

A yellow circular logo with the text "QUANTUM UNIVERSE" and "ATTRACT. WORKSHOP" in black. Below the text is a stylized drawing of a horseshoe magnet with two plug-like ends. The background of the slide is a dark blue space scene with white stars and streaks of light. On the right side, there are several circular diagrams representing different physical concepts: a spiral galaxy, a nebula, a solar system with a central star and planets, and a cluster of particles.

QUANTUM UNIVERSE

ATTRACT.
WORKSHOP

Alignment limit and strong first-order electroweak phase transition in 2HDM

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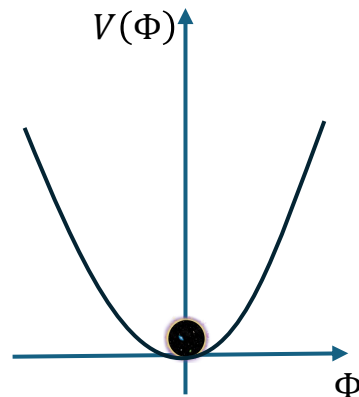
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SM explains mass generation : EWPT

- A single scalar SU(2) doublet $\Phi = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix}$,
- With $\mu^2 < 0, \lambda > 0$,

$$V(\Phi) = -\mu^2 (\Phi^\dagger \Phi) + \frac{\lambda}{4} (\Phi^\dagger \Phi)^2$$

BEH mechanism



**EW symmetry spontaneously broken to
 $SU(2)_L \times U(1)_Y \rightarrow U(1)_{EM}$**

- Mass to particles (Fermions and Gauge Bosons) and scalar Higgs boson

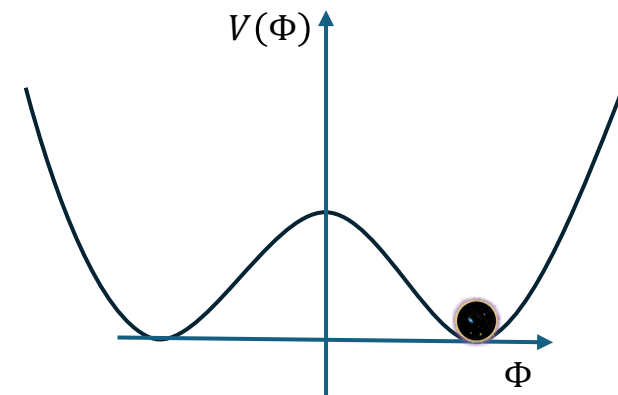


Fig 1. Smooth transition to the EW minimum

Thermal Evolution of the Universe

At finite-temperature:

$$V_{eff} = \underbrace{V_{tree} + V_{CW} + V_{CT}}_{T\text{-independent}} + \underbrace{V_T + V_{daisy}}_{\text{thermal corrections}}$$

V_{tree} : Tree-level scalar potential

V_{CW} : One-loop Coleman Weinberg potential (Ref.2)

V_{CT} : UV-finite counter-term potential

V_T : One-loop thermal corrections

V_{daisy} : Resummation of the daisy diagrams

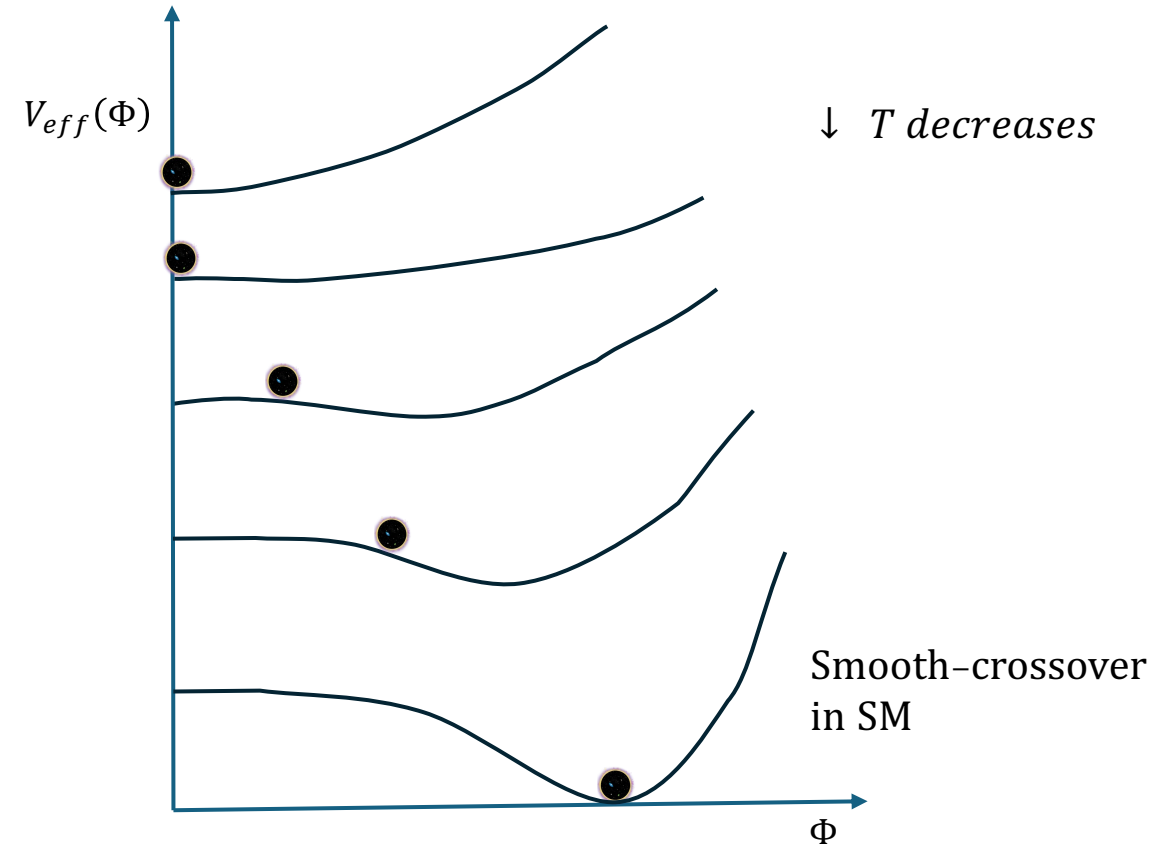


Fig 2.1: Finite-temperature potential: thermal evolution of vacuum as a smooth cross-over

First-order Phase Transition

- At very high temperatures:
EW symmetry preserved at $v = 0$
- Temperature decreased :
 - > Universe cooled -> develops non-zero vev
 - > EW symmetry breaks at T_n -> sudden change of vev
 - > Nucleation of vacuum bubbles

Strength of transition: $\frac{v_n}{T_n} \geq 1$

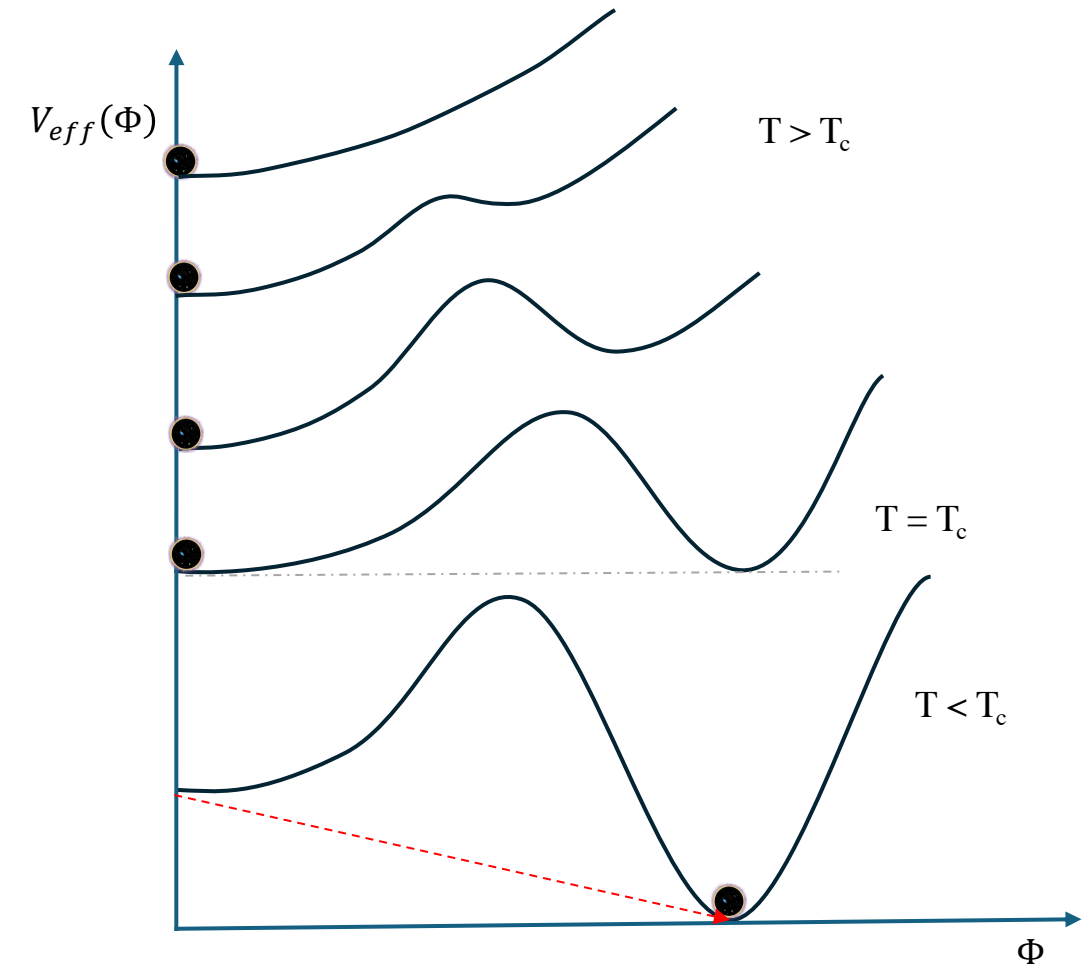
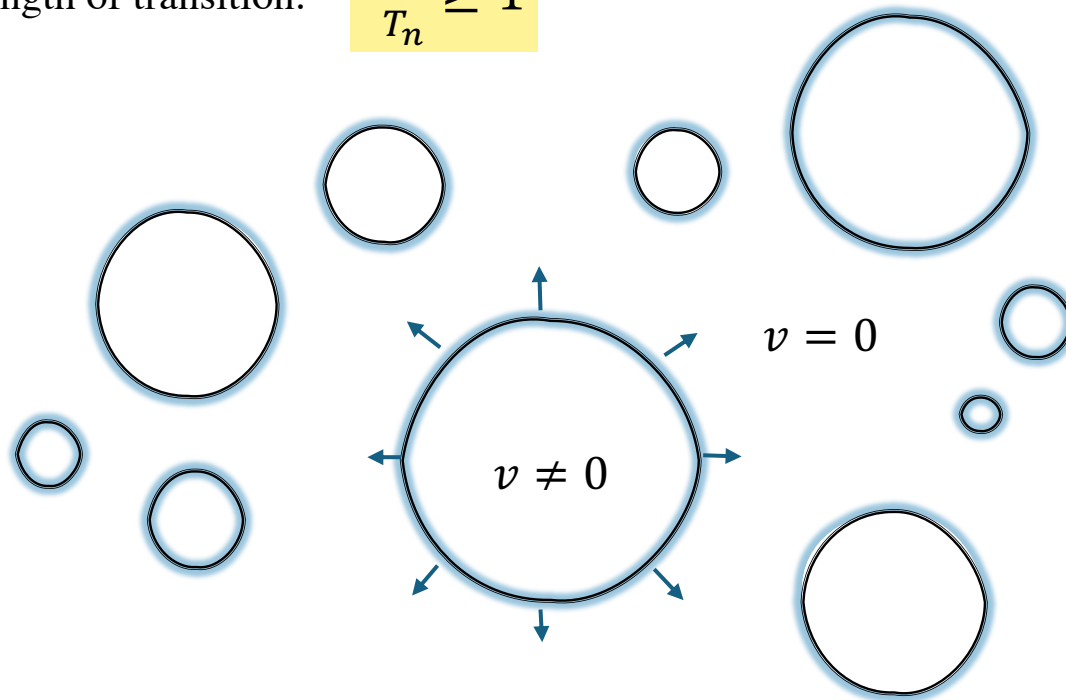


Fig 2.2: First-order phase transition of the universe

What are we looking for?



❑ Matter-Antimatter asymmetry

-> *SM prediction*: $6 \cdot 10^{-19}$; Observed: $6 \cdot 10^{-10}$

-> Sakharov conditions

(*B, C and CP violation + Out-of-Thermal-Equilibrium Process*)

❑ EW baryogenesis -> Strong first-order phase transition

✓ *BSM (extended) Higgs sector required!*

Two-Higgs-Doublet Model (2HDM)

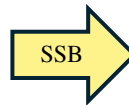
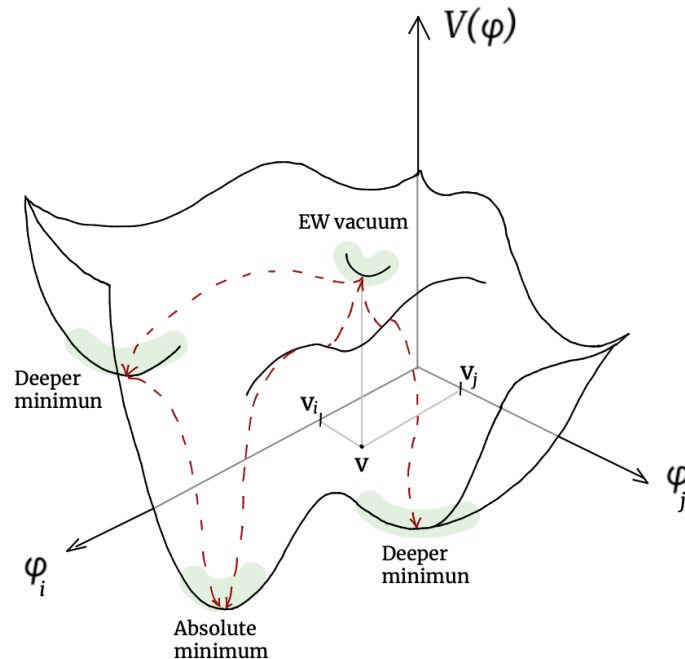
Z_2 symmetry:

$$\begin{aligned}\Phi_1 &\rightarrow \Phi_1 \\ \Phi_2 &\rightarrow -\Phi_2\end{aligned}$$

I, II, III, IV



$$\begin{aligned}V = & m_{11}^2(\Phi_1^\dagger\Phi_1) + m_{22}^2(\Phi_2^\dagger\Phi_2) - m_{12}^2(\Phi_1^\dagger\Phi_2 + \Phi_2^\dagger\Phi_1) + \frac{\lambda_1}{2}(\Phi_1^\dagger\Phi_1)^2 + \frac{\lambda_2}{2}(\Phi_2^\dagger\Phi_2)^2 \\ & + \lambda_3(\Phi_1^\dagger\Phi_1)(\Phi_2^\dagger\Phi_2) + \lambda_4(\Phi_1^\dagger\Phi_2)(\Phi_2^\dagger\Phi_1) + \frac{\lambda_5}{2}[(\Phi_1^\dagger\Phi_2)^2 + (\Phi_2^\dagger\Phi_1)^2].\end{aligned}$$

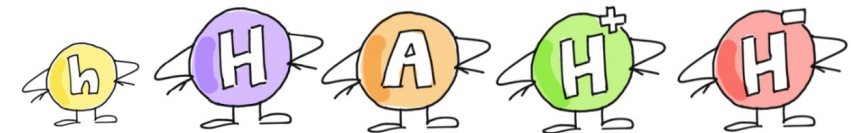


$$\langle \Phi_1 \rangle = \begin{pmatrix} 0 \\ \frac{v_1}{\sqrt{2}} \end{pmatrix}$$

$$\langle \Phi_2 \rangle = \begin{pmatrix} 0 \\ \frac{v_2}{\sqrt{2}} \end{pmatrix}$$

In mass basis:

- 5 physical scalar fields (h, H, A, H^\pm)
diagonalised by α and β ,
- 3 would-be-Goldstone bosons (G^0, G^\pm)
- Non-trivial structure of the scalar potential



Art credit: Kateryna Radchenko

Fig 3: Representation of minima structure of 2HDM (or BSM) scalar potential

Constraints on 2HDM Parameter Space

Free parameters:

$\tan \beta, \cos (\beta - \alpha), v, m_h, m_H, m_{H^\pm}, m_A, m_{12}^2$

Enters in self-couplings \rightarrow barrier shape

Alignment limit:

- $\cos (\beta - \alpha) \rightarrow 0$
- Light CP-even Higgs h has SM-like couplings of h_{125} at the lowest order
- Possible in both decoupling and non-decoupling regime

Theoretical constraints:

- Vacuum stability and boundedness-from-below
- Perturbative unitarity

Experimental constraints:

- Electroweak precision observables
- Constraints are from the limits from searches for additional Higgs bosons
- Measurements of the properties of h_{125}
- Flavour Constraints

Mass splittings

Key Regions:

1. $m_A = m_{H^\pm}$
2. $m_H = m_{H^\pm}$

- Larger mass splittings $m_A - m_H$ driven by the quartic couplings in the scalar potential correlate to stronger first-order phase transitions.
- For region 1: mass splitting range: 100 – 200 GeV
- For region 2: mass splitting range: 150 – 400 GeV
- This dependence is constrained by perturbative unitarity.

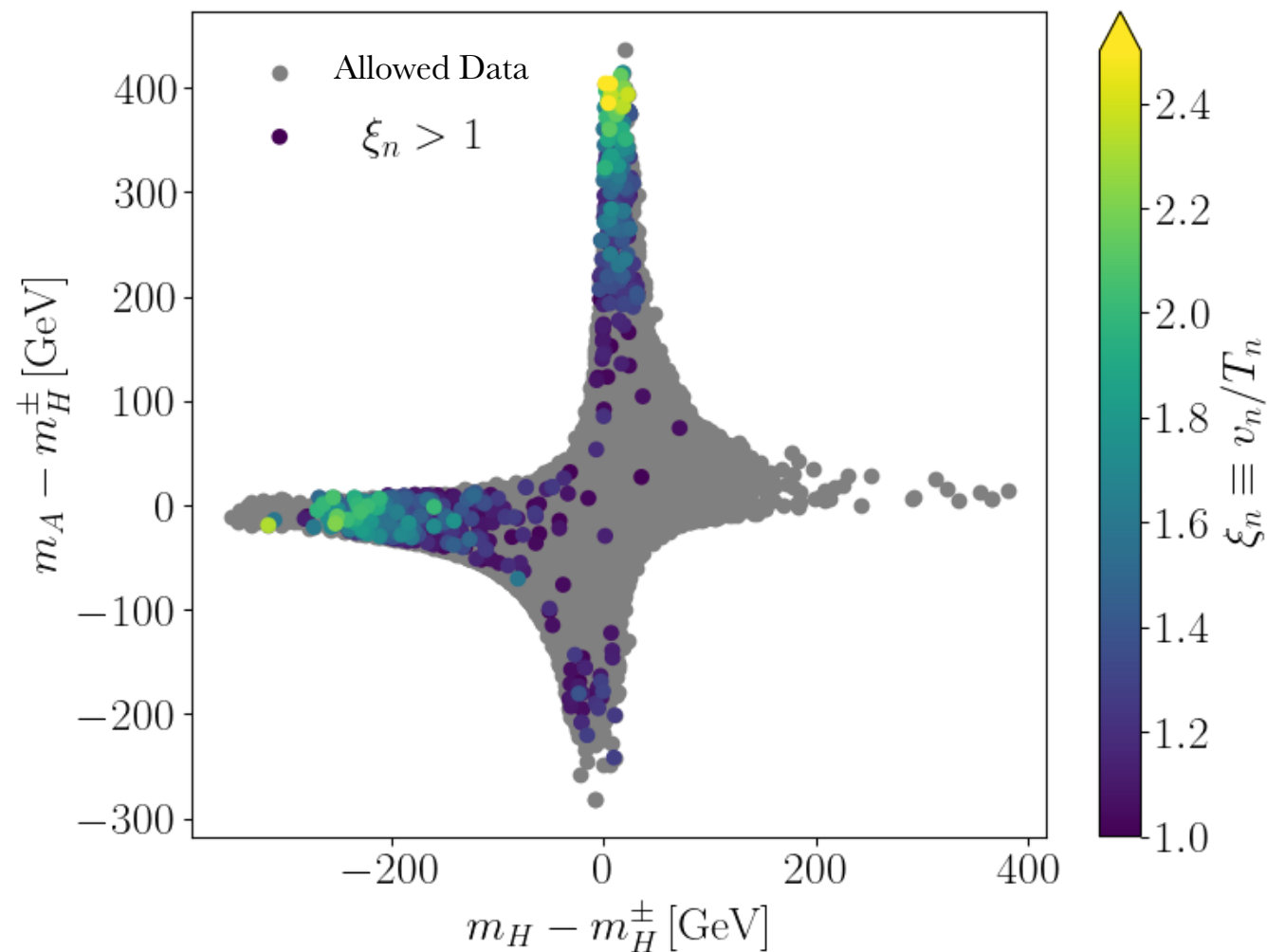


Fig 4: SFOEWPT strength depends on the mass splittings, here the free parameters vary as $1 \leq \tan \beta \leq 50$ and $150 \leq m_H \leq 1400$ GeV.

SFOEWPT in (m_H , $\cos(\beta-\alpha)$) plane

For low mass values m_H around 200 GeV:

- SFOEWPT occurs for wide range of $\cos(\beta-\alpha)$ values
- Both CP-even neutral Higgs bosons h and H take part in the EW phase transition \rightarrow enhanced thermal corrections \rightarrow relatively higher strengths

For higher mass values m_H :

- SFOEWPT occurs for $\cos(\beta-\alpha) \rightarrow 0$, i.e. occurs close to the alignment limit $\rightarrow h$ drives the phase transition

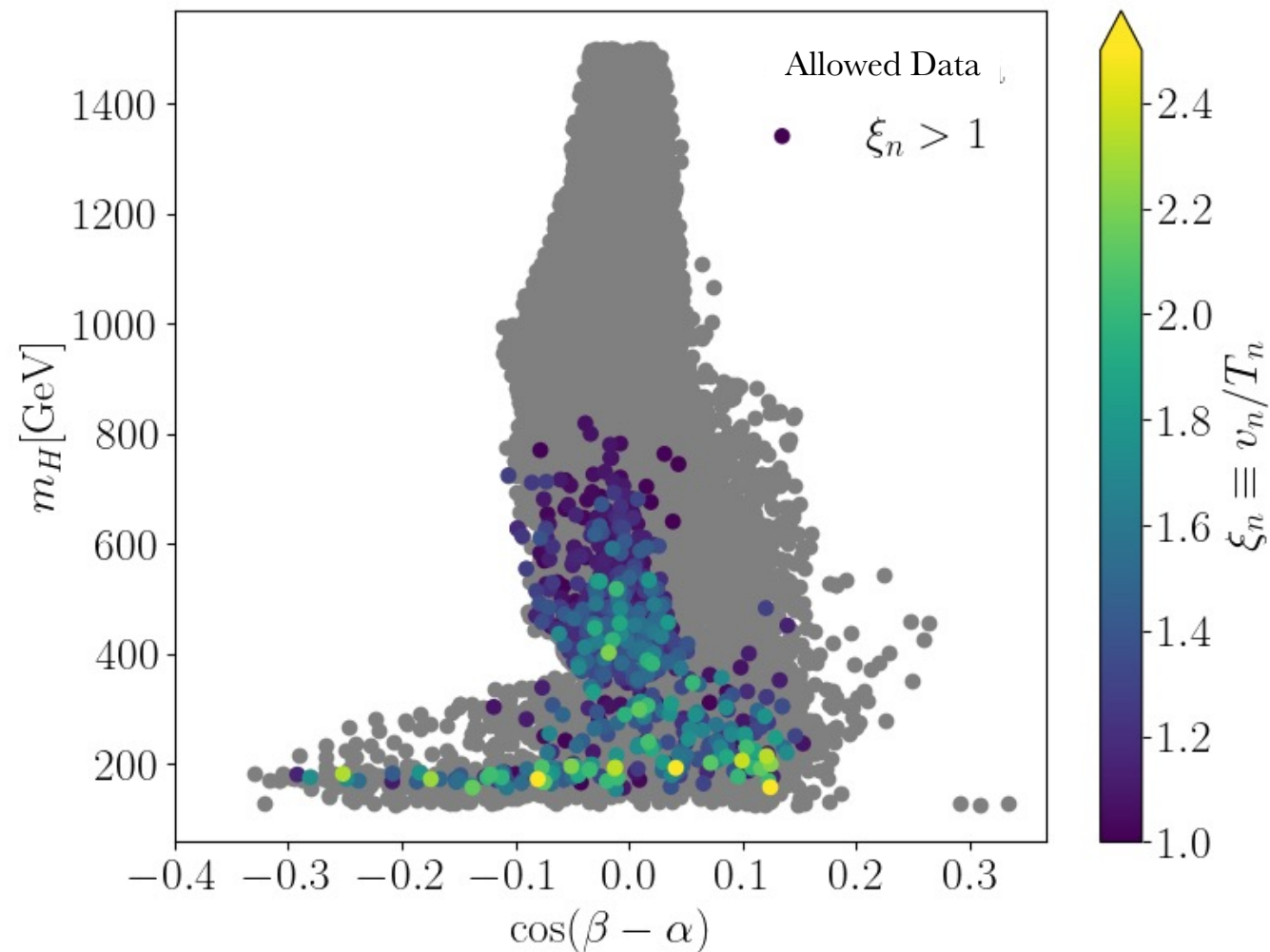


Fig 5: SFOEWPT for CP-even heavy Higgs mass values, here the free parameters vary as $1 \leq \tan \beta \leq 50$

SFOEWPT vs $\cos(\beta-\alpha)$ for lower m_H

SFOEWPT occurs for wide range of $\cos(\beta-\alpha)$, strength higher far from the alignment limit

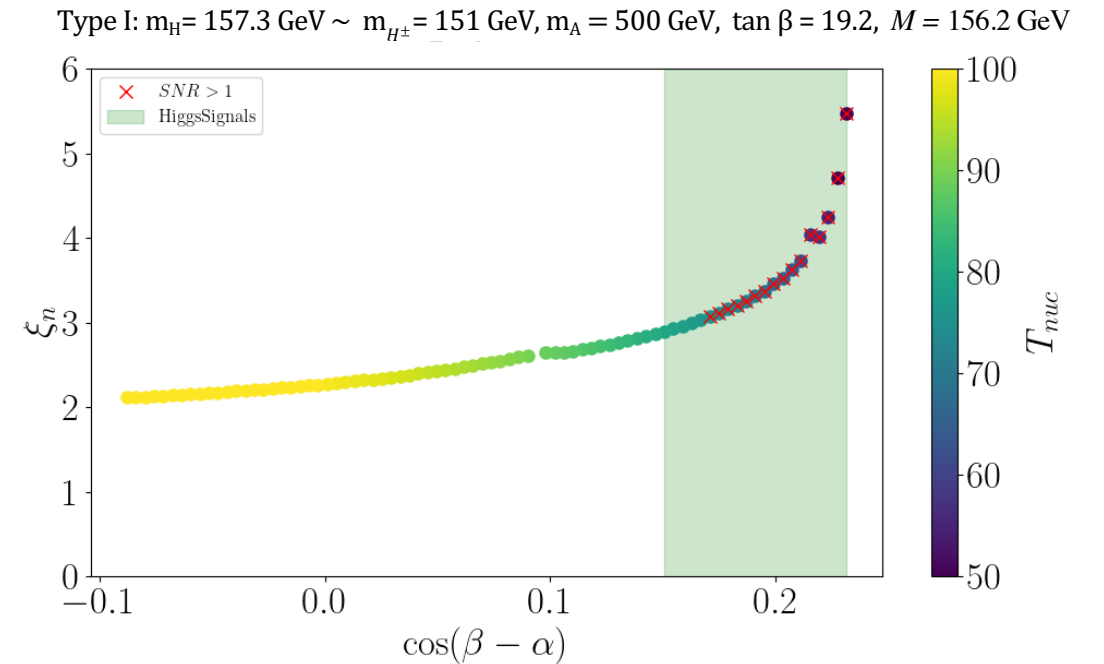
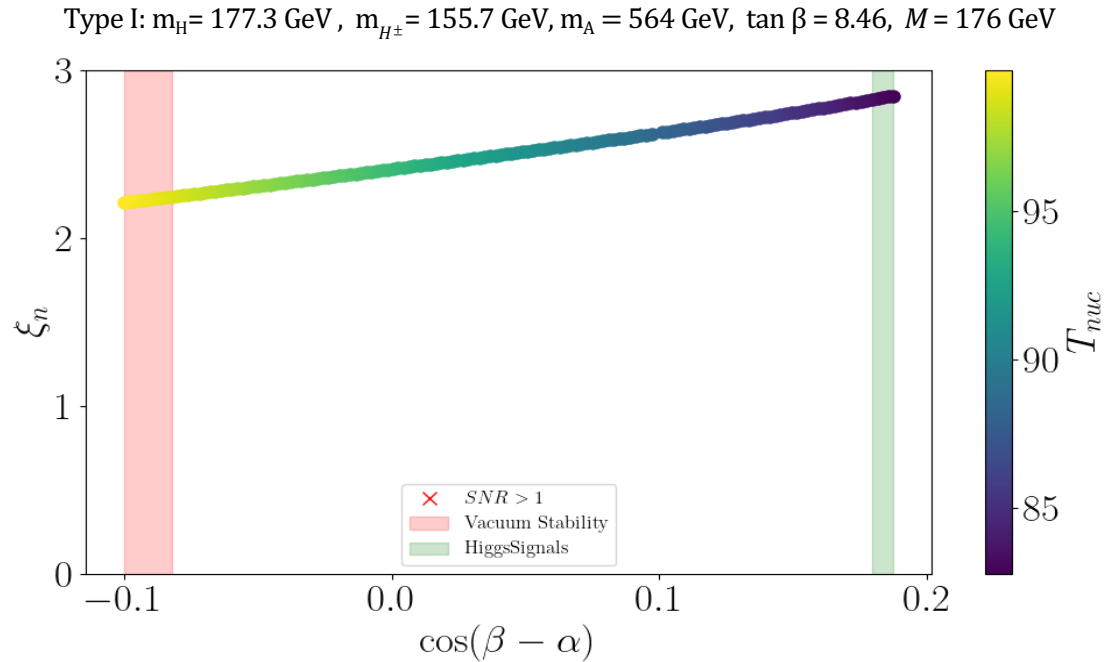


Fig 6a and b: Variation of strength with $\cos(\beta-\alpha)$ for lower CP-even Higgs masses to correlate the SFOEWPT parameter space close to as well as away from alignment limit

SFOEWPT vs $\cos(\beta - \alpha)$ for higher m_H

SFOEWPT occurs strongest for $\cos(\beta - \alpha) \rightarrow 0$, i.e. highest strength at the alignment limit

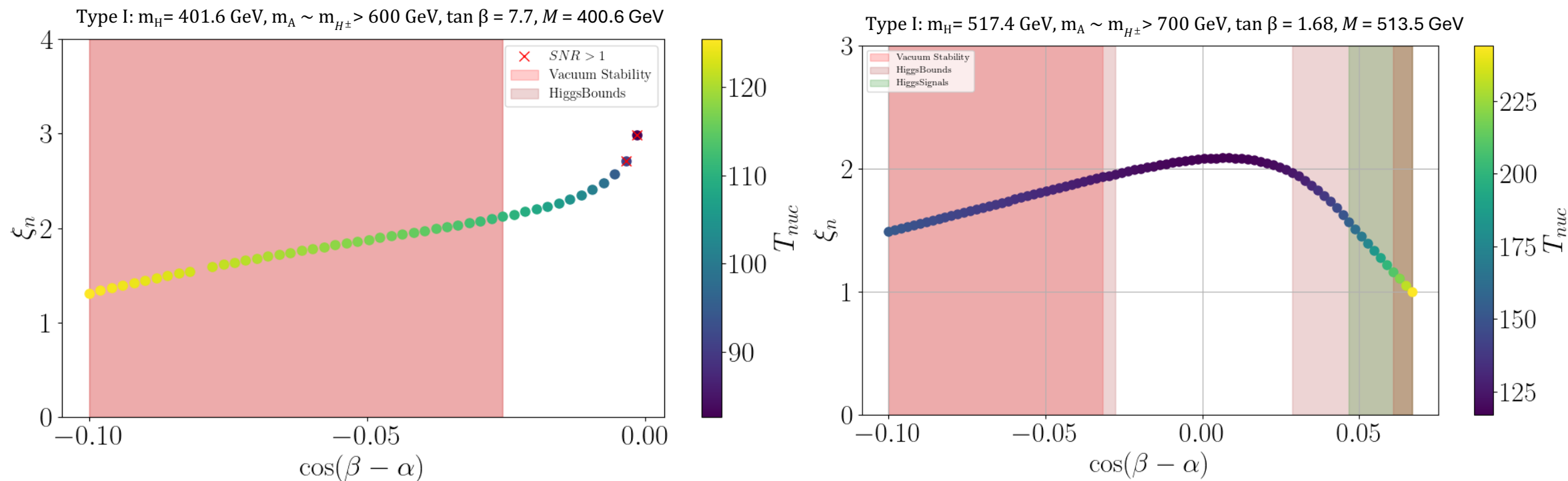


Fig 7a and b: Variation of strength with $\cos(\beta - \alpha)$ for higher CP-even Higgs masses to correlate the SFOEWPT parameter space close to alignment limit

Conclusions

- Extended Higgs sectors can feature a SFOEWPT, necessary for EW baryogenesis, can potentially lead to detectable GW signals
- The parameter region giving rise to a SFOEWPT in the 2HDM has a mass splitting between m_A and m_H of about 200 GeV
- For high values of m_H (up to 800 GeV): SFOEWPT occurs close to the alignment limit
- For low values of m_H (~ 200 GeV): SFOEWPT possible for larger deviations from the alignment limit
- Qualitatively, for SFOEWPT parameter points with higher m_H , the strength of phase transition increases with $\cos(\beta-\alpha)$, and is highest near the alignment limit, whereas for the case of lower m_H , the strength is higher as we go away from the alignment limit.

Next Steps:

- Extend the current results for the types II, III and IV of 2HDM
- Check for κ_λ variation for the above parameter regions showing SFOEWPT in correlation to the alignment limit
- Investigate the prospects for possible GW signals

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