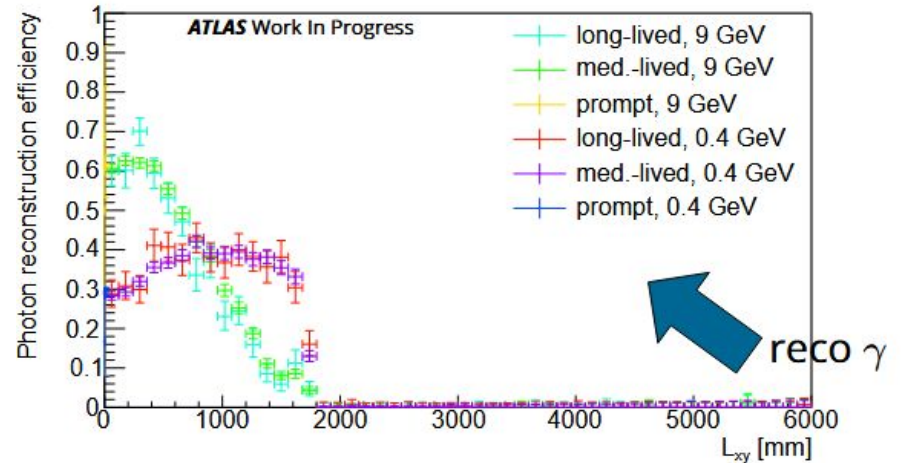
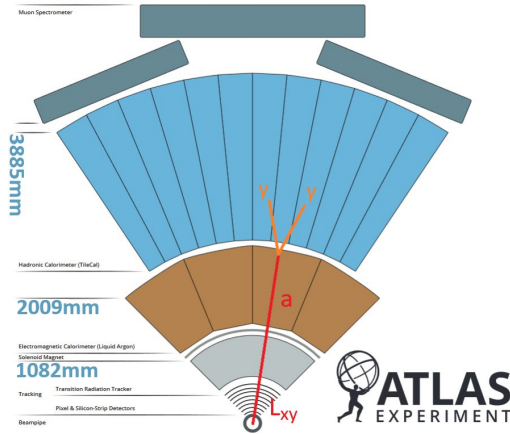


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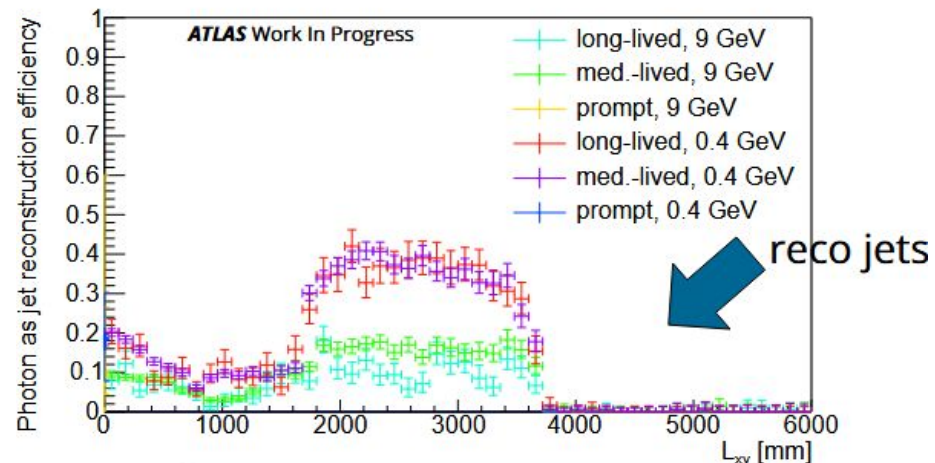
Displaced Showers in the ATLAS Detector



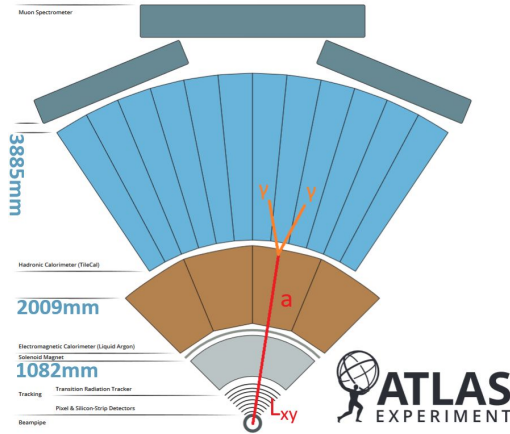
Photon Reconstruction Efficiency



- Reconstruction efficiency of ALP decay photons as function of L_{xy}
- At HCAL: photon reco efficiency = 0

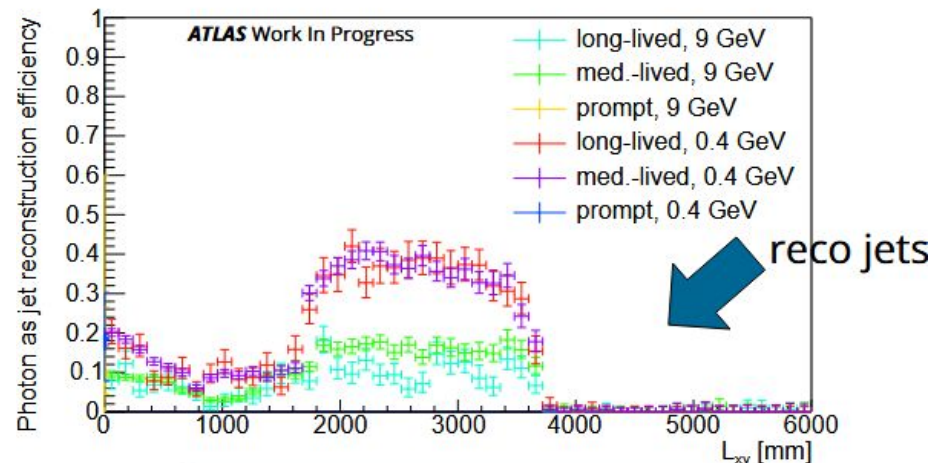
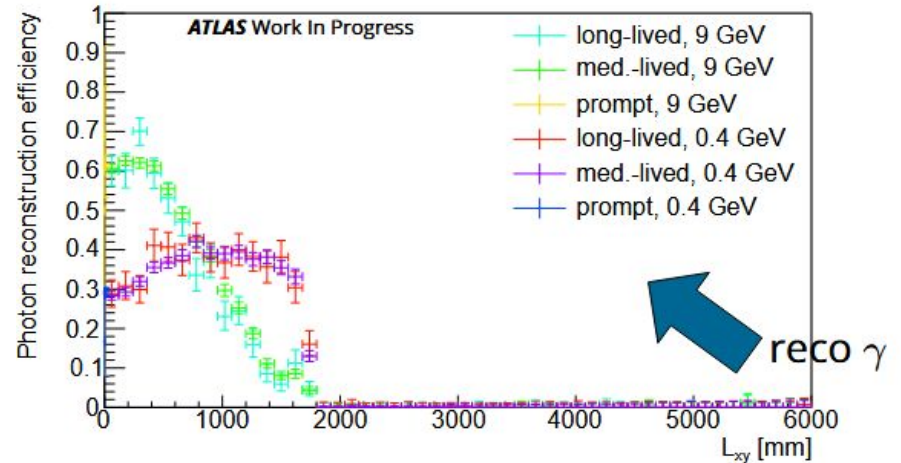


Photon Reconstruction Efficiency



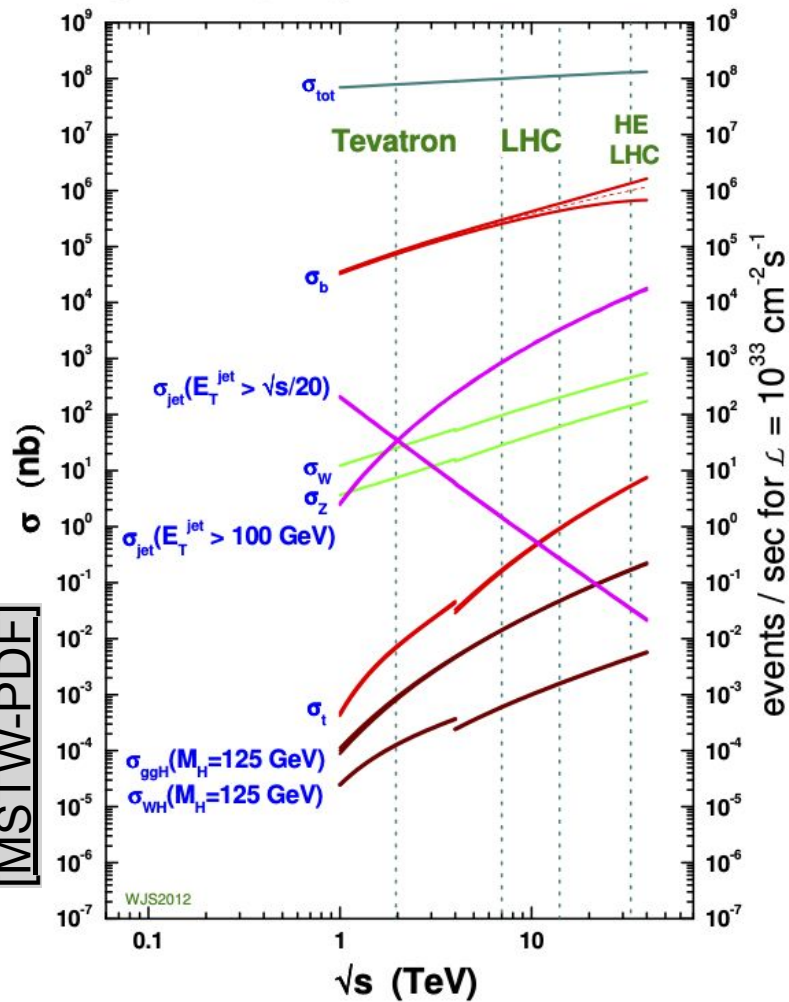
- Reconstruction efficiency of ALP decay photons as function of L_{xy}
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For large L_{xy} we need jets!



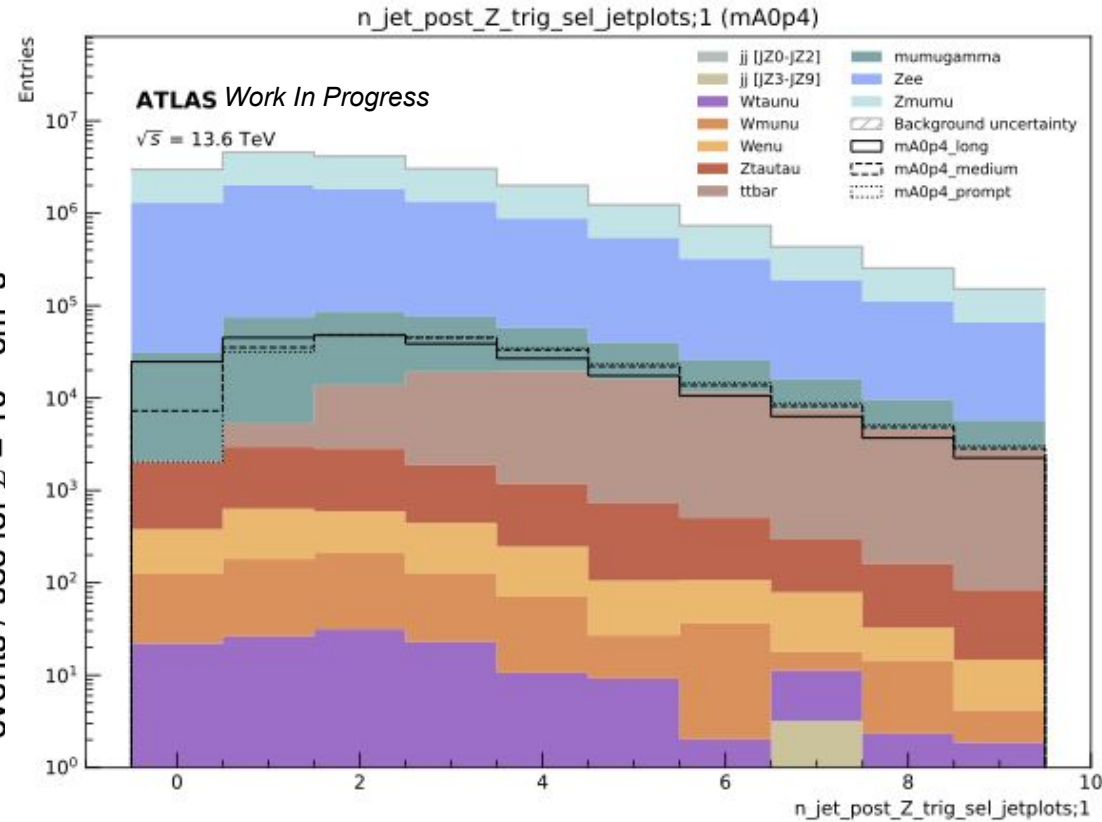
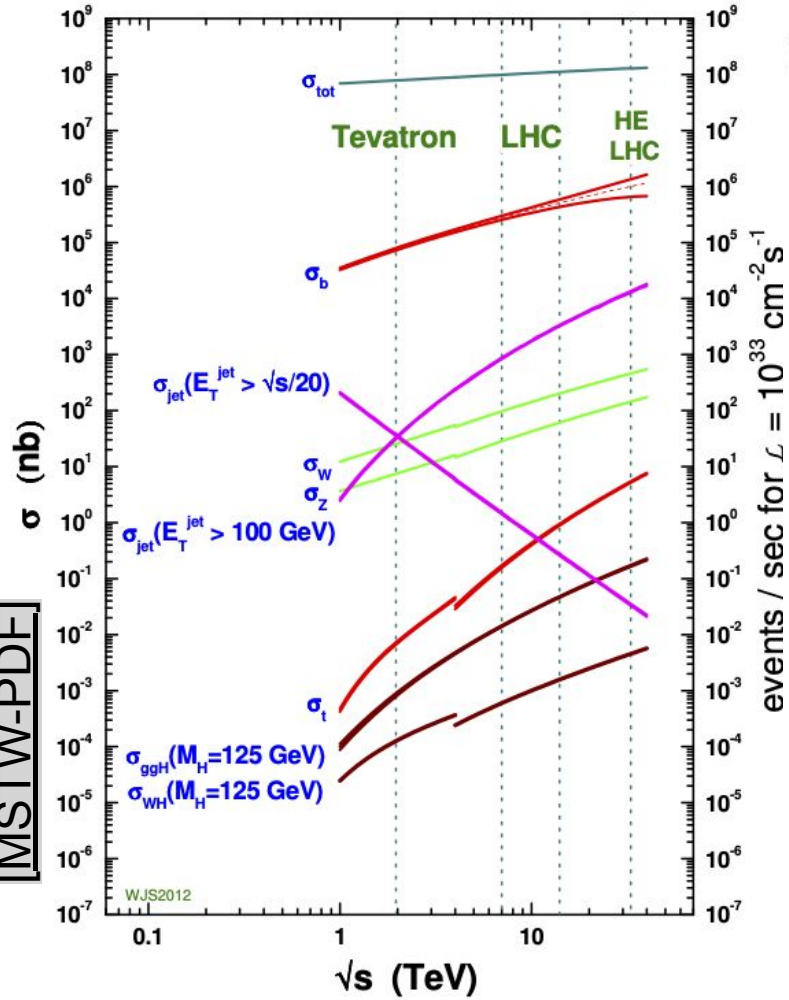
Background Jets

proton - (anti)proton cross sections



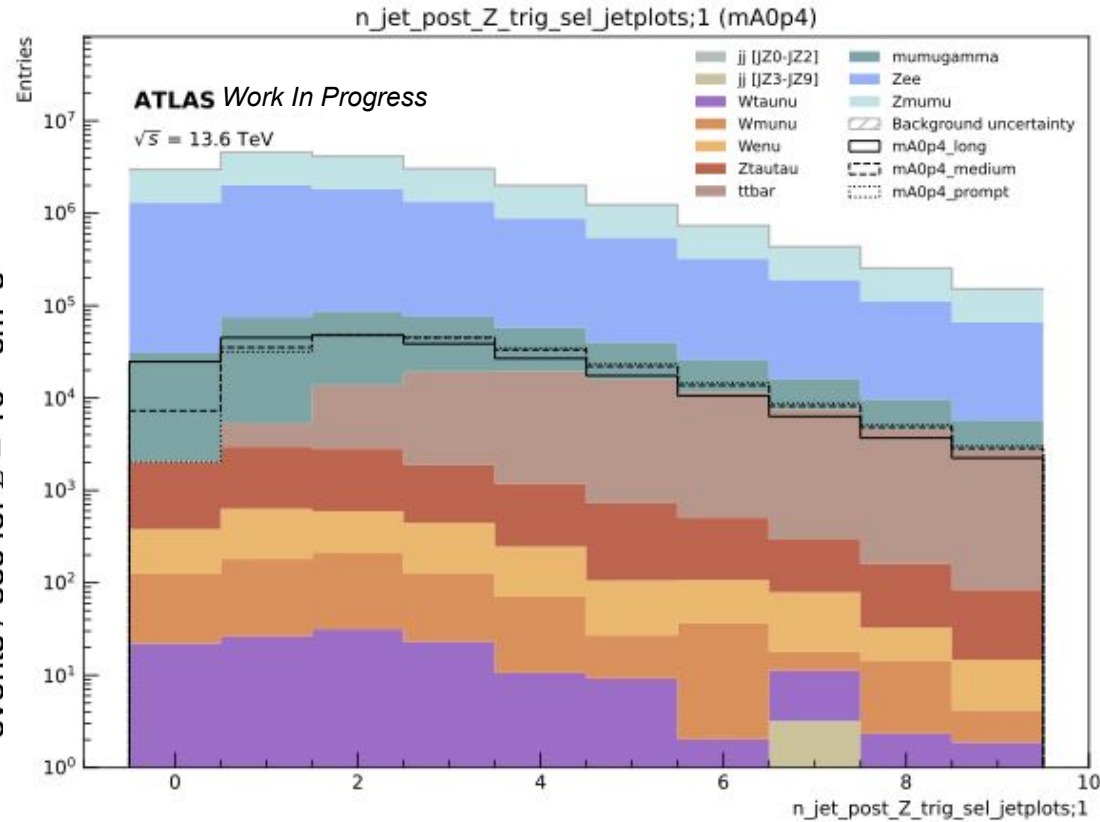
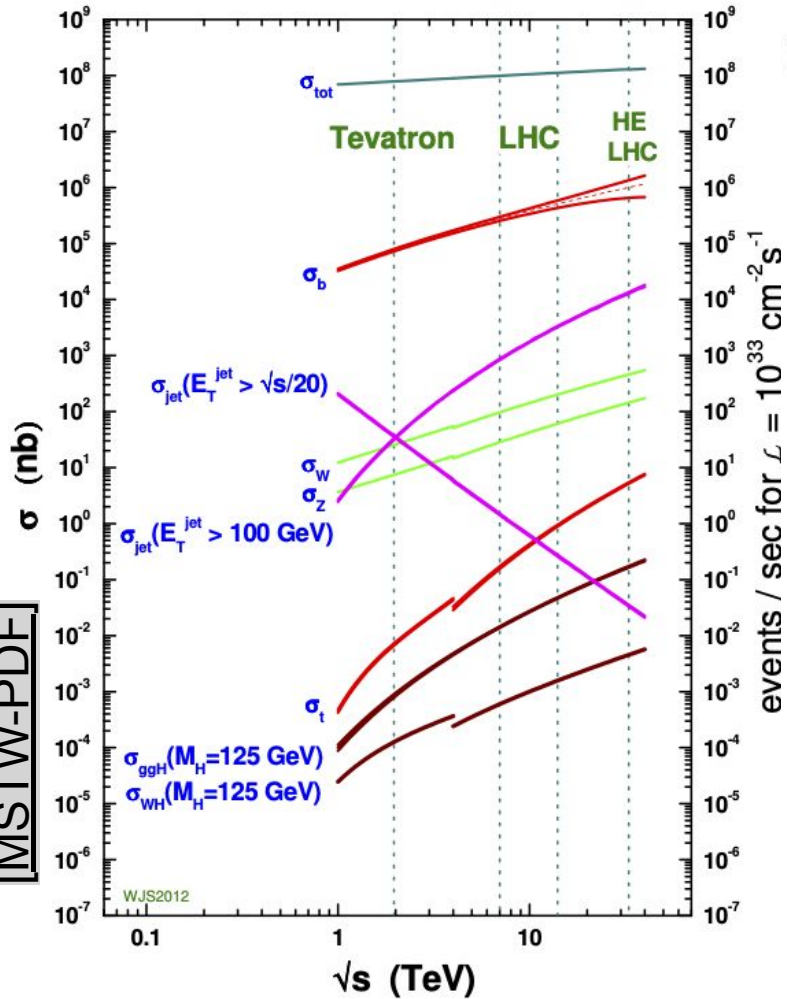
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Background Jets

proton - (anti)proton cross sections



Lots of background jets!

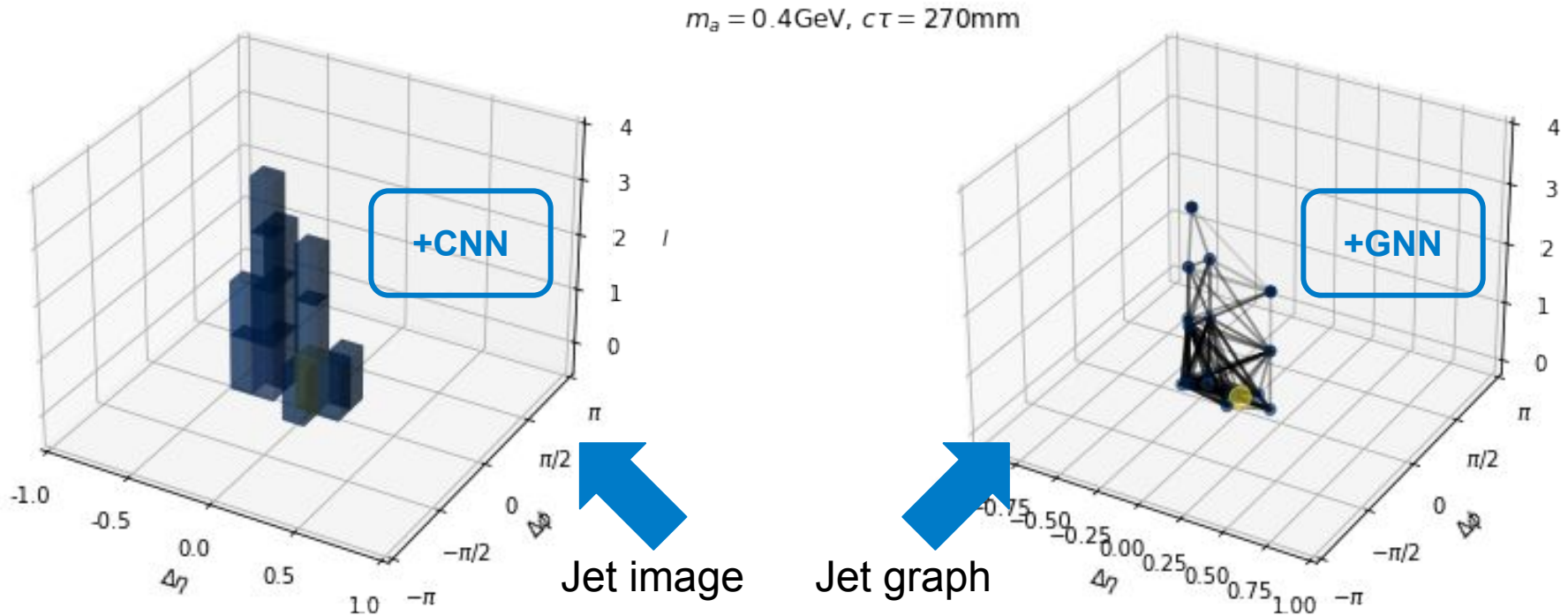
Jet Tagging with CNNs and GNNs

How to identify displaced photons as jets?

2 approaches

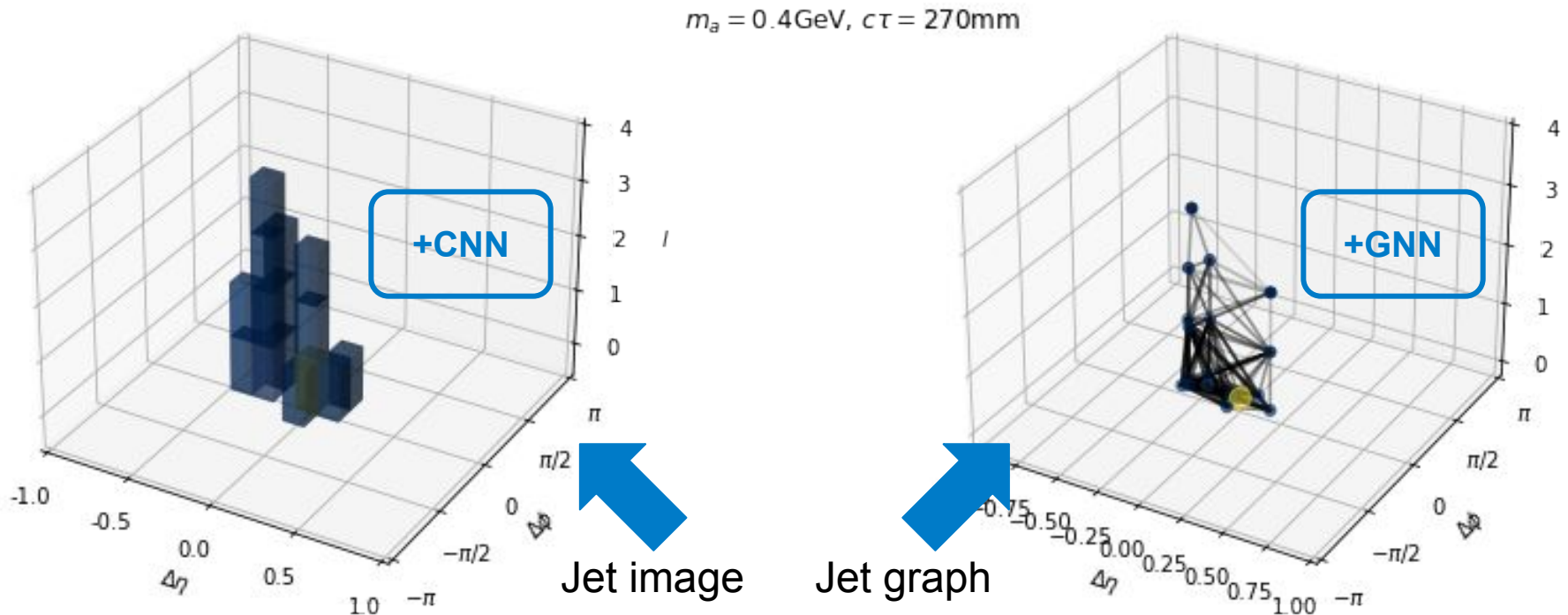
How to identify displaced photons as jets?

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How to identify displaced photons as jets?

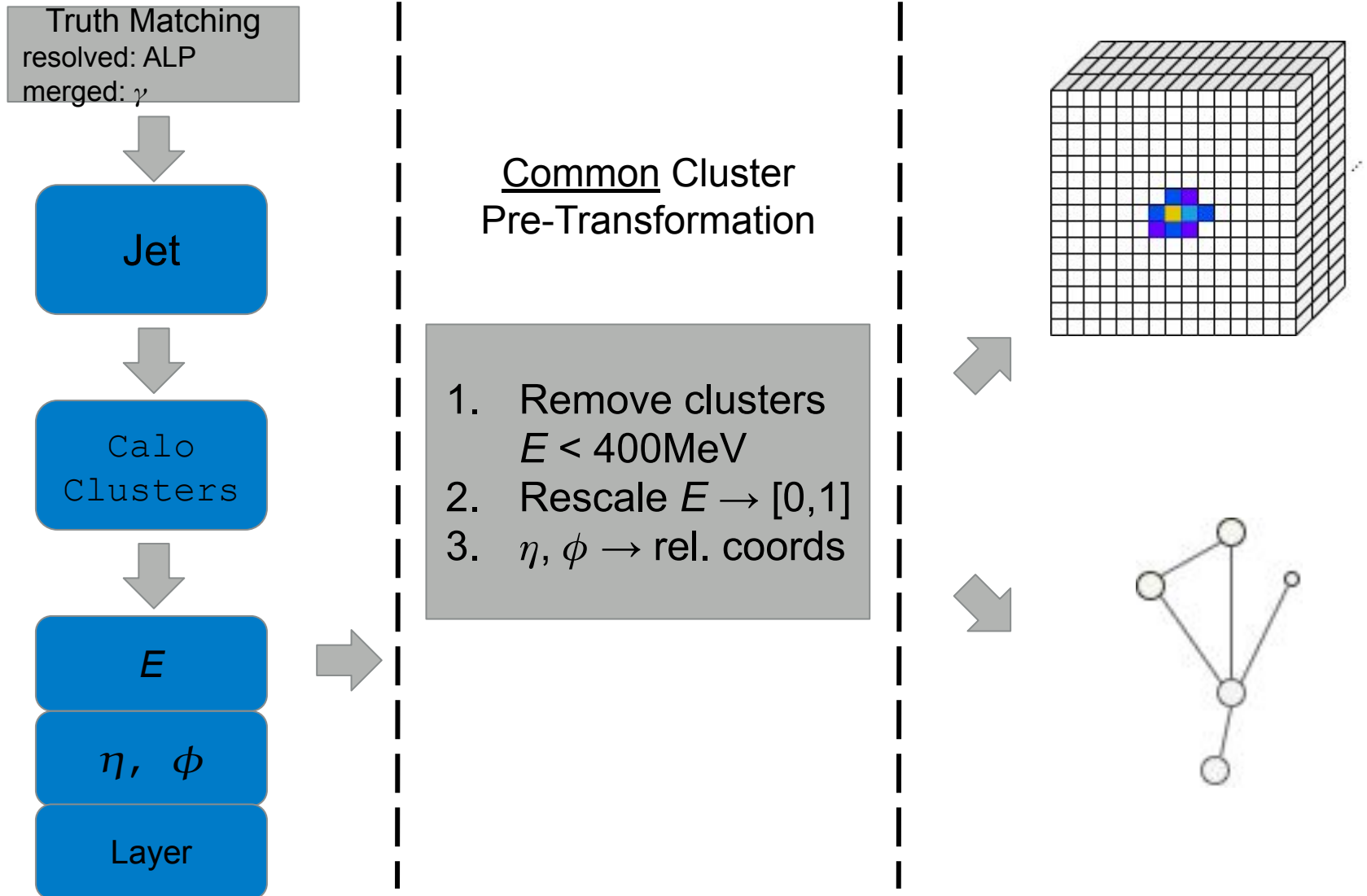
2 approaches



Which one performs better on the task?

Input Processing

Common Pre-Transformation



Input Processing Recipe

Common Pre-Transformation

Truth Matching
resolved: ALP
merged: γ



Jet



Calo
Clusters

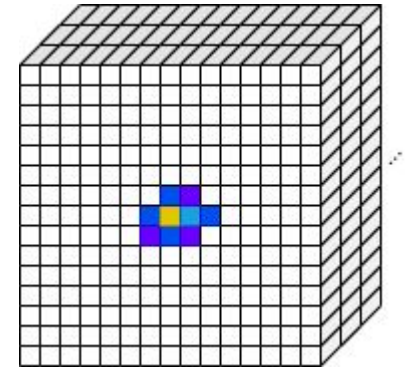
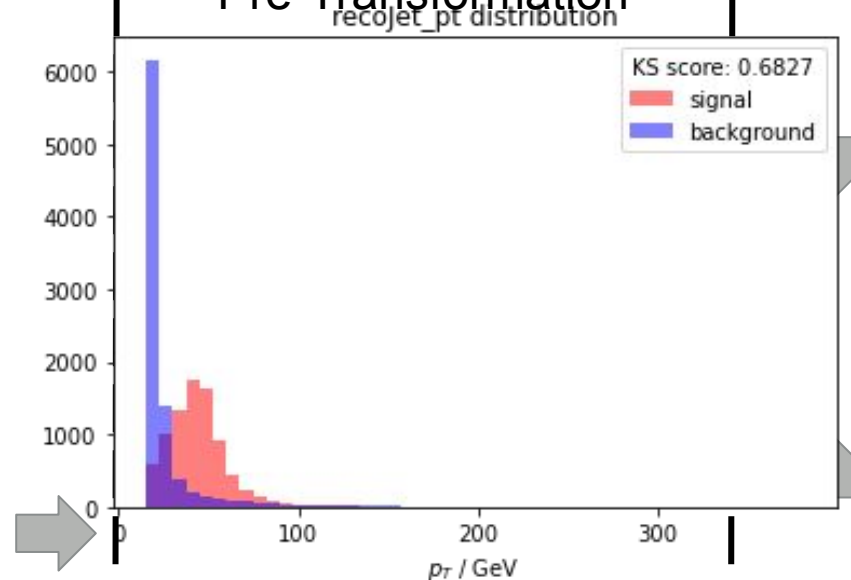


E

η, ϕ

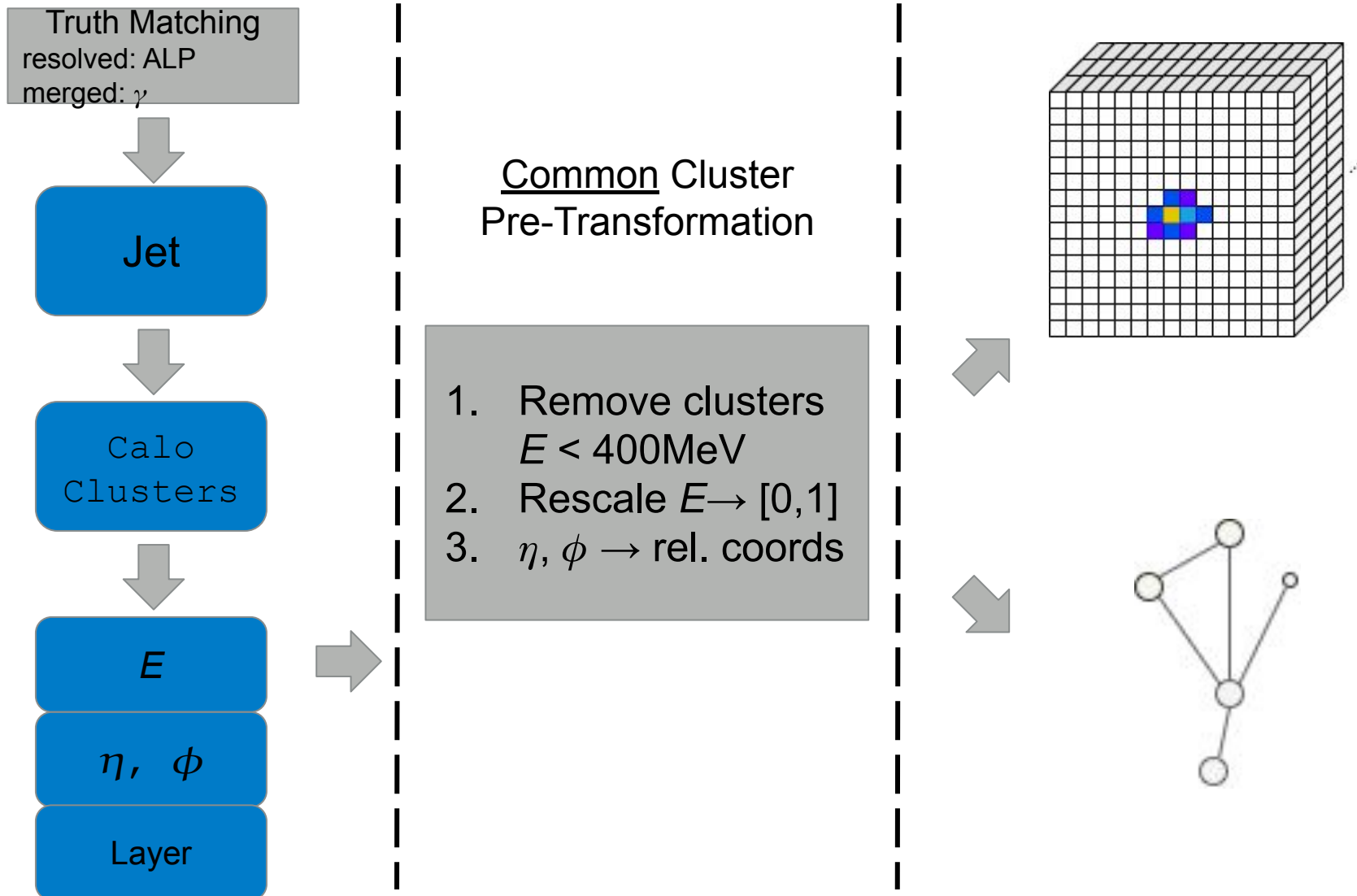
Layer

Common Cluster Pre-Transformation



Input Processing

Common Pre-Transformation



Input Processing

Common Pre-Transformation

Shower Development is Consistent Across Regions with Identical Material Structure

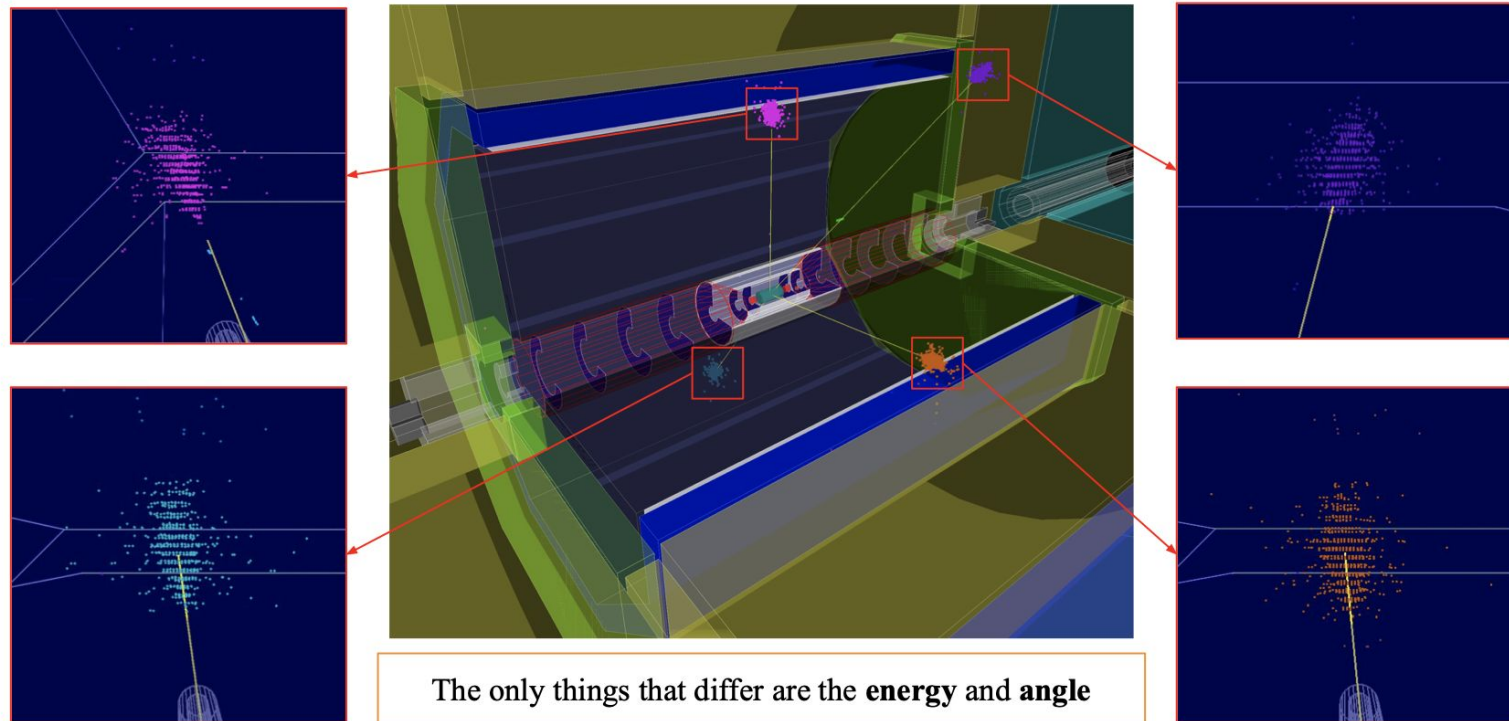
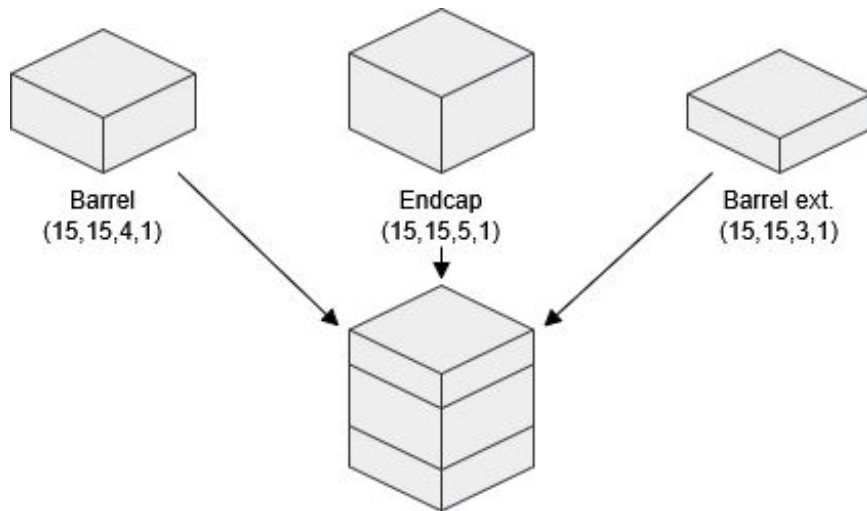
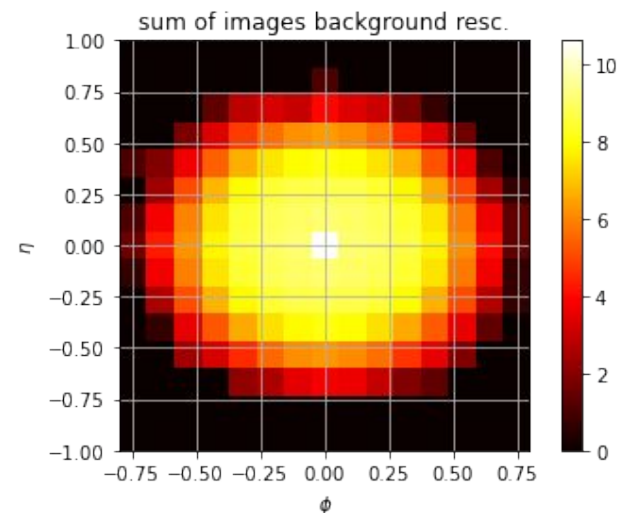
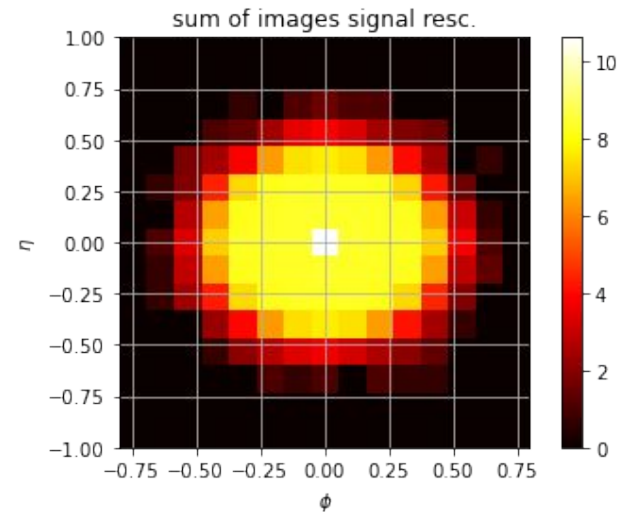


Image Building

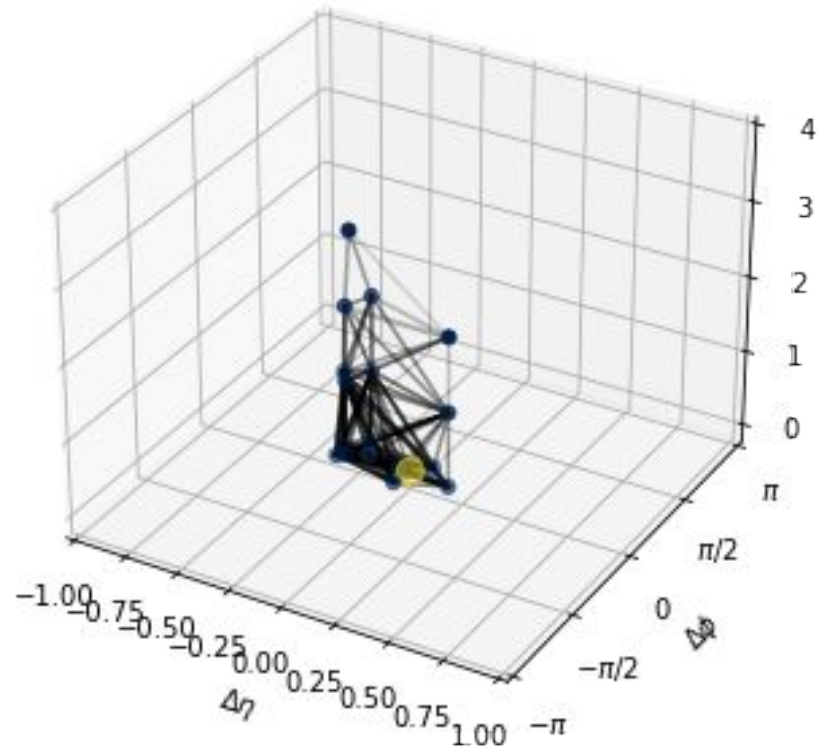


- Images binned in $\eta \times \phi = 15 \times 15$ bins
- Sph. symmetry, high E in center ✓
- Bound to fixed-size inputs
- Barrel, Endcap, Barrel ext. stacked



Graph Building

- Nodes = $[E, \eta, \phi, \ell]$ of clusters
- Edges = ΔR between clusters
 - $\Delta \ell = 0$: if $\Delta R \leq 0.6$
 - $\Delta \ell = 1$: if $\Delta R \leq 0.6$



Model Architectures

CNN:

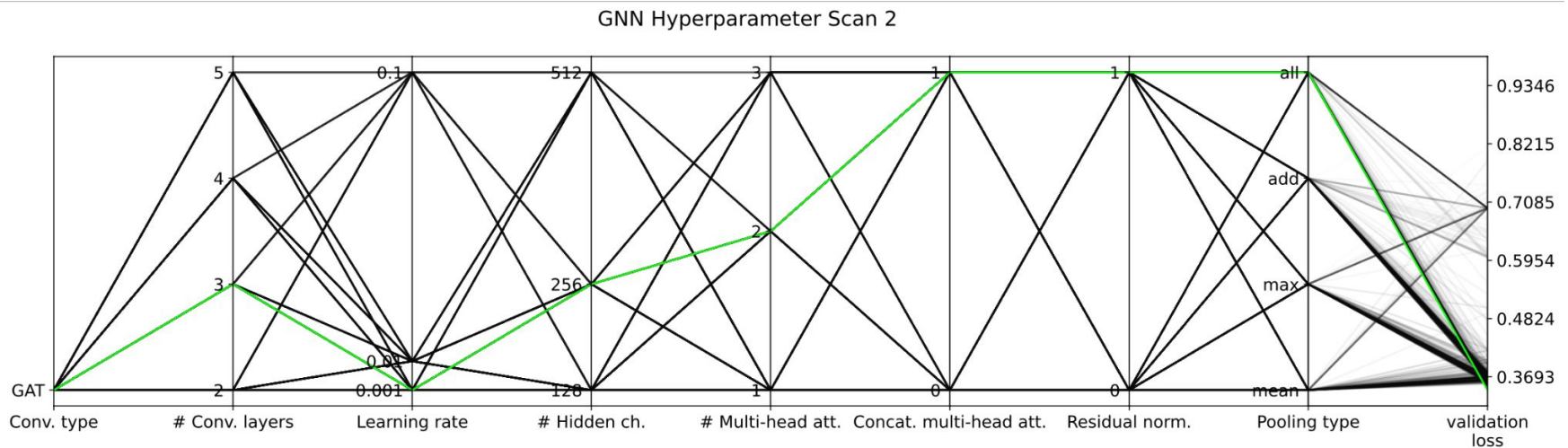
- Tensorflow
- Based on CNN developed for dark photon analysis ([arXiv:2206.12181](https://arxiv.org/abs/2206.12181))
- Conv3D + Pooling, Dense layers

GNN:

- PyTorch + pytorch_geometric
- Based loosely on jetgraphs library
- Convolution/Attention Graph Filters + Pooling

Hyperparameter Optimization

- Fully parallelized grid scans performed for CNN+GNN
- For 1 benchmark signal (0.4GeV, medium lifetime)



green line: best hyperparam configuration

& similar for the CNN

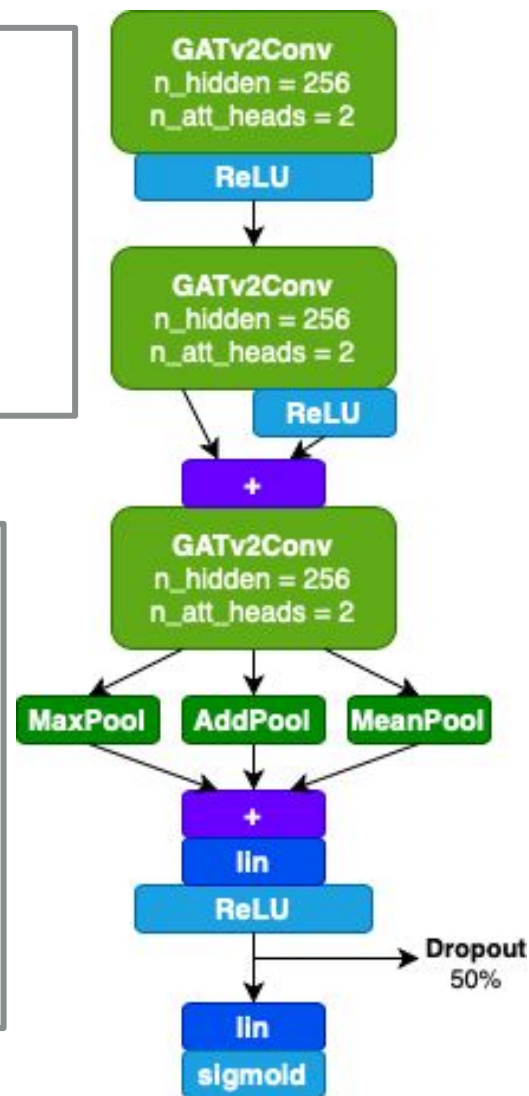
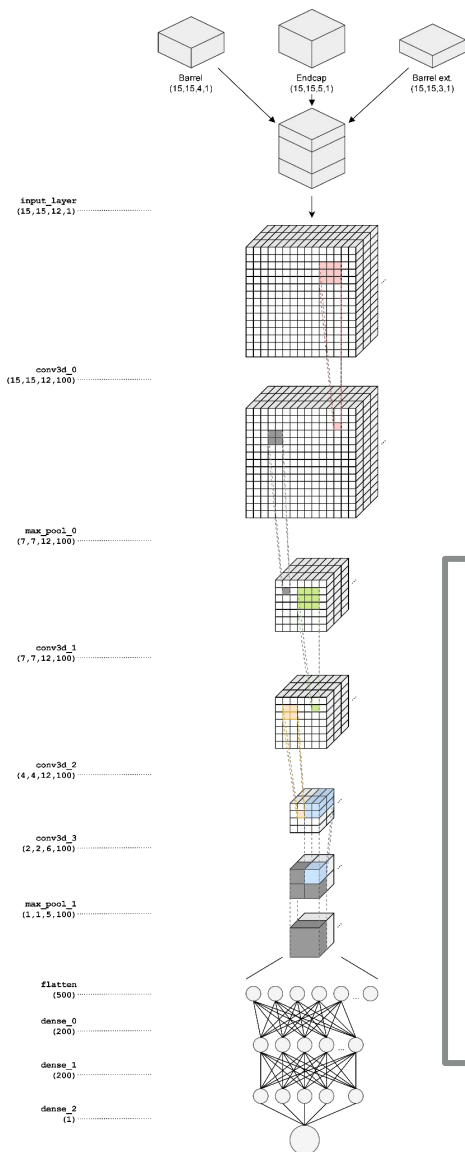
Model Architectures after Optimization

← CNN

- input [15,15,12,1]
- 3× Conv3D
- MaxPooling
- 3× Dense (n_nodes=200)
- ~1M parameters

GNN →

- 3× GATv2Conv (2 att. heads, 256 hidden ch.)
- Feed-forward at every 2nd layer
- Max+Mean+Add pooling combined
- ~1.3M parameters



Deploying to Analysis

C++ Framework Integration

- **Python** framework good for training & evaluation
- For analysis big scale inference: Implement model into ATLAS data processing framework (**TopCPToolkit**) **C++**
- CaloClusters can be **huge** in data! (esp. for backgrounds with many jets)
→ Only save output score for each jet & drop clusters afterwards!

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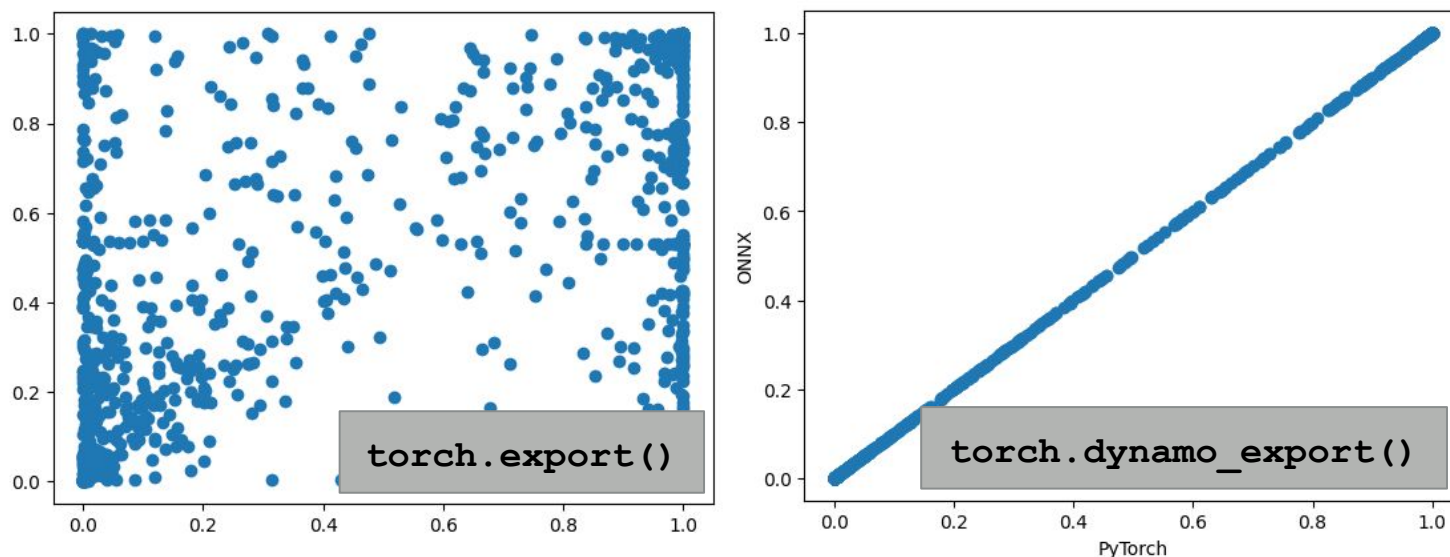
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- **TopCPToolkit** can talk to ONNX models ✅
- To accomplish:
 1. Export models to ONNX
 2. Rewrite model-inputs processing in C++ Algorithm

ONNX Export

Part I: The scatter_reduce fight

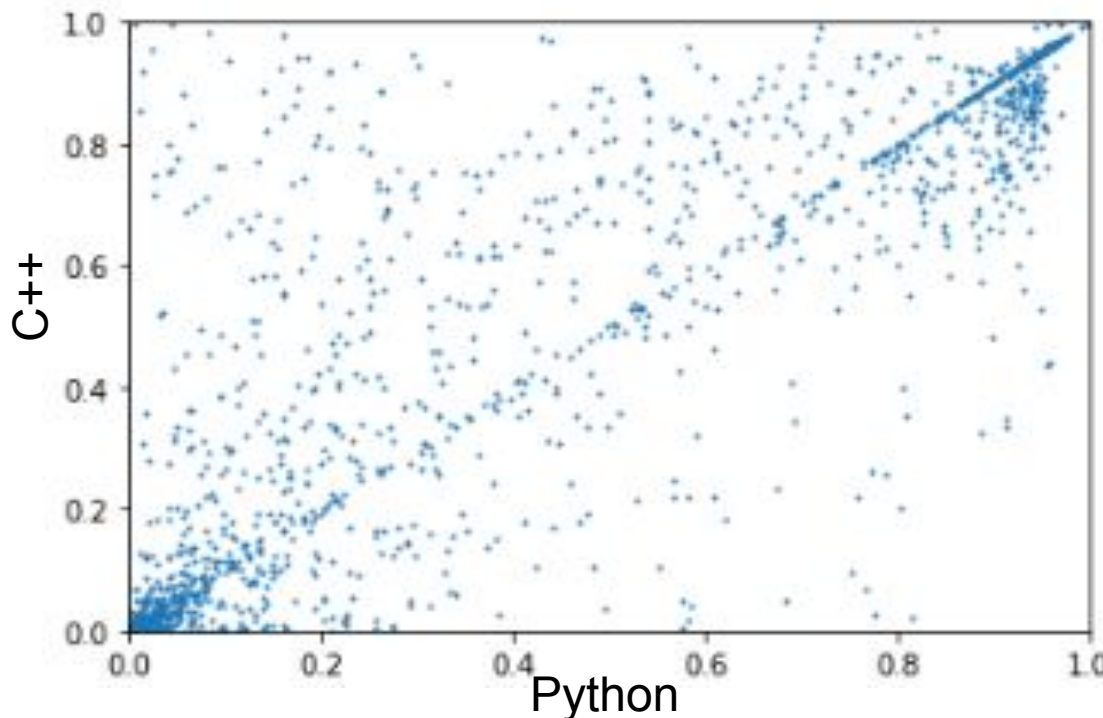
- ONNX does not support `torch.scatter_reduce` with `include_self=False`
- E.g. GATv2Conv layers in GNN from `torch_geometric` use this
- Luckily, there was a long-awaited fix to `torch_geometric` lib ([commit](#)) 
- GNN: variable-sized inputs (“dynamic axes”):



Input Building

Part II: Attack of the <vector<vector<vector<vector<...>>>>

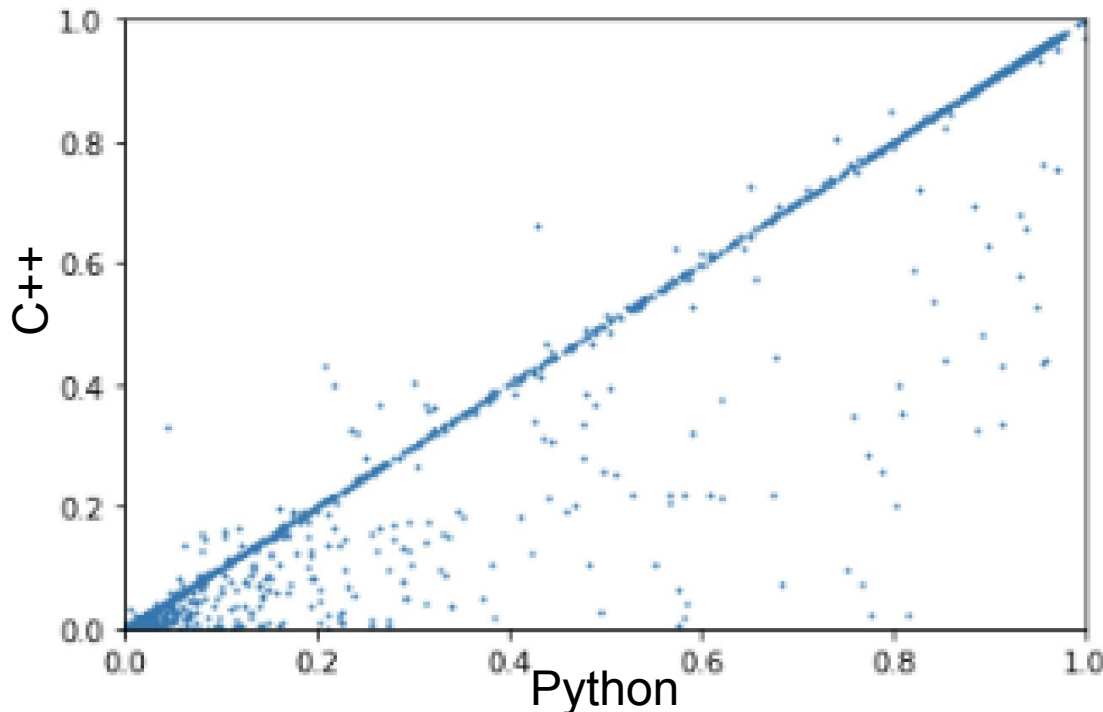
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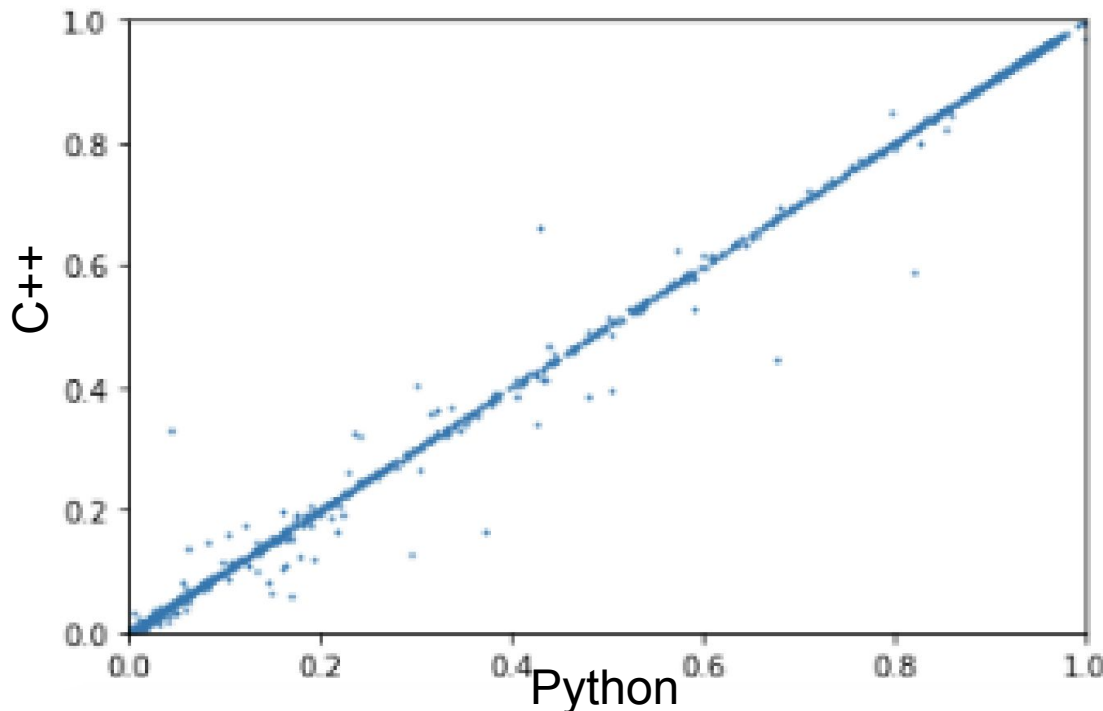
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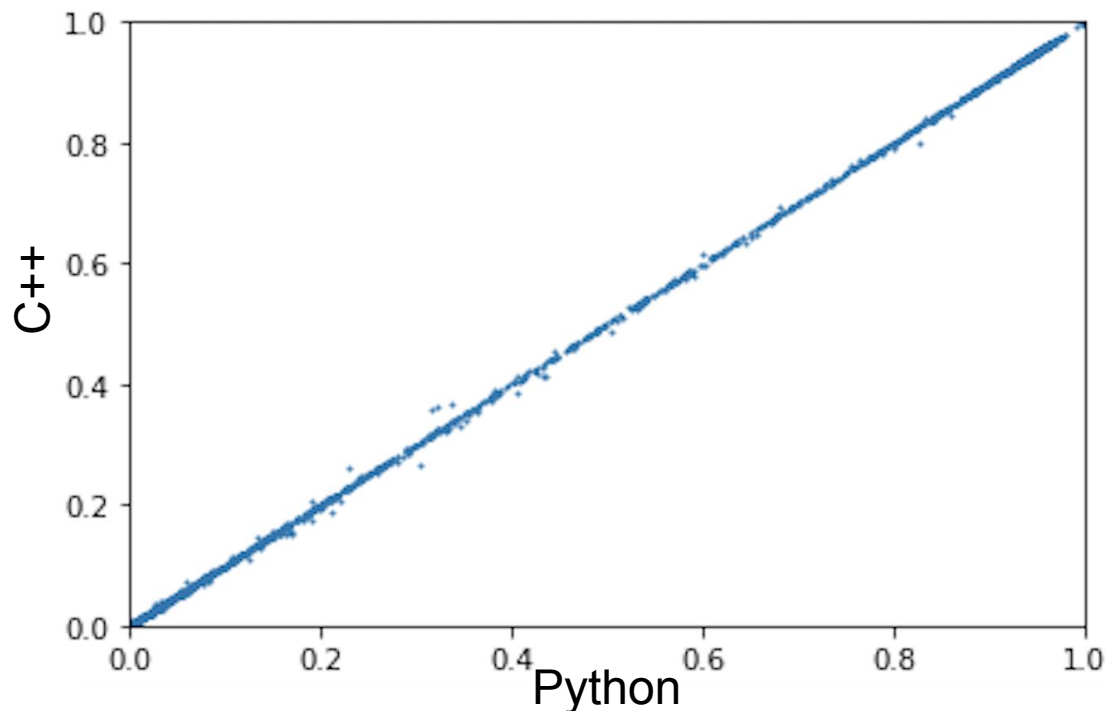
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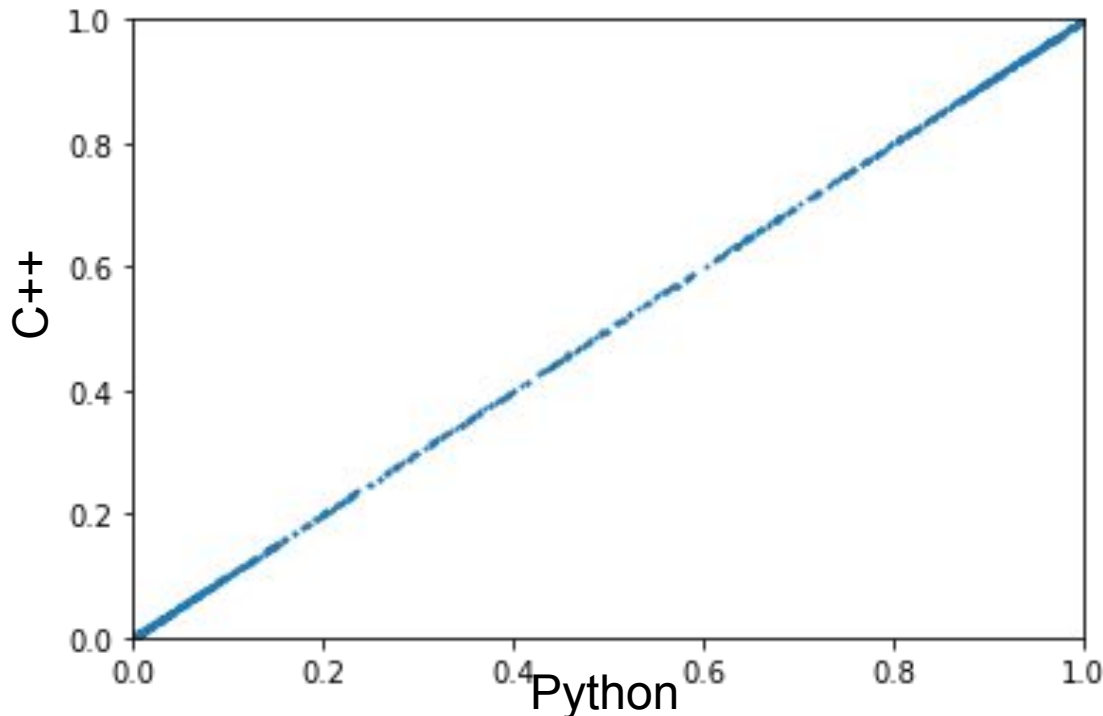
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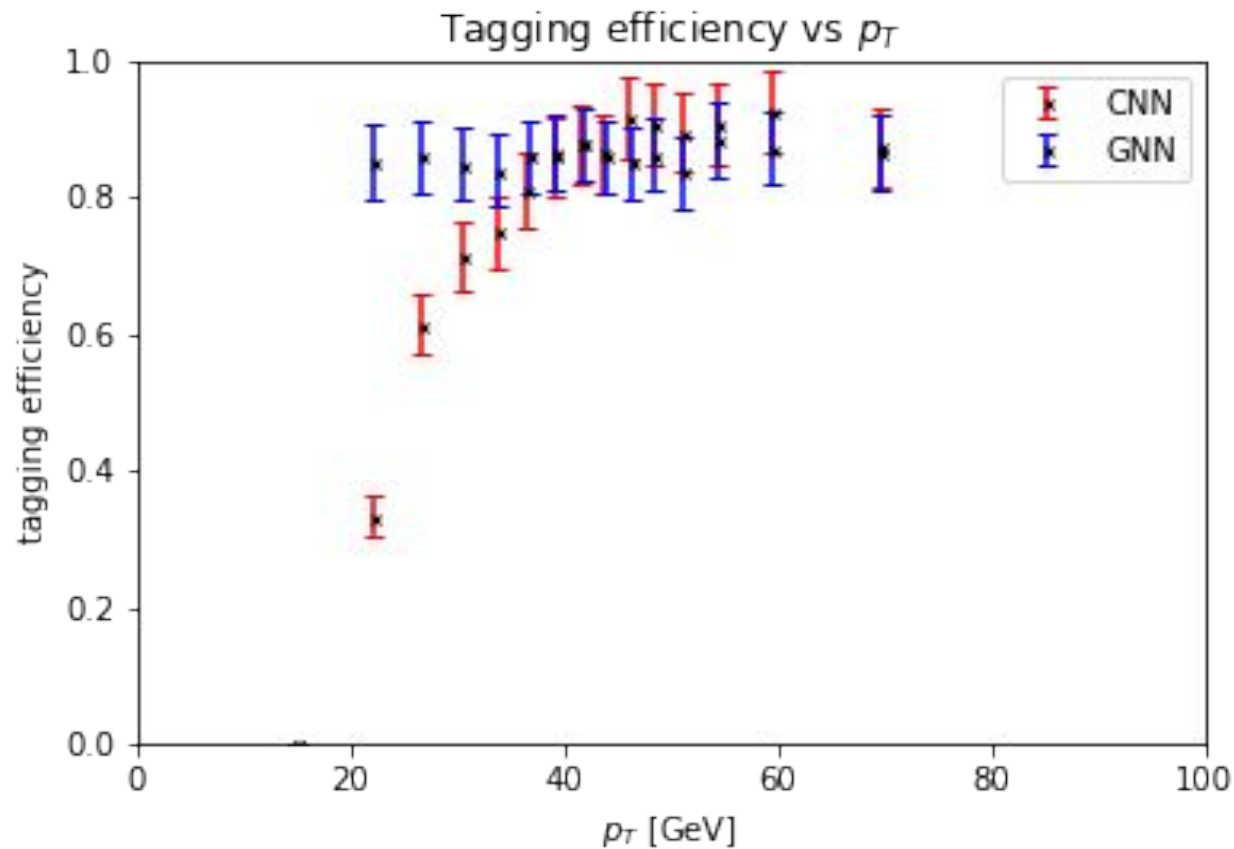
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Recent Results

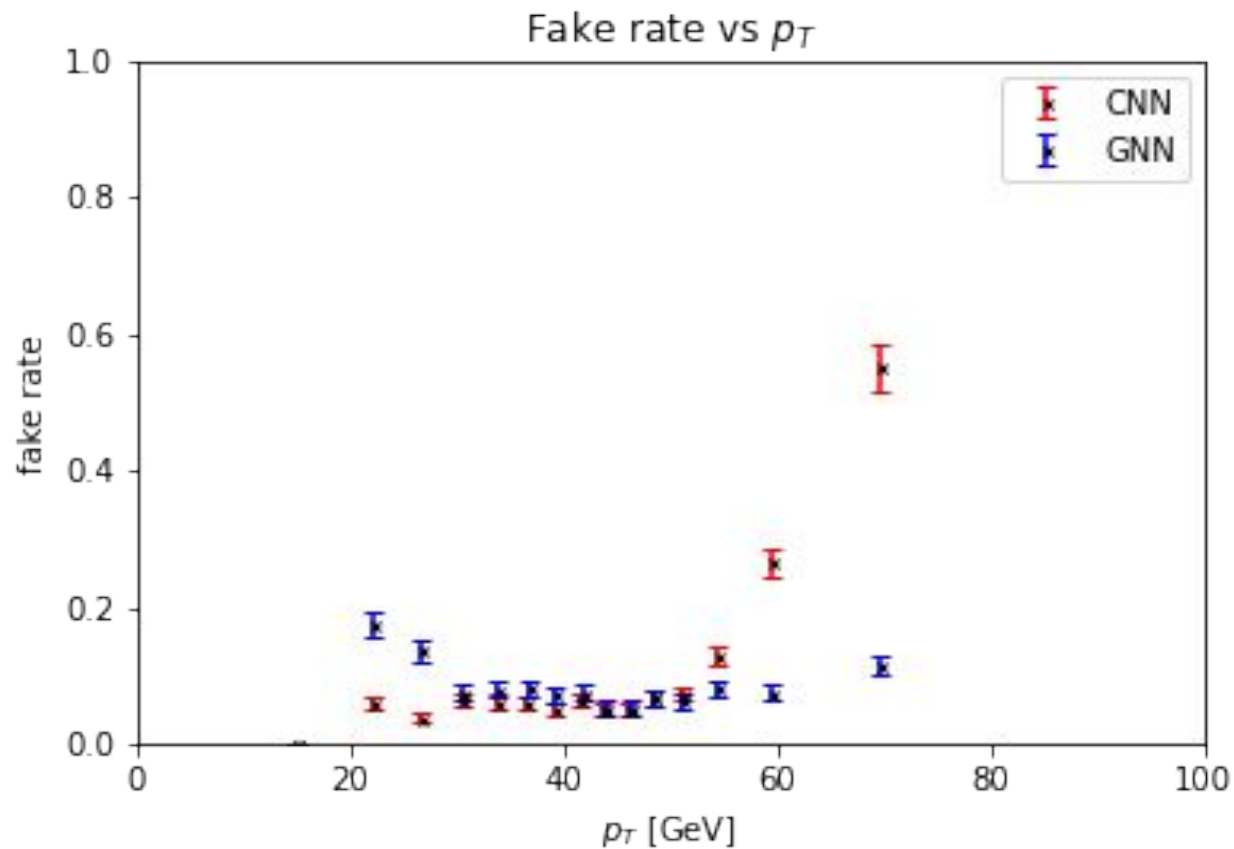
CNN vs. GNN

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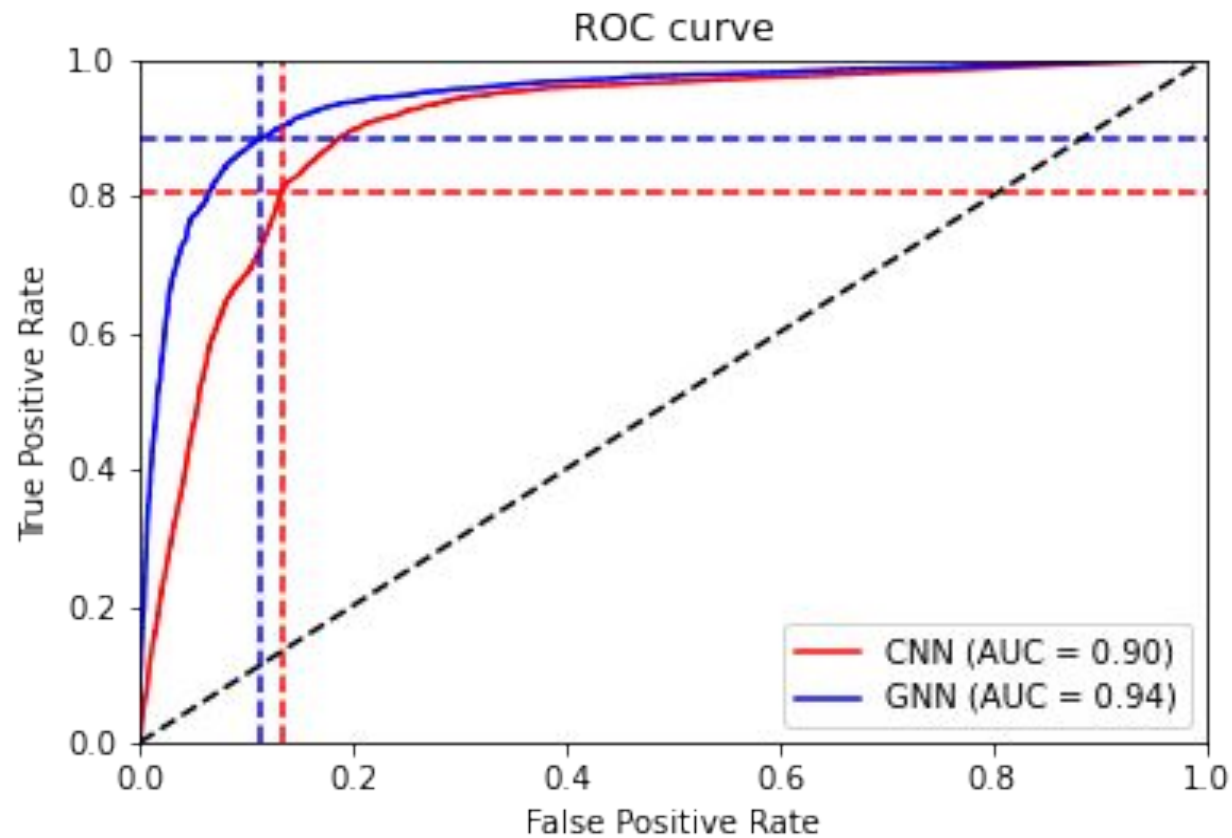
- (0.4GeV, medium lifetime sample)

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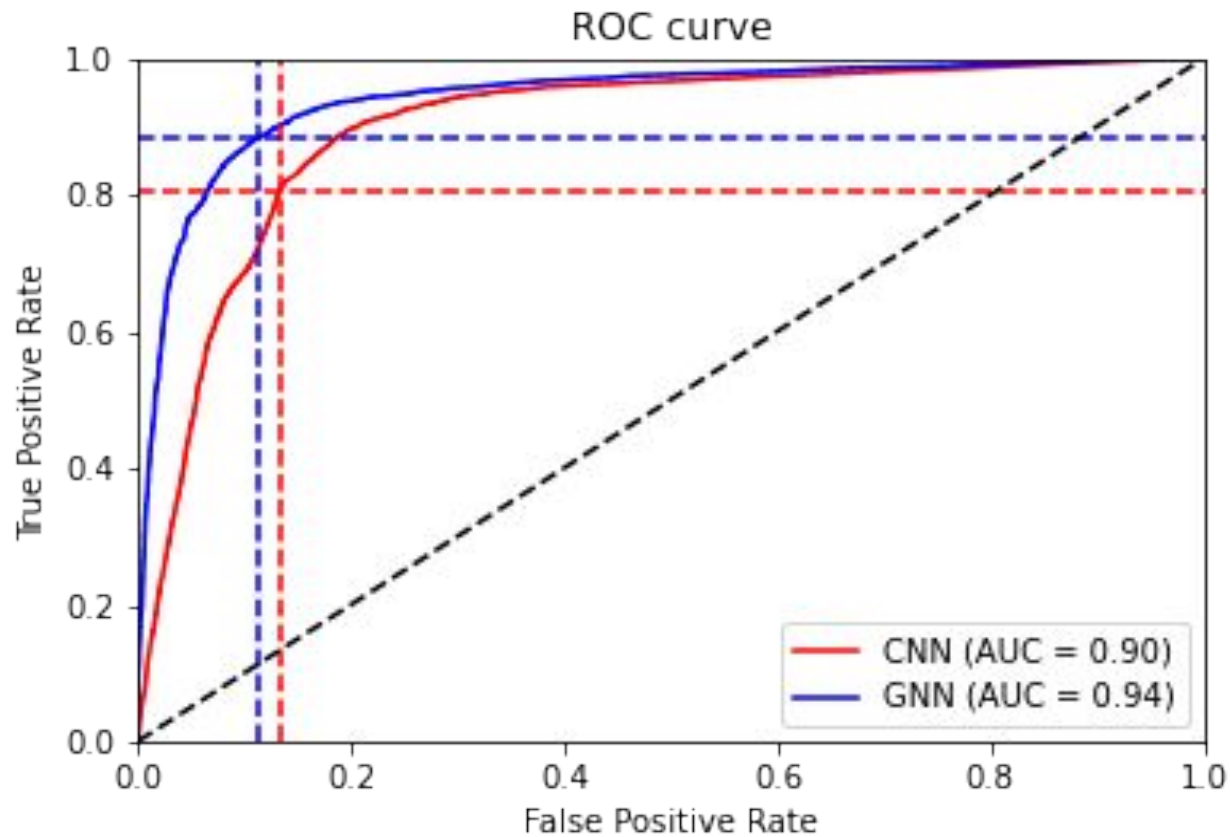
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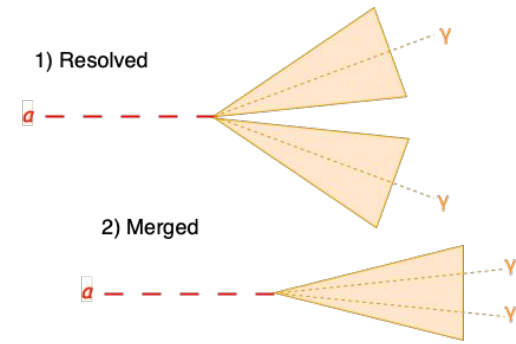
CNN vs. GNN



GNN seems to perform better than CNN!

GNN Signal Topology Dependence

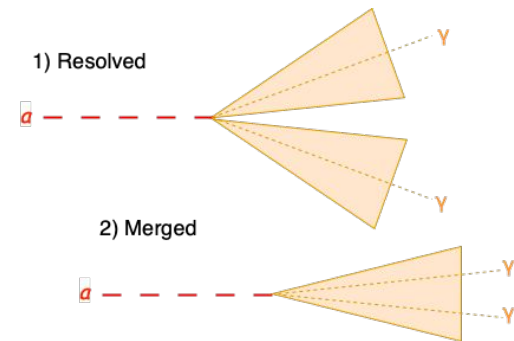
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Separate GNNs for merged/resolved topology? 

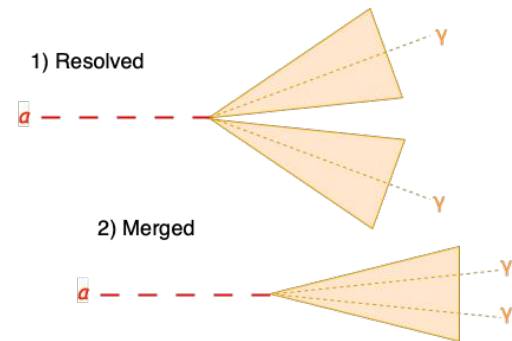


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- Train 2 GNNs on **merged** (0.4GeV) and **resolved** (9GeV) scenario → cross-tests!

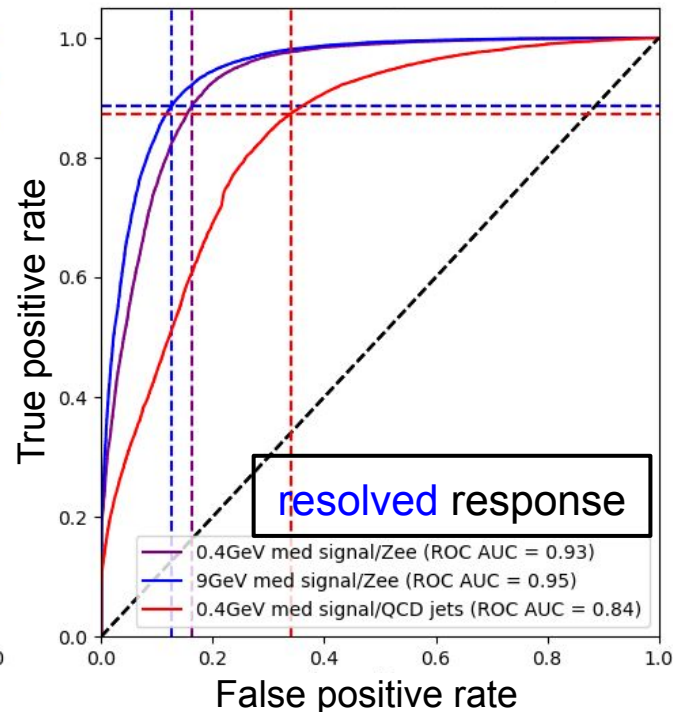
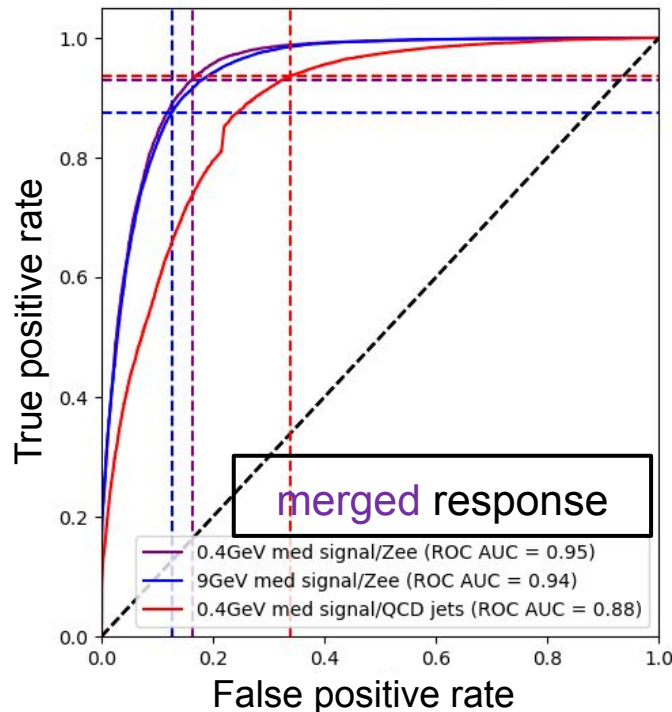
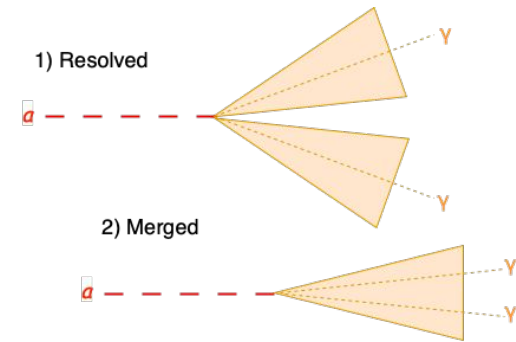


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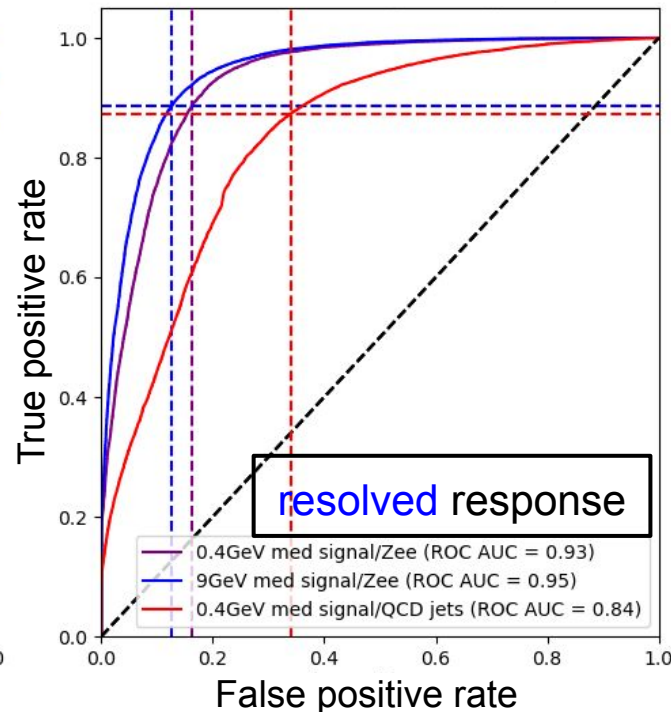
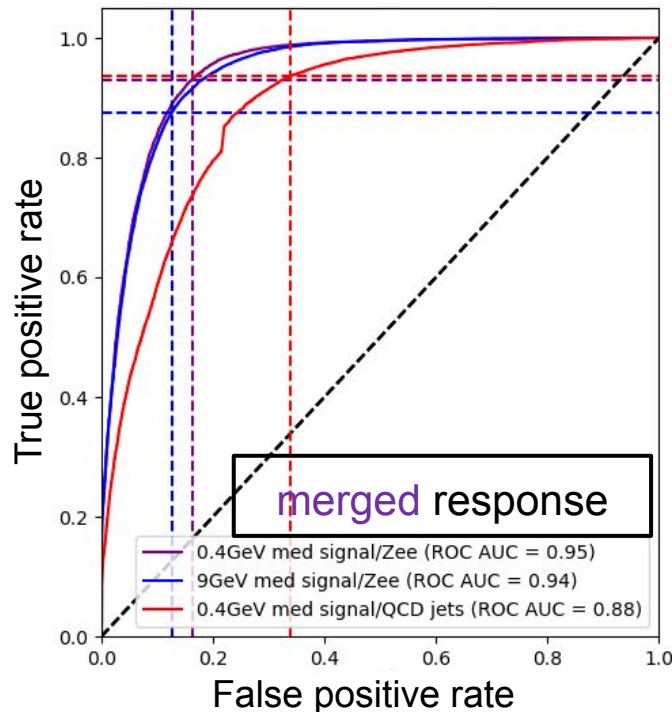
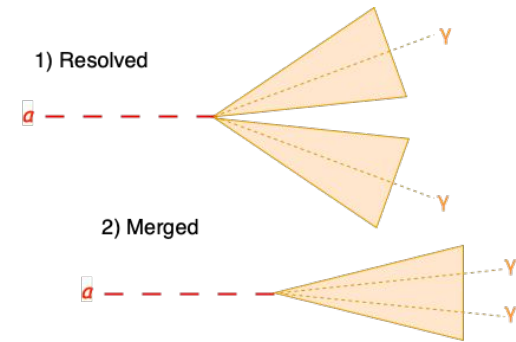


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- Both GNNs perform well on both test datasets!
- ⇒ no need for separate GNNs for **merged/resolved**!

Can train 1 GNN for both scenarios!

GNN Prompt Photon Background


- Other signal topologies:  - what about other backgrounds?

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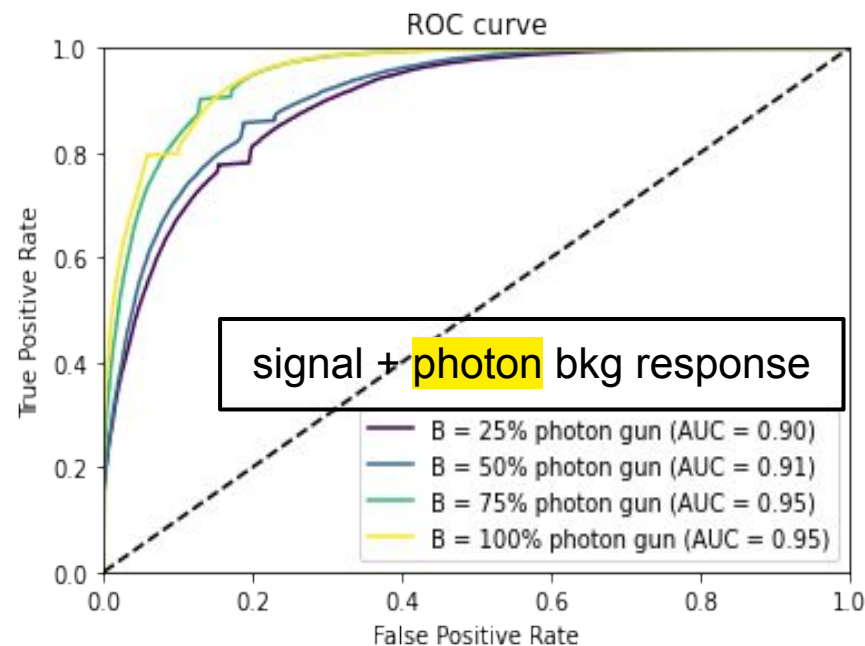
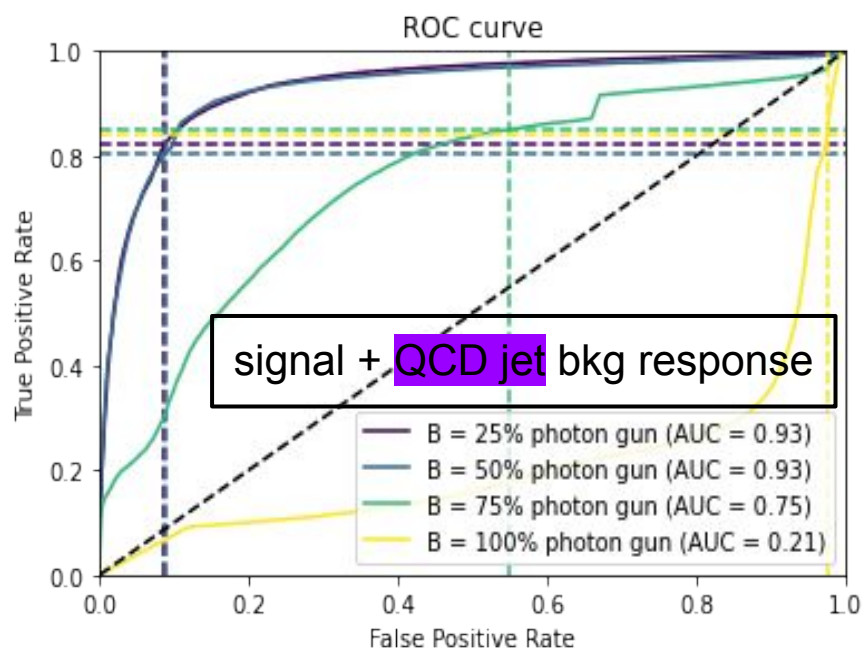
- Idea: Train separate GNNs on 4 datasets different backgrounds (same signal)
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- Here: Test on 100% **QCD jets** and 100% **photon gun**

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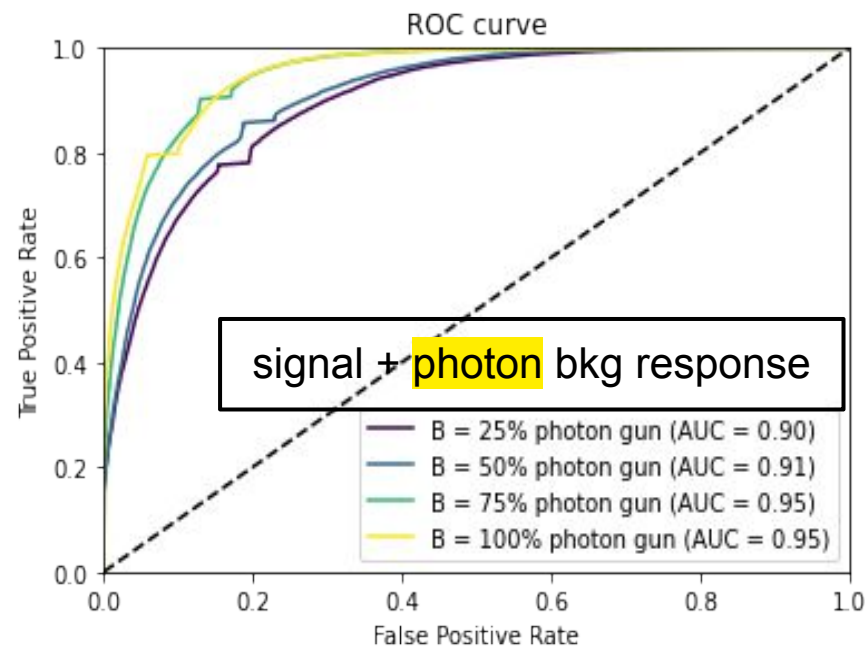
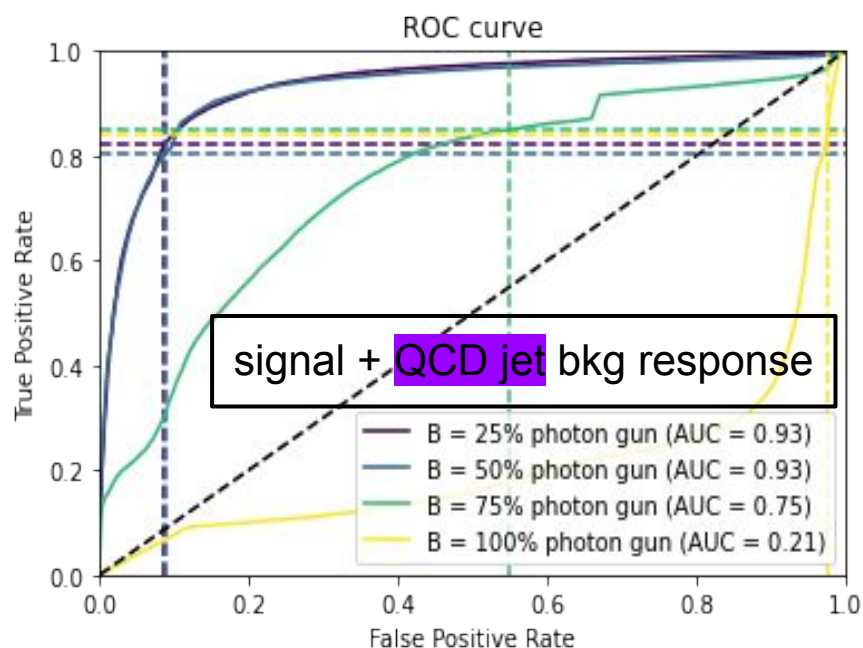


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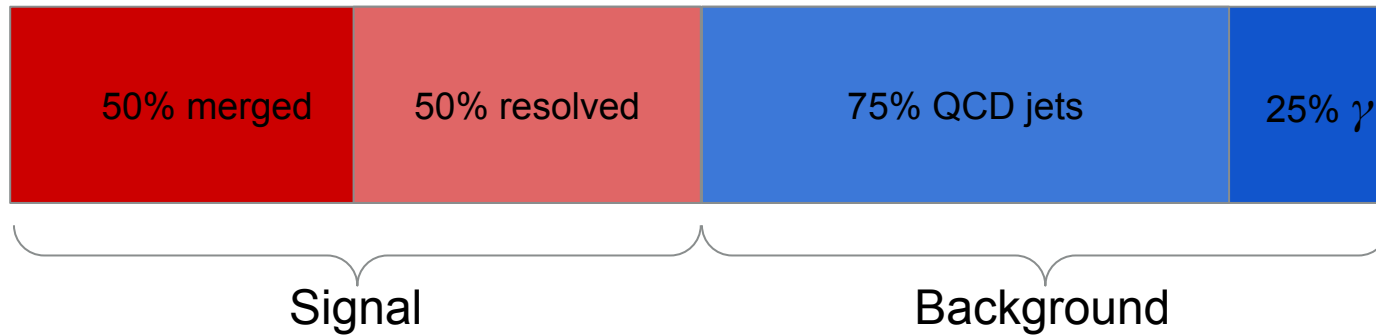


- ⇒ Low number of prompt photons to achieve for good separation
- ⇒ Trade-off: ✨“Perfect mixture”✨ for training around **25%** - **50%** prompt γ

Combining what we have learned

The Ultimate Training

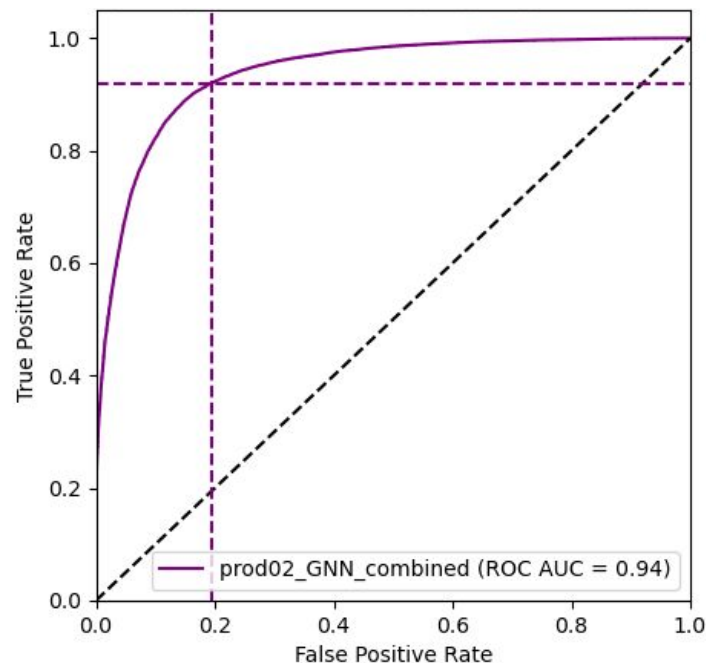
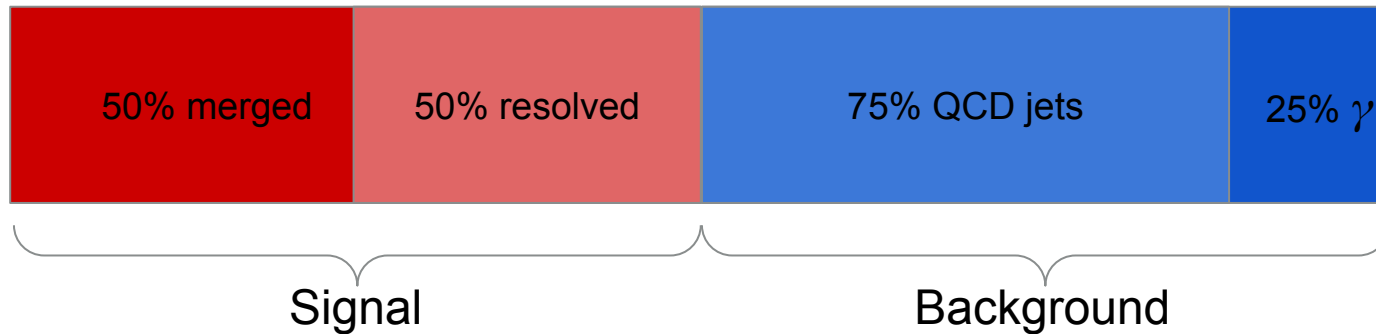
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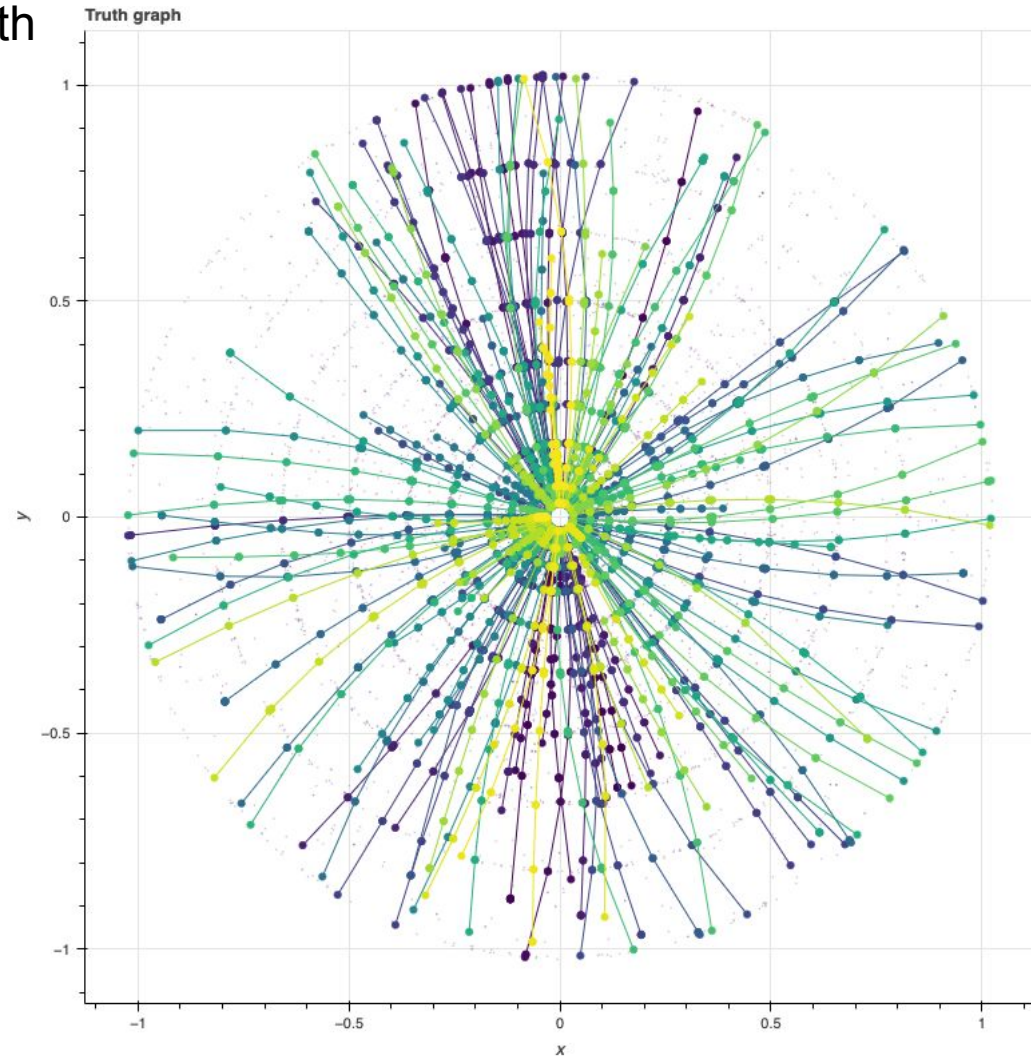
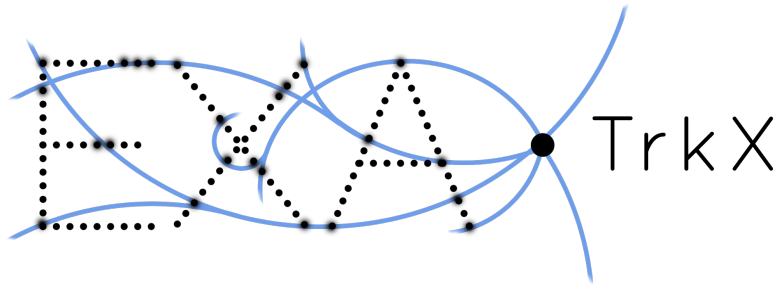
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- No need to have separate GNNs for different signal topologies
- With ~low fraction of training data, GNN can achieve good separation against prompt photons
- Next: Evaluating signal model dependence of GNN tagger

What else can you do with GNNs?

(A Cliffhanger)

- Exa.TrkX: advanced tracking with GNNs in HEP
- Side-project with Federico and Thomas
- Tracking at future muon collider experiments (e.g. MAIA)
- Incl. timing (4D Tracking)

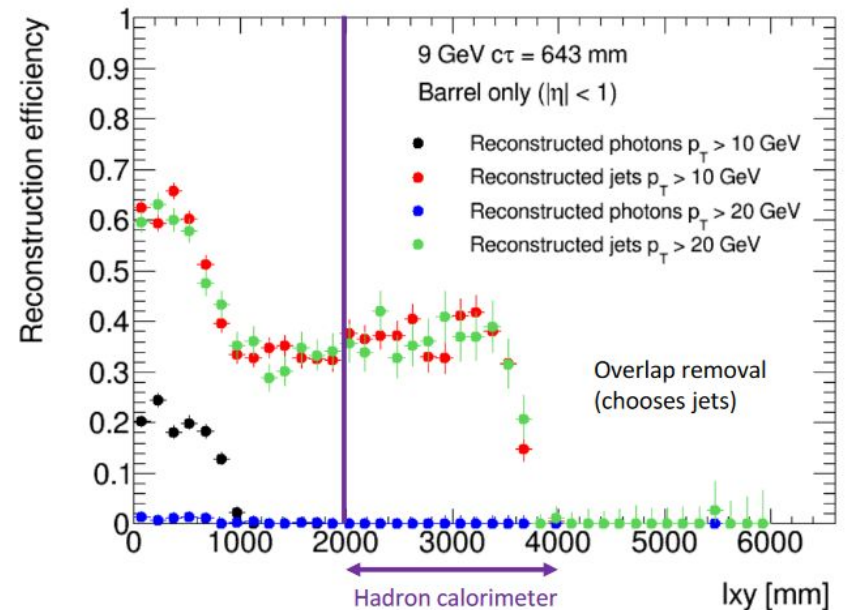
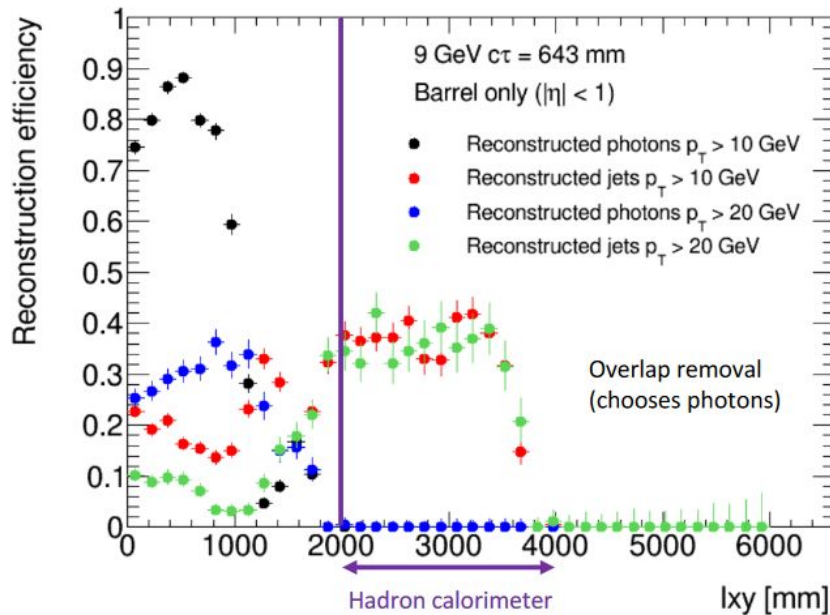


Thank you!

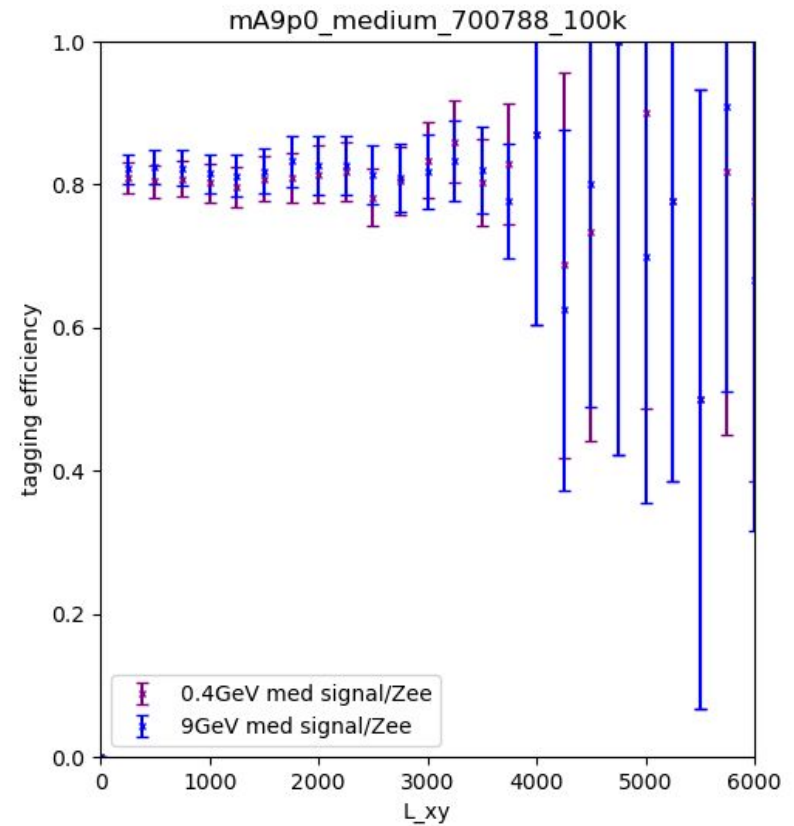
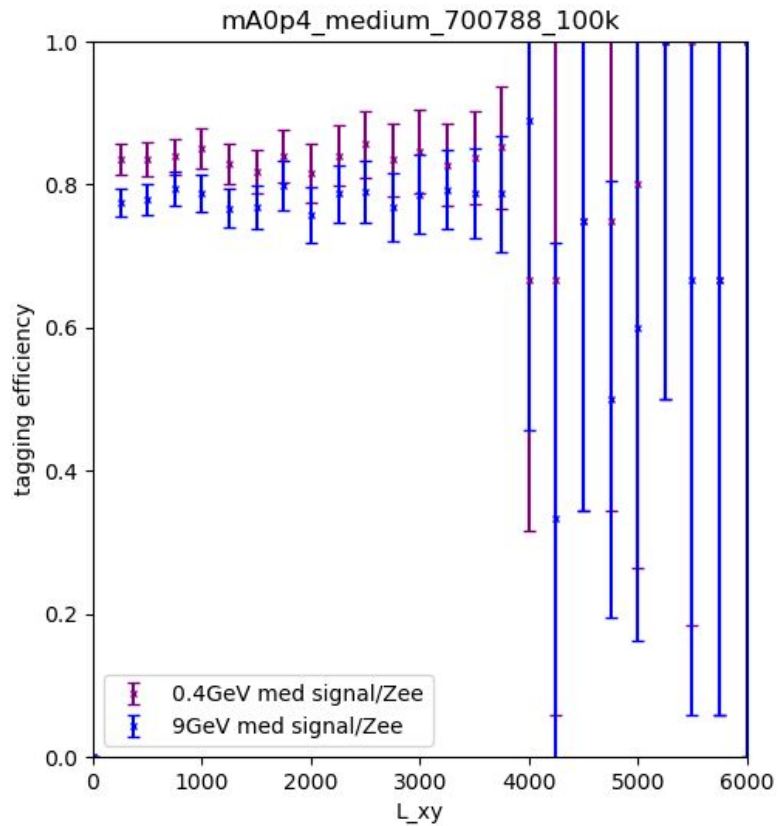
Backup

Photon Reconstruction Efficiency

Overlap Removal



Tagging efficiency L_{xy}



Output Score Distribution Features

