

DESY FH SEMINAR

OCTOBER 24TH 2025

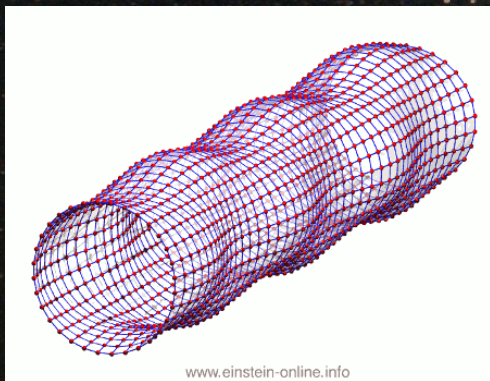
EINSTEIN TELESCOPE - THE 3RD GENERATION OF LASER INTERFEROMETRIC GRAVITATIONAL WAVE DETECTION IN EUROPE

O. GERBERDING



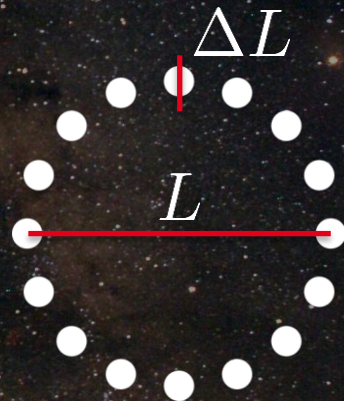
STRAIN FROM A BINARY

$$h = \frac{\Delta L}{L} \propto \frac{1}{r} m^{5/3} f^{2/3} \quad \bullet \quad \bullet$$



GRAVITATIONAL WAVES

- CONSEQUENCE OF GENERAL RELATIVITY
- SOURCED BY AN ACCELERATED QUADRUPOLE MOMENT
- PROPAGATE WITH THE SPEED OF LIGHT
- TWO POLARISATIONS: + & x
- STRETCH & COMPRESS SPACE-TIME TRANSVERSALLY





STRAIN FROM A BINARY

$$h = \frac{\Delta L}{L} \propto \frac{1}{r} m^{5/3} f^{2/3}$$



HULSE-TAYLOR BINARY-DISCOVERY 1974
ORBITAL DECAY FITS GRAVITATIONAL WAVES
NOBEL PRIZE 1993

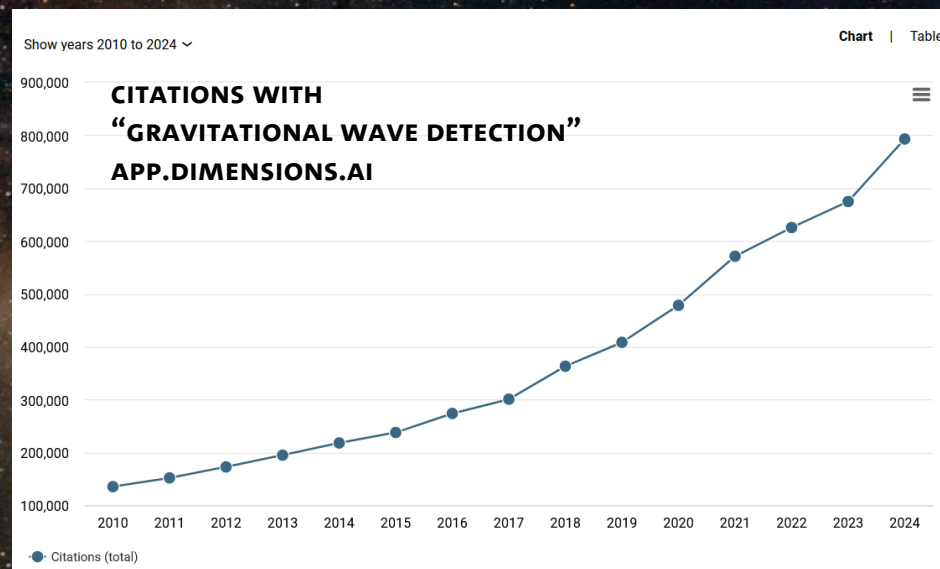
GRAVITATIONAL WAVE ASTRONOMY



2017 Nobel Prize in Physics

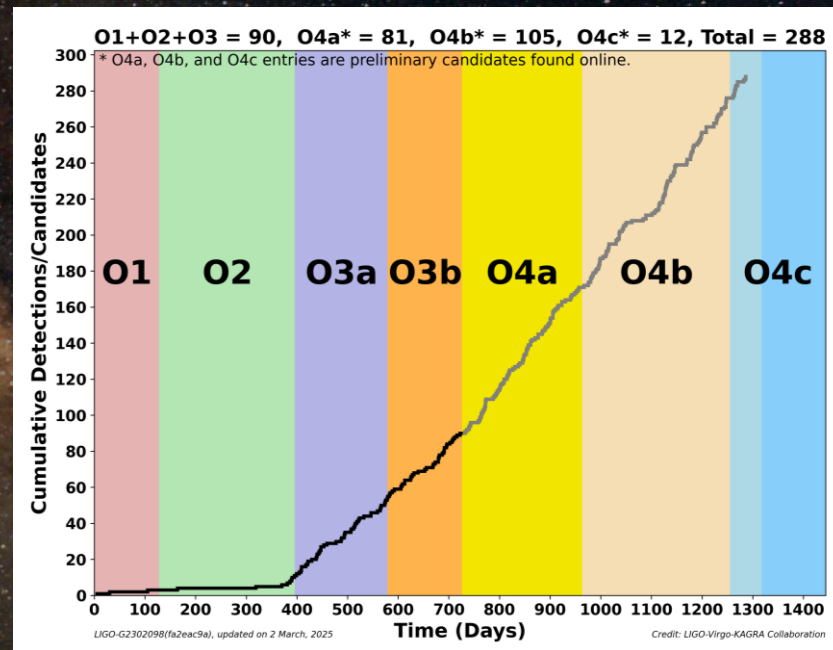
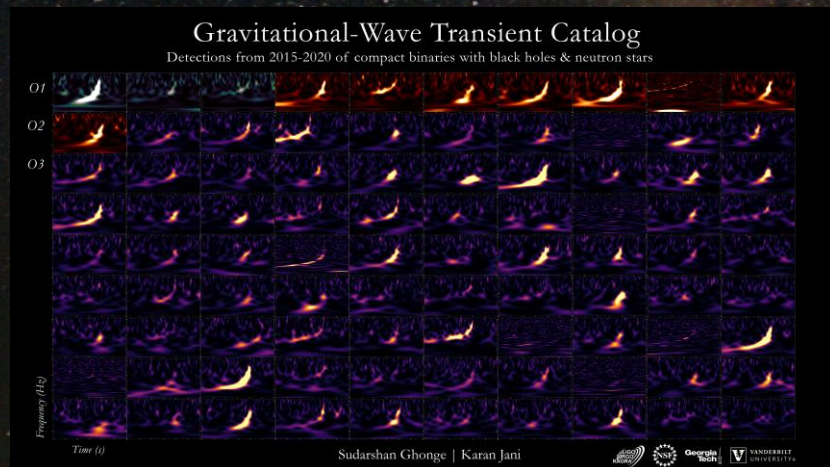
FIRST DIRECT DETECTION BY LIGO IN 2015
LASER INTERFEROMETRY MEASURES STRAIN
NOBEL PRIZE 2017

A NEW ERA IN GRAVITATIONAL PHYSICS



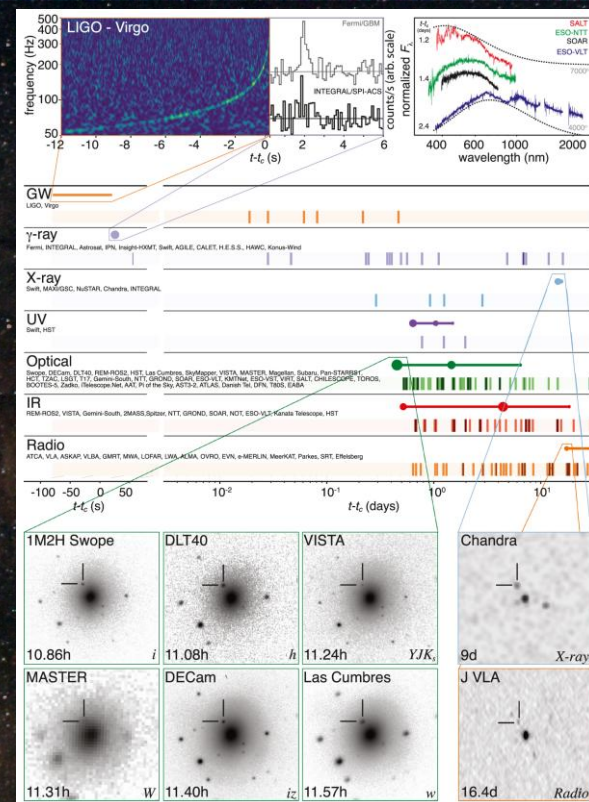
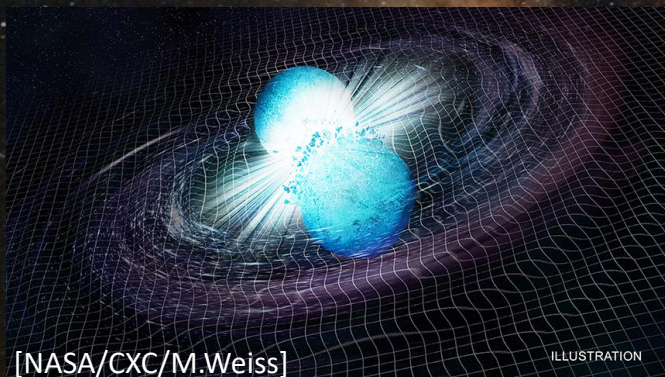
A NEW ERA IN GRAVITATIONAL PHYSICS

- Almost **daily** gravitational wave BBH, BNS & BHNS detections during observing runs



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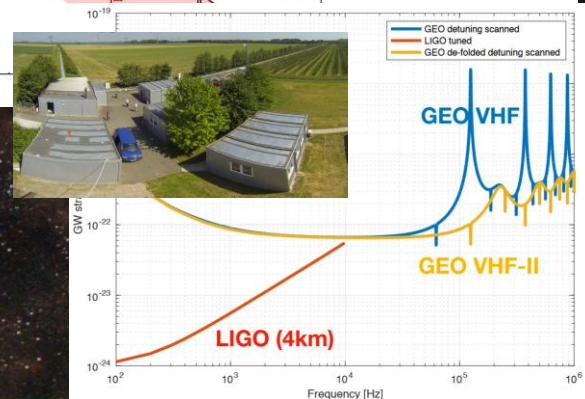
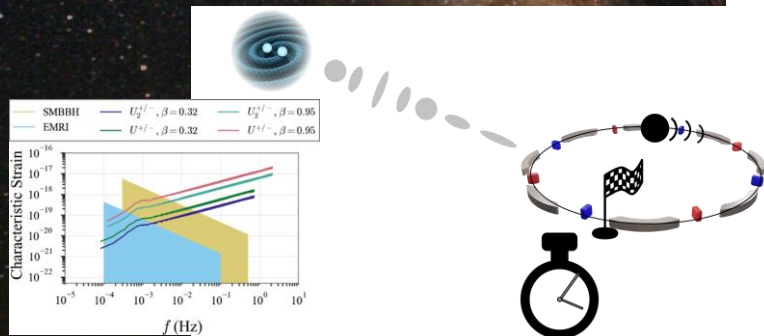
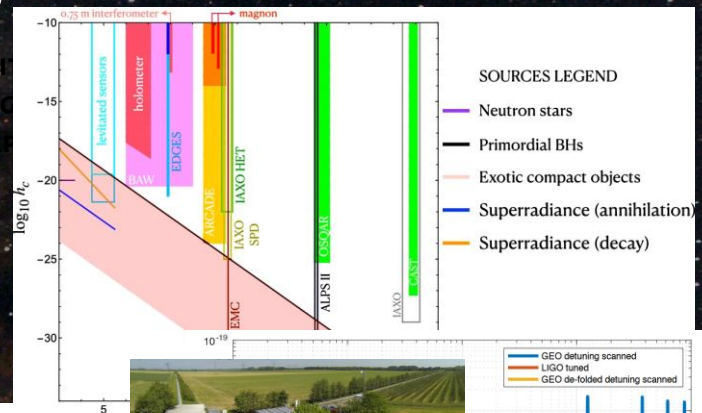
- Almost **daily** gravitational wave BBH, BNS & BHNS detections during observing runs
- Multi-messenger astronomy (GW170817, Kilonova)
- Precision tests of GR, black hole mass function, GW speed, Hubble constant via standard siren, mass-gap objects, Hawking area theorem



Aggarwal, Nancy, et al, *Living reviews in relativity* 24 (2021): 1-74.

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- New GWD concepts from mHz to THz



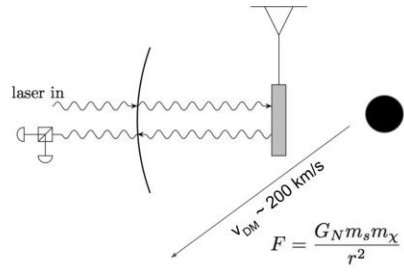
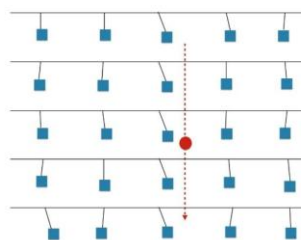
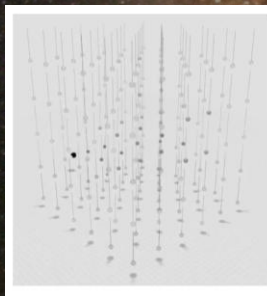
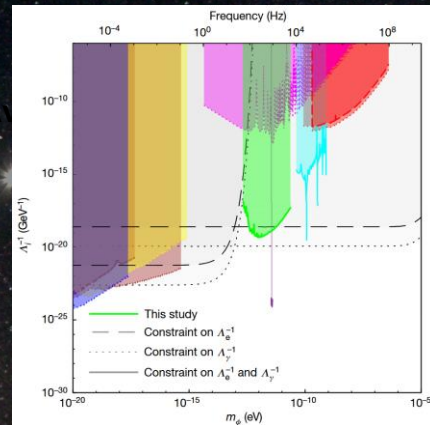
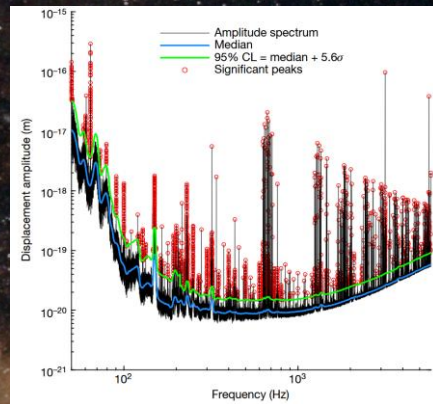
Schmirander, Thorben, et al. *Physical Review D* 110.8 (2024): 082002.

J. Hough et al., Talk at MG17, 2024, [Indico Link](#)

Vermeulen, Sander M., et al., Nature 600.7889 (2021): 424-428.

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- Gravitational dark matter detection

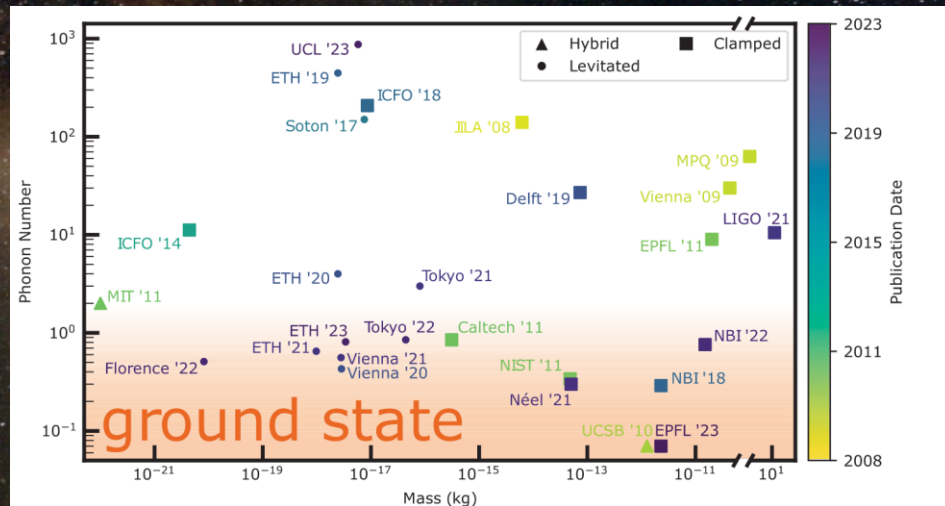
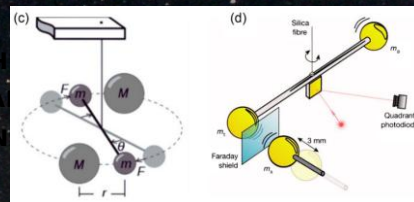


Attanasio, Alaina, et al. arXiv preprint arXiv:2203.07242 (2022).

A NEW ERA IN GRAVITATIONAL PHYSICS

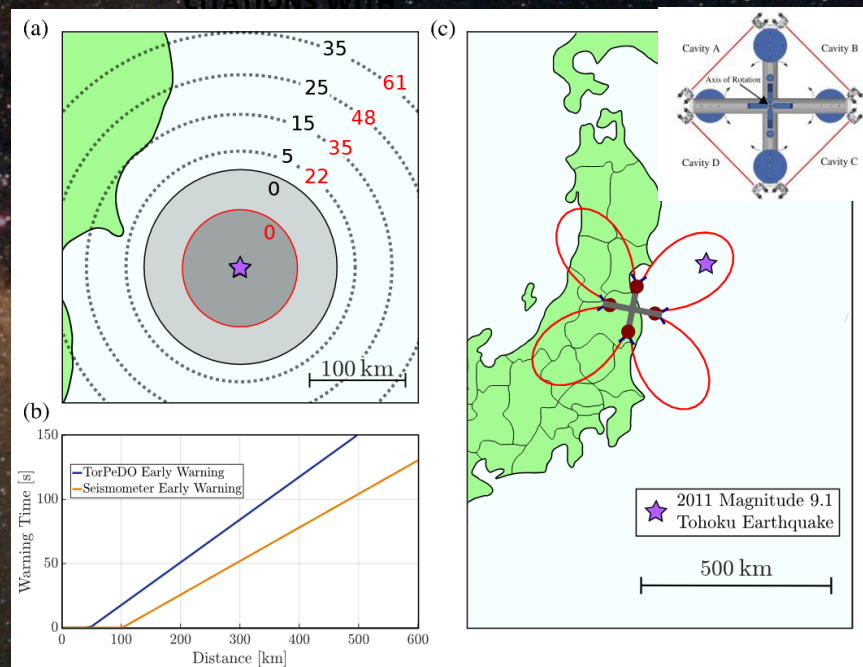
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- Table-top quantum-gravity experiments with macroscopic entangled masses

CITATIONS WITH
 "GRAVITATIONAL
 APP.DIMENSION"



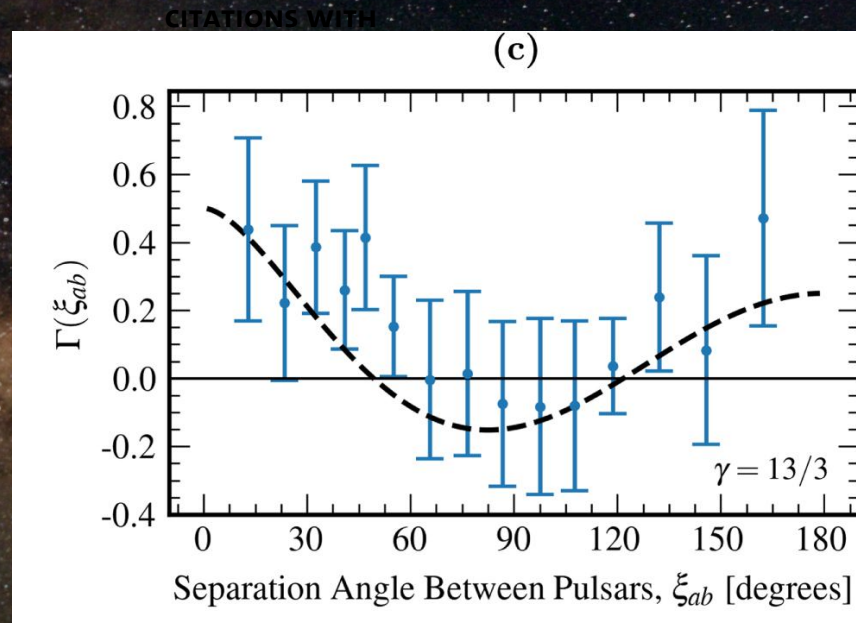
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- Table-top quantum-gravity experiments with macroscopic entangled masses
- Gravitational sensing of Earth



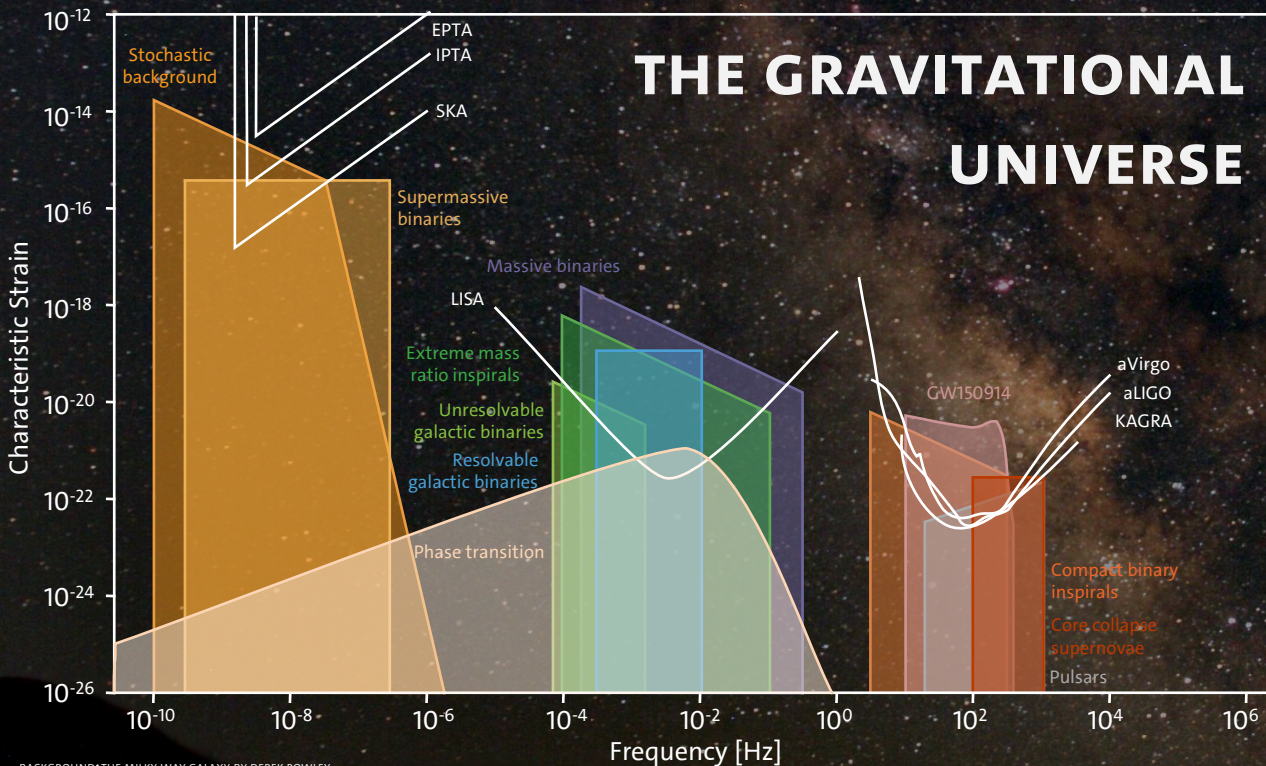
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- Table-top quantum-gravity experiments with macroscopic entangled masses
- Gravitational sensing of Earth
- Evidence for nano Hz GWs from Pulsar timing arrays



2ND GENERATION

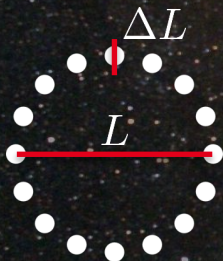
ADAPTED FROM GWPLOTTER.COM BY C. MOORE, R. COLE, AND C. BERRY.



GROUND BASED GRAVITATIONAL WAVE DETECTION VIA LASER INTERFEROMETERS

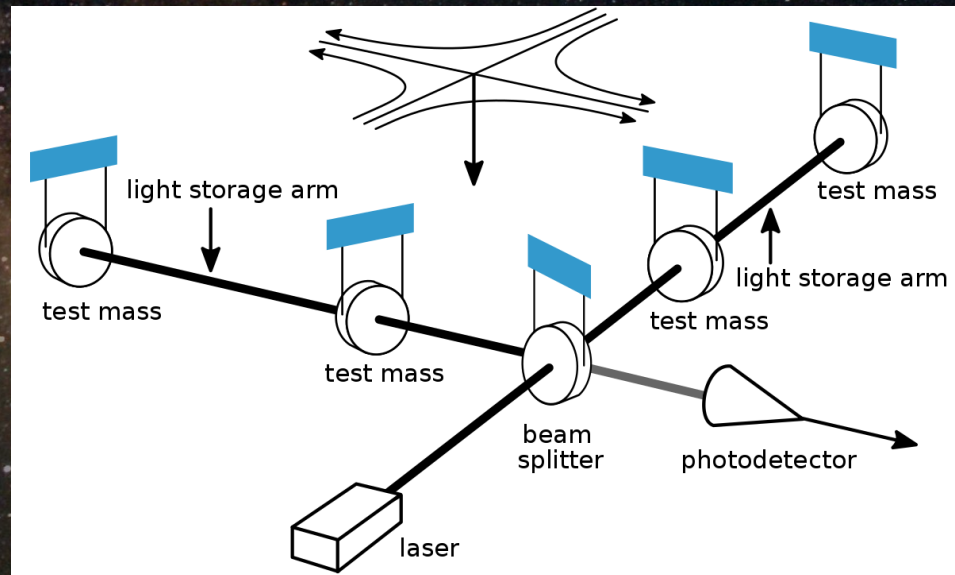
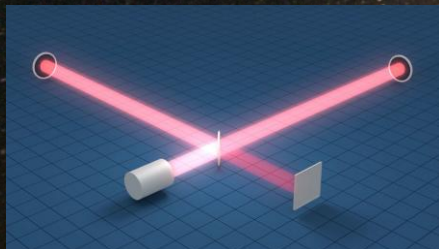
Free masses

$$h = \frac{\Delta L}{L}$$



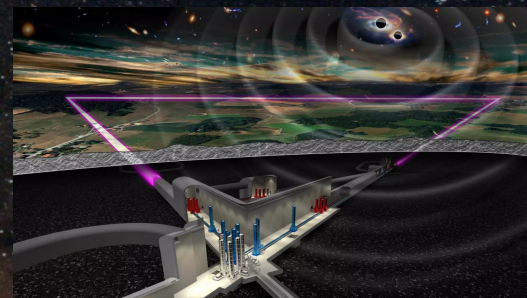
Laser displacement measurements

$$\Delta\varphi = \frac{2\pi\Delta L}{\lambda} \approx \Delta L \cdot 10^7$$

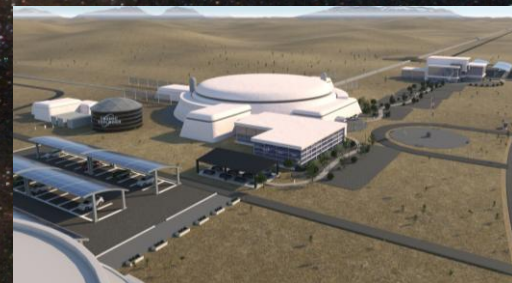


3RD GENERATION

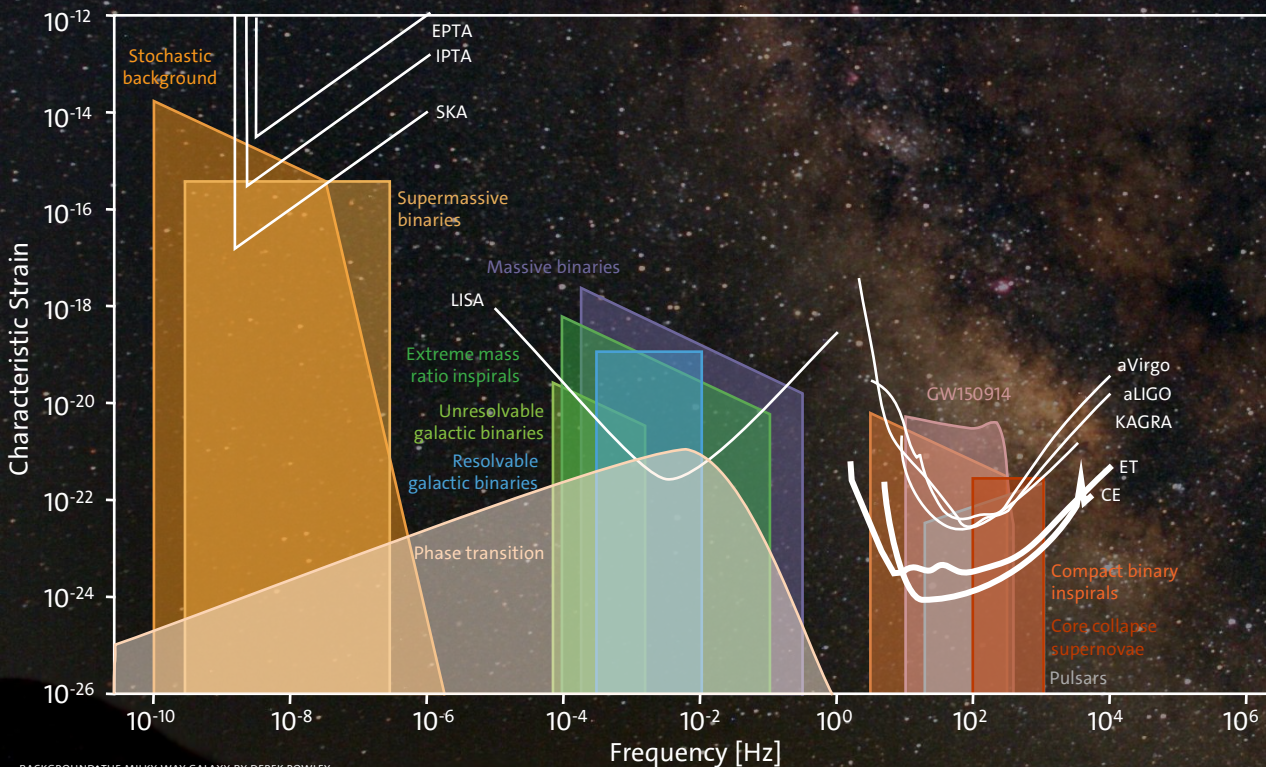
EINSTEIN TELESCOPE



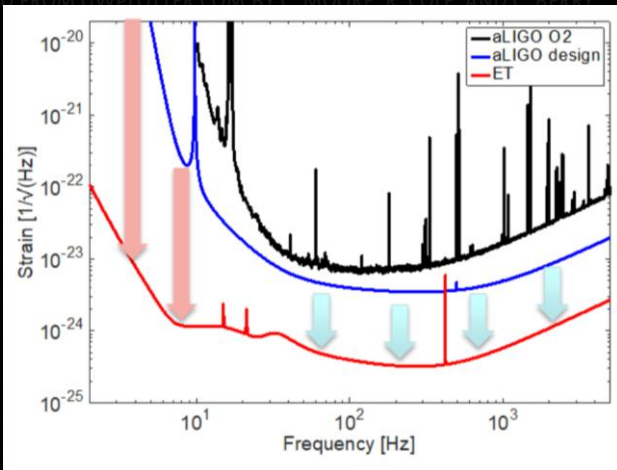
COSMIC EXPLORER



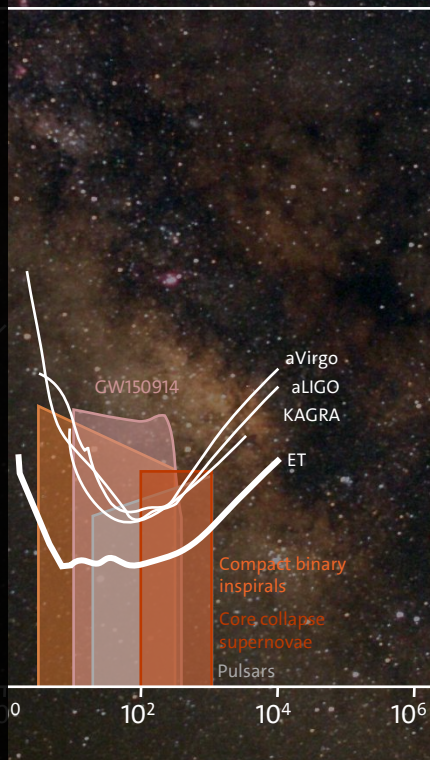
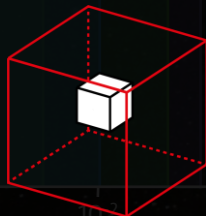
ADAPTED FROM GWPLOTTER.COM BY C. MOORE, R. COLE, AND C. BERRY.



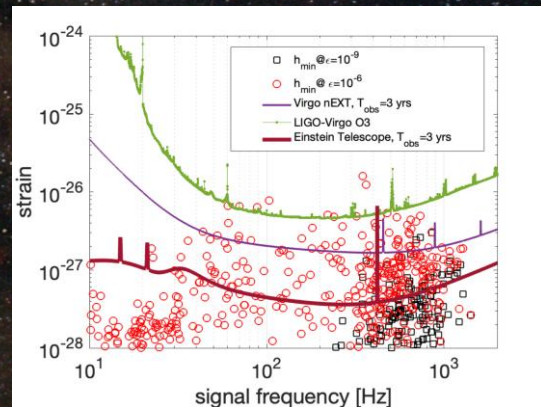
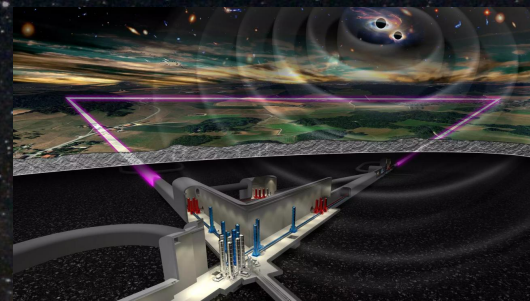
x10 broadband sensitivity increase



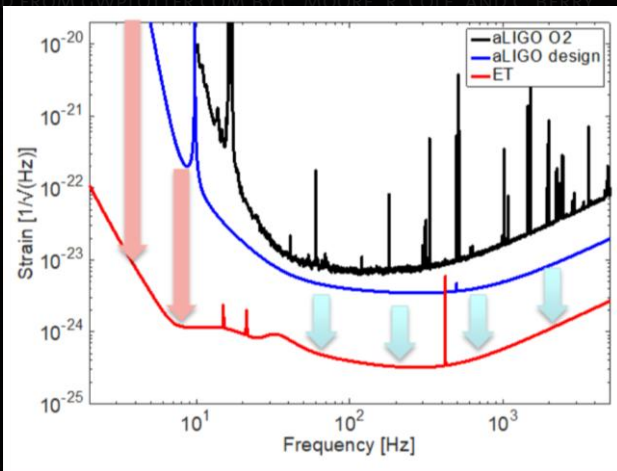
- x1000 more detections
- higher SNR (precision)
- Continuous GWs from known NSs



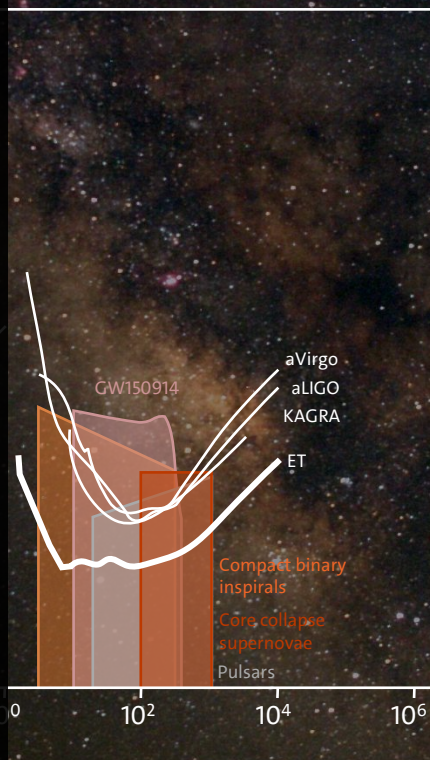
EINSTEIN TELESCOPE



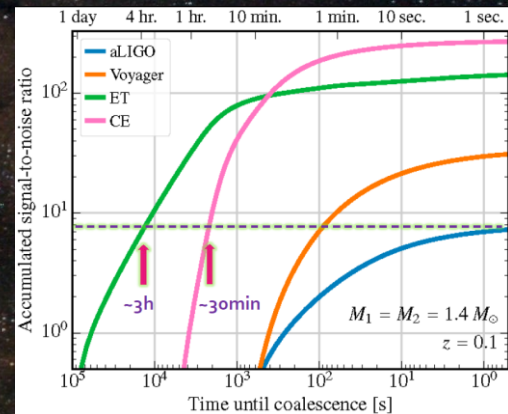
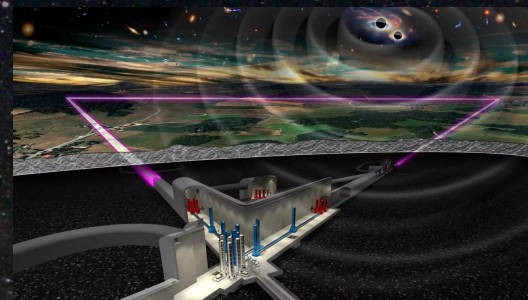
$\times 10^6$ noise reduction below 10 Hz



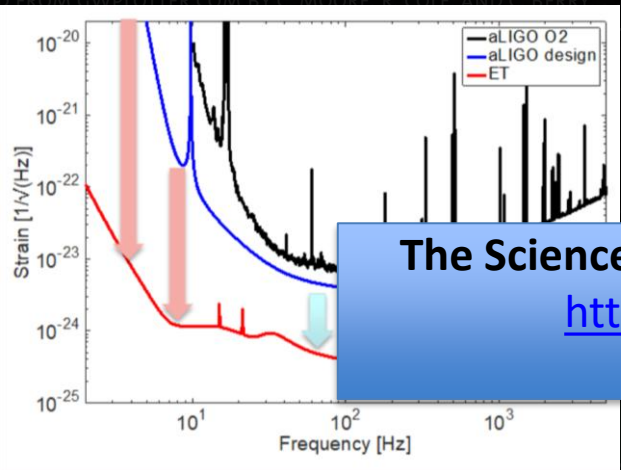
- Higher redshift detections
- Higher mass events
- Early warning for Multi-Messenger Astronomy



EINSTEIN TELESCOPE



$\times 10^6$ noise reduction below 10 Hz

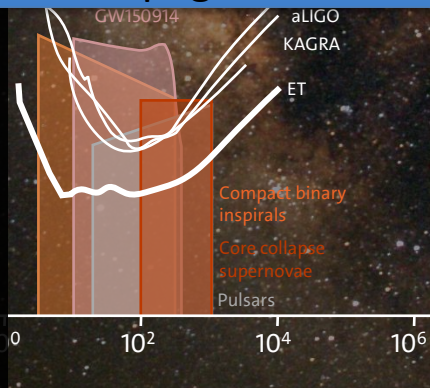


The Science of the Einstein Telescope – Blue book

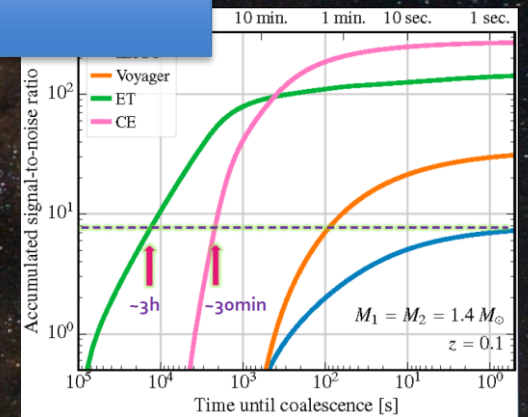
<https://arxiv.org/abs/2503.12263>

899 pages

- Higher redshift detections
- Higher mass events
- Early warning for Multi-Messenger Astronomy

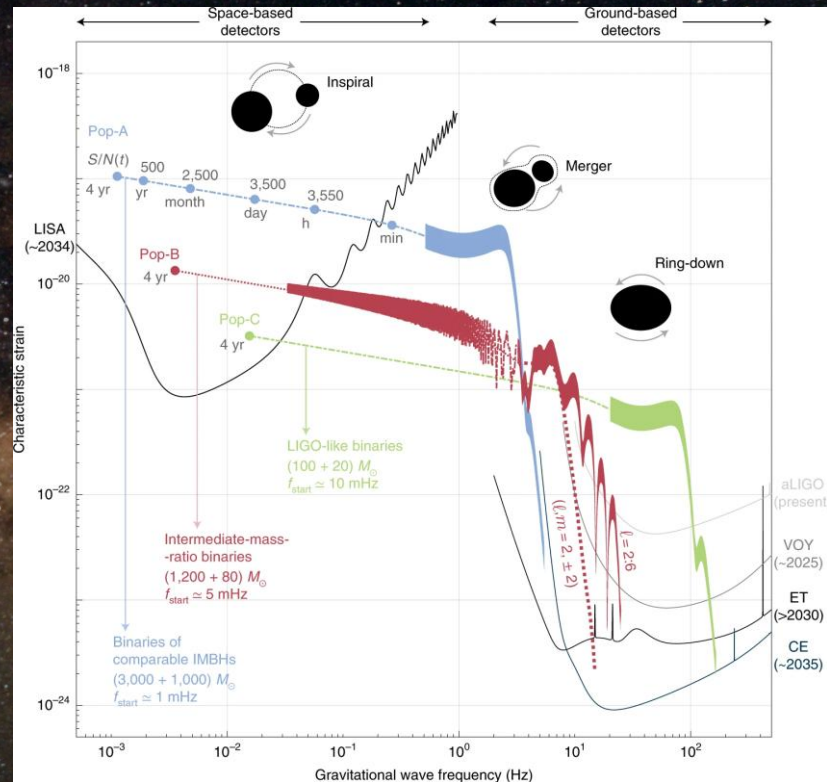
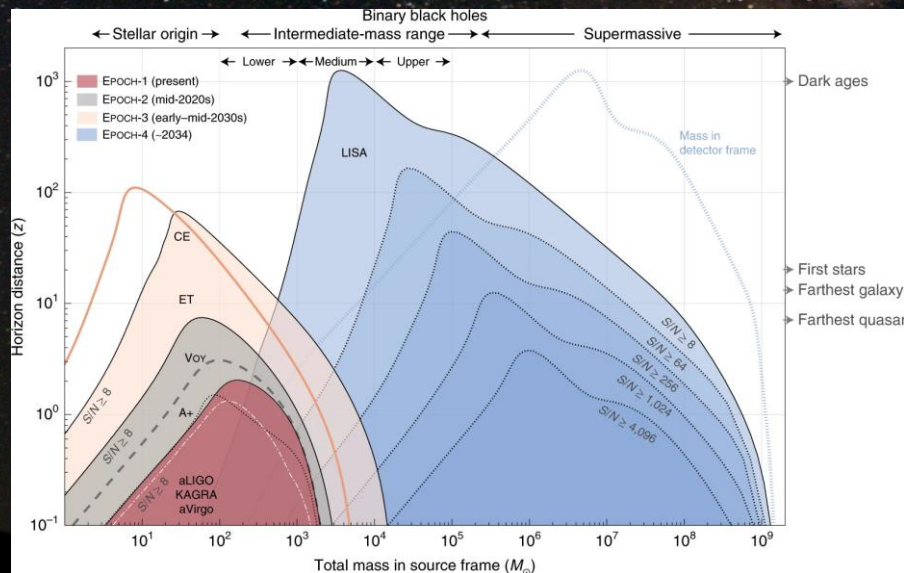


EINSTEIN TELESCOPE

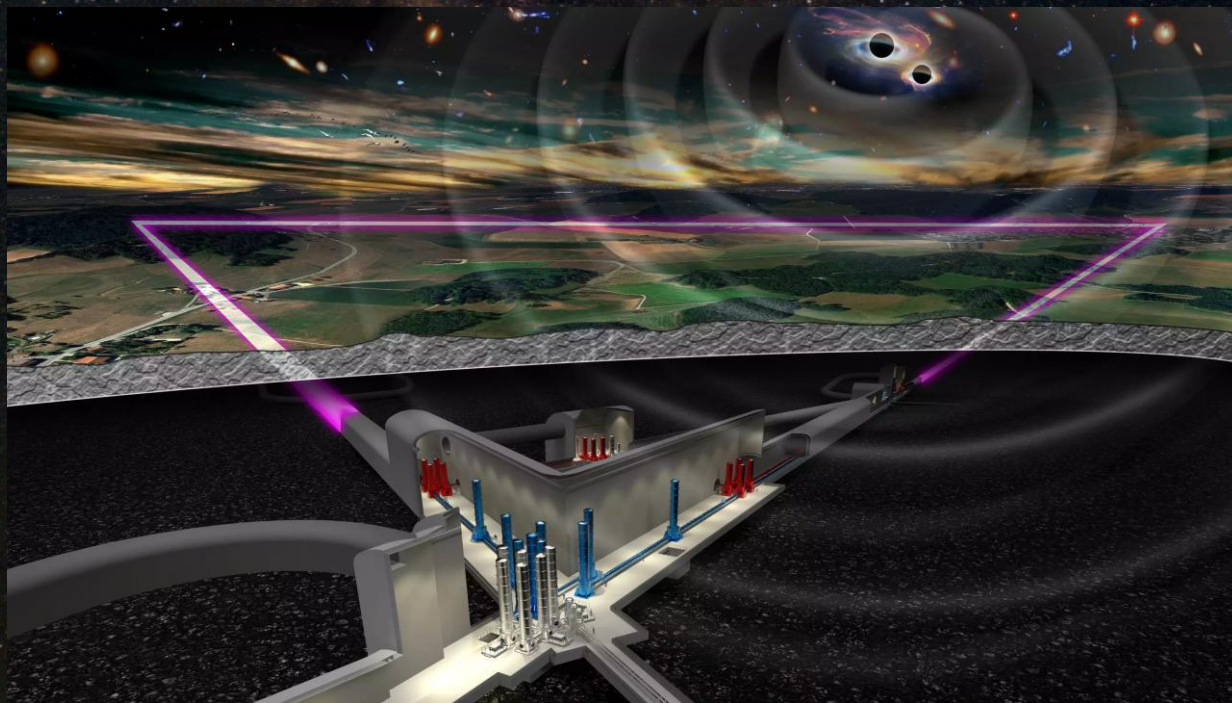


REACH & MULTI-BAND

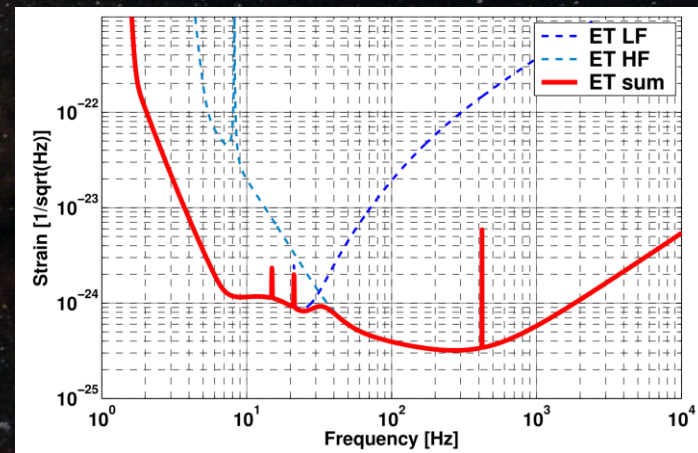
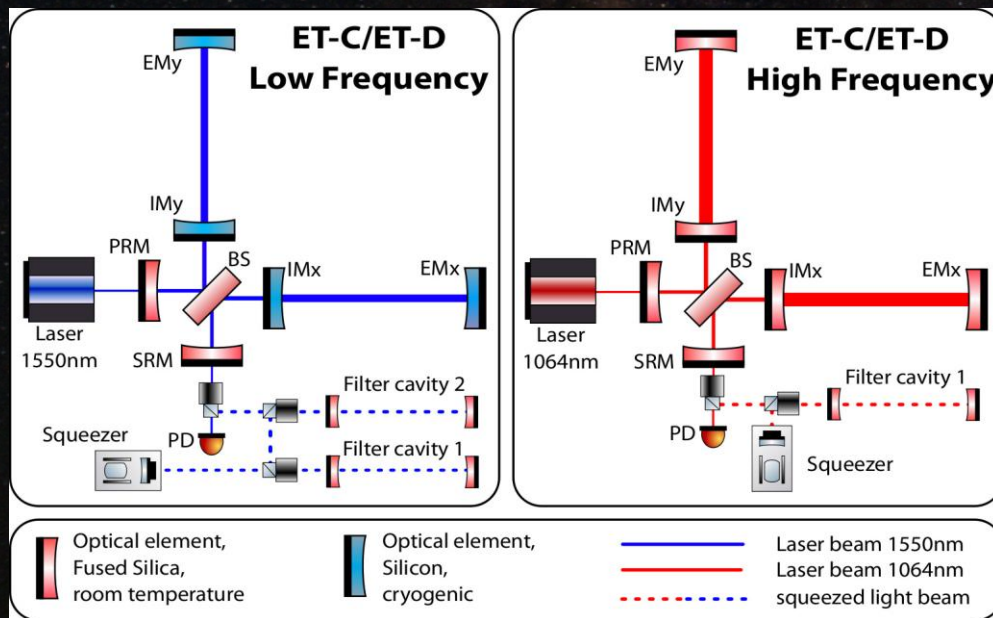
LISA + ET



THE EINSTEIN TELESCOPE

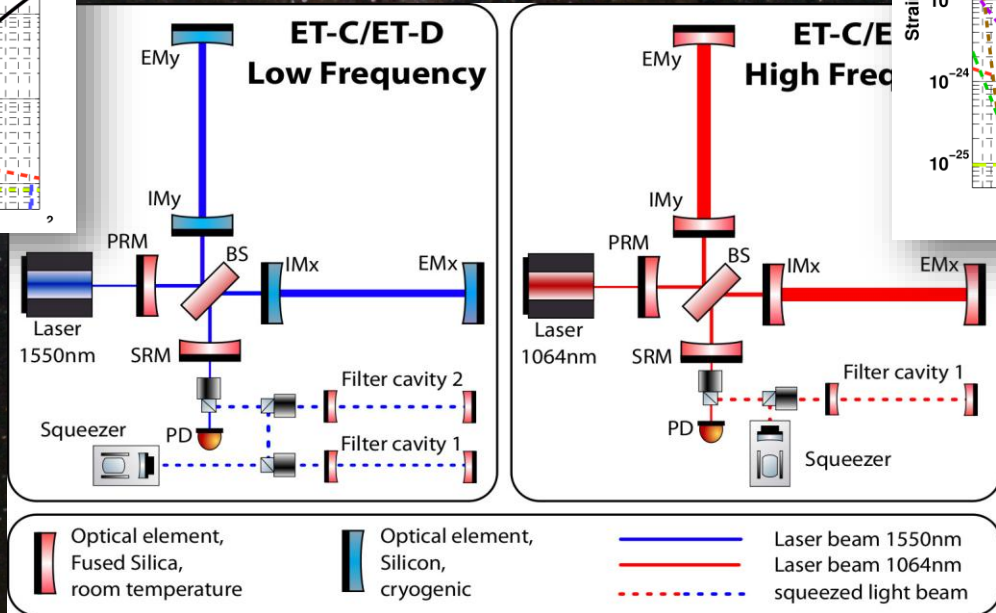
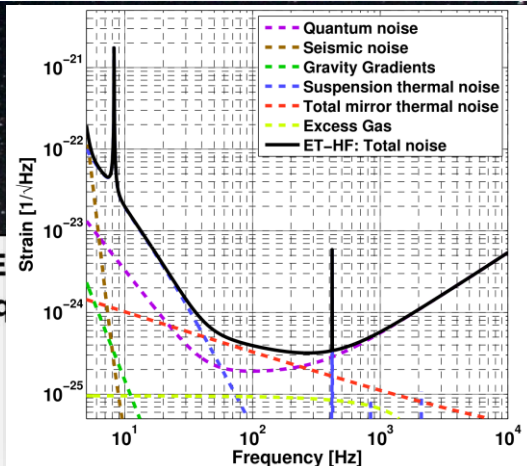
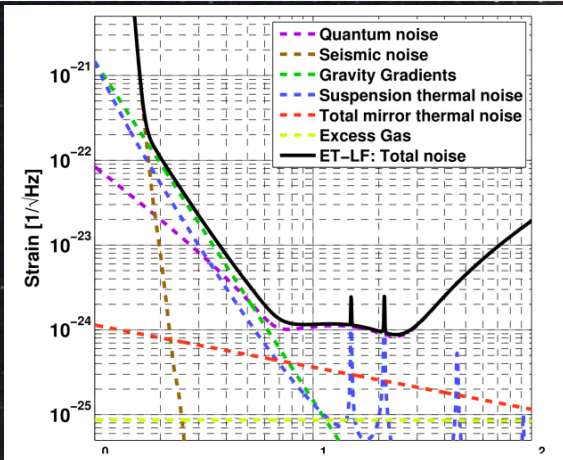


THE EINSTEIN TELESCOPE



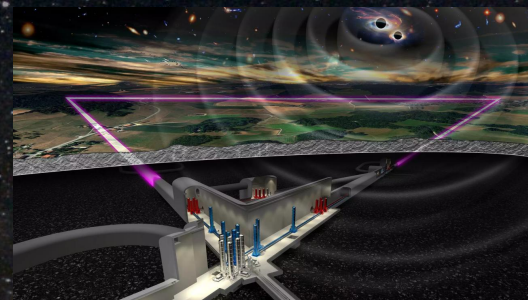
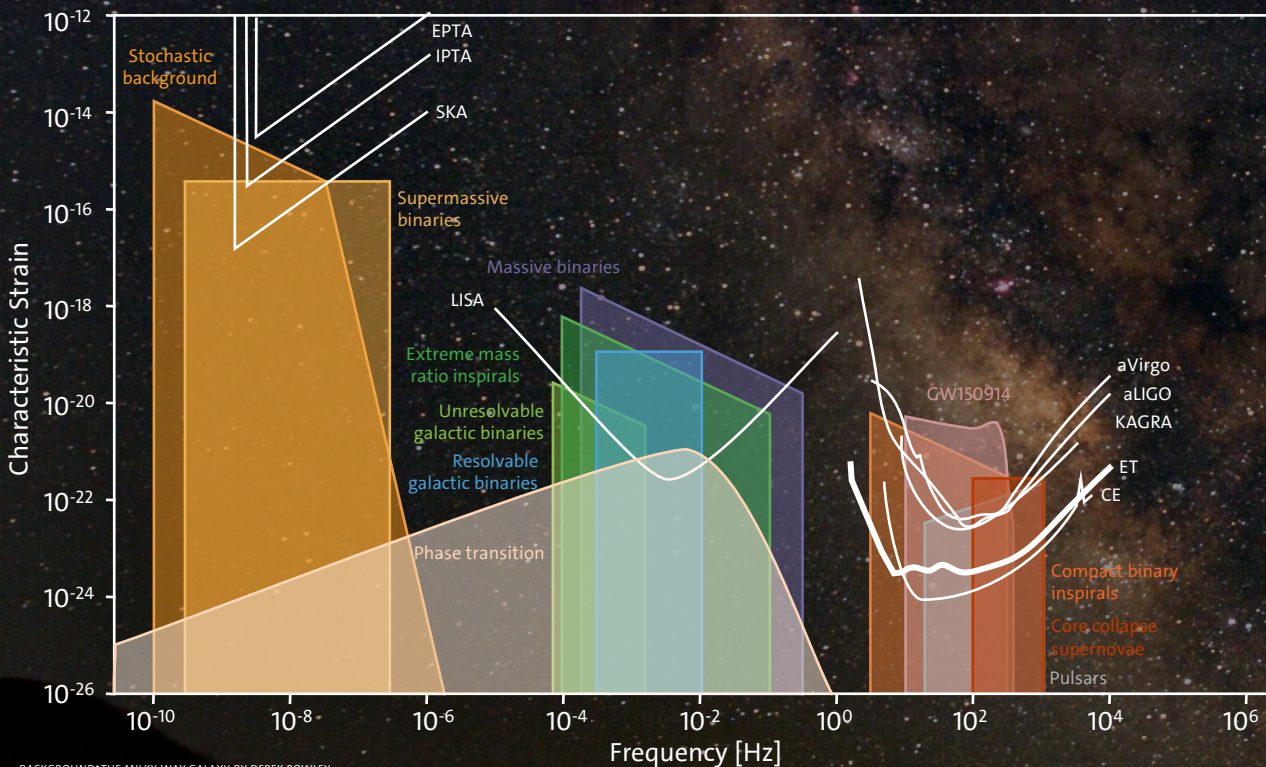
Einstein Telescope Conceptual Design Study 2011 ET-0106C-10

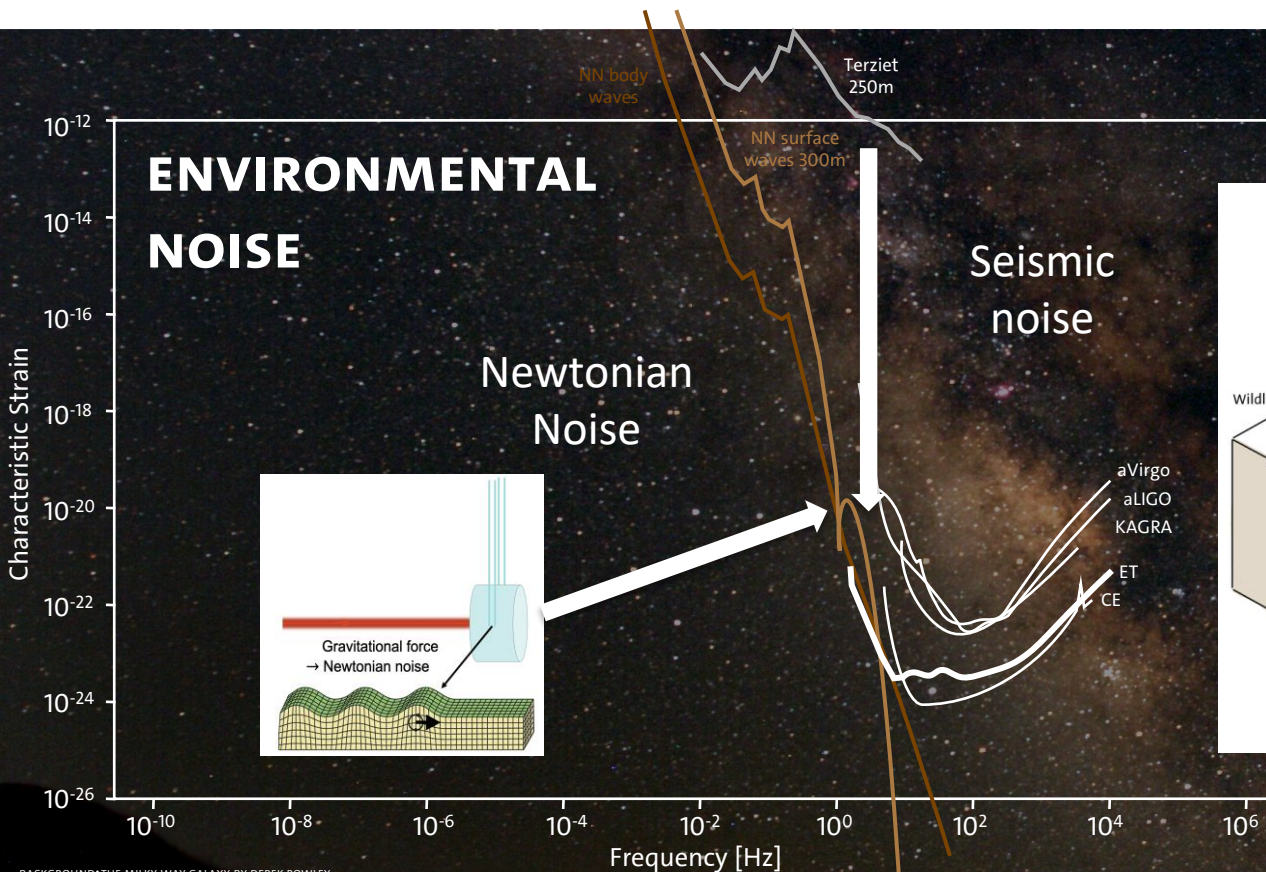
EINSTEIN TELESCOPE SENSITIVITY



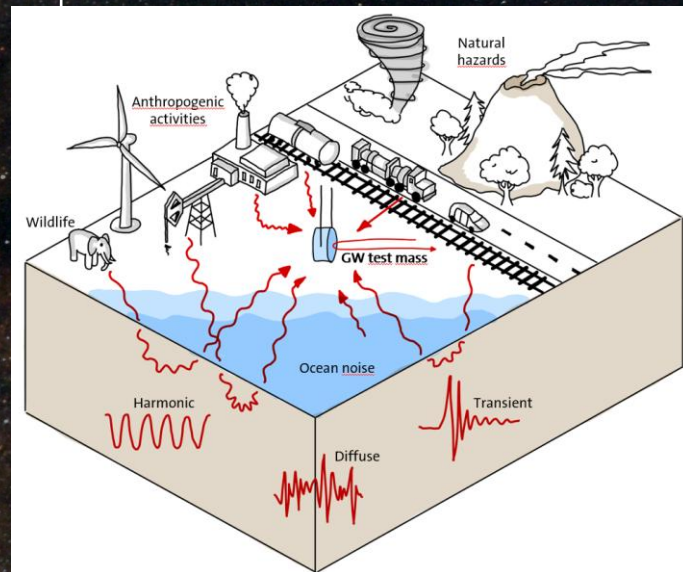
ET LOW FREQUENCY

ADAPTED FROM GWPLOTTER.COM BY C. MOORE, R. COLE, AND C. BERRY.



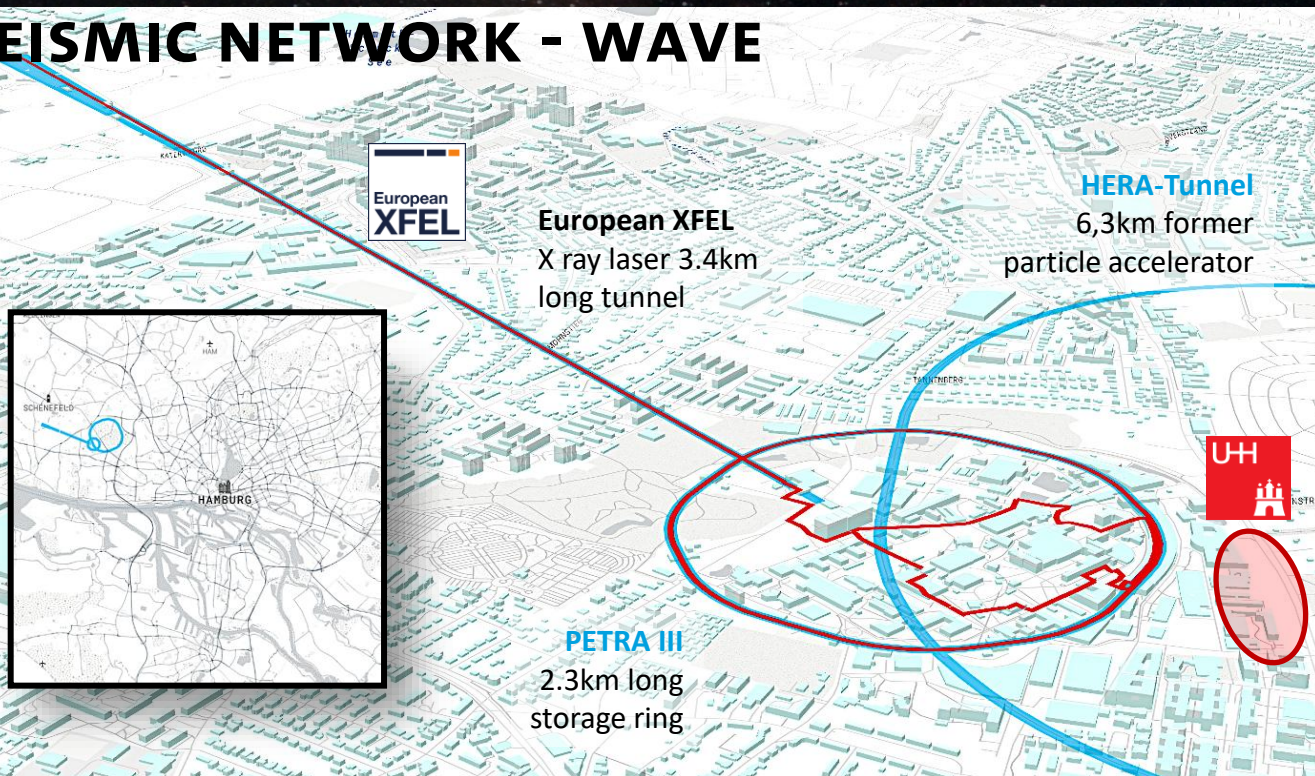


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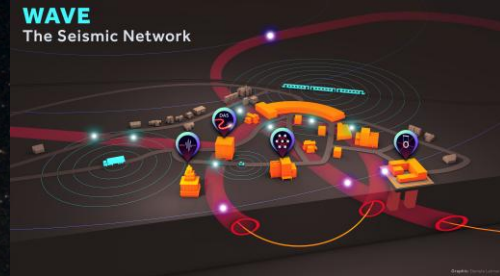




SEISMIC NETWORK - WAVE



WAVE The Seismic Network



19 km glass fibre
are turned into
19.000 sensors

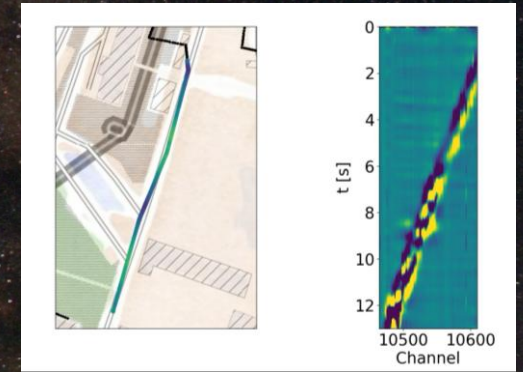
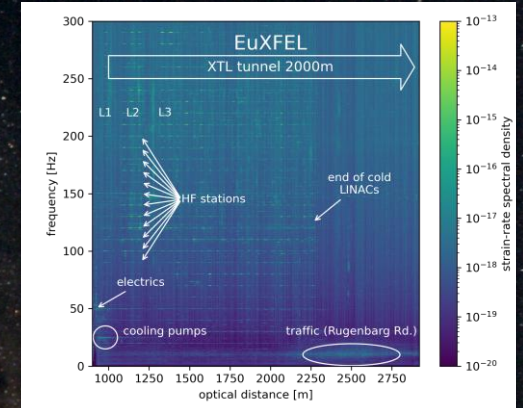
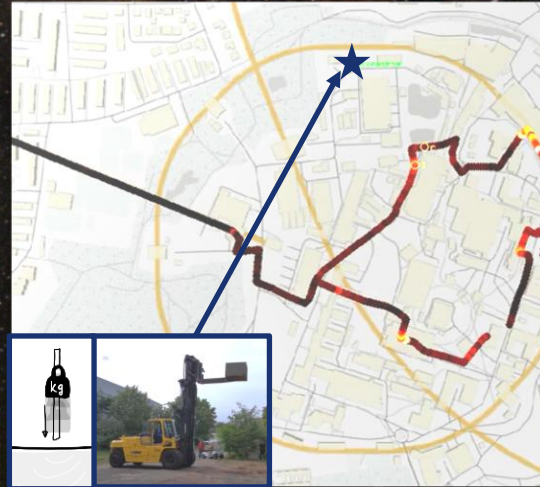
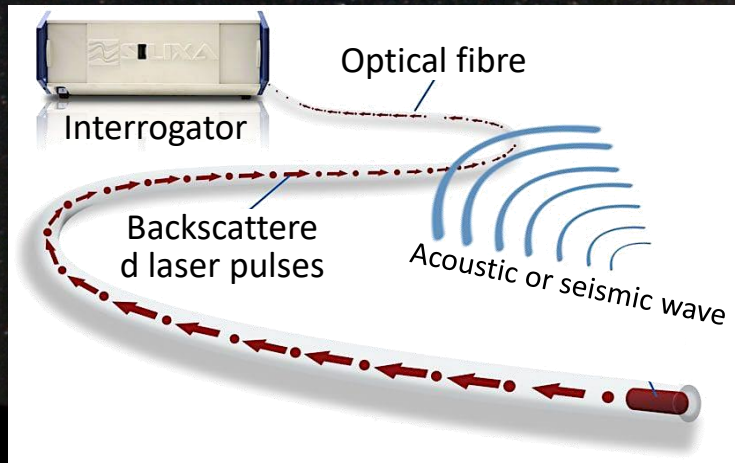
**Campus-wide
seismic sensor
network**

wave-hamburg.eu



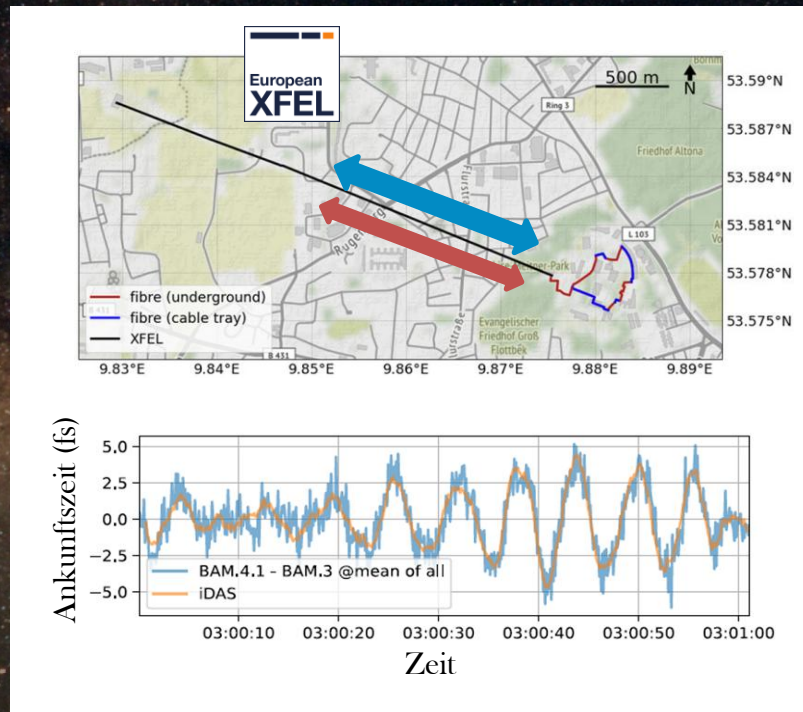
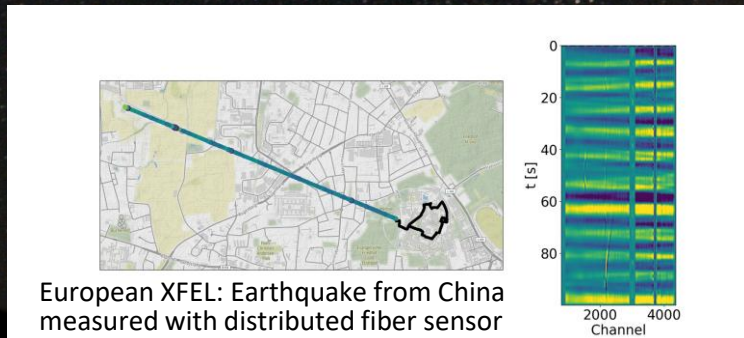
SEISMIC NETWORK - WAVE

- Distributed acoustic sensing (DAS)
- Sensitive to all infrastructure noise
- Sensitive to traffic, events, temperature, airplanes and more

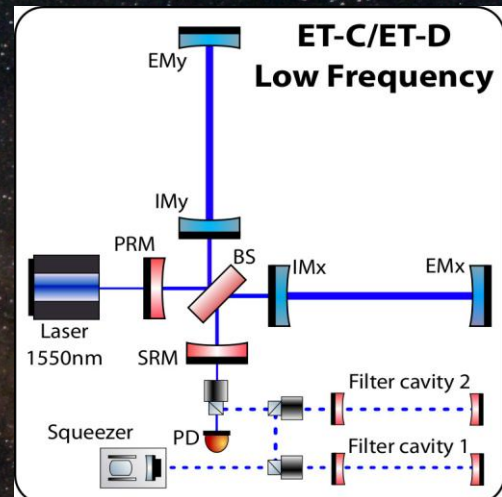
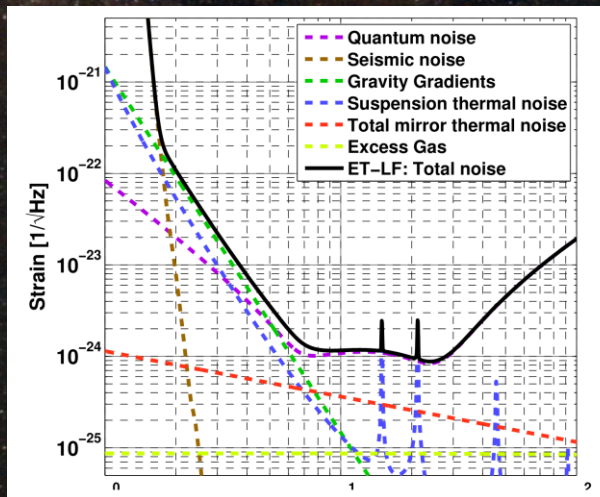
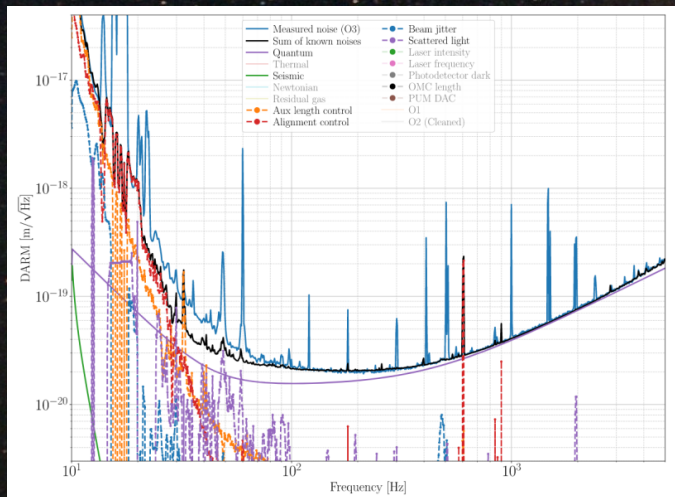


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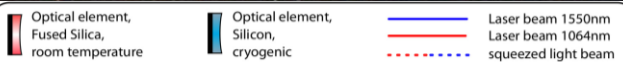
- Directional detection of long wavelength signals from e.g. Earthquakes
- Identified ocean microseism as limiting noise in European XFEL timing system
 -> can now correct fs variations using DAS



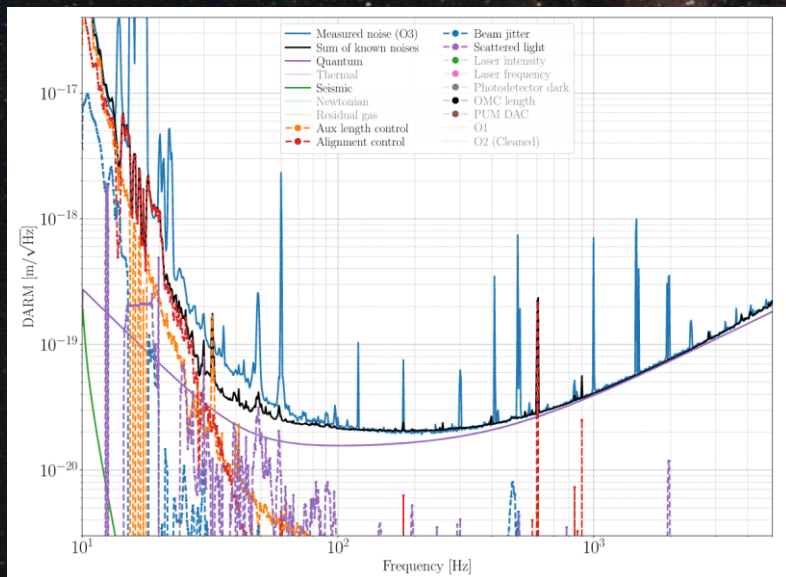
GOAL FOR ET-LF: QUANTUM-THERMAL NOISE LIMIT



Fundamental technical noise

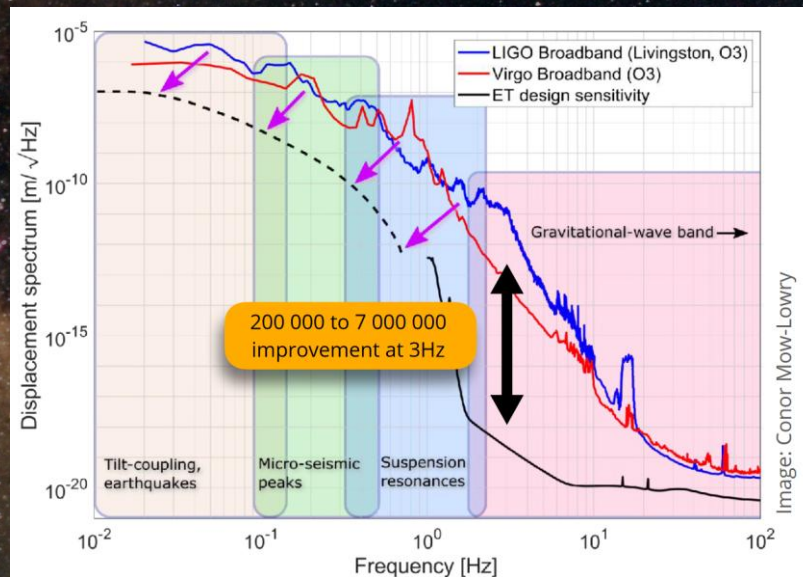

 Optical element, Fused Silica, room temperature
 Optical element, Silicon, cryogenic
 Laser beam 1550nm
 Laser beam 1064nm
 squeezed light beam

INTRINSIC LOW-FREQUENCY INSTRUMENT NOISE



LIGO O3 noise budget

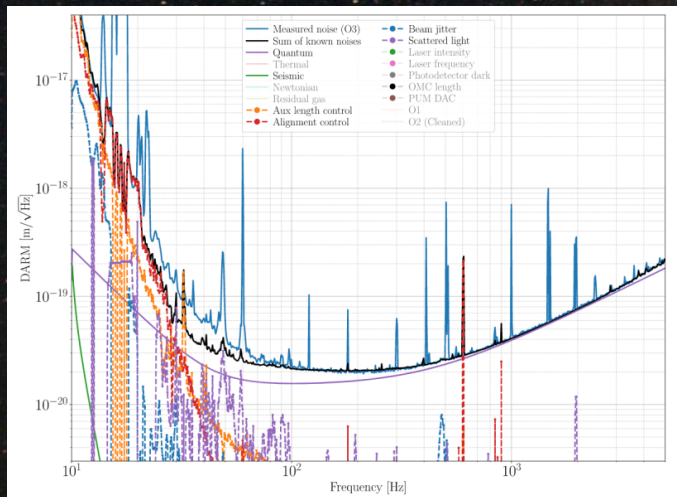
https://ccahilla.github.io/lho_noisebudget.svg



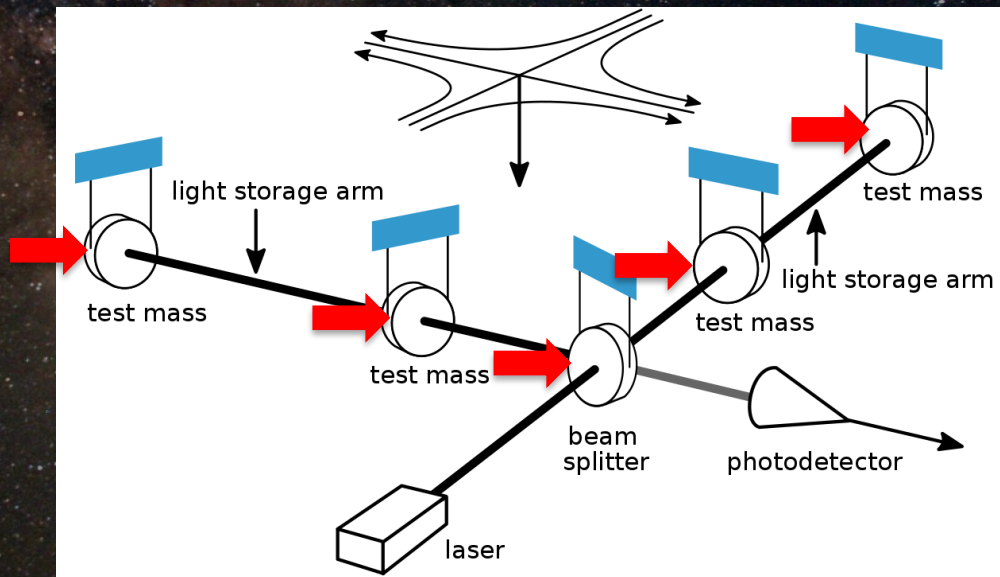
Einstein Telescope

Image: Conor Mow-Lowry

INSTRUMENT SCIENCE CHALLENGE: **CONTROL NOISE**



Fundamental technical noise

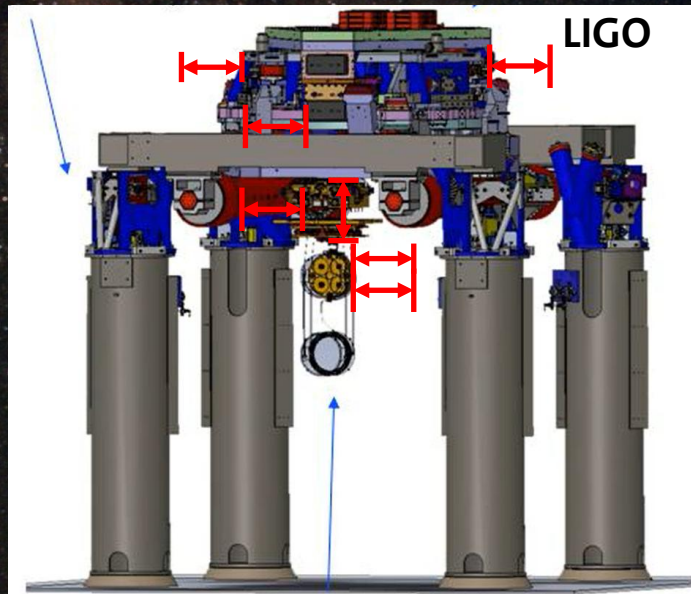


SEISMIC ISOLATION

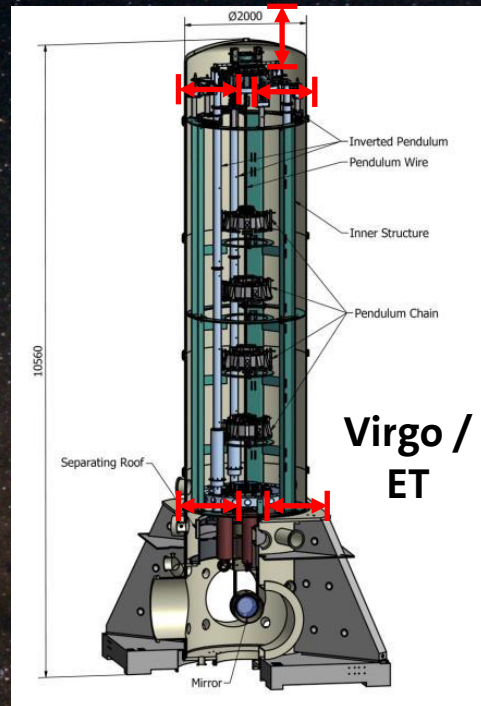
- Active & passive isolation
- Various control loops to
 - Suppress seismic noise
 - Damp resonances
 - Align the mirrors
 - Position the mirrors
- Limited by local displacement and position sensor noise

Goal: $<10\text{fm}/\text{rtHz}$

Active platform



Passive suspension chain

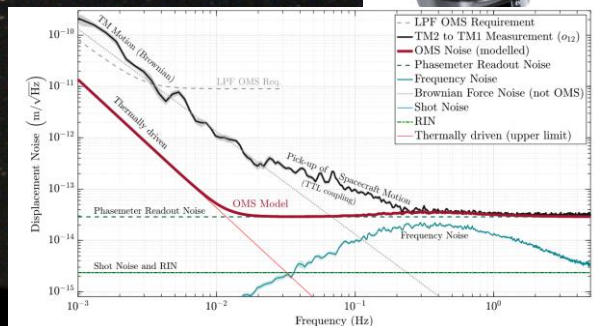
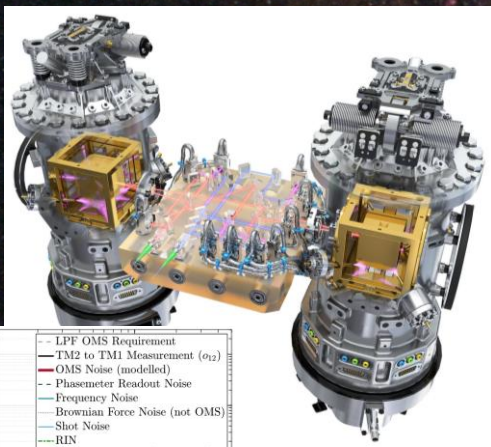


COMPACT INTERFEROMETRIC SENSORS

LISA PATHFINDER

Test mass readout

30fm/rHz
displacement noise



LIGO

10X SIZE REDUCTION
10X MORE SENSORS

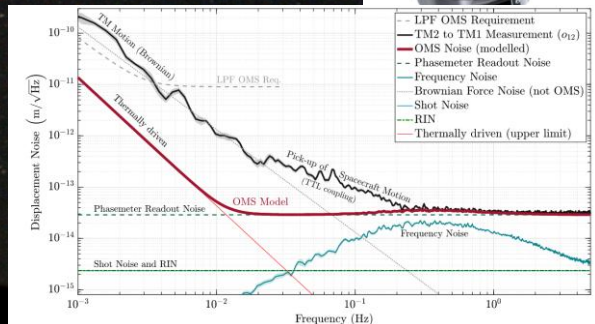
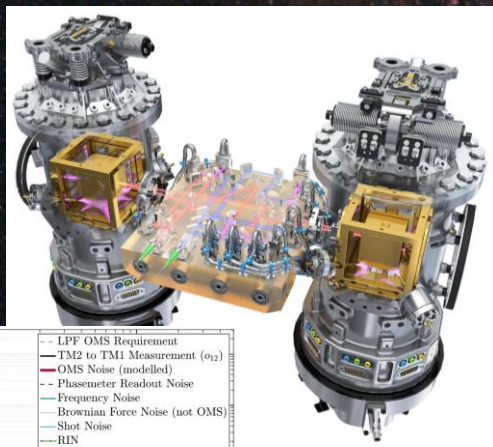


COMPACT INTERFEROMETRIC SENSORS

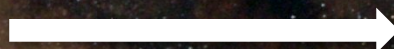
LISA PATHFINDER

Test mass readout

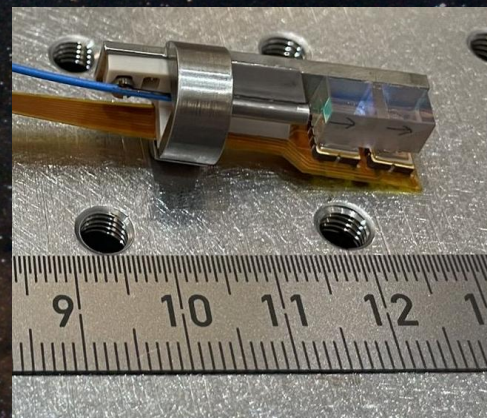
30fm/rHz
displacement noise



10X SIZE REDUCTION
10X MORE SENSORS



LIGO

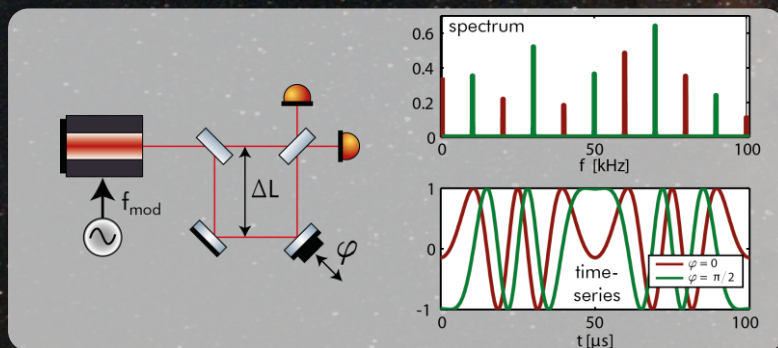


COBRI

- 10fm/rHz displacement noise
- absolute ranging

COMPLEXITY FROM OPTICS TO DIGITAL

Deep frequency modulation interferometry DFMI



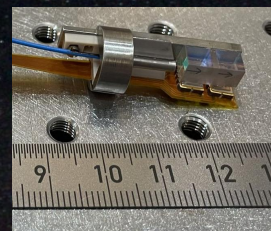
$$i(t) = i_{\text{DC}} + i_{\text{AC}} \cos(\varphi + m \cos(\omega_m t + \psi))$$

Estimate parameters
from non-linear function in real time

Realtime laser control
& stabilisation



FPGA & CPU
processing



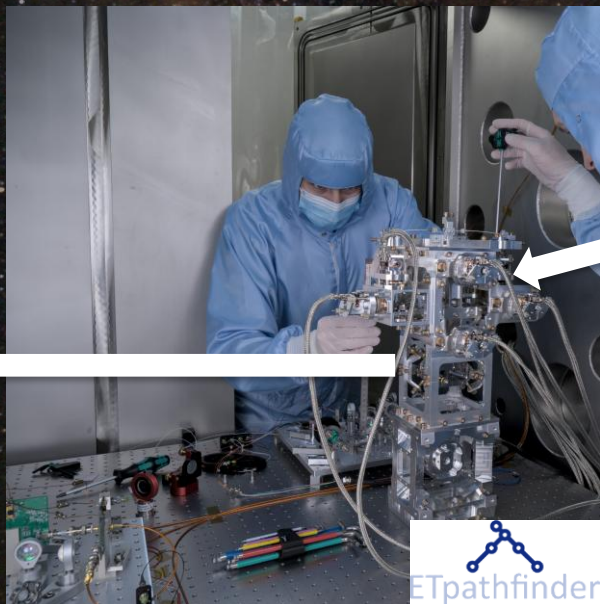
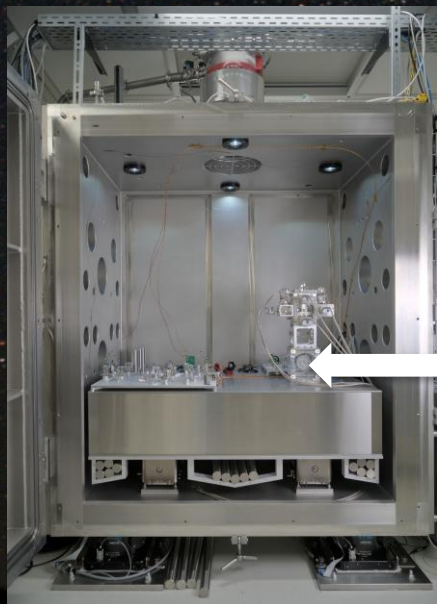
COBRI

- G. Heinzel et al., Optics Express, 18, 19, (2010)
- O. Gerberding, Optics Express, 23, 11, (2015)
- K.-S. Isleif et al., Optics Express, 24, 2, (2016)
- K.-S. Isleif et al., PRApplied 12, 034025 (2019)
- Gerberding O. et al., Sensors 21.5 (2021)
- Smetana J. et al., PRApplied 18.3 (2022)
- Eckhardt T. et al., Metrology 2.1 (2022)
- Eckhardt T. et al., Scientific Reports 14.1 (2024)

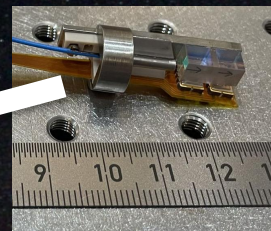
LIGO

TEST IN RELEVANT ENVIRONMENT

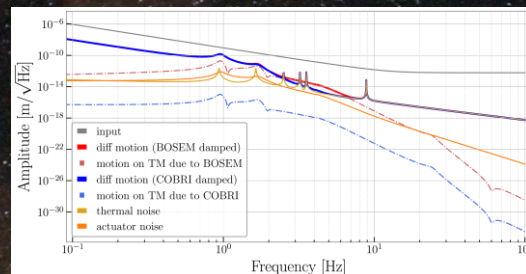
Inertial testbed



LIGO



COBRI



L.
Weickhardt
et al. In
preparation

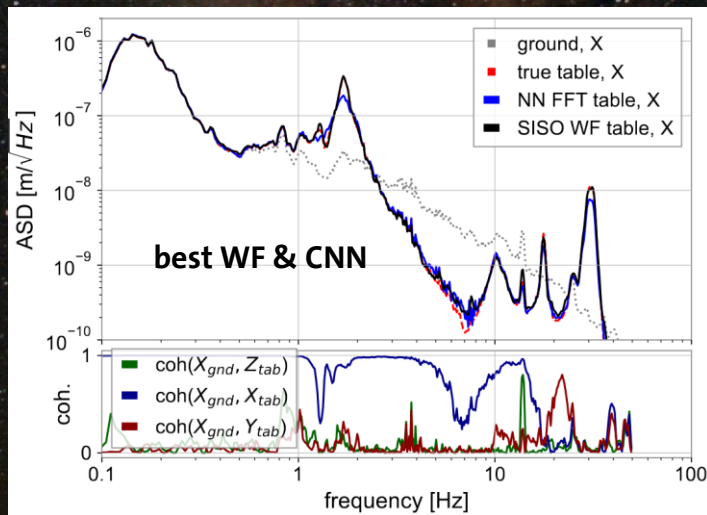
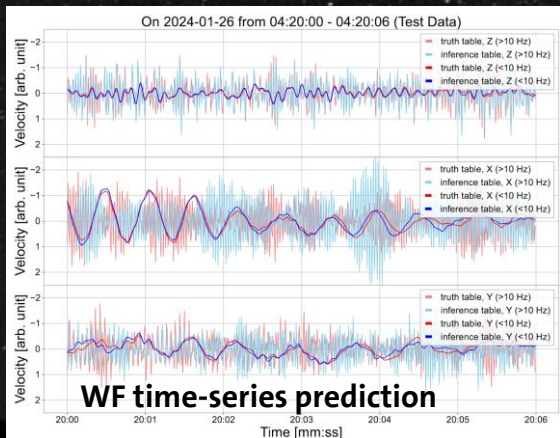
HAM Relay Triple Suspensions

Demonstrate performance

MACHINE LEARNING FOR PREDICTION & CONTROL

Predict the table motion (3 DOF) with seismometer on ground

Can convolutional neural networks outperform (optimal) Wiener Filters for such a mostly linear, mechanical system??





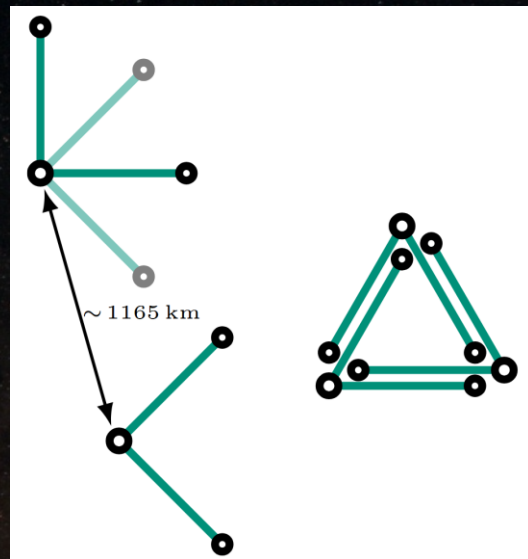
EINSTEIN TELESCOPE SITES AND CONFIGURATIONS

- There are now three candidate sites for the Einstein Telescope: Sardinia (Italy), Euregio-Meuse-Rhine (EMR) region (Netherlands, Belgium, Germany) and Lusatia (Germany)
- Two major configurations are being studied:
 - Triangle/ Δ : 3 co-located detectors (6 interferometers)
 - 2L: 2 detectors on separated sites (4 interferometers)
- The configurations have various implications with regards to cost, risk, noise, science output and politics

The current goal is to come to a decision on the sites and the configurations around early 2027. Ongoing site studies are planned to conclude by the end of 2026.

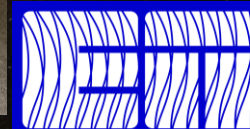
1st detector operation is currently scheduled around the end of 2030s, assuming approval is happening shortly after site decision.

By Tjonnie Li, <https://indico.nikhef.nl/event/6639/attachments/9753/14692/TjonnieLi.pdf>



EINSTEIN TELESCOPE IN GERMANY

- The current German government has shown great interest in the Einstein Telescope
- A preparatory phase for the Einstein Telescope was selected for the FIS Prioritisation of the federal government
- We are currently preparing for this funding
- Germany has a very active community and can play a leading role in the project going forward (independent of the chosen configuration)



EINSTEIN
TELESCOPE

CREDIT:

- **GWD GROUP
+ ALUMNI**
- **COLLABORATORS**
- **FUNDING AGENCIES**



**EINSTEIN
TELESCOPE**



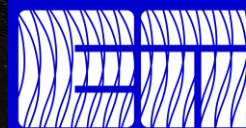
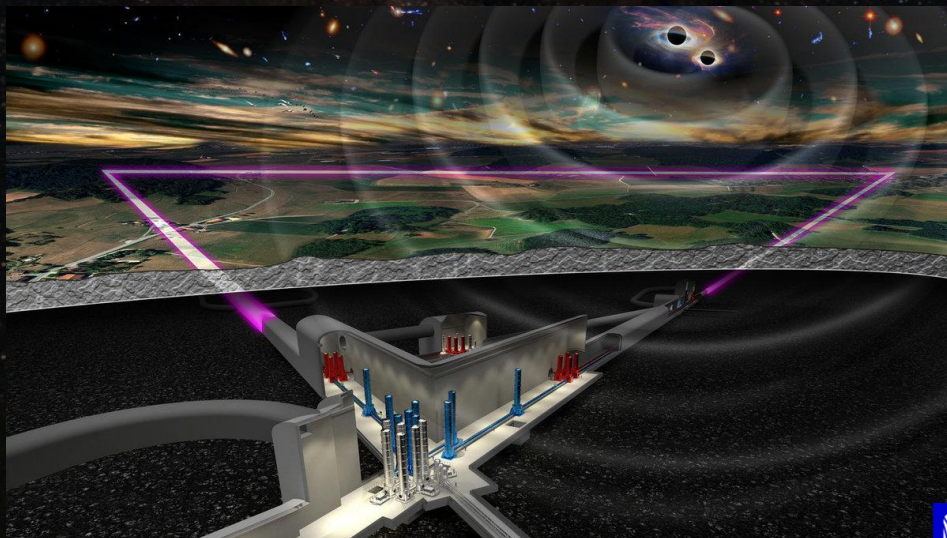
**LIGO
Scientific
Collaboration**



**CLUSTER OF EXCELLENCE
QUANTUM UNIVERSE**



Thank you for your attention!



EINSTEIN
TELESCOPE