

# The Higgs sector (alternatives to susy)

*Implications of the Early  
LHC for cosmology  
DESY, April 18-20, 2012*

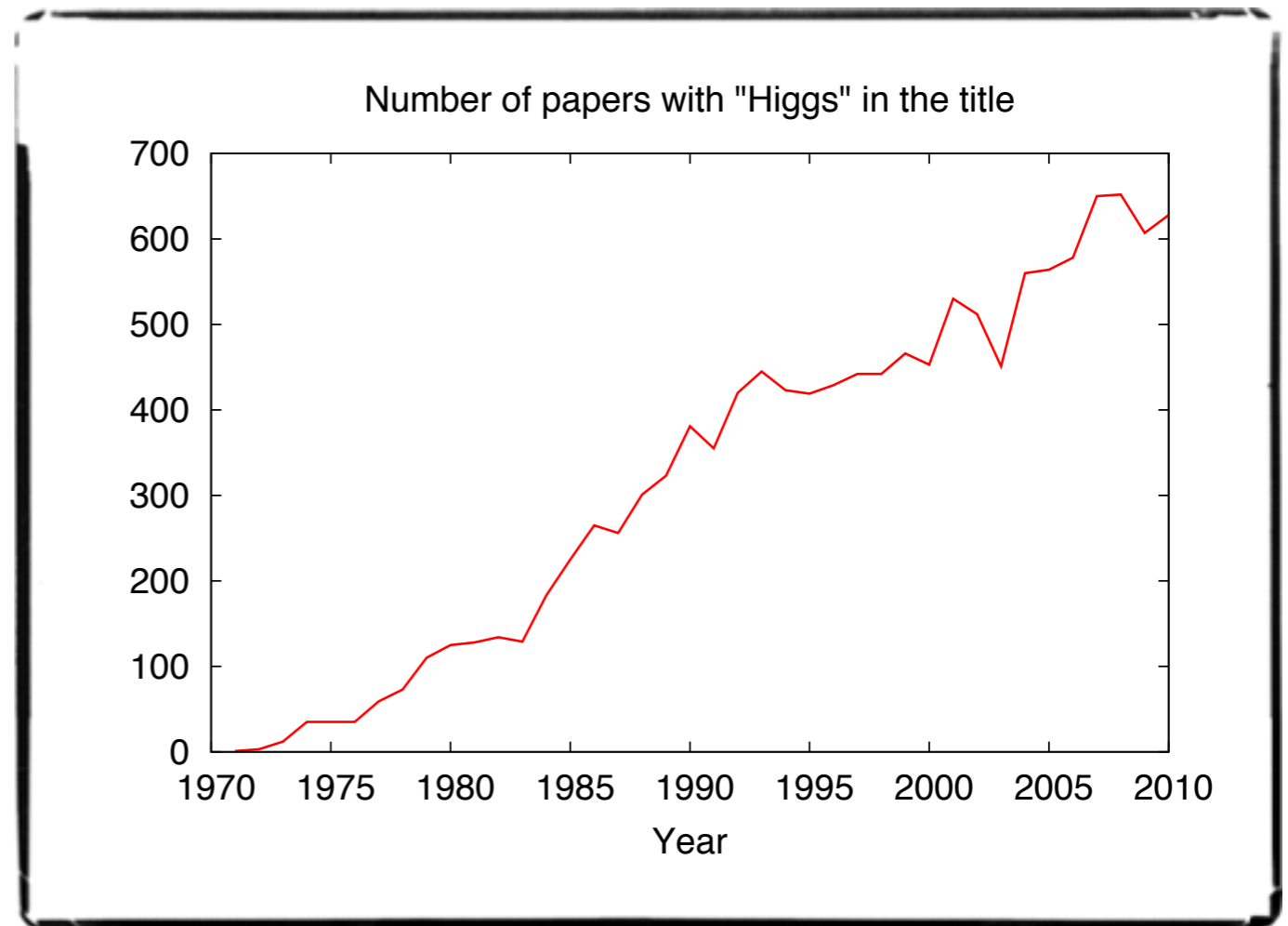
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# Higgs = "raison d'être" of LHC

- $\approx 500$  physics papers over the last 5 years have an introduction starting like "*the (main) goal of the LHC is to discover the Higgs boson*"
- $\approx 11'000$  papers in Spires contain "Higgs" in their title



with even a bigger peak since last Dec.!

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- $\approx 11'000$  papers in Spire contain "Higgs" in their title
- $\approx 3 \times 10^6$  references in google ( $14 \times 10^6 \approx 1\%$  of k€ requested by the German banks to the Greek government)
- ... no Nobel prize (so far)

## Reasons of a success

- last missing piece of the SM?
- at the origin of the masses of elementary particles?
- unitarization of WW scattering amplitudes
- screening of gauge boson self-energies

"Higgs = emergency tire of the SM"

# The UV behavior of the weak Goldstone

symmetry breaking: new phase with more degrees of freedom

massive  $W^\pm, Z$ : 3 physical polarizations=eaten Goldstone bosons  $\frac{SU(2)_L \times SU(2)_R}{SU(2)_V}$

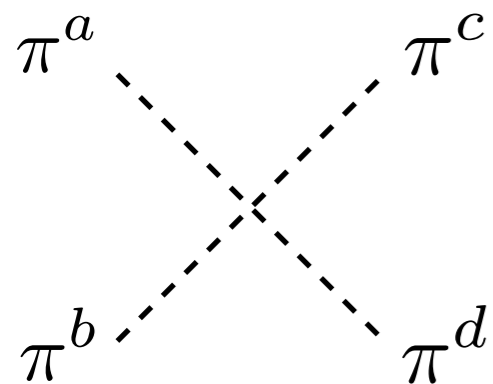
$\Rightarrow$  UV behavior of these Goldstone's?  $\Leftarrow$

$$\mathcal{L}_{\text{mass}} = m_W^2 W_\mu^+ W^{\mu-} + \frac{1}{2} m_Z^2 Z_\mu Z^\mu = \frac{v^2}{4} \text{Tr} (D_\mu \Sigma^\dagger D_\mu \Sigma)$$

$\Sigma = e^{i\sigma^a \pi^a / v}$   
Goldstone of  
 $SU(2)_L \times SU(2)_R / SU(2)_V$

$$\mathcal{L}_{\text{mass}} = \frac{1}{2} (\partial_\mu \pi^a)^2 - \frac{1}{6v^2} ((\pi^a \partial_\mu \pi^a)^2 - (\pi^a)^2 (\partial_\mu \pi^a)^2) + \dots$$

contact interaction growing with energy



$$\mathcal{A}(\pi^a \pi^b \rightarrow \pi^c \pi^d) = \mathcal{A}(s, t, u) \delta^{ab} \delta^{cd} + \mathcal{A}(t, s, u) \delta^{ac} \delta^{bd} + \mathcal{A}(u, t, s) \delta^{ad} \delta^{bc}$$

$$\mathcal{A}(s, t, u) = \frac{s}{v^2} \quad \text{Weinberg's LET}$$

the behavior of this amplitude is not consistent above  $4\pi v$  ( $\approx 1-3 \text{ TeV}$ )

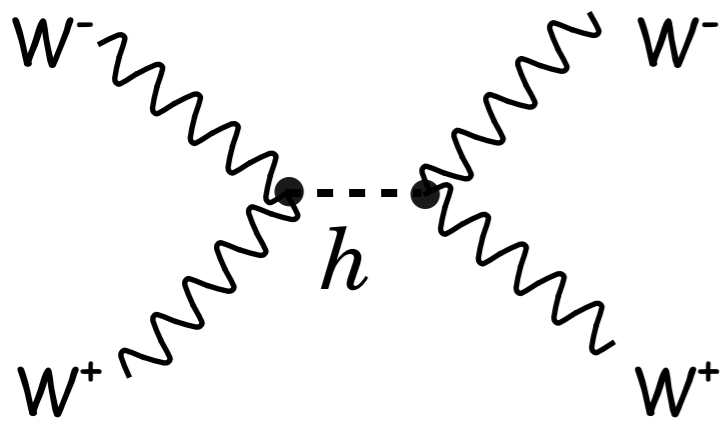
Lee, Quigg & Thacker '77

# What is the SM Higgs?

A single scalar degree of freedom neutral under  $SU(2)_L \times SU(2)_R / SU(2)_V$

$$\mathcal{L}_{\text{EWSB}} = \frac{v^2}{4} \text{Tr} (D_\mu \Sigma^\dagger D_\mu \Sigma) \left( 1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} \right) - \lambda \bar{\psi}_L \Sigma \psi_R \left( 1 + c \frac{h}{v} \right)$$

'a', 'b' and 'c' are arbitrary free couplings



$$\mathcal{A} = \frac{1}{v^2} \left( s - \frac{a^2 s^2}{s - m_h^2} \right)$$

growth cancelled for  
 $a = 1$   
 restoration of  
 perturbative unitarity

Cornwall, Levin, Tiktopoulos '73

Contino, Grojean, Moretti, Piccinini, Rattazzi '10

$$\Sigma = e^{i\sigma^a \pi^a / v}$$

Goldstone of  $SU(2)_L \times SU(2)_R / SU(2)_V$

$$D_\mu \Sigma \approx W_\mu$$

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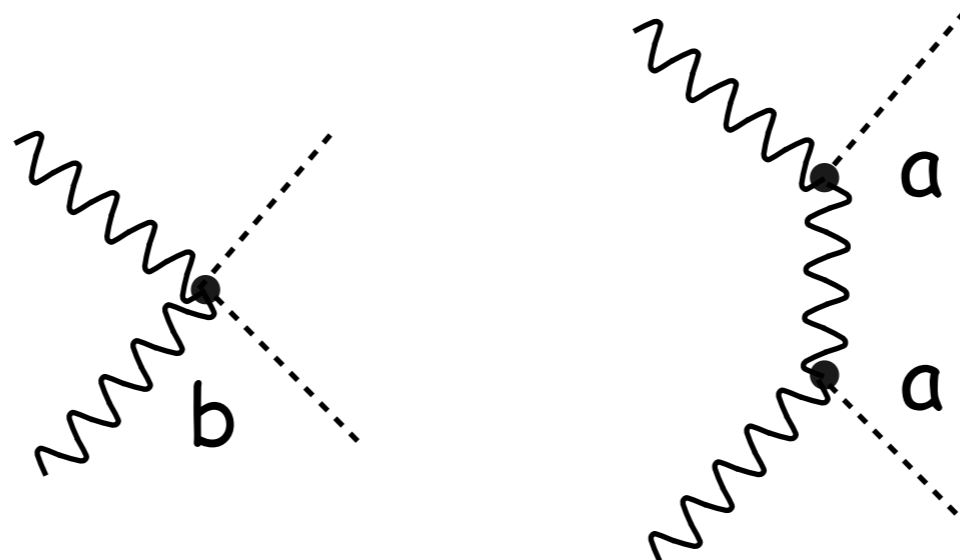
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For  $a=1$ : perturbative unitarity in elastic channels  $WW \rightarrow WW$

For  $b = a^2$ : perturbative unitarity in inelastic channels  $WW \rightarrow hh$

Cornwall, Levin, Tiktopoulos '73

Contino, Grojean, Moretti, Piccinini, Rattazzi '10



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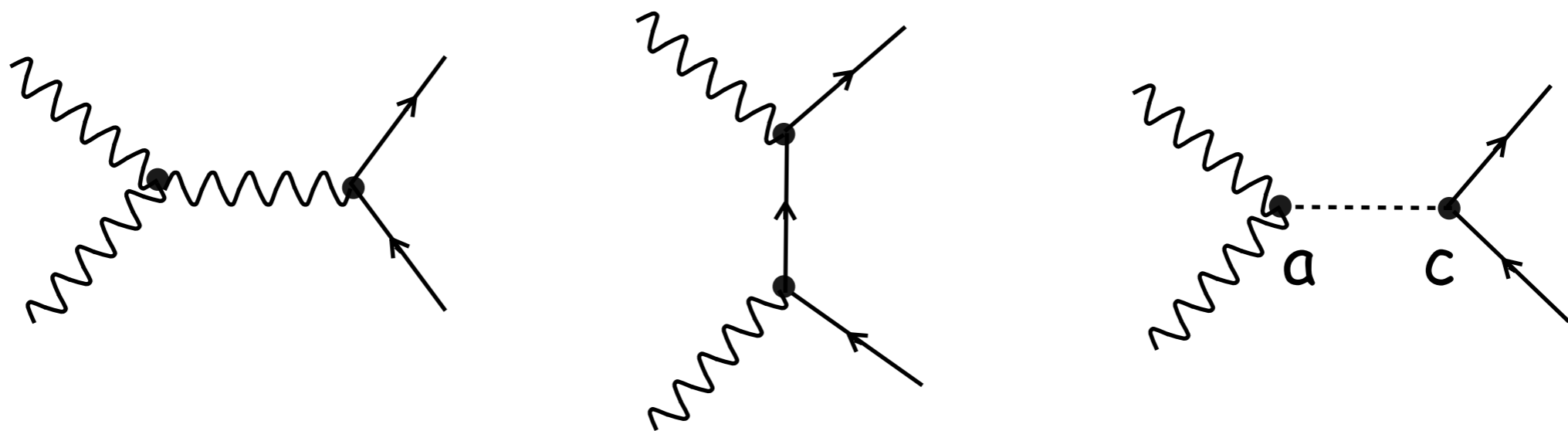
For  $a=1$ : perturbative unitarity in elastic channels  $WW \rightarrow WW$

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For  $ac=1$ : perturbative unitarity in inelastic  $WW \rightarrow \psi \psi$

Cornwall, Levin, Tiktopoulos '73

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For  $ac=1$ : perturbative unitarity in inelastic  $WW \rightarrow \psi \psi$

'a=1', 'b=1' & 'c=1' define the SM Higgs

Higgs properties depend on a single unknown parameter ( $m_H$ )

$\mathcal{L}_{\text{EWSB}}$  can be rewritten as  $D_\mu H^\dagger D_\mu H$

$$H = \frac{1}{\sqrt{2}} e^{i\sigma^a \pi^a / v} \begin{pmatrix} 0 \\ v + h \end{pmatrix}$$

$h$  and  $\pi^a$  (ie  $W_L$  and  $Z_L$ ) combine to form a linear representation of  $SU(2)_L \times U(1)_Y$



# What is a composite Higgs?

A  $\sigma$  particle that combines with  $W_L$  and  $Z_L$  to form a  $SU(2)$  doublet

*renormalizable level = uniqueness*

$SU(2)_L \times U(1)_Y$  linearly realized  $\Leftrightarrow$  Standard Model  $\Leftrightarrow a=b=c=1$

*non-renormalizable level*

$SU(2)_L \times U(1)_Y$  linearly realized &  $a, b, c \neq 1 \Leftrightarrow$  Composite Higgs

deviations of Higgs couplings originate from higher dimensional operators

$$\underbrace{(\partial_\mu |H|^2)^2 \quad |H|^2 \bar{\psi} H \psi}_{\text{relevant for composite Higgs models}} \quad \underbrace{|H|^2 B_{\mu\nu} B^{\mu\nu} \quad |H|^2 G_{\mu\nu} G^{\mu\nu}}_{\text{irrelevant for composite Higgs models (not for dilaton)}}$$

relevant for

composite Higgs models

irrelevant

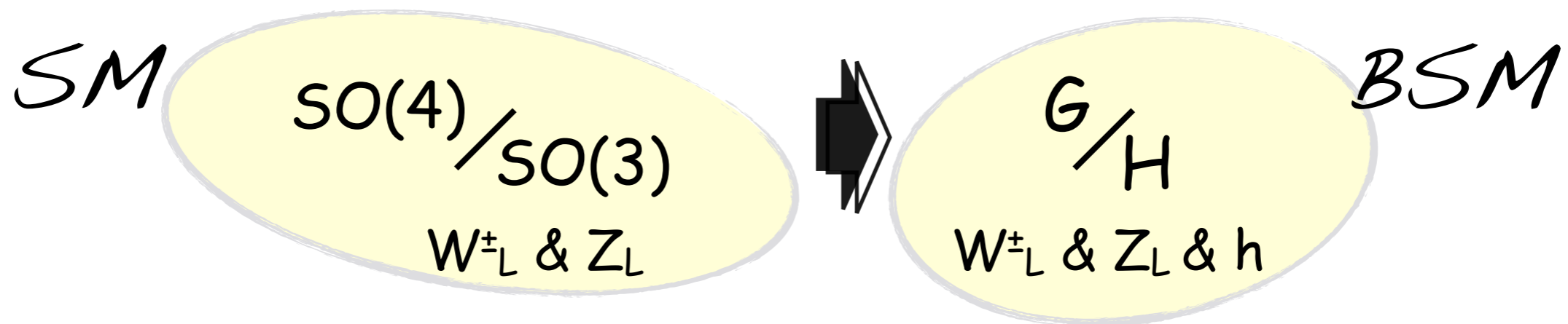
for composite Higgs models (not for dilaton)

# Higgs as a PGB: a natural extension of SM

One solution to the hierarchy pb:

Higgs transforms non-linearly under some global symmetry

Higgs=Pseudo-Goldstone boson (PGB)



Examples:  $SO(5)/SO(4)$ : 4 PGBs =  $W^\pm_L, Z_L, h$

Minimal Composite Higgs Model

Agashe, Contino, Pomarol '04

$SO(6)/SO(5)$ : 5 PGBs =  $H, a$

Next MCHM

Gripaios, Pomarol, Riva, Serra '09

$SU(4)/Sp(4, \mathbb{C})$ : 5 PGBs =  $H, s$

$SO(6)/SO(4) \times SO(2)$ : 8 PGBs =  $H_1 + H_2$

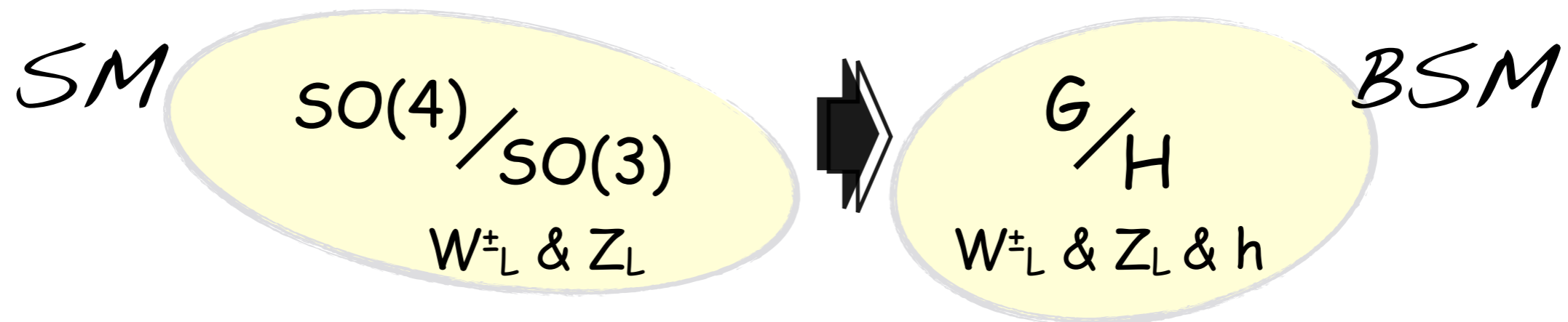
Minimal Composite  
Two Higgs Doublets

Mrazek, Pomarol, Rattazzi, Serra, Wulzer '11

# Higgs as a PGB: a natural extension of SM

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*How can we tell the difference with the SM Higgs?*

# SILH Effective Lagrangian

(strongly-interacting light Higgs)

Giudice, Grojean, Pomarol, Rattazzi '07

■ extra Higgs leg:  $H/f$

■ extra derivative:  $\partial/m_\rho$

## ■ Genuine strong operators (sensitive to the scale $f$ )

$$\frac{c_H}{2f^2} \left( \partial^\mu |H|^2 \right)^2$$

$$\frac{c_T}{2f^2} \left( H^\dagger \overleftrightarrow{D}^\mu H \right)^2$$

custodial breaking

$$\frac{c_y y_f}{f^2} |H|^2 \bar{f}_L H f_R + \text{h.c.}$$

$$\frac{c_6 \lambda}{f^2} |H|^6$$

## ■ Form factor operators (sensitive to the scale $m_\rho$ )

$$\frac{i c_W}{2m_\rho^2} \left( H^\dagger \sigma^i \overleftrightarrow{D}^\mu H \right) (D^\nu W_{\mu\nu})^i$$

$$\frac{i c_B}{2m_\rho^2} \left( H^\dagger \overleftrightarrow{D}^\mu H \right) (\partial^\nu B_{\mu\nu})$$

$$\frac{i c_{HW}}{m_\rho^2} \frac{g_\rho^2}{16\pi^2} (D^\mu H)^\dagger \sigma^i (D^\nu H) W_{\mu\nu}^i$$

$$\frac{i c_{HB}}{m_\rho^2} \frac{g_\rho^2}{16\pi^2} (D^\mu H)^\dagger (D^\nu H) B_{\mu\nu}$$

minimal coupling:  $h \rightarrow \gamma Z$

loop-suppressed strong dynamics

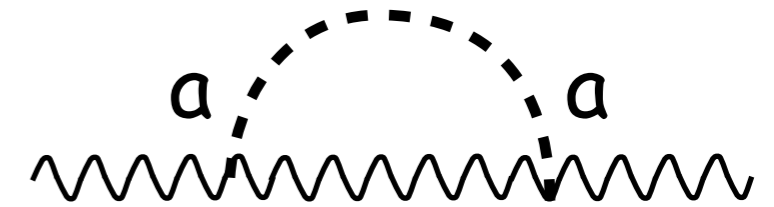
$$\frac{c_\gamma}{m_\rho^2} \frac{g_\rho^2}{16\pi^2} \frac{g^2}{g_\rho^2} H^\dagger H B_{\mu\nu} B^{\mu\nu}$$

$$\frac{c_g}{m_\rho^2} \frac{g_\rho^2}{16\pi^2} \frac{y_t^2}{g_\rho^2} H^\dagger H G_{\mu\nu}^a G^{a\mu\nu}$$

Goldstone sym.

# Deformation of the SM Higgs: EW constraints

The parameter 'a' controls the size of the one-loop IR contribution to the LEP precision observables



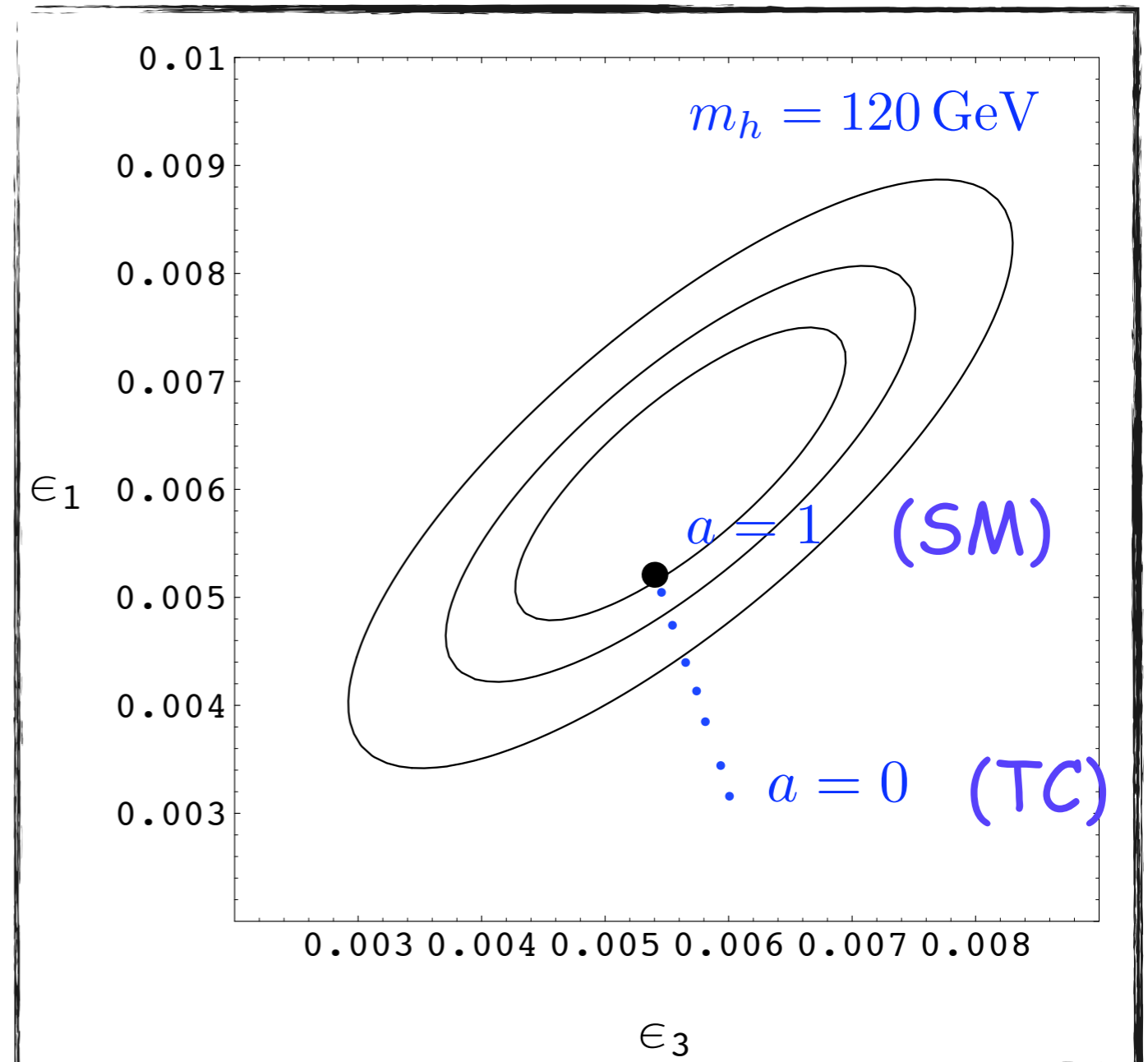
$$\epsilon_{1,3} = c_{1,3} \log(m_Z^2/\mu^2) - c_{1,3} a^2 \log(m_h^2/\mu^2) - c_{1,3} (1 - a^2) \log(m_\rho^2/\mu^2) + \text{finite terms}$$

$$c_1 = + \frac{3}{16\pi^2} \frac{\alpha(m_Z)}{\cos^2 \theta_W}$$

$$c_3 = - \frac{1}{12\pi} \frac{\alpha(m_Z)}{4 \sin^2 \theta_W}$$

$$\Delta\epsilon_{1,3} = -c_{1,3} (1 - a^2) \log(m_\rho^2/m_h^2)$$

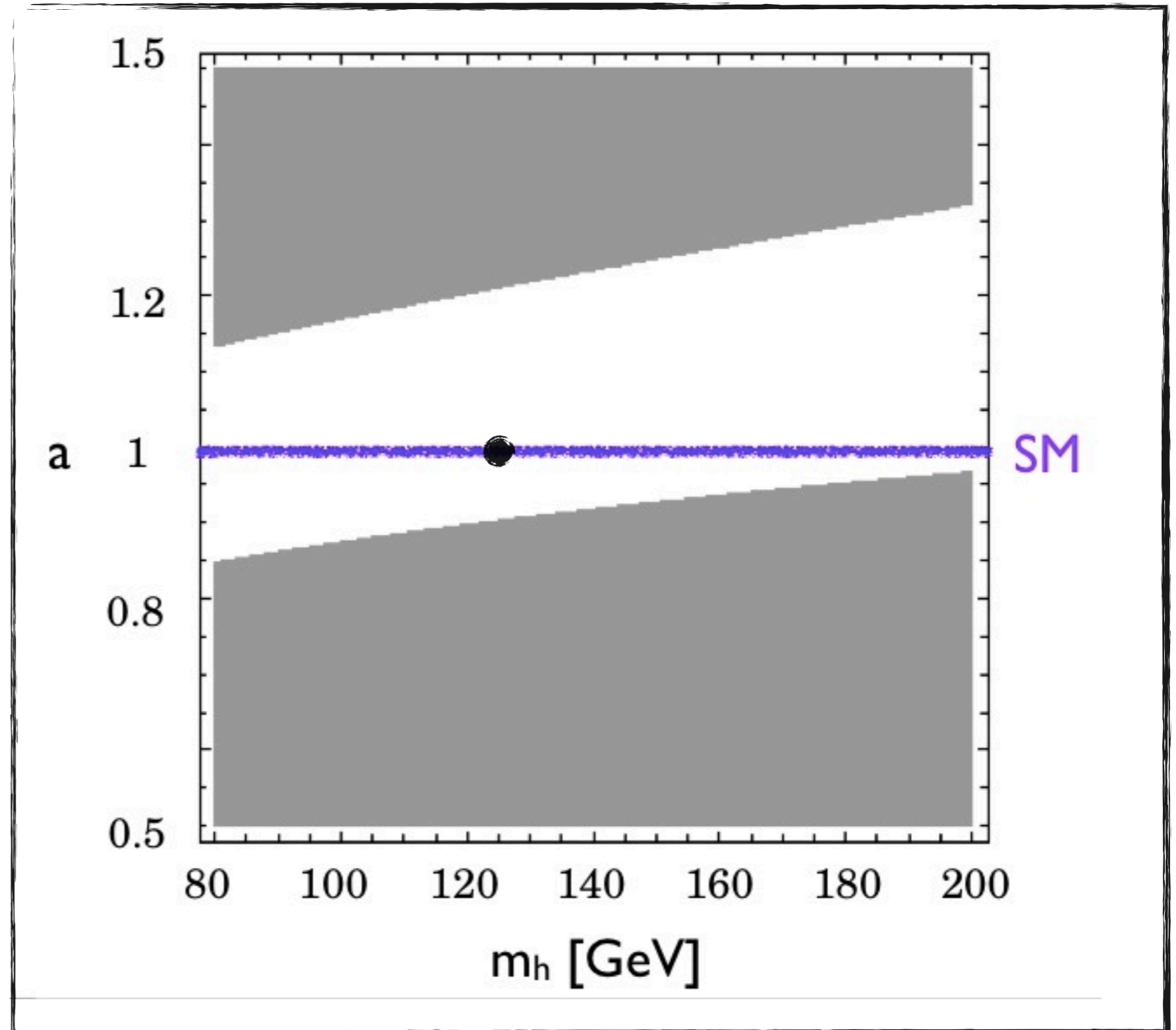
Barbieri, Bellazzini, Rychkov, Varagnolo '07



# EW data constraints on 'a'

EW fit with SM degrees of freedom + (composite) Higgs

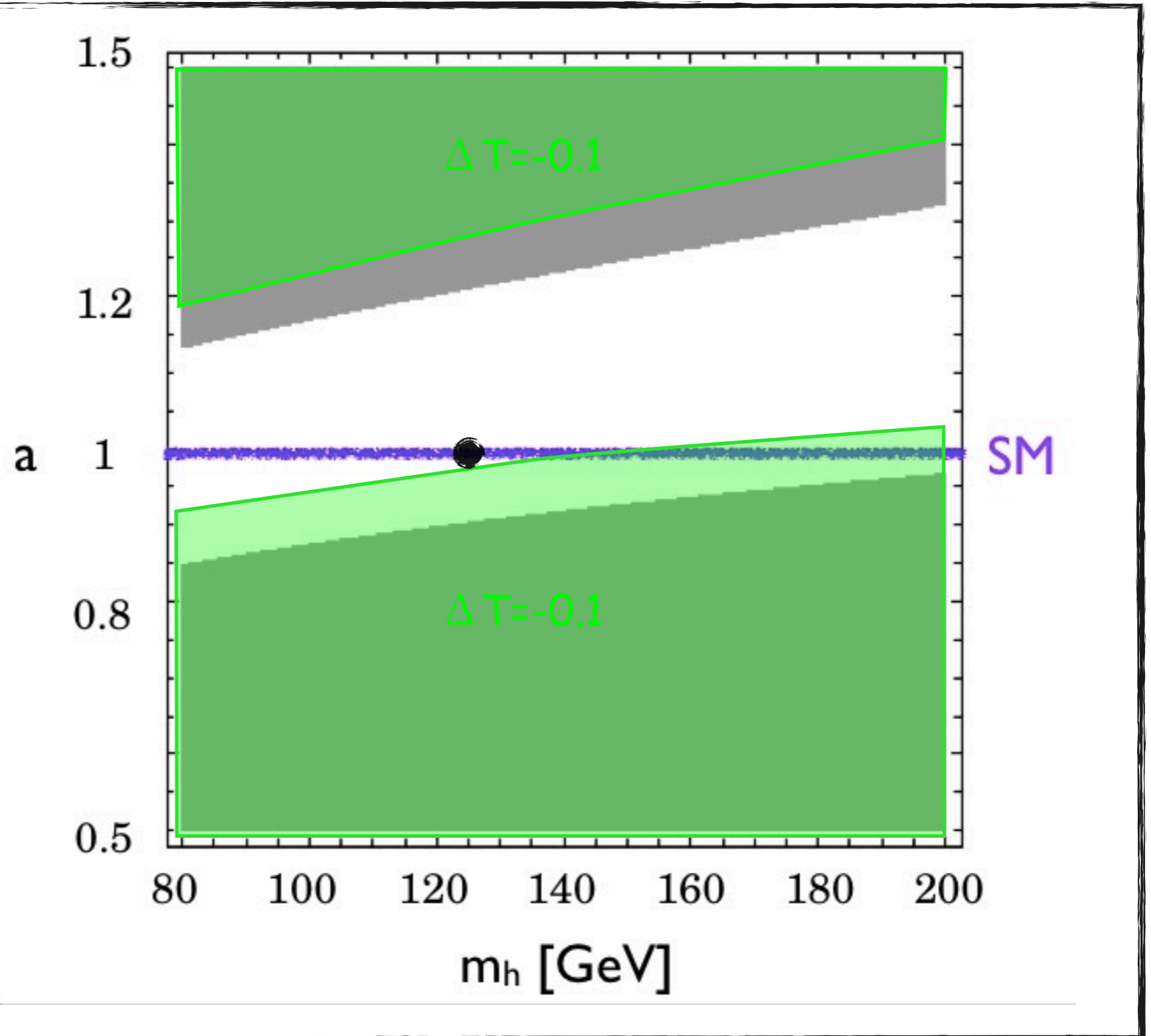
EW data require less than 15-20% deviations in the couplings of the Higgs to gauge bosons



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*note:*  
additional UV contributions to S and T can modify the preferred values of couplings

# EW data constraints on 'a'

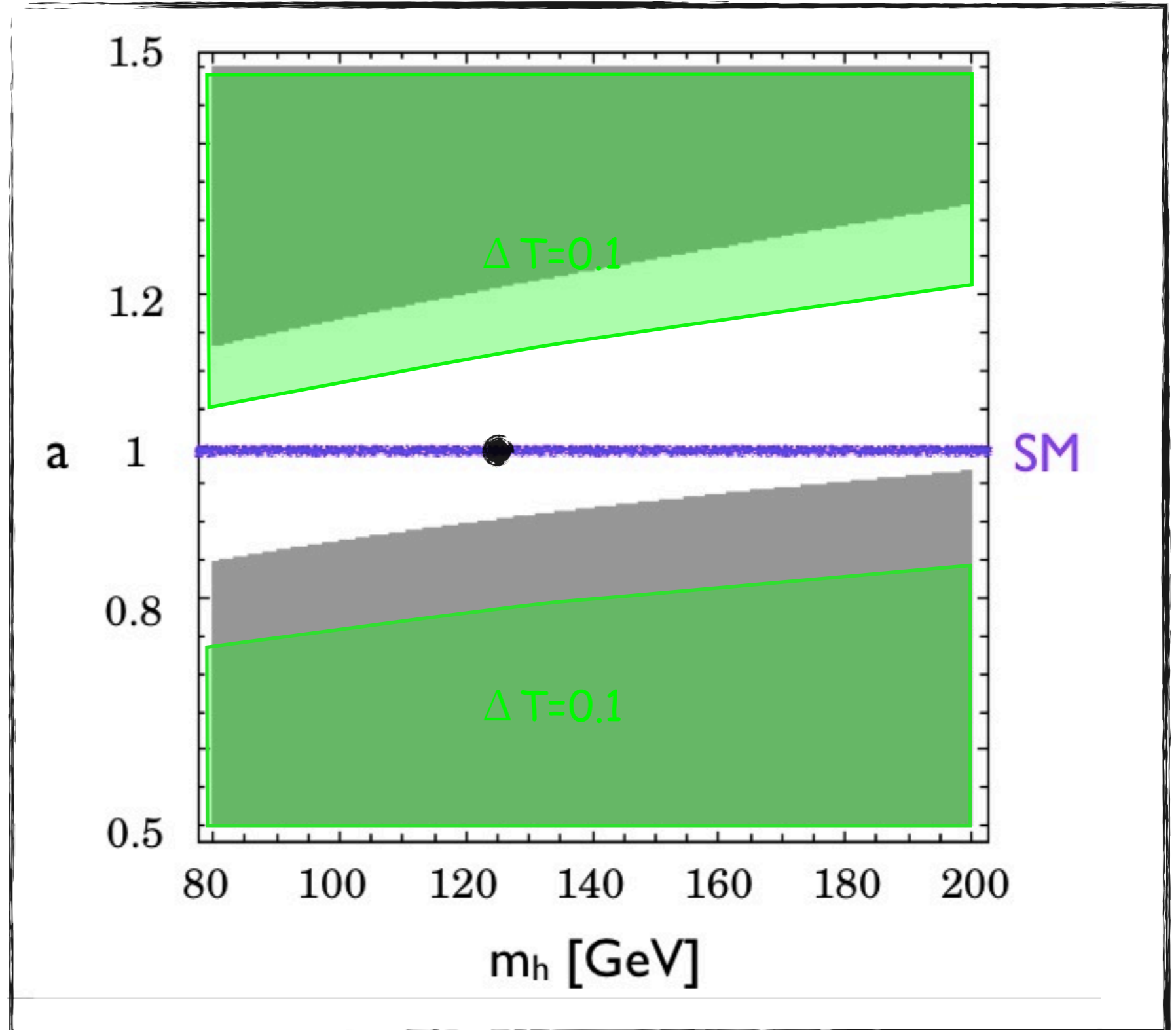
EW fit with SM degrees of freedom + (composite) Higgs

EW data require less than 15-20% deviations in the couplings of the Higgs to gauge bosons

EW data don't constraint the other Higgs couplings

*note:*

additional UV contributions to S and T can modify the preferred values of couplings





# Flavor Constraints

$$\left(1 + \frac{c_{ij}|H|^2}{f^2}\right) y_{ij} \bar{f}_{Li} H f_{Rj} = \left(1 + \frac{c_{ij}v^2}{2f^2}\right) \frac{y_{ij}v}{\sqrt{2}} \bar{f}_{Li} f_{Rj}$$

mass terms

$$\left(1 + \frac{3c_{ij}v^2}{2f^2}\right) \frac{y_{ij}v}{\sqrt{2}} h \bar{f}_{Li} f_{Rj}$$

Higgs fermion interactions

mass and interaction matrices are not diagonalizable simultaneously  
if  $c_{ij}$  are arbitrary

$\Rightarrow$  FCNC

Composite Higgs set-up:  $c$  is flavor universal  
(except may be for the top)

$\Rightarrow$  Minimal flavor violation built in

# Direct Searches

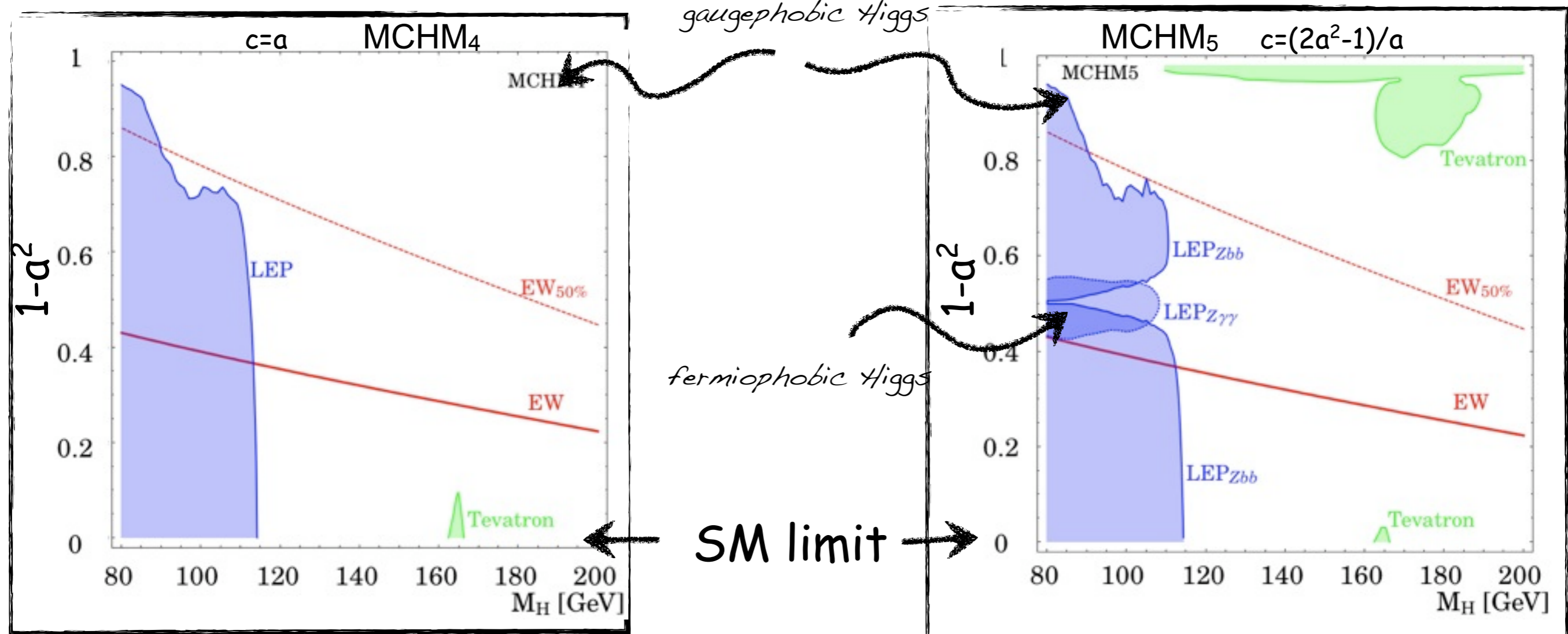
$$\mathcal{L}_{\text{EWSB}} = \frac{v^2}{4} \text{Tr} (D_\mu \Sigma^\dagger D_\mu \Sigma) \left( 1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} \right) - \lambda \bar{\psi}_L \Sigma \psi_R \left( 1 + c \frac{h}{v} \right)$$

SM 'a=1', 'b=1' & 'c=1'

Current EW data constrain only 'a'

Direct searches constrain also 'c'

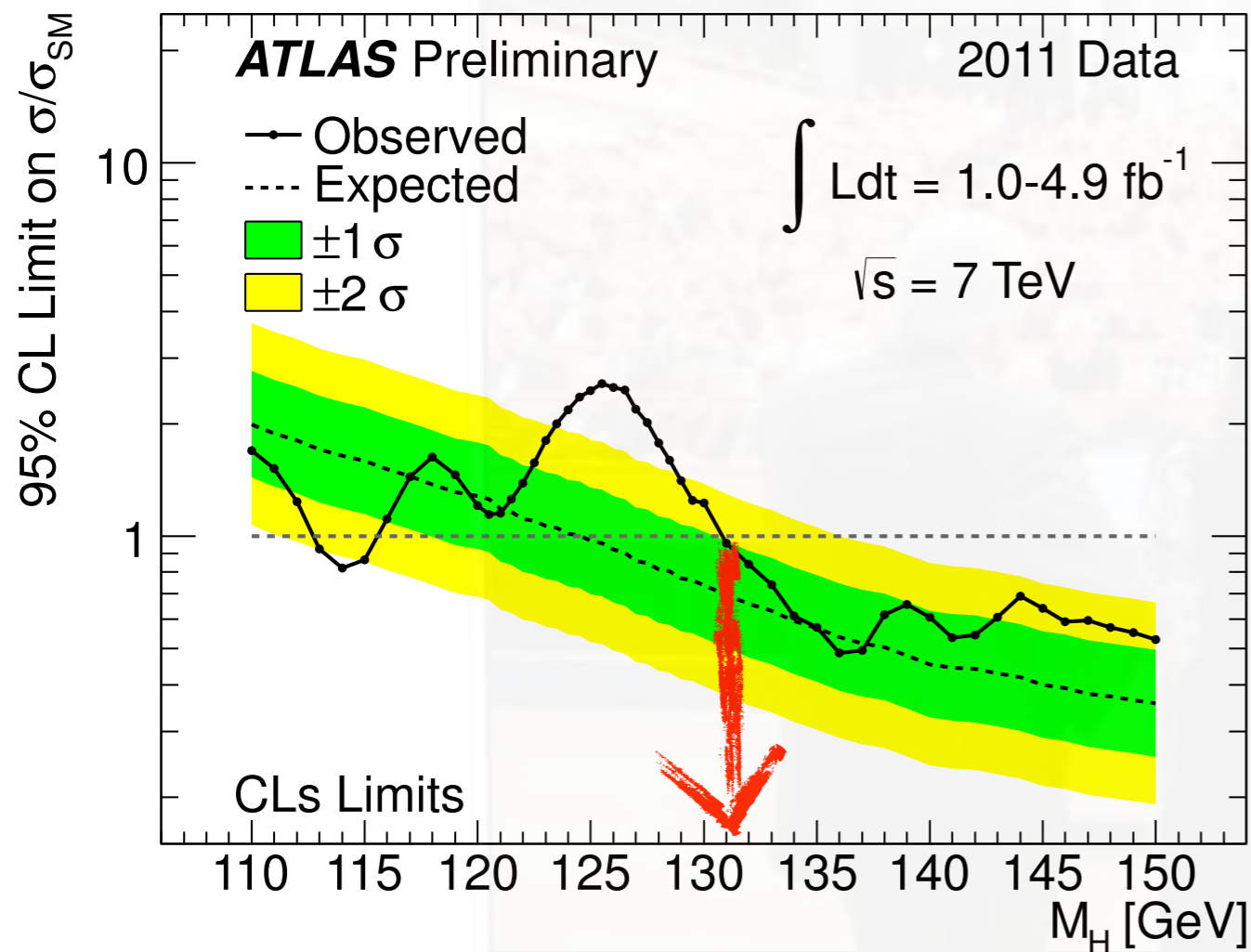
Espinosa, Grojean, Muehlleitner '10



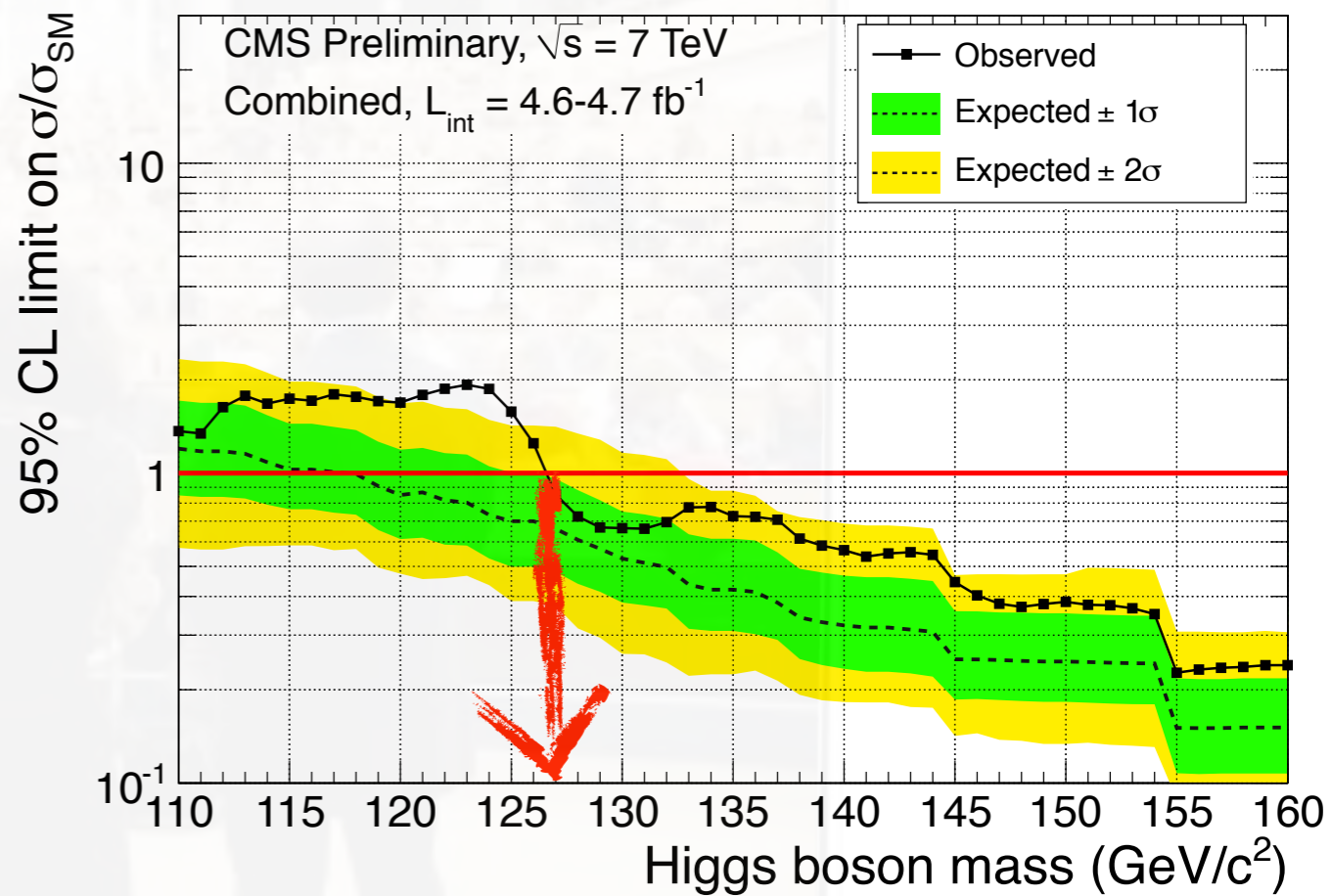
# Higgs bounds: news from last December

ATLAS-CONF-2011-163

CMS PAS HIG-11-032



$\approx 131 \text{ GeV}$

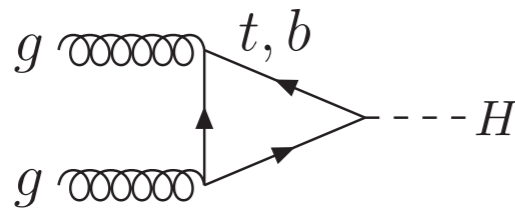


$\approx 127 \text{ GeV}$

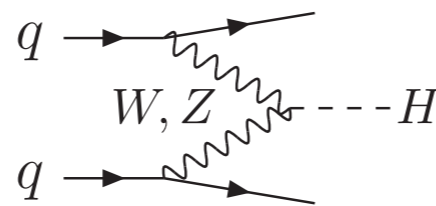
# Rescaling Higgs Searches

- Higgs couplings modified w.r.t. SM but same kinematics
- Background processes unaffected

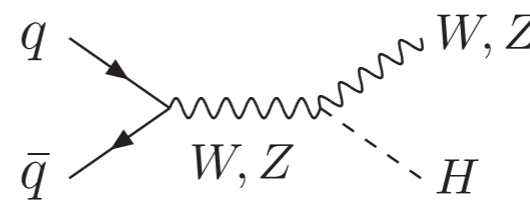
$\Downarrow$                        $\Downarrow$                        $\Downarrow$   
 simple rescaling of SM searches



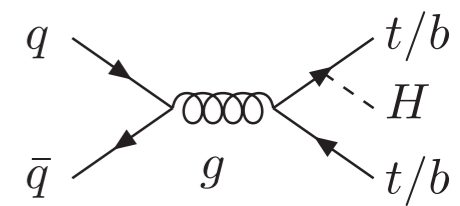
$c^2$



$a^2$



$a^2$



$c^2$

$$\frac{\sigma_{NLO}}{\sigma_{SM}} \frac{\sigma_{SM}}{\sigma_{NLO}}$$

$$\Gamma(H \rightarrow f\bar{f}) = c^2 \Gamma^{SM}(H \rightarrow f\bar{f}),$$

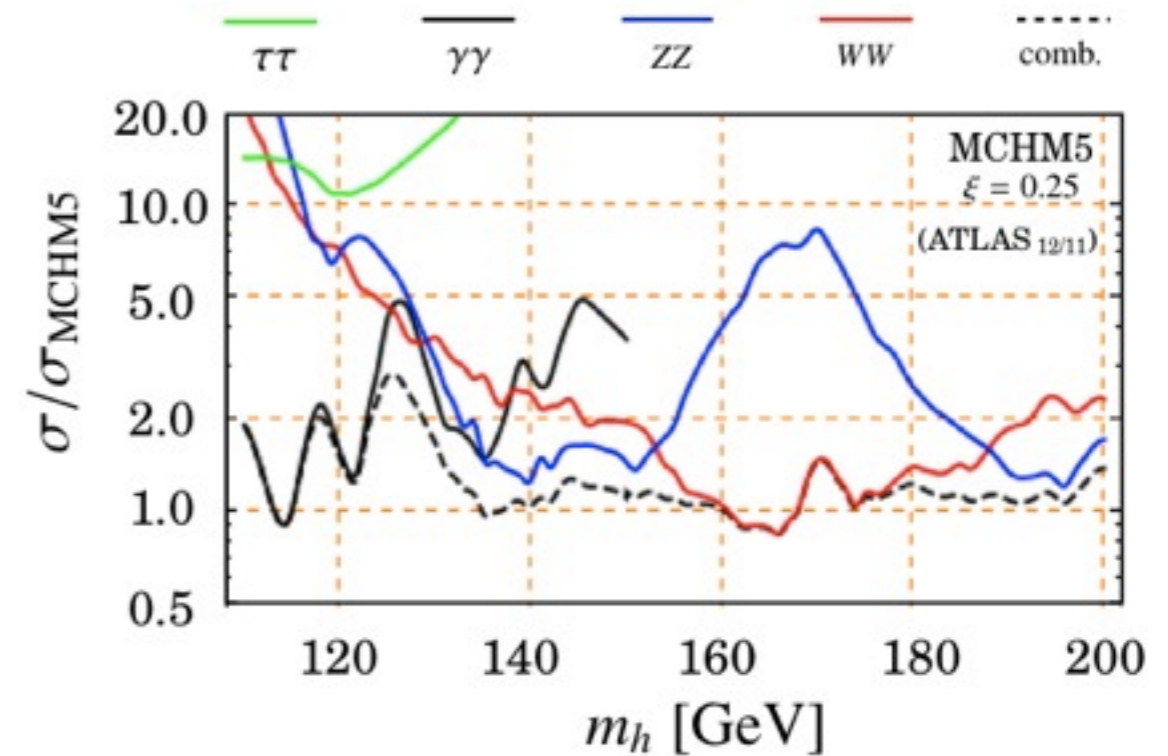
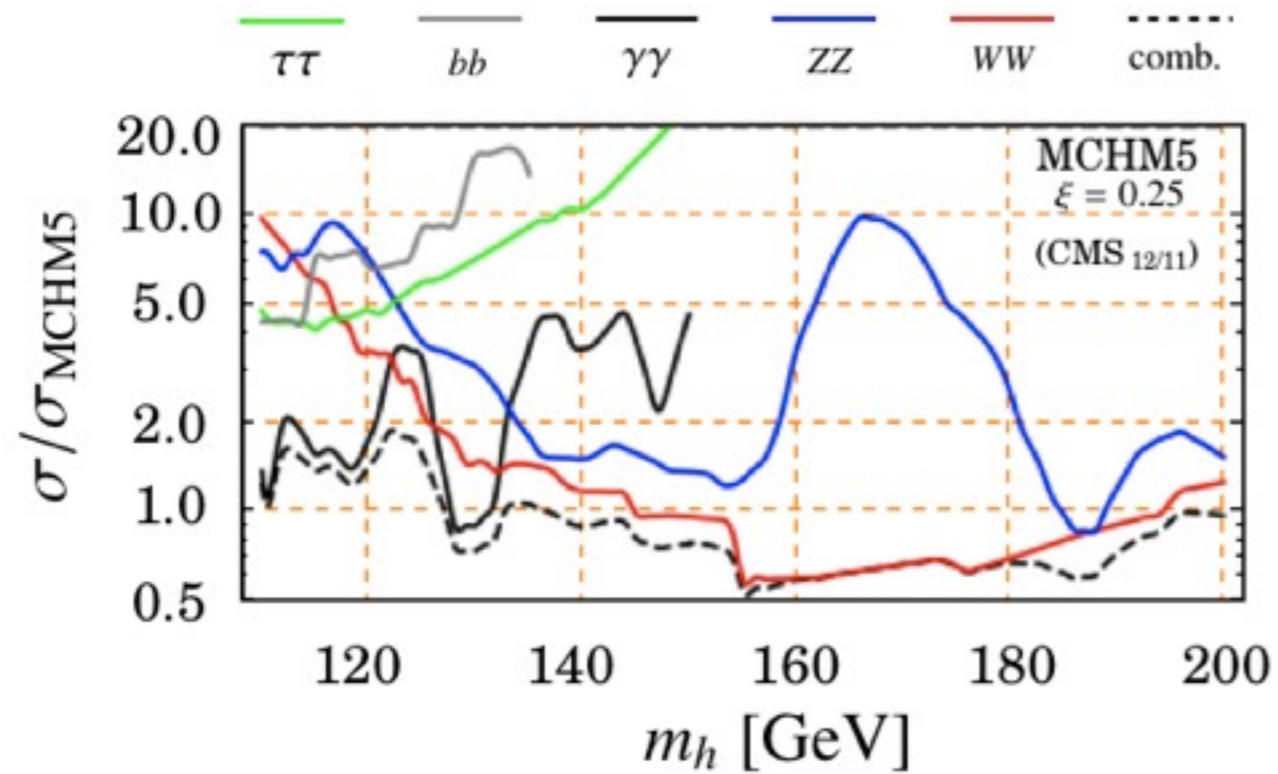
$$\Gamma(H \rightarrow VV) = a^2 \Gamma^{SM}(H \rightarrow VV),$$

$$\Gamma(H \rightarrow gg) = c^2 \Gamma^{SM}(H \rightarrow gg),$$

$$\Gamma(H \rightarrow \gamma\gamma) = \frac{(cI_\gamma + aJ_\gamma)^2}{(I_\gamma + J_\gamma)^2} \Gamma^{SM}(H \rightarrow \gamma\gamma),$$

# Rescaling Higgs Searches

Espinosa, Grojean, Muehlleitner '10



each search channel is rescaled individually  
all the channels are then combined

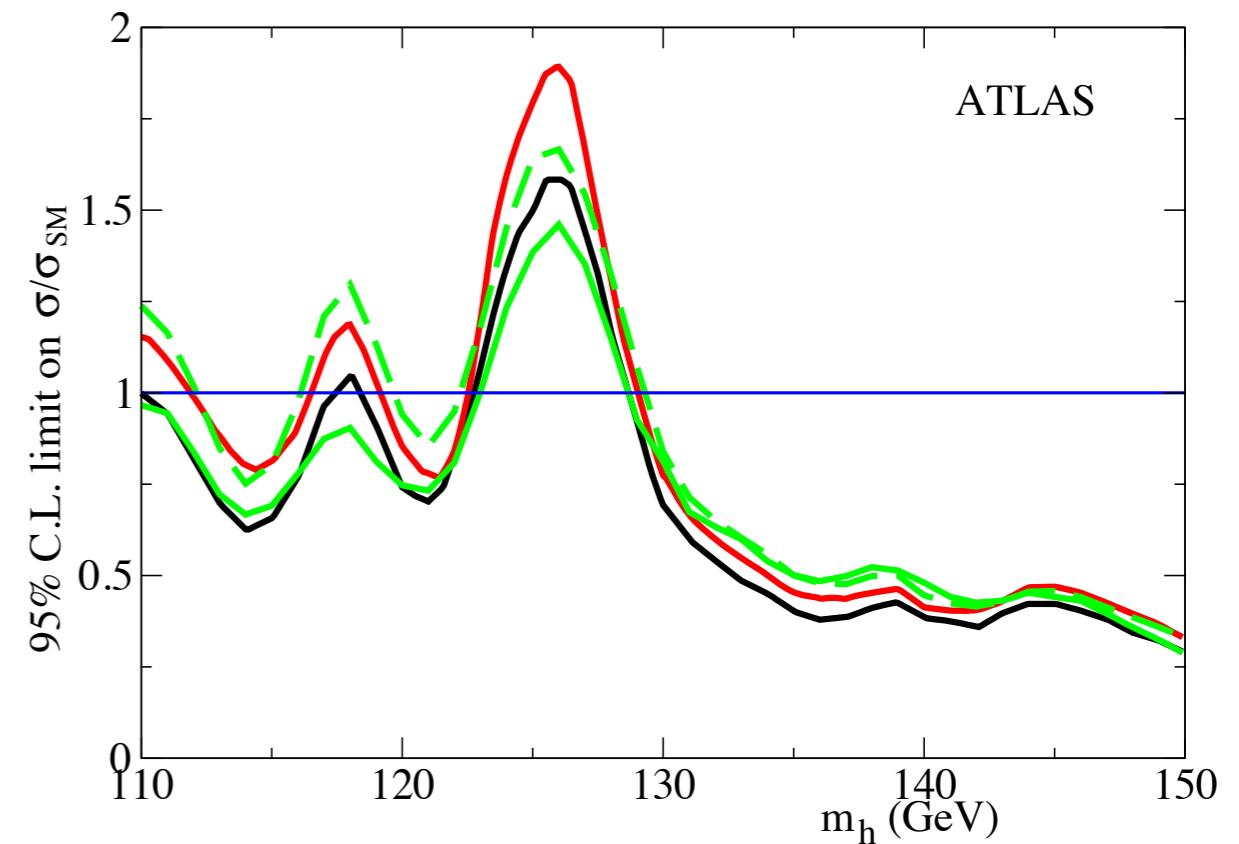
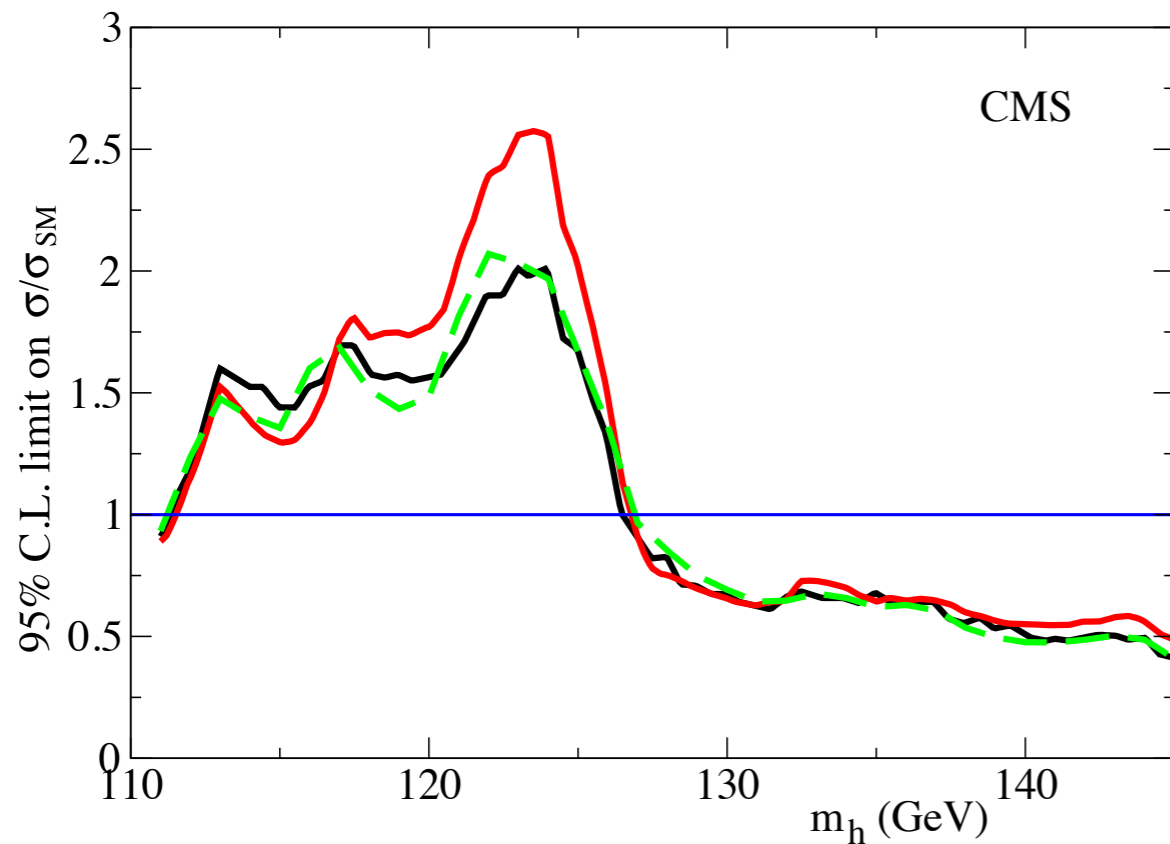
# Rescaling Higgs Searches

Espinosa, Grojean, Muehlleitner '10

Espinosa, Grojean, Muehlleitner, Trott '12

## How robust is our TH combination?

### Let's look at the SM ( $a=c=1$ )



official

improved combination  
in quadrature

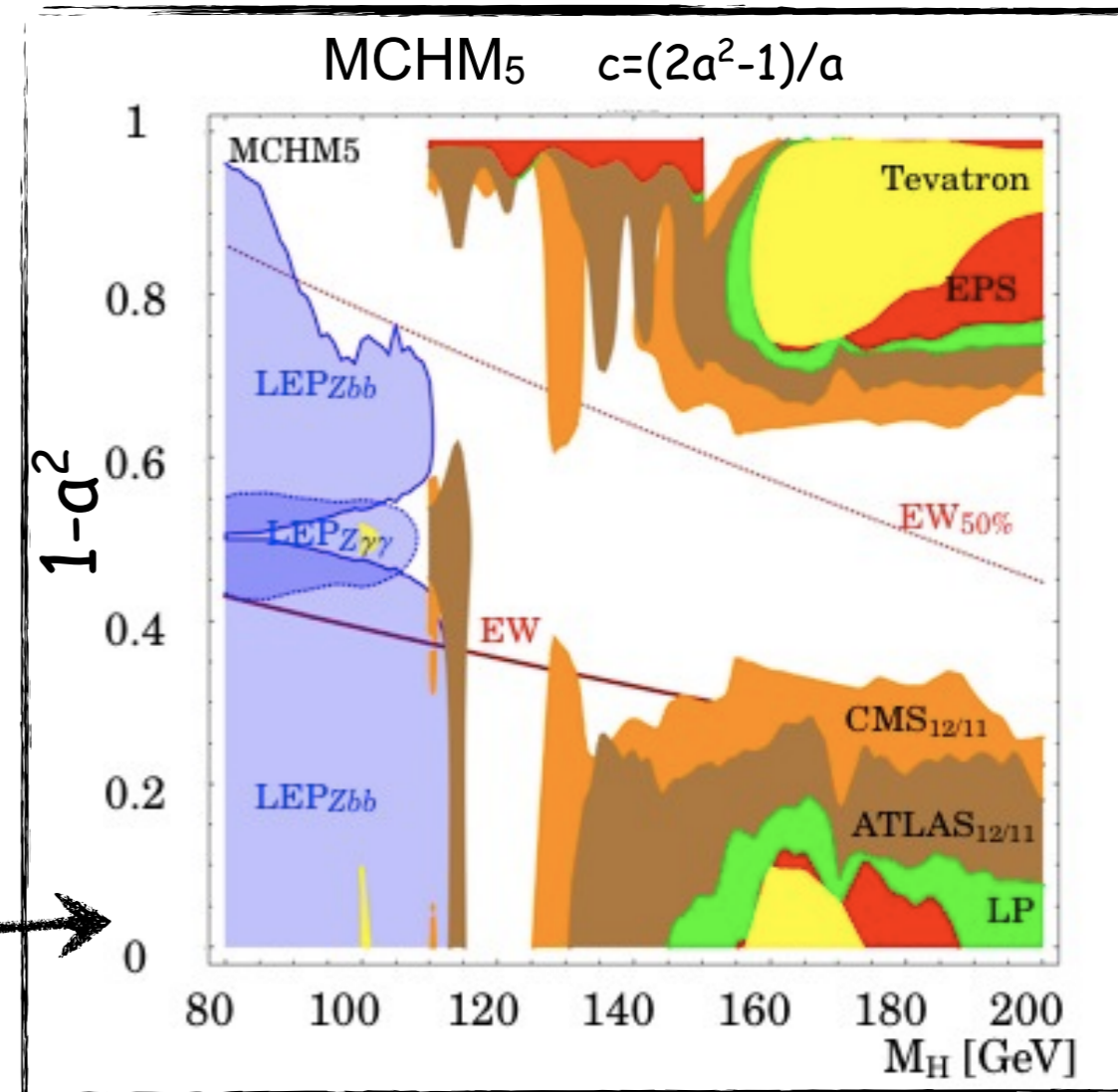
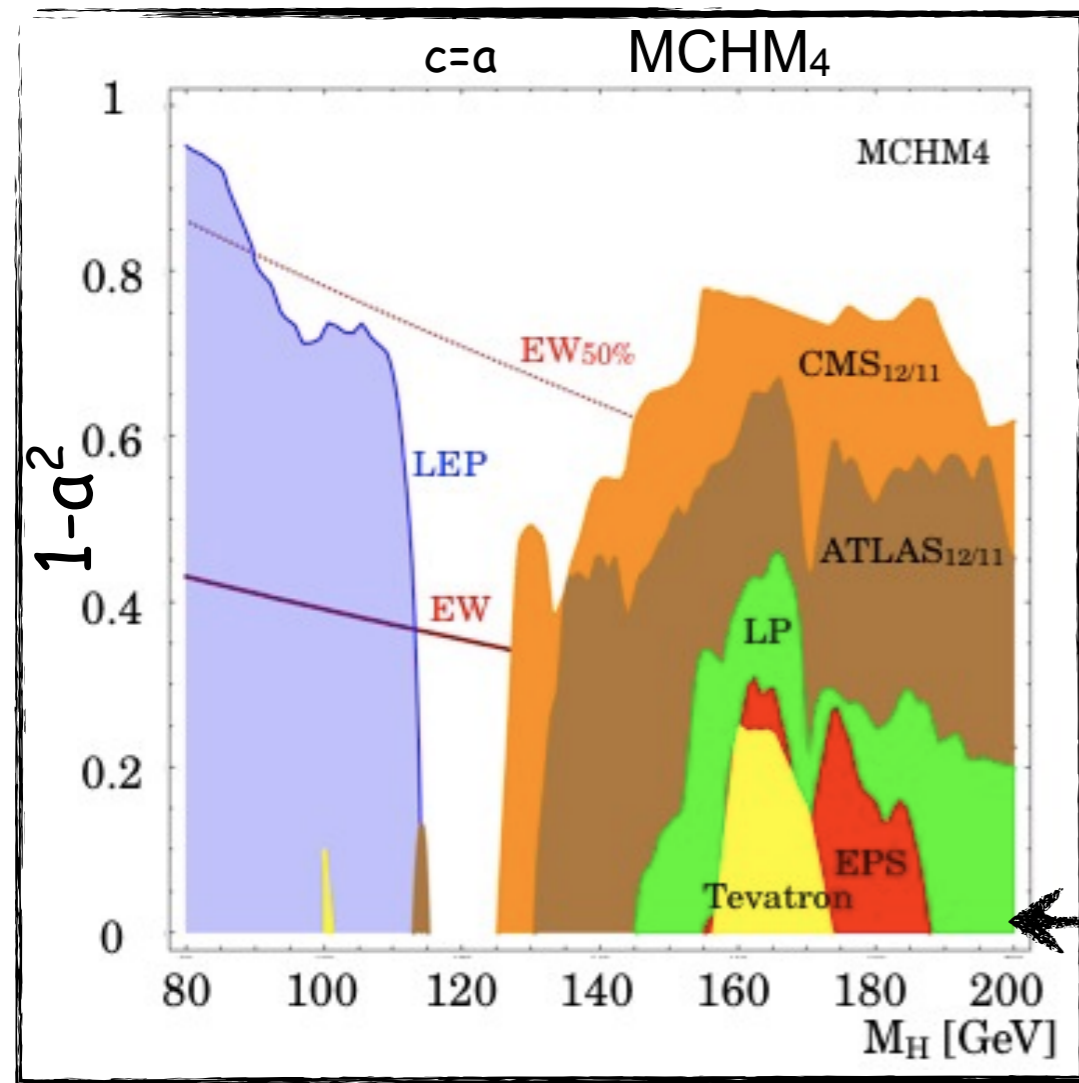
improved combination  
with reconstructed likelihood

See also [Azatov, Contino, Galloway '12](#)

# Deformation of the SM Higgs: current constraints

the SM exclusion bounds are easily rescaled in the  $(m_H, a)$  plane

Espinosa, Grojean, Muehlleitner '11



← SM limits →

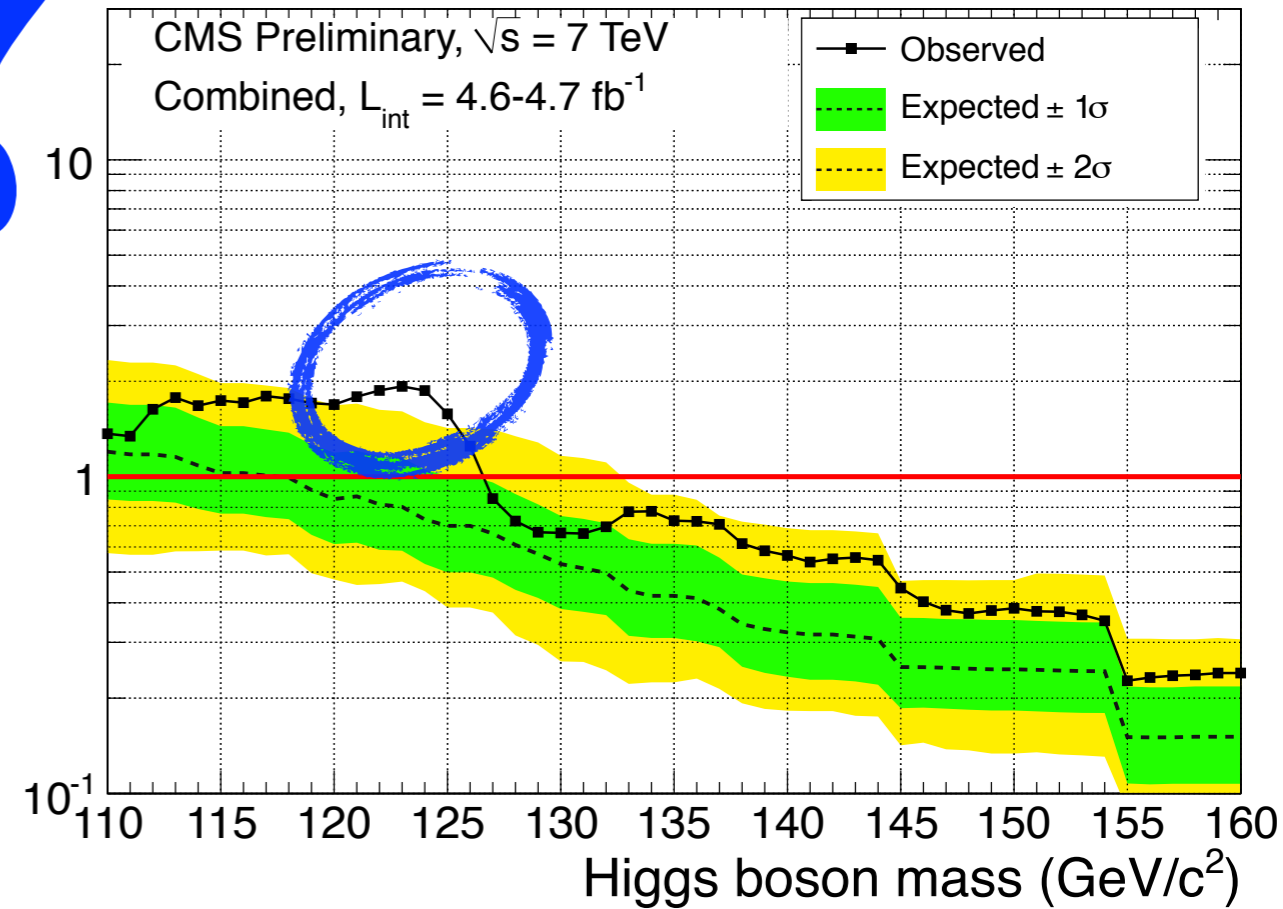
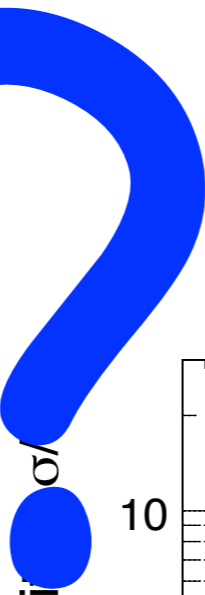
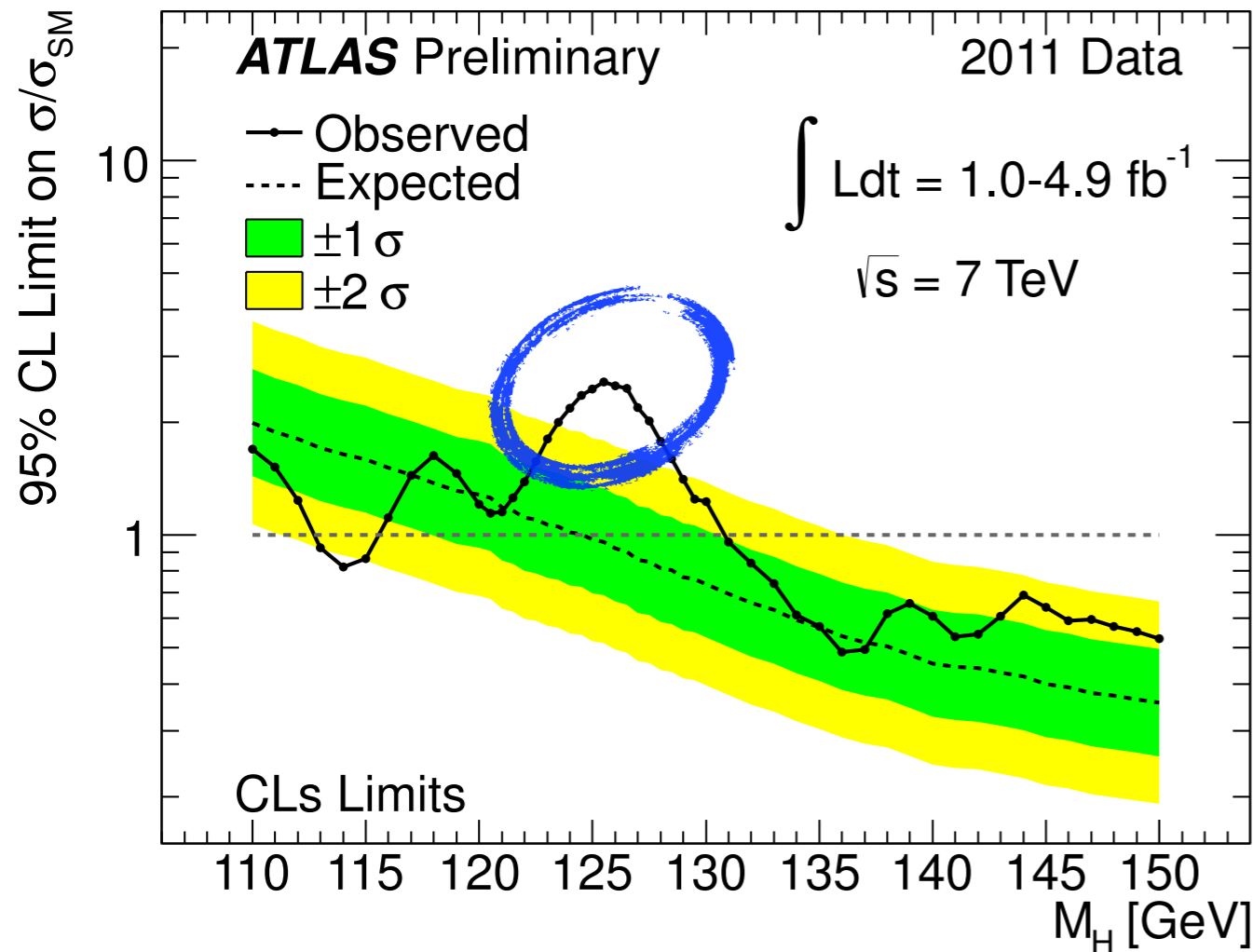
LHC tsunami!

the LHC can do much more than simply excluding the SM Higgs

# Higgs bounds: news from last December

ATLAS-CONF-2011-163

CMS PAS HIG-11-032



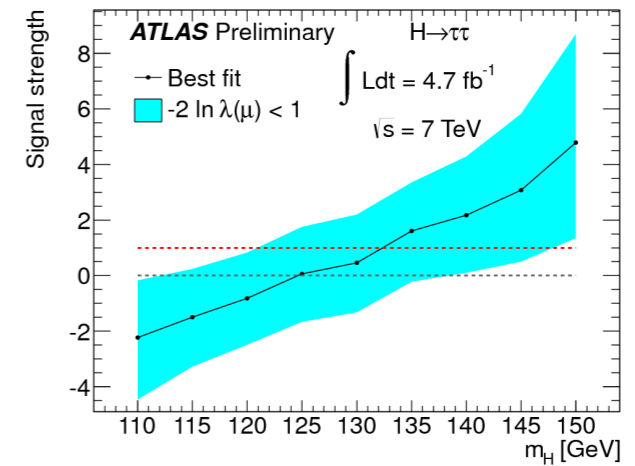
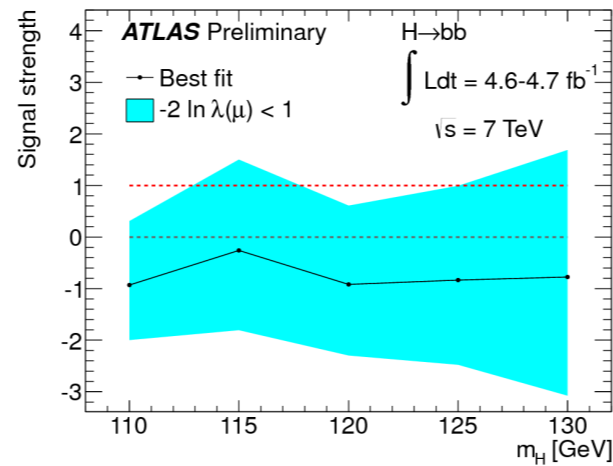
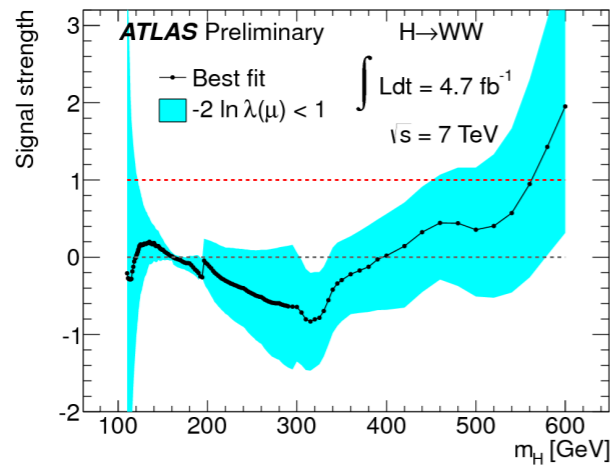
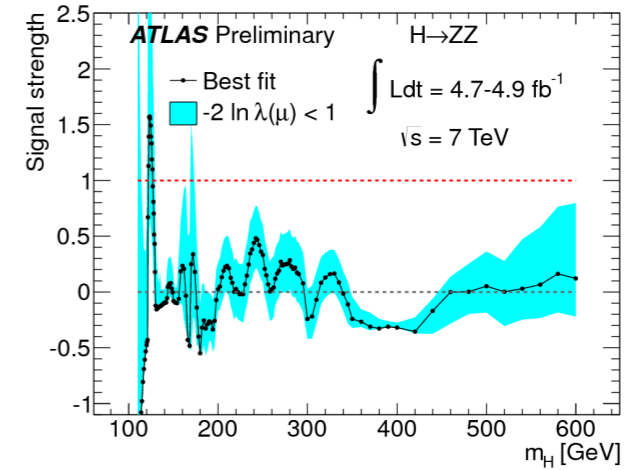
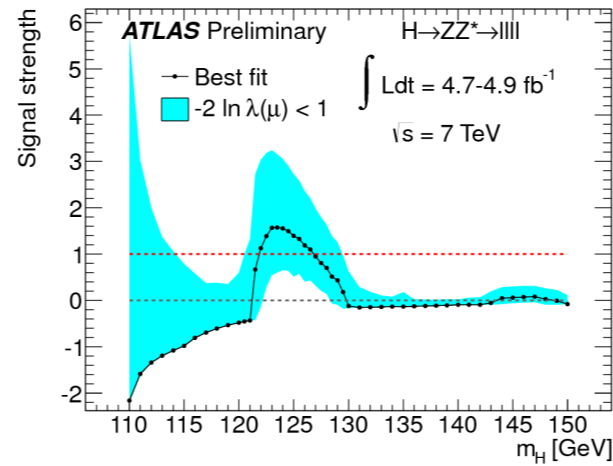
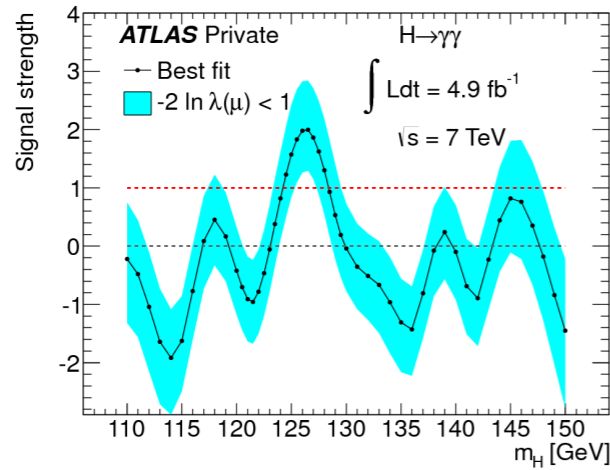
a 120-130 GeV higgs is very interesting (from the exp. point of view)  
 since many competing decay channels



# Various Search Channels

signal strength

$$\mu = \frac{\sigma \times BR}{(\sigma \times BR)_{SM}}$$



# Various Search Channels

signal strength  $\mu = \frac{\sigma \times BR}{(\sigma \times BR)_{SM}}$

Channel [Exp]	$m_h$ [GeV] (Local Significance)	$\mu$ ( $\mu_L$ )	Scaling to SM
$pp \rightarrow \gamma\gamma$ [ATLAS]	$126.5 \pm 0.7$ (2.8 $\sigma$ ) [26]	$2_{-0.7}^{+0.9}$ [27] (2.6)	$\sim c^2 \text{Br}_{\gamma\gamma}[a, c]$
$pp \rightarrow Z Z^* \rightarrow \ell^+ \ell^- \ell^+ \ell^-$ [ATLAS]	$126 \pm \sim 2\%$ (2.1 $\sigma$ ) [26]	$1.2_{-0.8}^{+1.2}$ [27] (4.9)	$\sim c^2 \text{Br}_{ZZ}[a, c]$
$pp \rightarrow W W^* \rightarrow \ell^+ \nu \ell^- \bar{\nu}$ [ATLAS]	$126 \pm \sim 20\%$ (1.4 $\sigma$ ) [26]	$1.2_{-0.8}^{+0.8}$ [27] (3.4)	$\sim c^2 \text{Br}_{WW}[a, c]$
$pp \rightarrow \gamma\gamma jj$ [CMS]	$124 \pm 3\%$ [10, 11]	$3.7_{-1.8}^{+2.5}$ [11]	$\sim a^2 \text{Br}_{\gamma\gamma}[a, c]$
$pp \rightarrow \gamma\gamma$ [CMS, b, $R_9^{\min} > 0.94$ ]	$124 \pm 3\%$ [10, 11]	$1.5_{-1.0}^{+1.1}$ [11]	$\sim c^2 \text{Br}_{\gamma\gamma}[a, c]$
$pp \rightarrow \gamma\gamma$ [CMS, b, $R_9^{\min} < 0.94$ ]	$124 \pm 3\%$ [10, 11]	$2.1_{-1.4}^{+1.5}$ [11]	$\sim c^2 \text{Br}_{\gamma\gamma}[a, c]$
$pp \rightarrow \gamma\gamma$ [CMS, e, $R_9^{\min} > 0.94$ ]	$124 \pm 3\%$ [10, 11]	$0.0^{+2.9}$ [11]	$\sim c^2 \text{Br}_{\gamma\gamma}[a, c]$
$pp \rightarrow \gamma\gamma$ [CMS, e, $R_9^{\min} < 0.94$ ]	$124 \pm 3\%$ [10, 11]	$4.1_{-4.1}^{+4.6}$ [11]	$\sim c^2 \text{Br}_{\gamma\gamma}[a, c]$
$pp \rightarrow Z Z^* \rightarrow \ell^+ \ell^- \ell^+ \ell^-$ [CMS]	$126 \pm 2\%$ (1.5 $\sigma$ ) [11, 28]	$0.5_{-0.7}^{+1.0}$ [10] (2.7)	$\sim c^2 \text{Br}_{ZZ}[a, c]$
$pp \rightarrow W W^* \rightarrow \ell^+ \nu \ell^- \bar{\nu}$ [CMS]	$126 \pm 20\%$ [10, 29]	$0.7_{-0.6}^{+0.4}$ [10] (1.8)	$\sim c^2 \text{Br}_{WW}[a, c]$
$pp \rightarrow b\bar{b}$ [CMS]	$124 \pm 10\%$ [10]	$1.2_{-1.7}^{+1.4}$ [10] (4.1)	$\sim a^2 \text{Br}_{b\bar{b}}[a, c]$
$pp \rightarrow \tau\bar{\tau}$ [CMS]	$124 \pm 20\%$ [10]	$0.8_{-1.7}^{+1.2}$ [10] (3.3)	$\sim c^2 \text{Br}_{\tau\bar{\tau}}[a, c]$

# Various Search Channels (after Moriond)

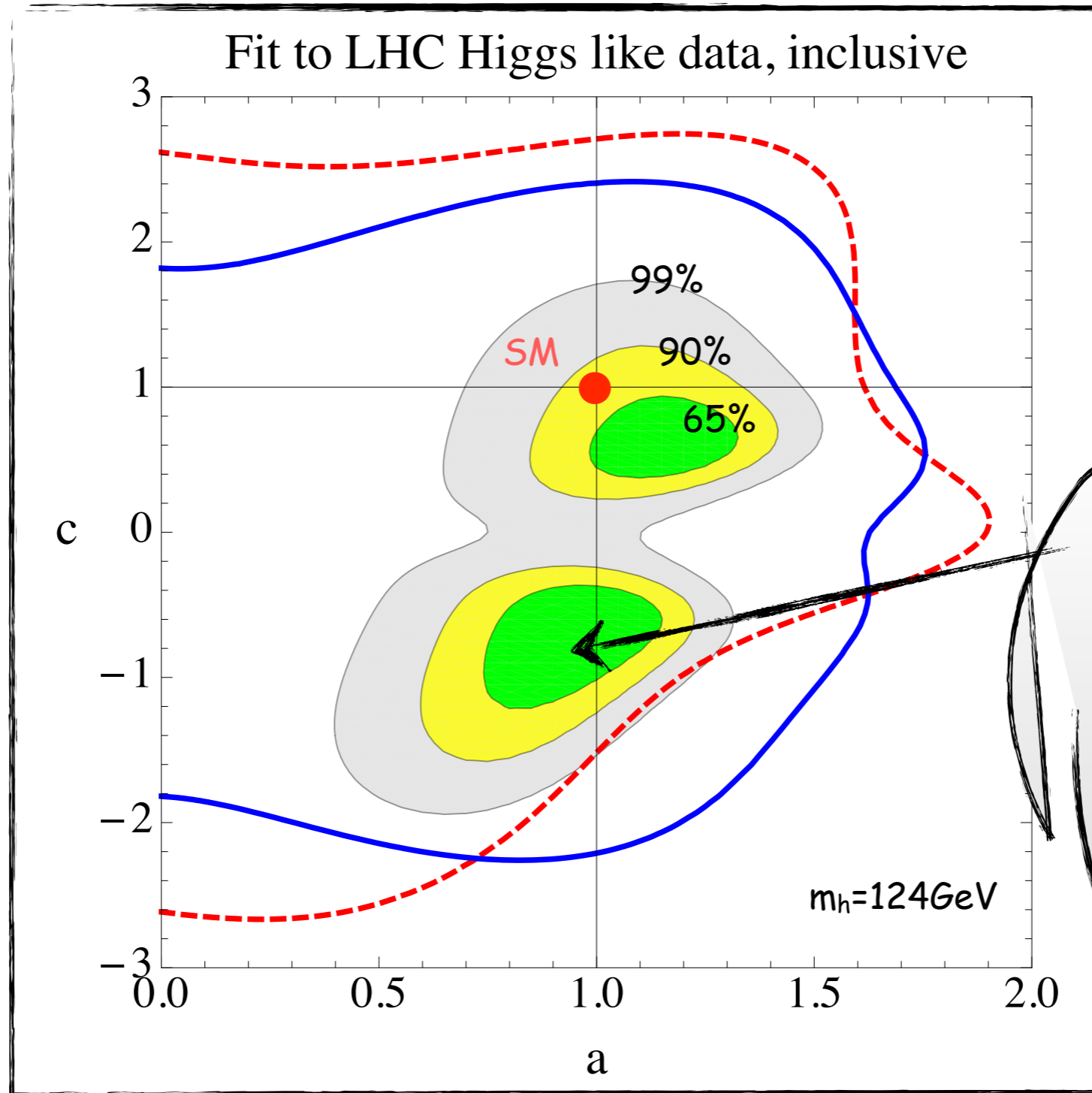
signal strength  $\mu = \frac{\sigma \times BR}{(\sigma \times BR)_{SM}}$

Channel [Exp]	$\mu_{119.5} (\mu_{119.5}^L)$	$\mu_{124} (\mu_{124}^L)$	$\mu_{125} (\mu_{125}^L)$
$pp \rightarrow \gamma \gamma$ [ATLAS]	$0.0_{-0.8}^{+0.6}$ (1.5)	$0.8_{-0.7}^{+0.8}$ (2.6)	$1.6_{-0.8}^{+0.9}$ (3.9)
$pp \rightarrow Z Z^* \rightarrow \ell^+ \ell^- \ell^+ \ell^-$ [ATLAS]	$-0.5^{+0.5??}$ (5.1)	$1.6_{-0.8}^{+1.4}$ (4.7)	$1.4_{-0.8}^{+1.3}$ (4.1)
$pp \rightarrow W W^* \rightarrow \ell^+ \nu \ell^- \bar{\nu}$ [ATLAS]	$0.0_{-1.3}^{+1.2}$ (2.4)	$0.1_{-0.7}^{+0.7}$ (1.6)	$0.1_{-0.6}^{+0.7}$ (1.4)
$pp \rightarrow \gamma \gamma$ [CMS]	$-1.1_{-0.6}^{+0.6}$ (1.3)	$1.5_{-0.7}^{+0.7}$ (3.5)	$1.6_{-0.6}^{+0.7}$ (3.0)
$pp \rightarrow Z Z^* \rightarrow \ell^+ \ell^- \ell^+ \ell^-$ [CMS]	$2.0_{-1.1}^{+1.6}$ (5.2)	$0.5_{-0.7}^{+1.1}$ (2.7)	$0.6_{-0.6}^{+0.9}$ (2.5)
$pp \rightarrow W W^* \rightarrow \ell^+ \nu \ell^- \bar{\nu}$ [CMS]	$0.9_{-0.7}^{+0.8}$ (2.5)	$0.6_{-0.7}^{+0.7}$ (1.8)	$0.4_{-0.6}^{+0.6}$ (1.5)
$pp \rightarrow b \bar{b}$ [CMS]	$0.4_{-1.6}^{+1.8}$ (4.1)	$1.2_{-1.8}^{+1.9}$ (5.0)	$1.2_{-1.7}^{+2.1}$ (5.2)
$pp \rightarrow \tau \bar{\tau}$ [CMS]	$0.2_{-1.1}^{+0.9}$ (3.6)	$0.4_{-1.2}^{+1.0}$ (3.9)	$0.6_{-1.2}^{+1.1}$ (4.1)
$pp \rightarrow \tau \bar{\tau}$ [ATLAS]	$-0.9_{-1.7}^{+1.7}$ (2.9)	$-0.1_{-1.8}^{+1.7}$ (3.4)	$0.1_{-1.8}^{+1.7}$ (3.5)
$p\bar{p} \rightarrow b \bar{b}$ [CDF&DØ]	$1.5_{-0.5}^{+0.6}$ (2.5)	$1.9_{-0.6}^{+0.8}$ (3.1)	$2.0_{-0.7}^{+0.8}$ (3.2)

Espinosa, Grojean, Muhlleitner, Trott '12

# Model independent fit to LHC data

Espinosa, Grojean, Muhlleitner, Trott '12



note: a fermiophobic Higgs is disfavored by data (mostly VBF channels)

"disfermiophilia"

the current data prefers "negative" coupling to fermions  
 $\approx$   
 positive interference between top and W in  $\gamma\gamma$  channel

Atlas 95%CL exclusion

CMS 95%CL exclusion

SM 82%CL away from best fit point

Two minima:

$(a,c) = (1.13, 0.58)$   
 $\chi^2 = 2.86$

$(a,c) = (0.96, -0.64)$   
 $\chi^2 = 1.96$

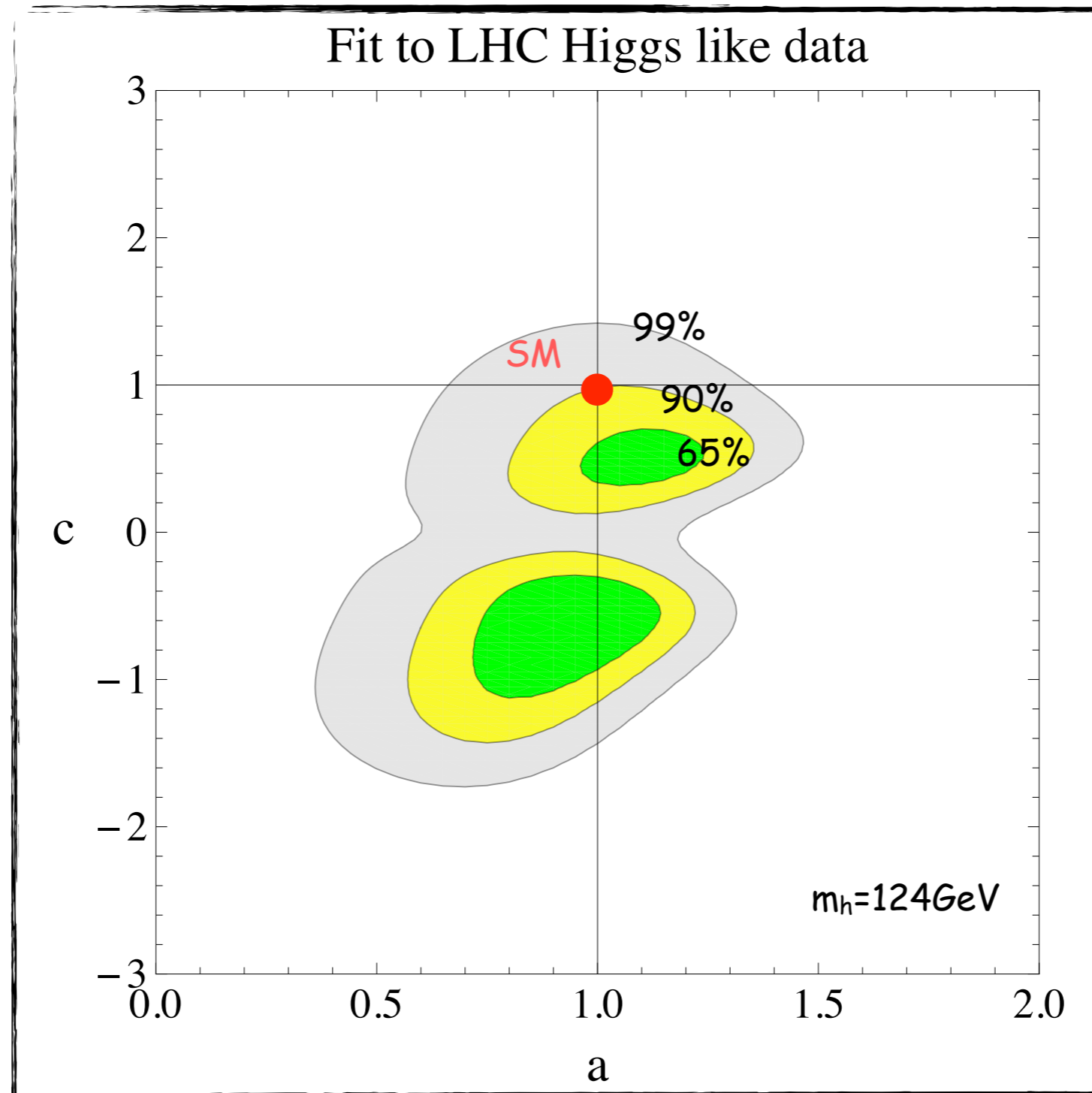
for similar analyses, see also

Azatov, Contino, Galloway '12

Carni, Falkowski, Kuflik, Volansky '12

# Model independent fit to (Moriond) LHC data

Espinosa, Grojean, Muhlleitner, Trott '12



note: a fermiophobic Higgs is disfavored by data (mostly VBF channels) at 96%CL

-----

Atlas 95%CL exclusion

—————

CMS 95%CL exclusion

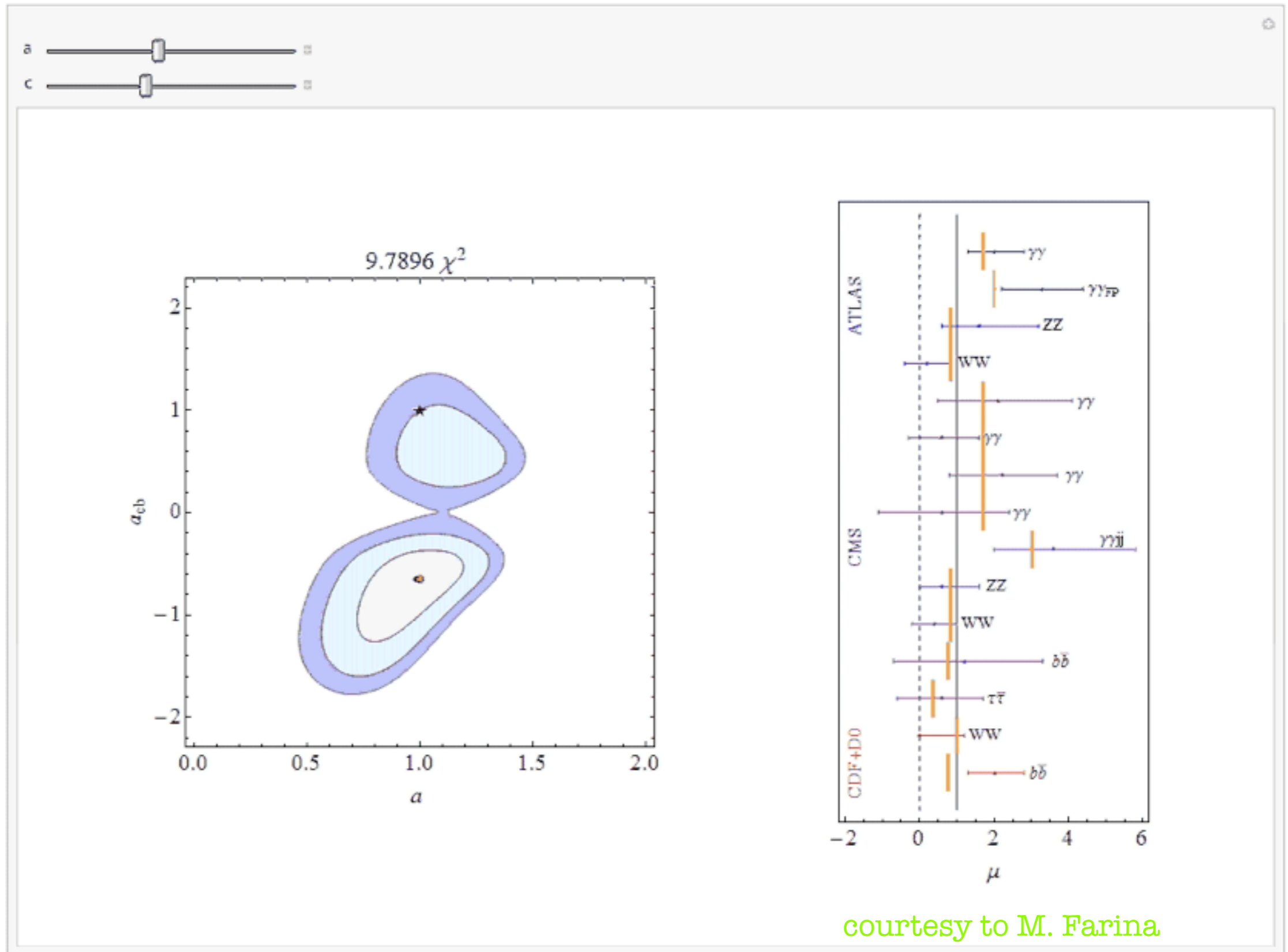
SM 94%CL  
away from  
best fit point

Azatov, Contino, Galloway '12

for similar analyses, see also

Carni, Falkowski, Kuflik, Volansky '12

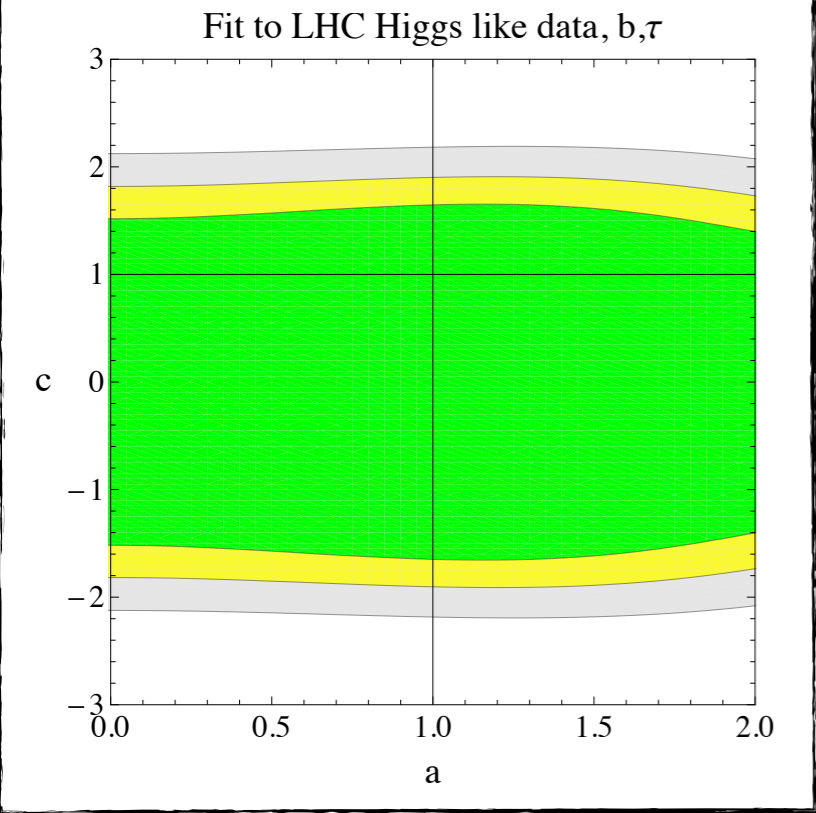
# Model independent fit to LHC data



# Which are the channels driving the fit?

Espinosa, Grojean, Muhlleitner, Trott '12

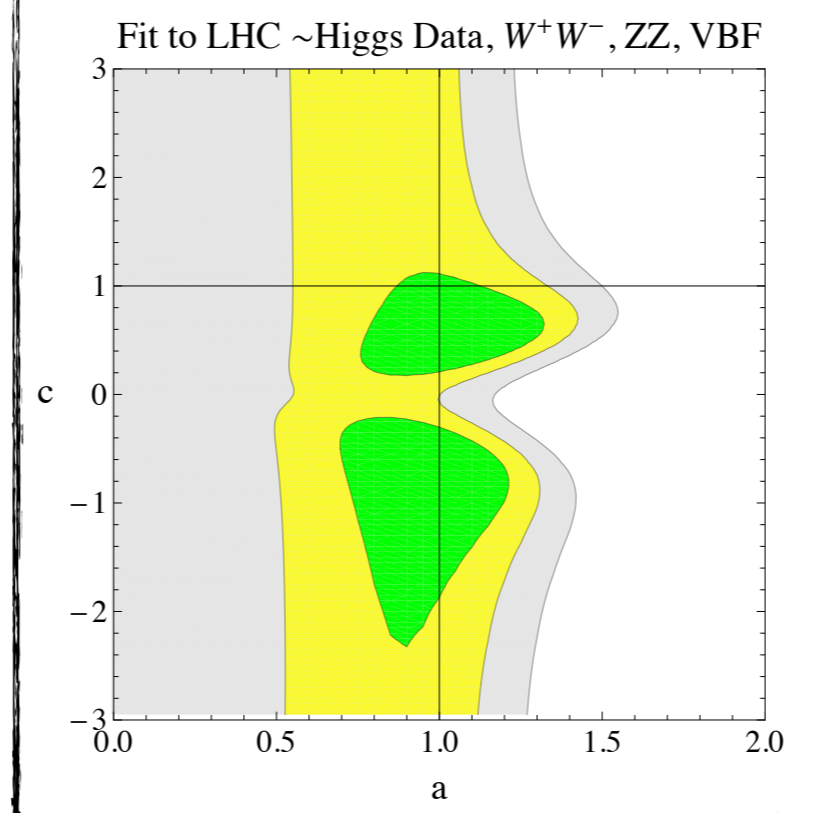
## fermion couplings



almost

no constraints at the moment

## gauge couplings

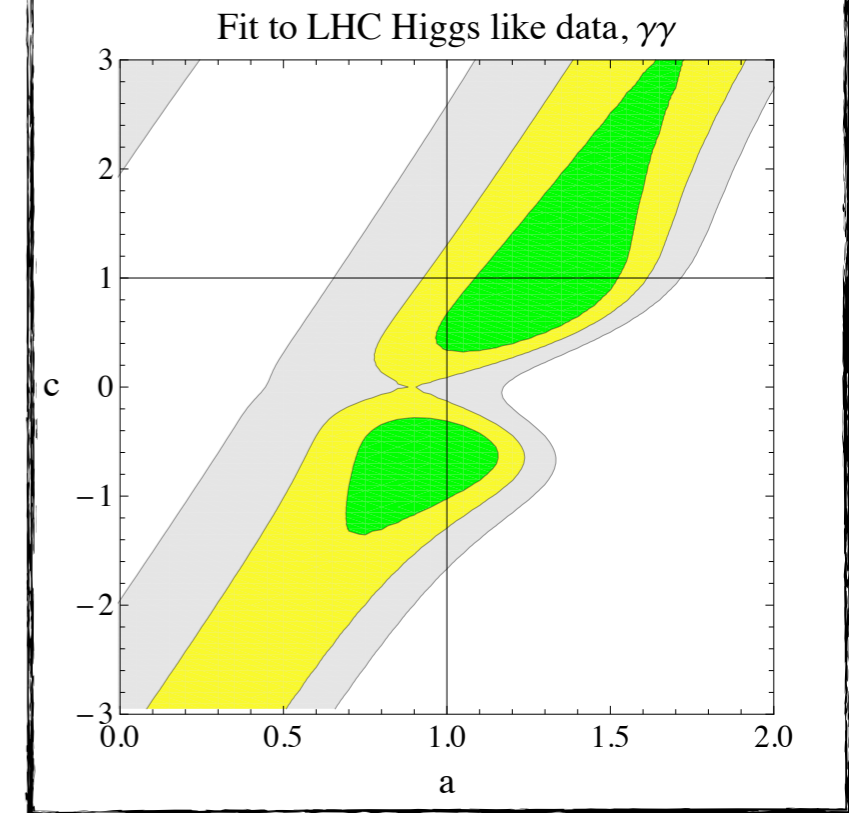


almost

$(a,c) \leftrightarrow (a,-c)$  symmetric

large  $a$  are disfavored

## both couplings

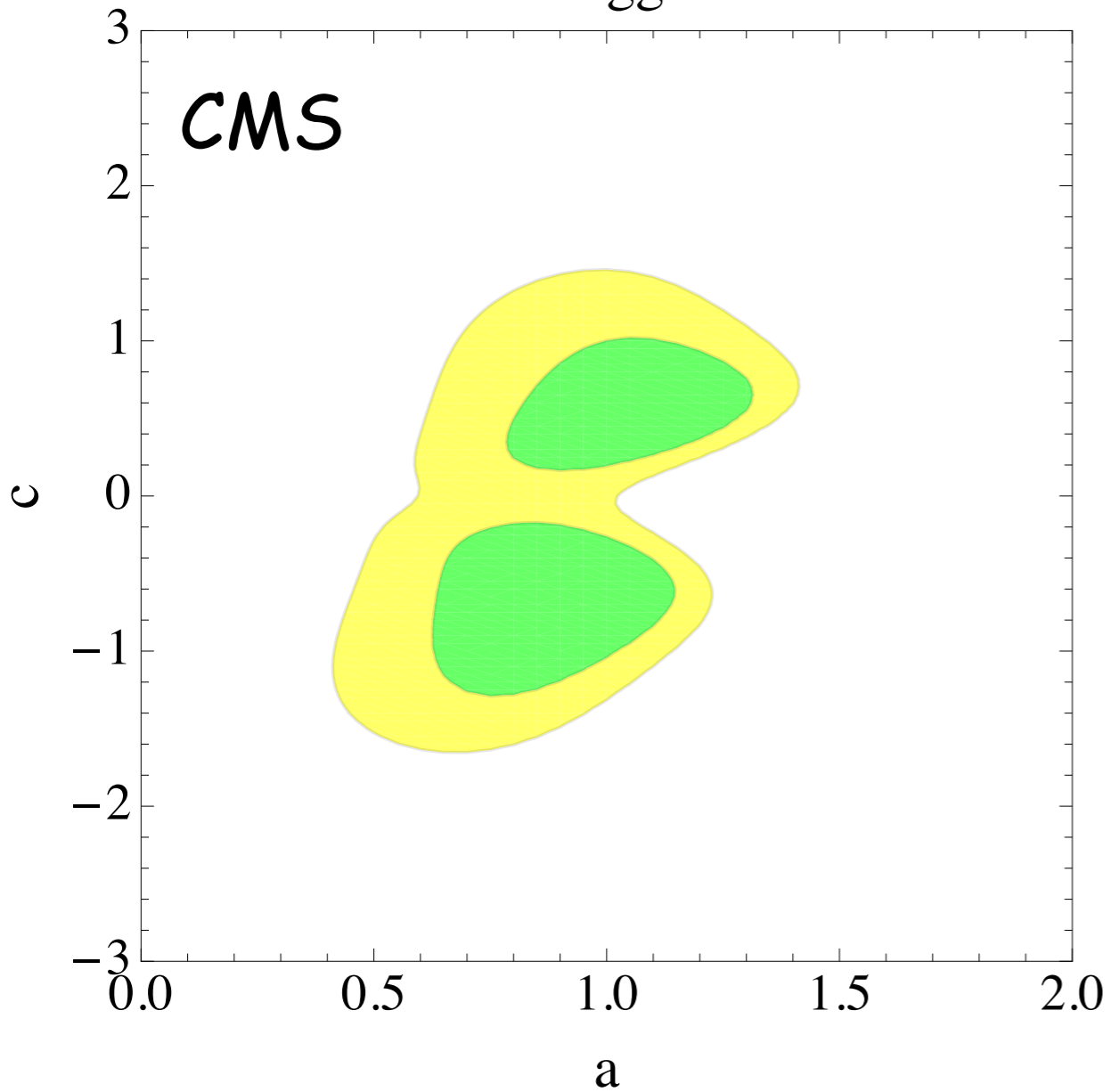


most constraining

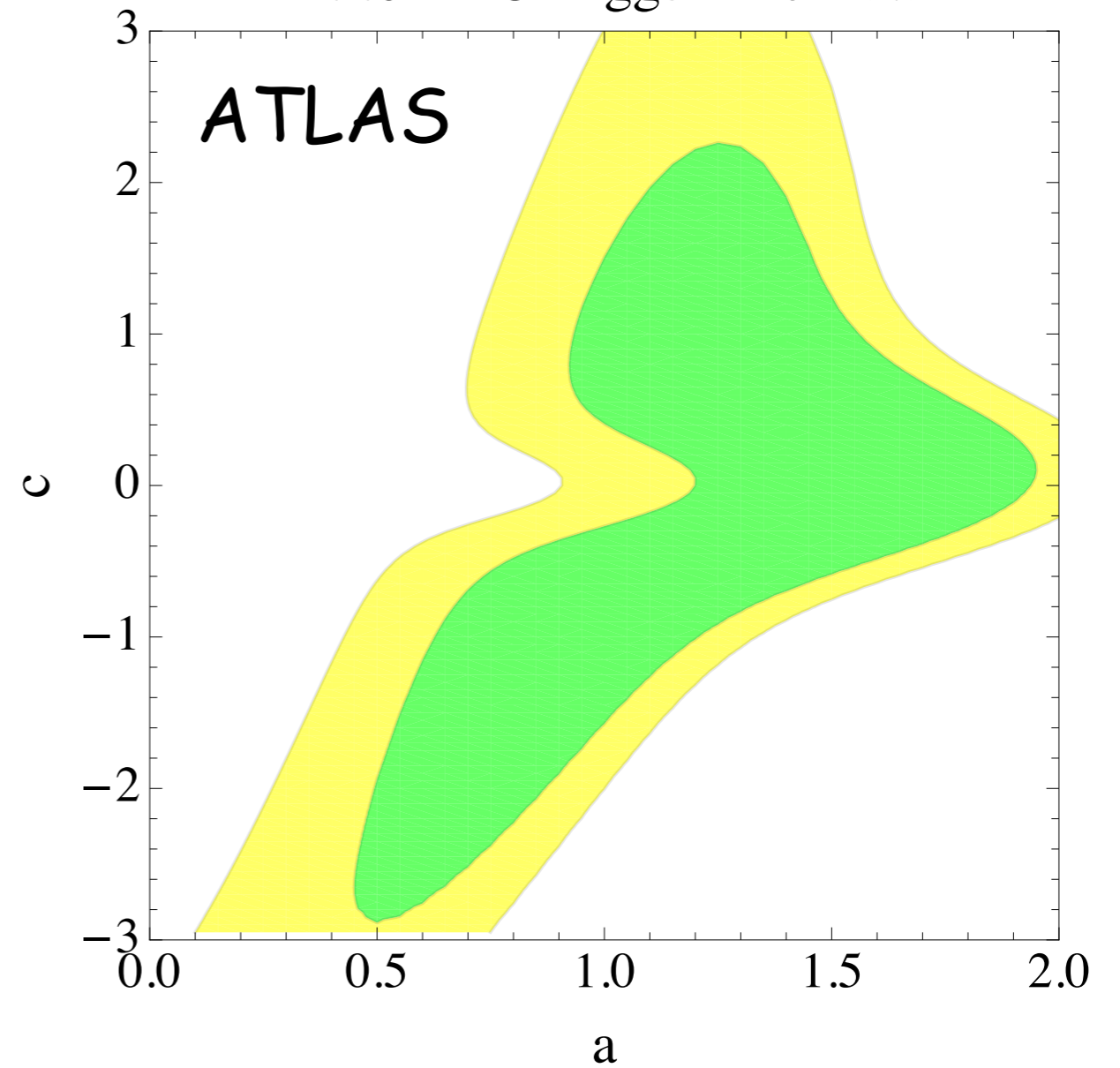
**note:** even if  $\Gamma(h \rightarrow \gamma\gamma)$  is not really modified (no operator  $|H|^2 B_{\mu\nu} B^{\mu\nu}$ ),  $BR_{\gamma\gamma}$  has strong dependence on 'a' and 'c'

# CMS vs ATLAS

Fit to LHC Higgs-like Data



Fit to LHC Higgs-like Data

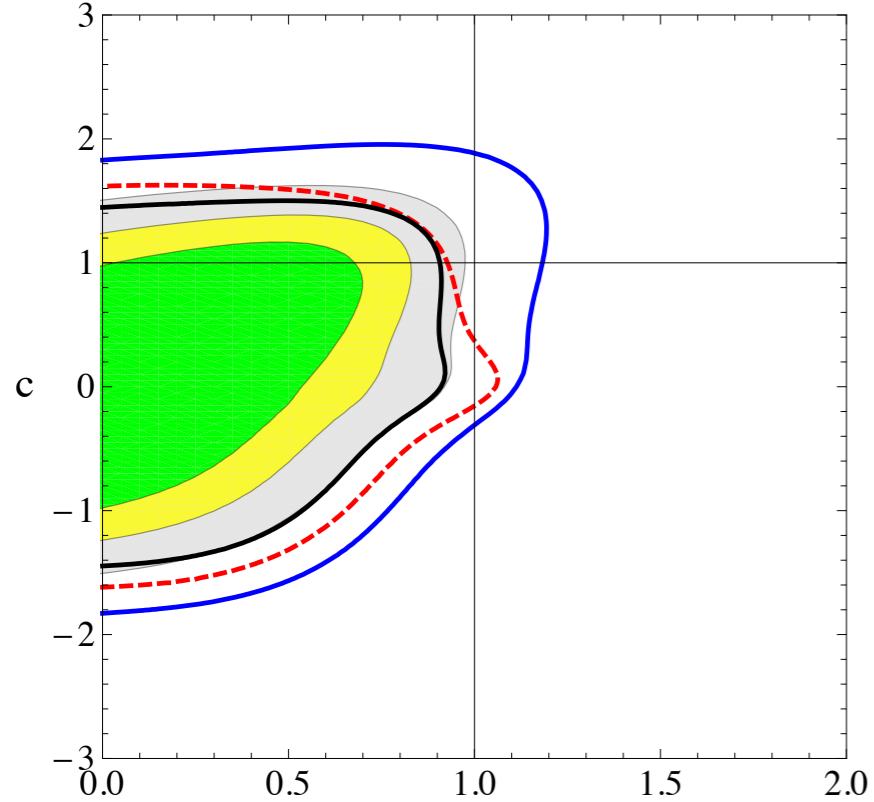


Espinosa, Grojean, Muhlleitner, Trott '12

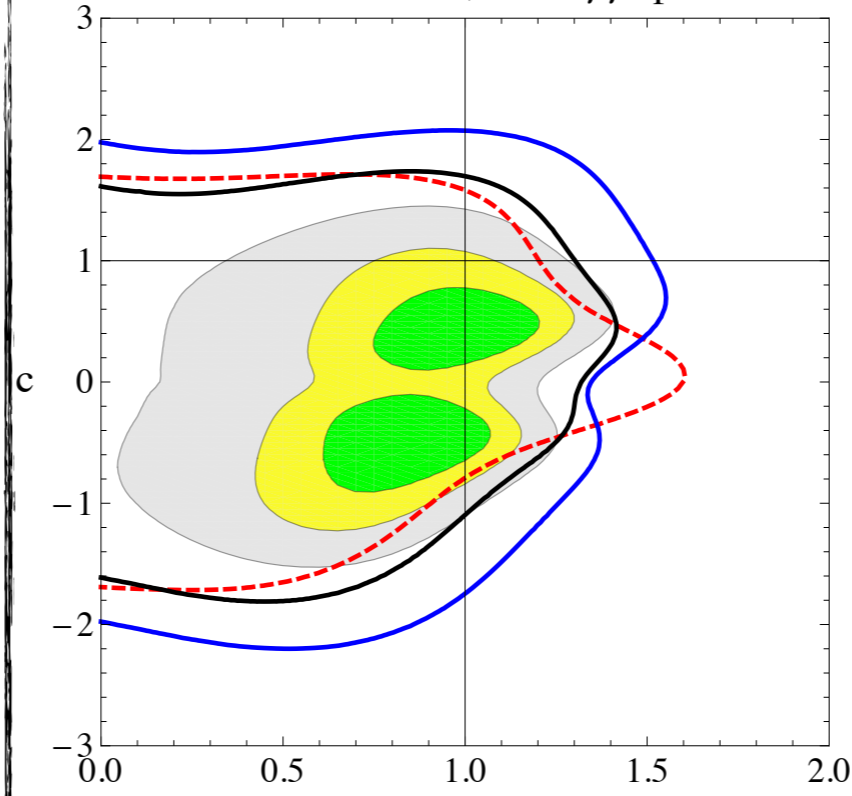


# Which Higgs mass?

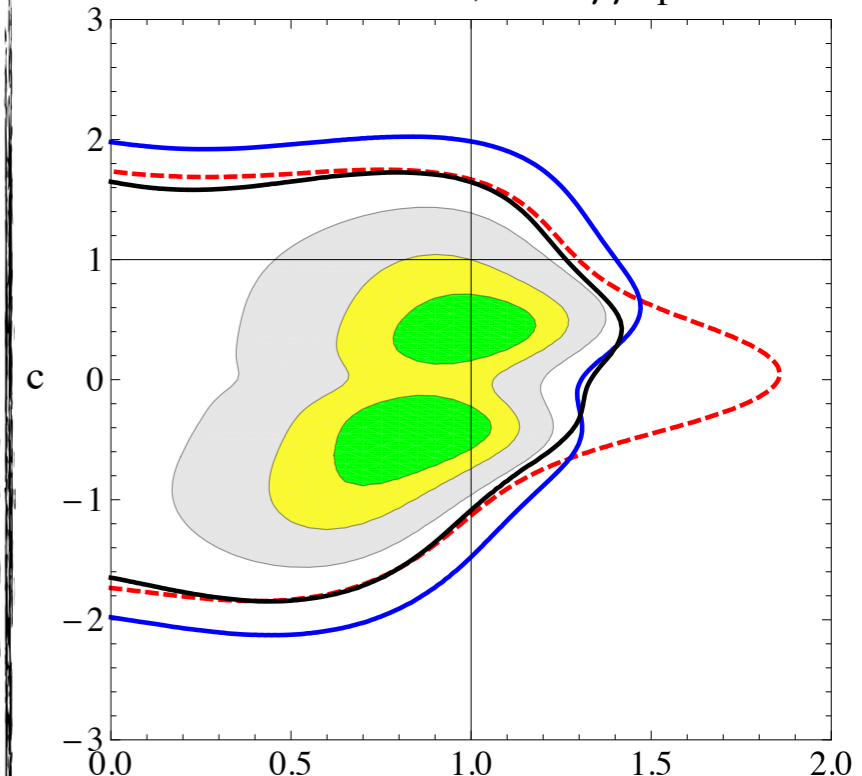
mh = 119.5 GeV



mh = 124 GeV, CMS  $\gamma\gamma$  split



mh = 125 GeV, CMS  $\gamma\gamma$  split



(119.5, 0.28, 0.49, 9.8)

(124, 0.87, -0.43, 4.1)

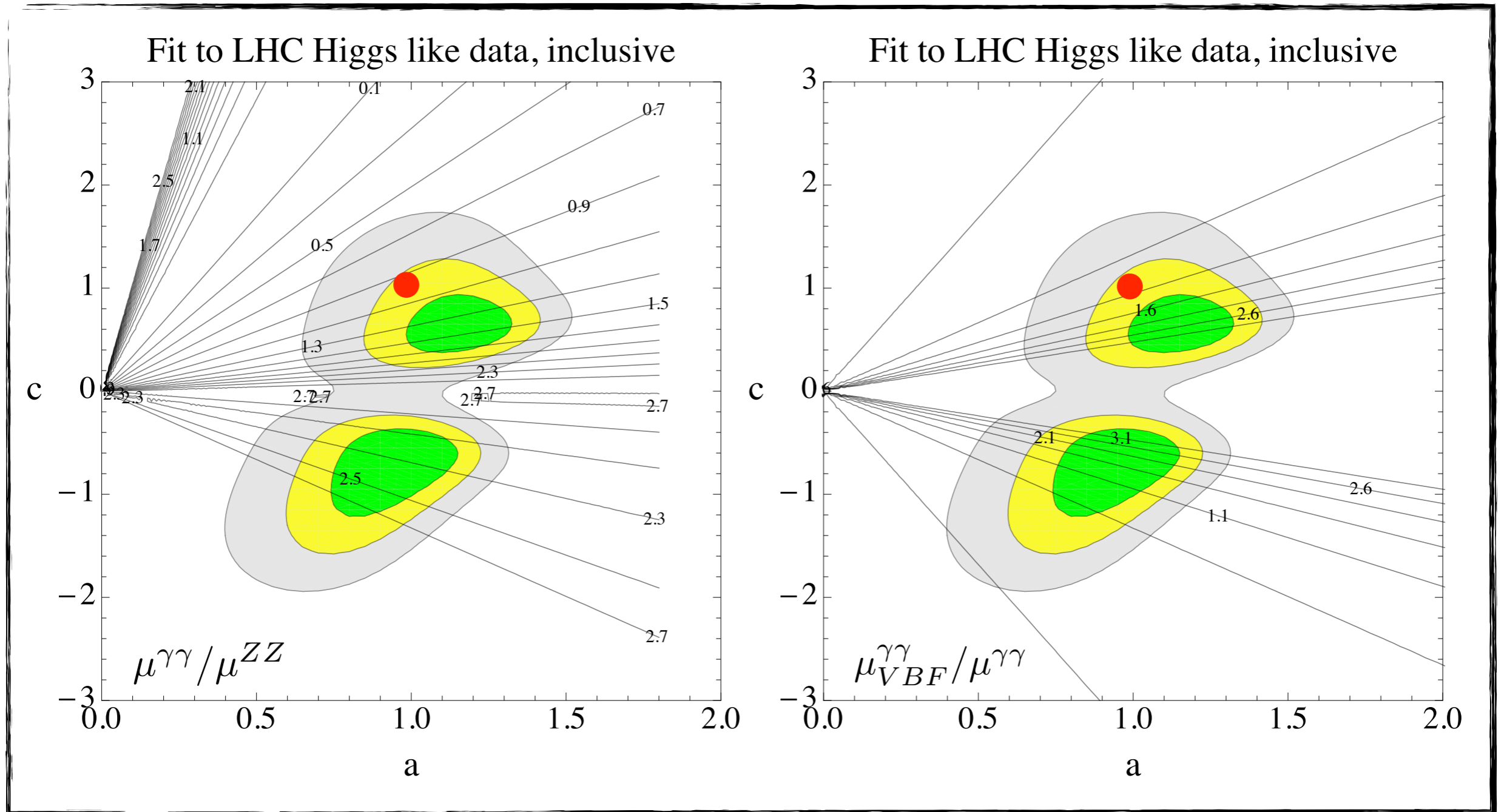
(125, 0.87, -0.42, 4.6)

$(m_h, a, c, \chi^2)$

Espinosa, Grojean, Muhlleitner, Trott '12

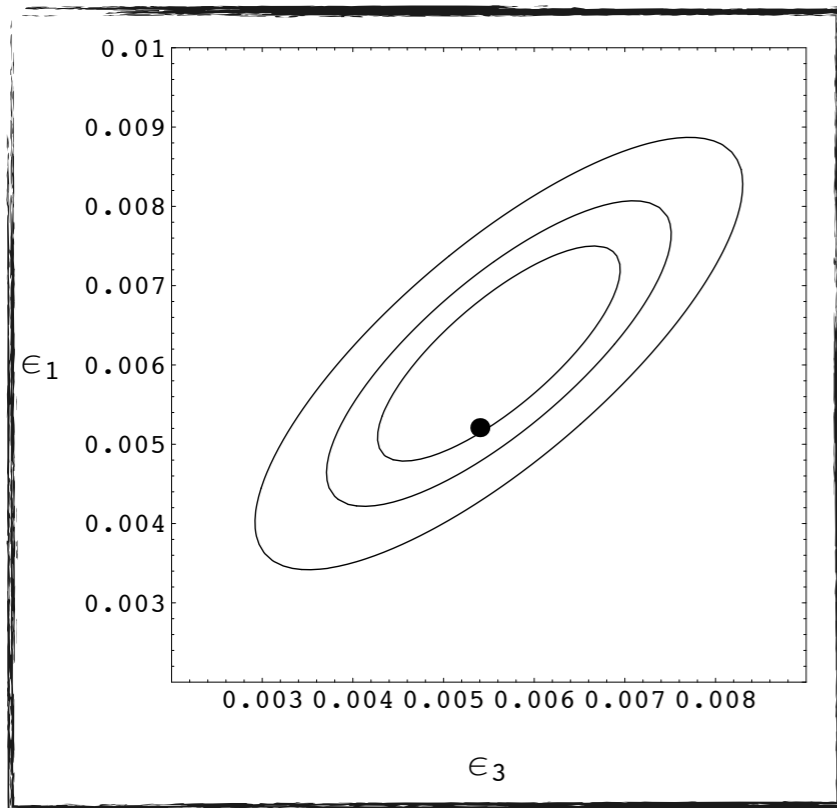
# How to distinguish the two minima

the  $(a,c) \leftrightarrow (a,-c)$  symmetry is broken in the  $\gamma\gamma$  channel



Espinosa, Grojean, Muhlleitner, Trott '12

# A tension between LHC and EW data



EW fit strongly suggests custodial symmetry

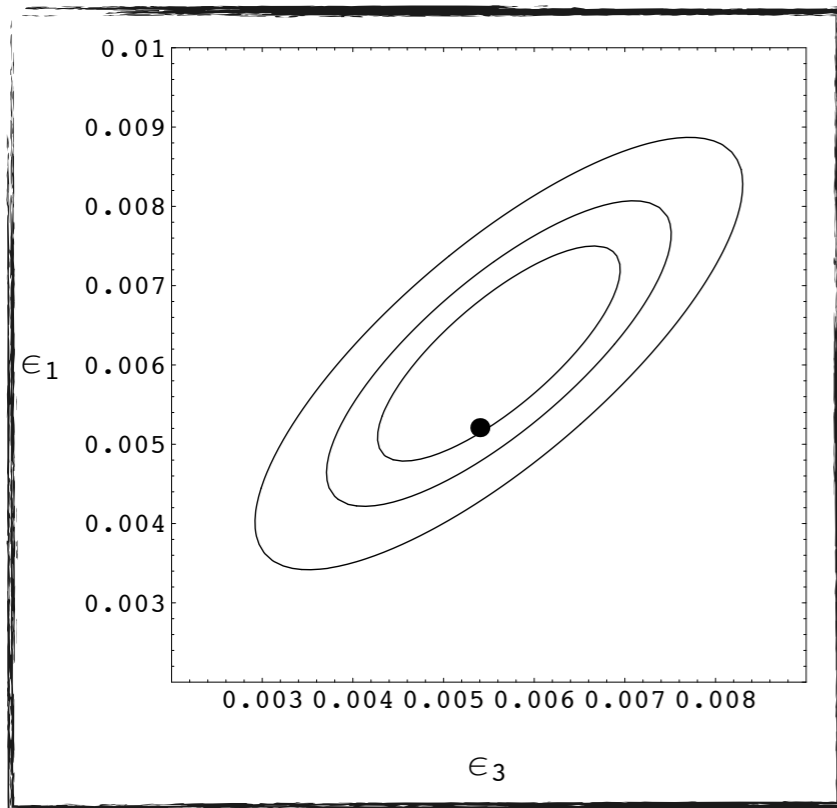
$$\Sigma = e^{i\sigma^a \pi^a / v}$$

Goldstone of  
 $SU(2)_L \times SU(2)_R / SU(2)_V$

$$\frac{v^2}{4} \text{Tr} (D_\mu \Sigma^\dagger D^\mu \Sigma) \Rightarrow \rho = 1 \quad \text{ie} \quad \epsilon_1 = \hat{T} = 0 \quad \checkmark$$

$$\frac{v^2}{4} \text{Tr}^2 (\Sigma^\dagger D_\mu \Sigma \sigma^3) \Rightarrow \rho = 2 \quad \text{ie} \quad \epsilon_1 = \hat{T} = 1 \quad \text{strongly disfavored}$$

# A tension between LHC and EW data



EW fit strongly suggests custodial symmetry

$$\Sigma = e^{i\sigma^a \pi^a / v}$$

Goldstone of  
 $SU(2)_L \times SU(2)_R / SU(2)_V$

$$\frac{v^2}{4} \text{Tr} (D_\mu \Sigma^\dagger D^\mu \Sigma) \Rightarrow \rho = 1 \text{ ie } \epsilon_1 = \hat{T} = 0 \quad \checkmark$$

$$\text{also } \Rightarrow \mu_{ZZ} = \mu_{WW} \quad \times$$

but

Channel [Exp]	$\mu_{119.5} (\mu_{119.5}^L)$	$\mu_{124} (\mu_{124}^L)$	$\mu_{125} (\mu_{125}^L)$
$pp \rightarrow Z Z^* \rightarrow \ell^+ \ell^- \ell^+ \ell^-$ [ATLAS]	$-0.5^{+0.5??} (5.1)$	$1.6^{+1.4}_{-0.8} (4.7)$	$1.4^{+1.3}_{-0.8} (4.1)$
$pp \rightarrow W W^* \rightarrow \ell^+ \nu \ell^- \bar{\nu}$ [ATLAS]	$0.0^{+1.2}_{-1.3} (2.4)$	$0.1^{+0.7}_{-0.7} (1.6)$	$0.1^{+0.7}_{-0.6} (1.4)$

1. has LHC identified a violation of the custodial symmetry?
2. if yes, how to reconcile LHC data with EW data?

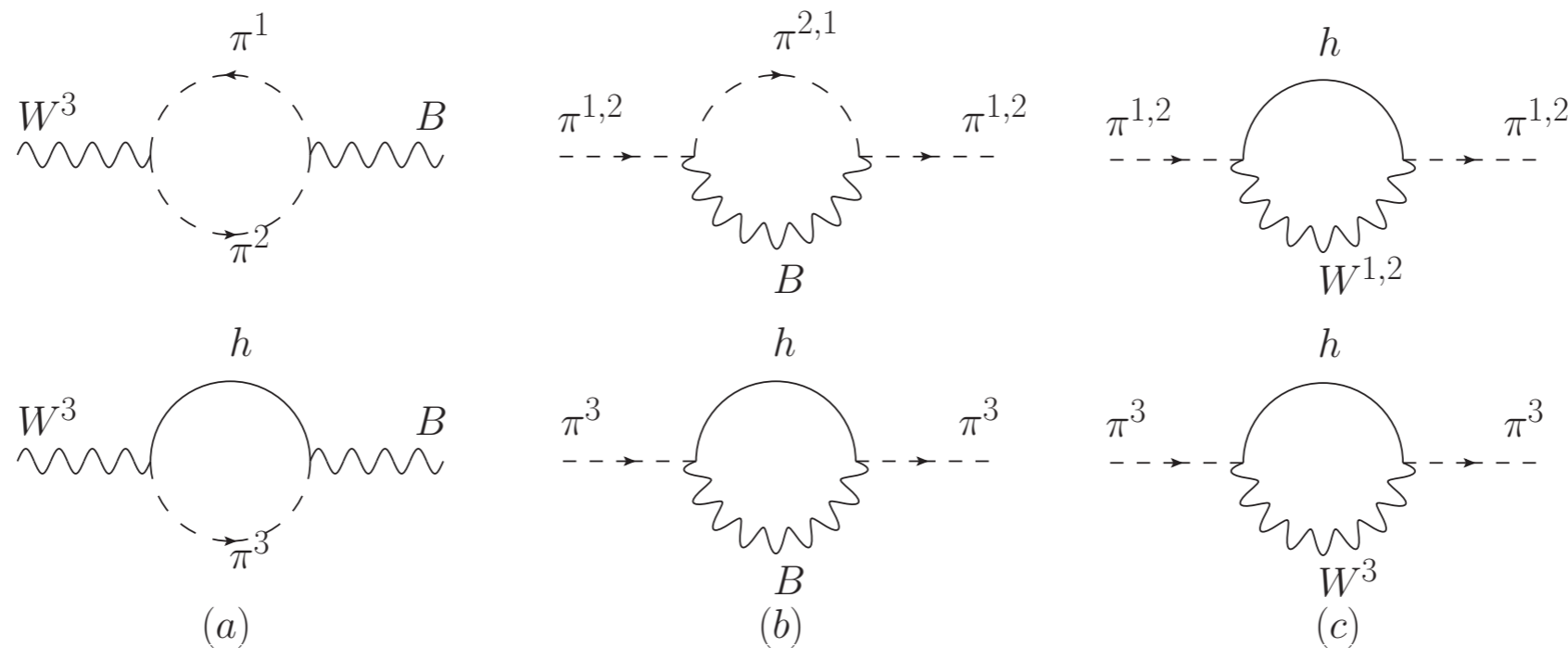
# DisZphilia or how to live with custodial breaking

Farina, Grojean, Salvioni 'to appear

$$\mathcal{L}_{cb} = -\frac{v^2}{8} \left( \text{Tr} \left[ \Sigma^\dagger D_\mu \Sigma \sigma^3 \right] \right)^2 \left( 0 + 2a_{cb} \frac{h}{v} + \dots \right)$$

no custodial breaking  
in the vacuum

custodial breaking  
Higgs interactions



$$\Delta\epsilon_1 = -\frac{3}{16\pi} \frac{\alpha(m_Z)}{\cos^2 \theta_W} \left[ 1 - (a + a_{cb})^2 + \left( \frac{g}{g'} \right)^2 (a^2 - (a + a_{cb})^2) \right] \log \left( \frac{\Lambda^2}{m_h^2} \right)$$

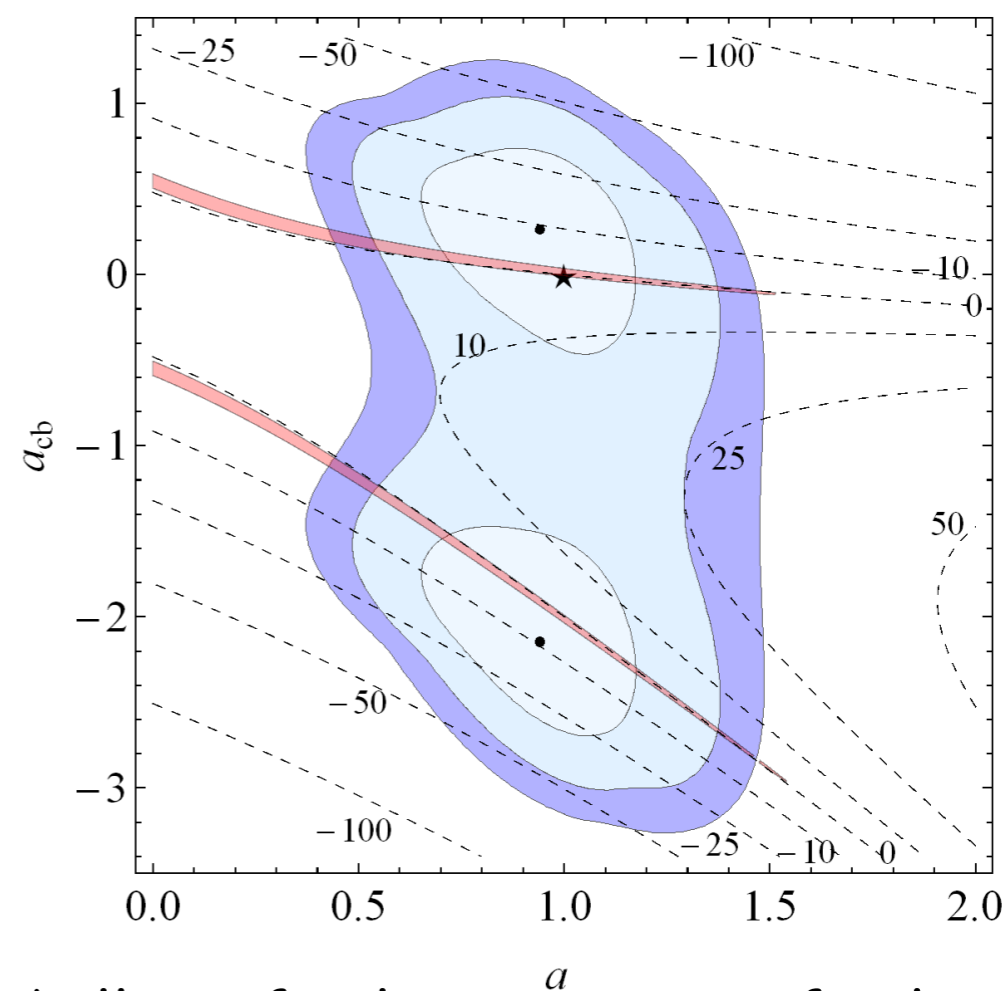
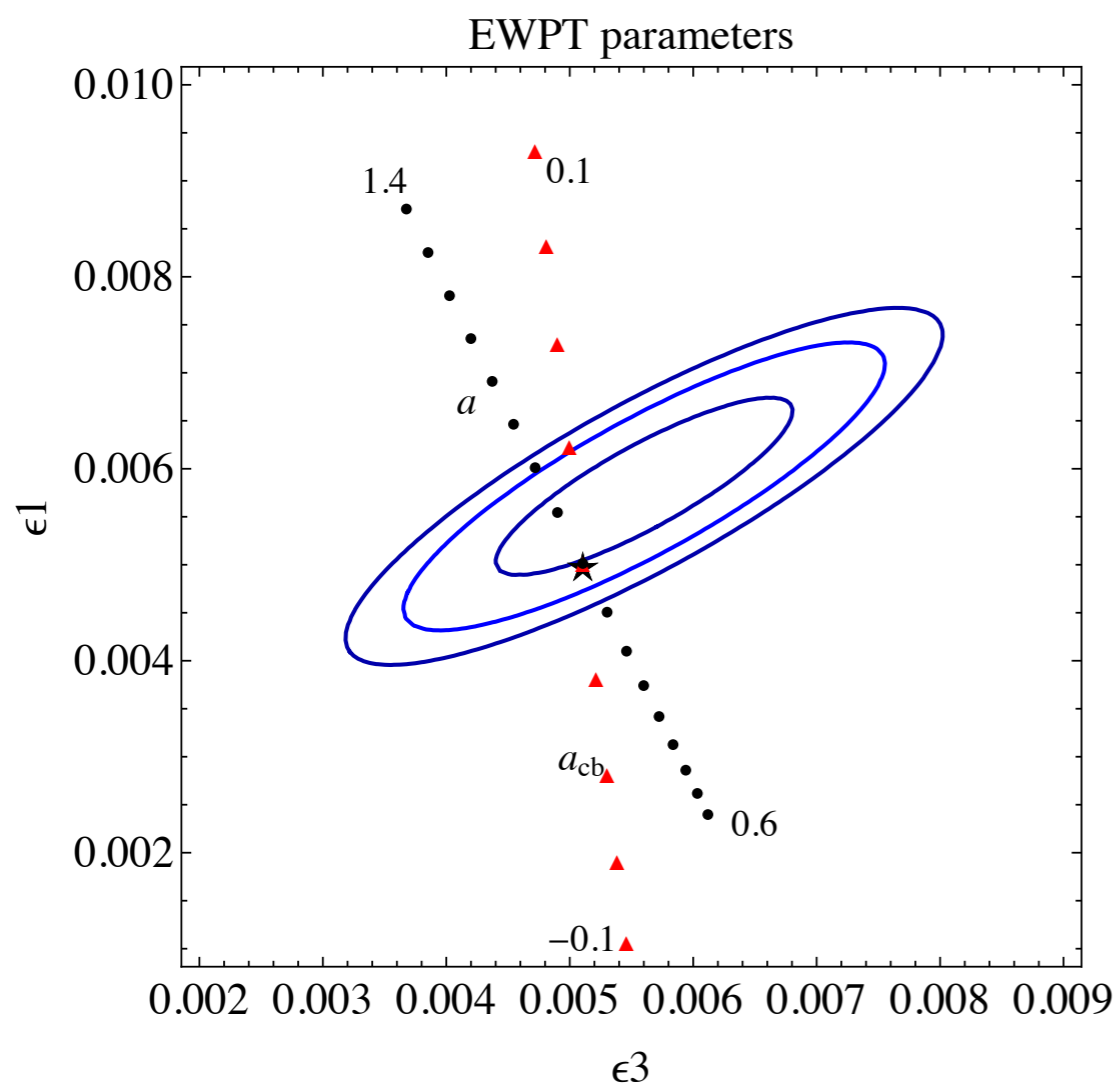
$$\Delta\epsilon_3 = +\frac{1}{48\pi} \frac{\alpha(m_Z)}{\sin^2 \theta_W} \left[ 1 - (a + a_{cb})^2 \right] \log \left( \frac{\Lambda^2}{m_h^2} \right)$$

Note: quadratic custodial breaking couplings will give  $\Lambda^2$  UV sensitivity in  $\epsilon_1$

# DisZphilia or how to live with custodial breaking

Farina, Grojean, Salvioni 'to appear

$$\mathcal{L}_{cb} = -\frac{v^2}{8} \left( \text{Tr} \left[ \Sigma^\dagger D_\mu \Sigma \sigma^3 \right] \right)^2 \left( 0 + 2a_{cb} \frac{h}{v} + \dots \right)$$



' $a_{cb}$ ' allows for larger range of values for 'a'

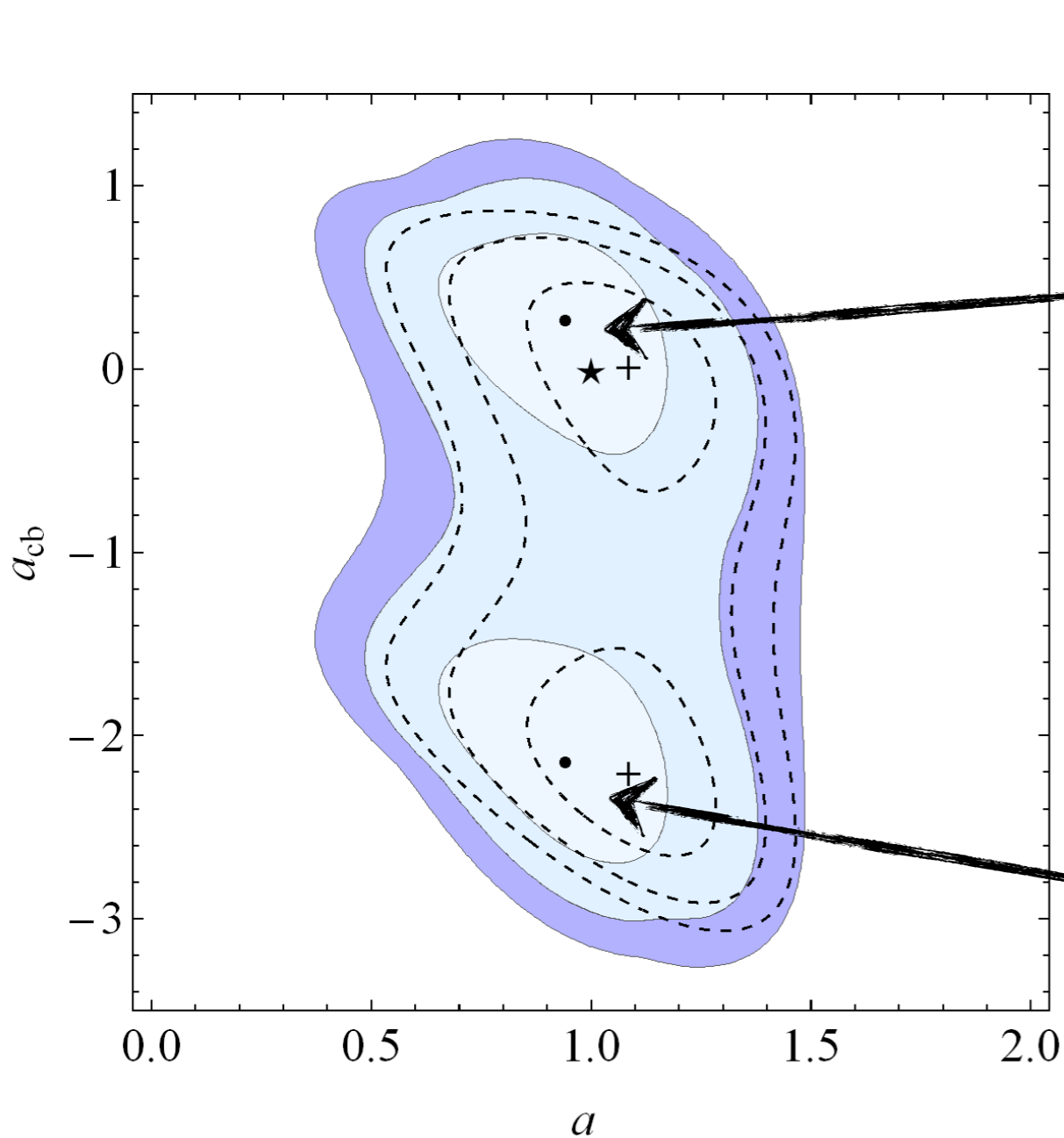
$$a_{cb}=0: 0.95 \leq a \leq 1.1$$

$$a_{cb} \neq 0: 0.4 \leq a \leq 1.5$$

# DisZphilia or how to live with custodial breaking

Farina, Grojean, Salvioni 'to appear

$$\mathcal{L}_{cb} = -\frac{v^2}{8} \left( \text{Tr} [\Sigma^\dagger D_\mu \Sigma \sigma^3] \right)^2 \left( 0 + 2a_{cb} \frac{h}{v} + \dots \right)$$



-----  $c = 1$   
 ————— marginalization over  $c$

$$2 \frac{h}{v} \left( m_W^2 W_\mu^+ W_\mu^- + \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right)$$

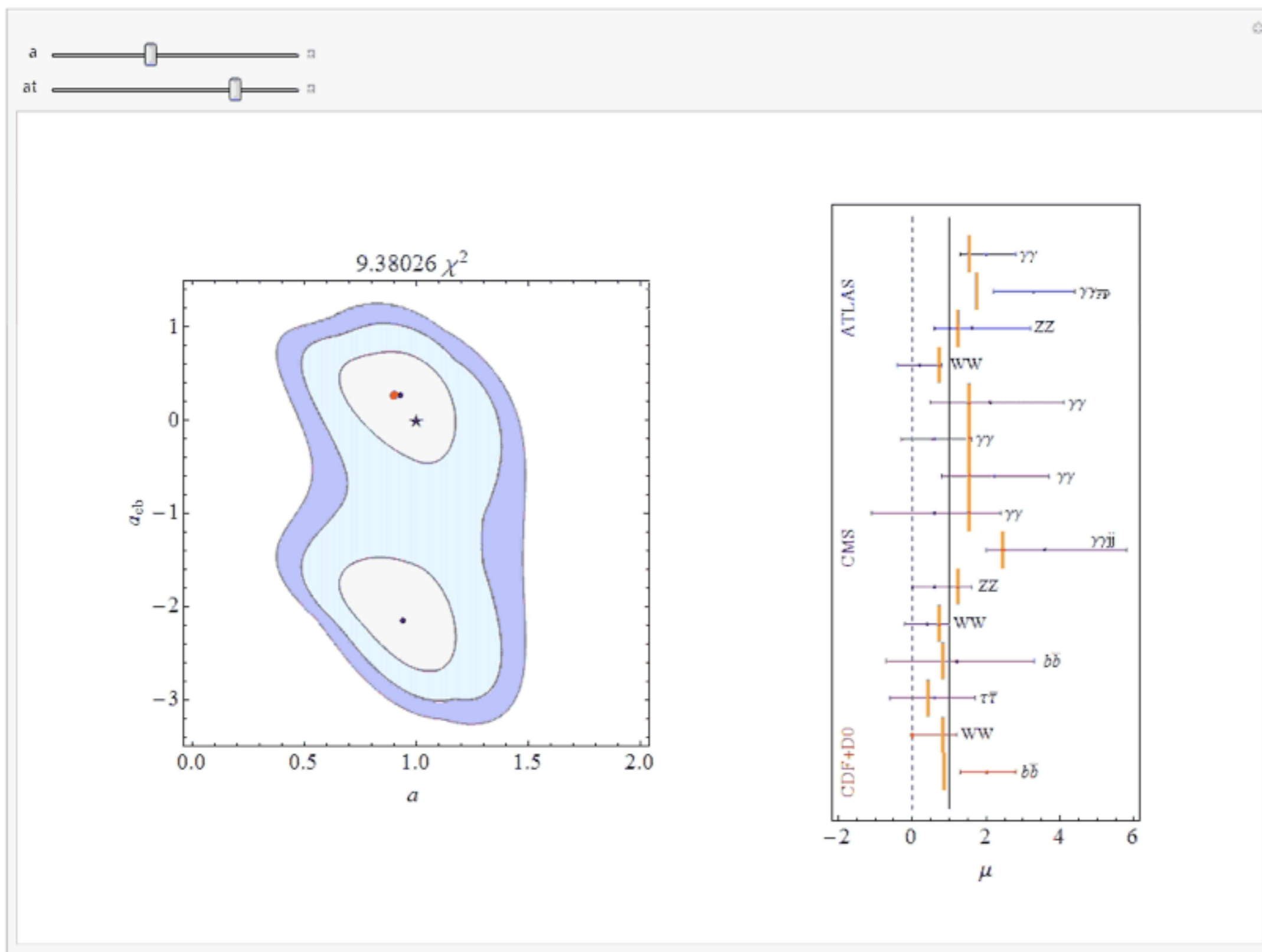
the two solutions can only be distinguished in the presence of interference with a single  $hZZ$  vertex

$$2 \frac{h}{v} \left( m_W^2 W_\mu^+ W_\mu^- - \frac{1}{2} m_Z^2 Z_\mu Z_\mu \right)$$

↑  
 "disZphilia"

# DisZphilia or how to live with custodial breaking

Farina, Grojean, Salvioni 'to appear



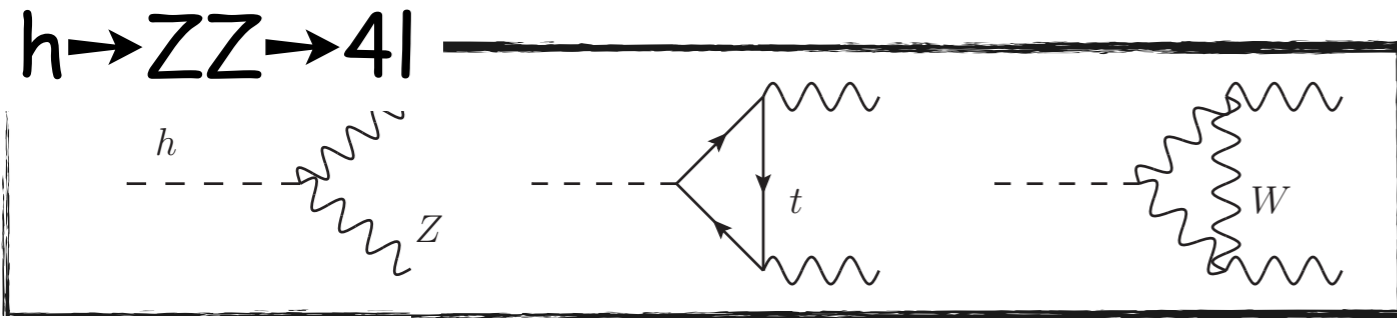


# Zphilia or DisZphilia?

difficult!

Farina, Grojean, Salvioni 'to appear

difference is physically relevant only in the presence of interference with single hZZ coupling

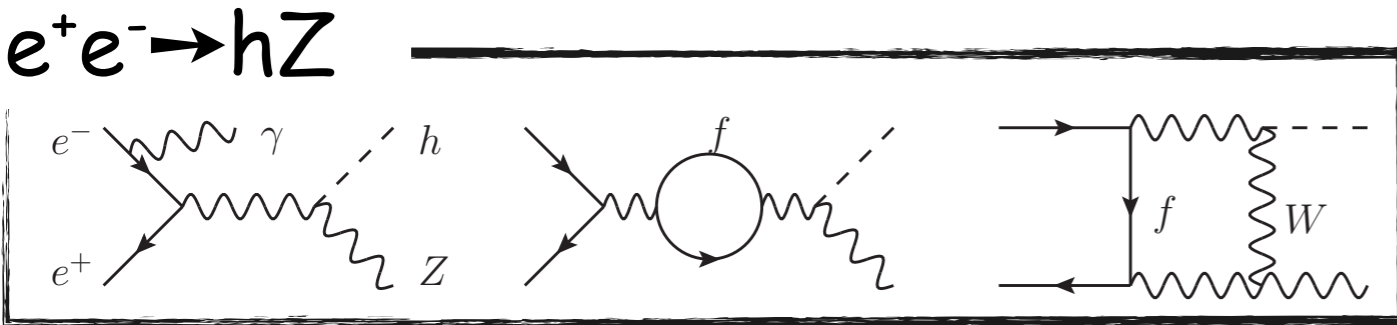


TH prediction

$$\Delta = \left| \frac{\Gamma_Z^+ - \Gamma_Z^-}{\Gamma_Z^+ + \Gamma_Z^-} \right| = \delta \approx 1\%$$

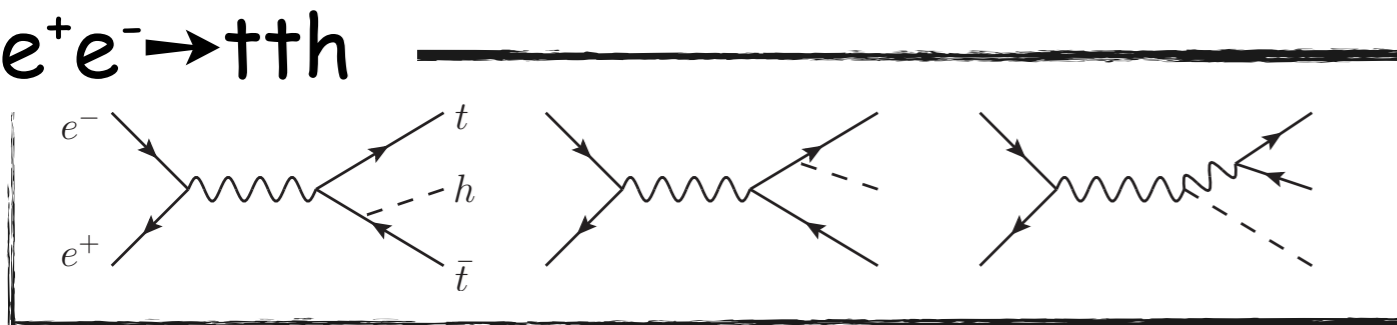
ILC ( $\sqrt{s}=800\text{GeV}$  and  $1\text{ab}^{-1}$ )

$\approx 1\%$



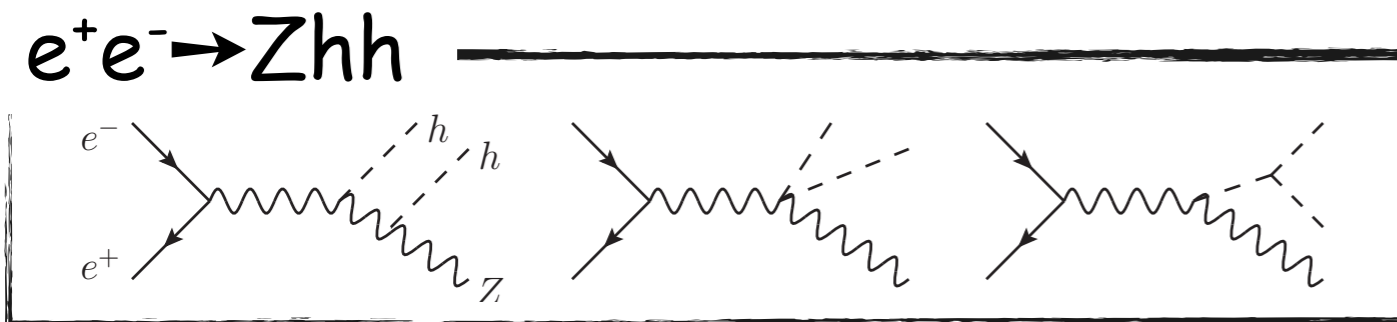
$$\Delta = \left| \frac{\sigma_+ - \sigma_-}{\sigma_+ + \sigma_-} \right| \approx 15\%$$

$\approx 5\%$



$$\Delta = \left| \frac{\sigma_+ - \sigma_-}{\sigma_+ + \sigma_-} \right| \lesssim 4\%$$

$\approx 10\%$



$$\Delta = \left| \frac{\sigma_+ - \sigma_-}{\sigma_+ + \sigma_-} \right| \approx 50\%$$

$\approx 10\%$



# Conclusions

EW interactions need Goldstone bosons to provide mass to  $W, Z$



EW interactions also need a UV moderator/new physics  
to unitarize  $WW$  scattering amplitude

We'll need another Gargamelle experiment  
to discover the still missing neutral current of the SM: the Higgs  
weak NC  $\Leftrightarrow$  gauge principle  
Higgs NC  $\Leftrightarrow$  ?

Strong EWSB w/o an elementary Higgs can be very similar to SM

it might take a long time to decipher the true dynamics of EWSB!