

DIRECT DETECTION OF DARK MATTER IN A MODEL WITH RADIATIVE NEUTRINO MASSES

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FLASY12

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INTERNATIONAL
MAX PLANCK
RESEARCH SCHOOL



FOR PRECISION TESTS
OF FUNDAMENTAL
SYMMETRIES



OUTLINE

- The Model
- Dark Matter
- Direct Detection

THE MODEL

Ma, Phys. Rev. D73 (2006) 077301

neutrino masses at 1 loop

flavor structure:

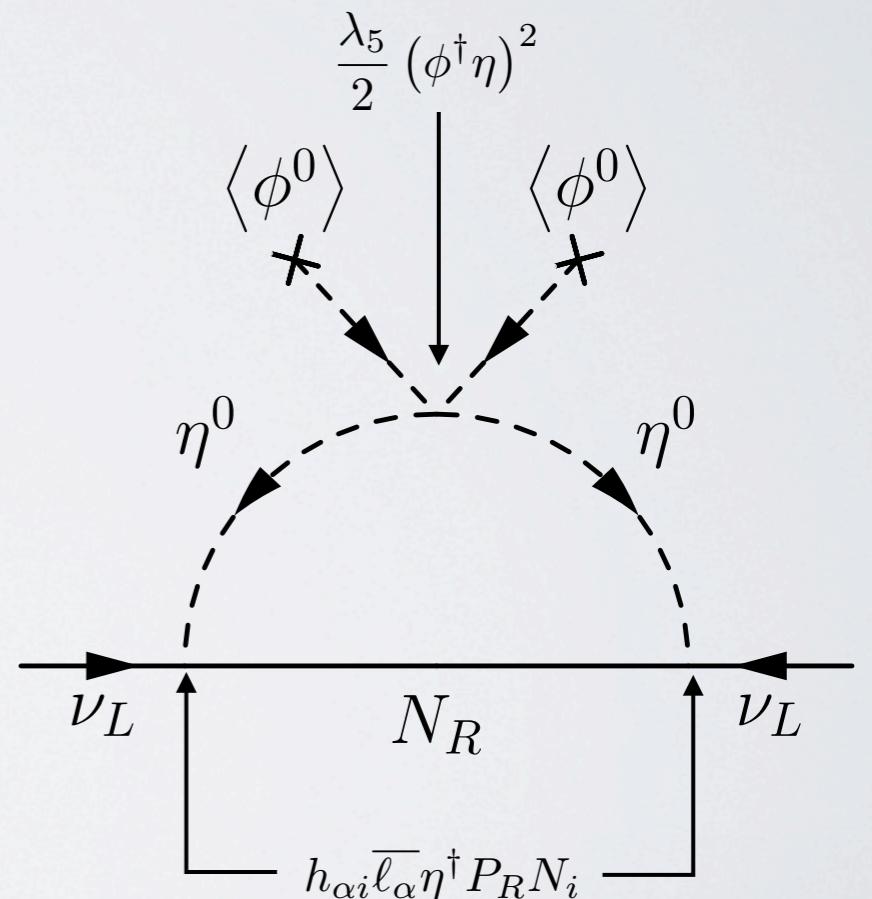
$$h_{\alpha i} = \begin{pmatrix} 0 & 0 & h'_3 \\ h_1 & h_2 & h_3 \\ h_1 & h_2 & -h_3 \end{pmatrix}$$

lepton-flavor violation (LFV):

$$\text{Br}(\ell_\alpha \rightarrow \ell_\beta \gamma) = \frac{3\alpha_{\text{em}}}{64\pi G_F^2 M_\eta^4} \left| \sum_{i=1}^3 h_{\alpha i}^* h_{\beta i} F_2 \left(\frac{M_i^2}{M_\eta^2} \right) \right|^2 \text{Br}(\ell_\alpha \rightarrow \ell_\beta \nu_\alpha \bar{\nu}_\beta)$$

$\text{Br}(\mu \rightarrow e \gamma) < 2.4 \times 10^{-12}$: $M_3 = 6000 \text{ GeV}$, $h_3 = 0.3$ (benchmark point)

independent parameters: $M_\eta, M_1, \delta \equiv M_2 - M_1, \xi \equiv \text{Im}(h_2^* h_1)$



THE MODEL

Ma, Phys. Rev. D73 (2006) 077301

neutrino masses at 1 loop

flavor structure:

$$h_{\alpha i} = \begin{pmatrix} \epsilon_1 & \epsilon_2 & h'_3 \\ h_1 & h_2 & h_3 \\ h_1 & h_2 & -h_3 \end{pmatrix} + \mathcal{O}(\epsilon^2) \quad \text{allowing } \sin \Theta_{13} \equiv \epsilon_3 \neq 0^\circ$$

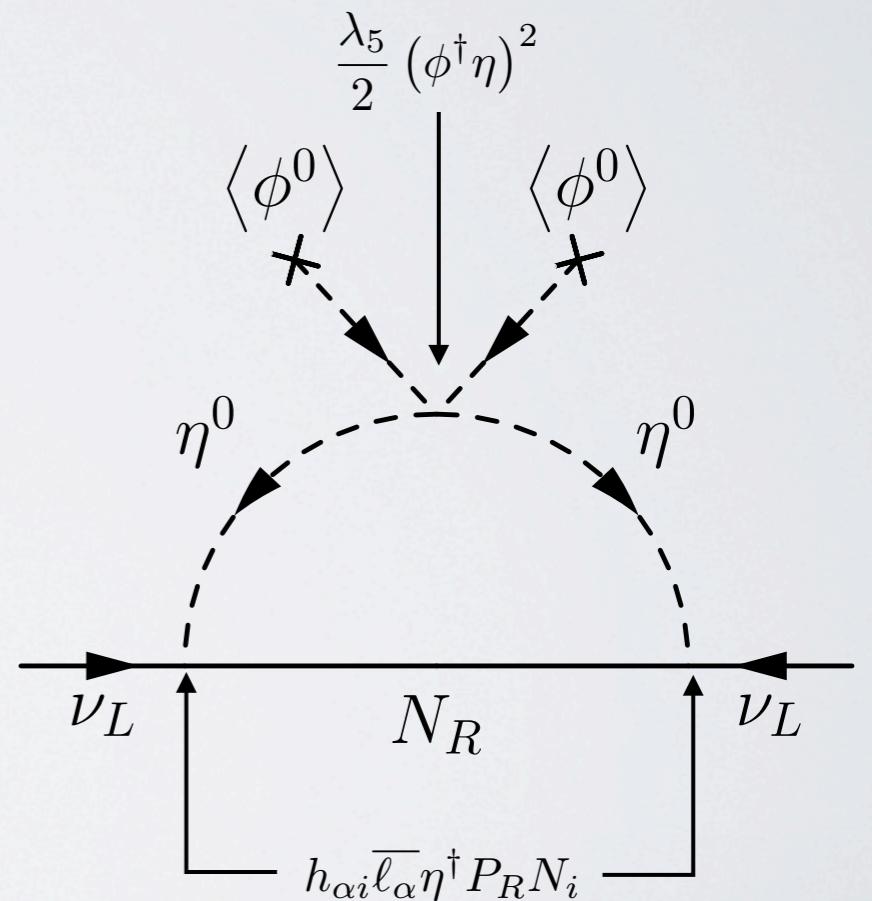
$$\epsilon_1 h_1 + \epsilon_1 h_1 \equiv P \epsilon_3$$

lepton-flavor violation (LFV):

$$\text{Br}(\ell_\alpha \rightarrow \ell_\beta \gamma) = \frac{3\alpha_{\text{em}}}{64\pi G_F^2 M_\eta^4} \left| \sum_{i=1}^3 h_{\alpha i}^* h_{\beta i} F_2 \left(\frac{M_i^2}{M_\eta^2} \right) \right|^2 \text{Br}(\ell_\alpha \rightarrow \ell_\beta \nu_\alpha \bar{\nu}_\beta)$$

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DARK MATTER

Dark Matter (DM) candidate (lightest odd particle):

N_1 with $m_{N_1} \lesssim m_{N_2} < m_{N_3}$

freeze-out: $\sigma_{\text{eff}} v = \left[\frac{\xi^2}{2\pi} \frac{M_1^2}{(M_\eta^2 + M_1^2)^2} \right] + \left[\frac{|h_1^2 + h_2^2|^2}{24\pi} \frac{M_1^2 (M_\eta^4 + M_1^4)}{(M_\eta^2 + M_1^2)^4} + \frac{\xi^2}{2\pi} \frac{M_1^2 (M_\eta^4 - 3M_\eta^2 M_1^2 - M_1^4)}{(M_\eta^2 + M_1^2)^4} \right] v^2 + \mathcal{O}(v^4)$

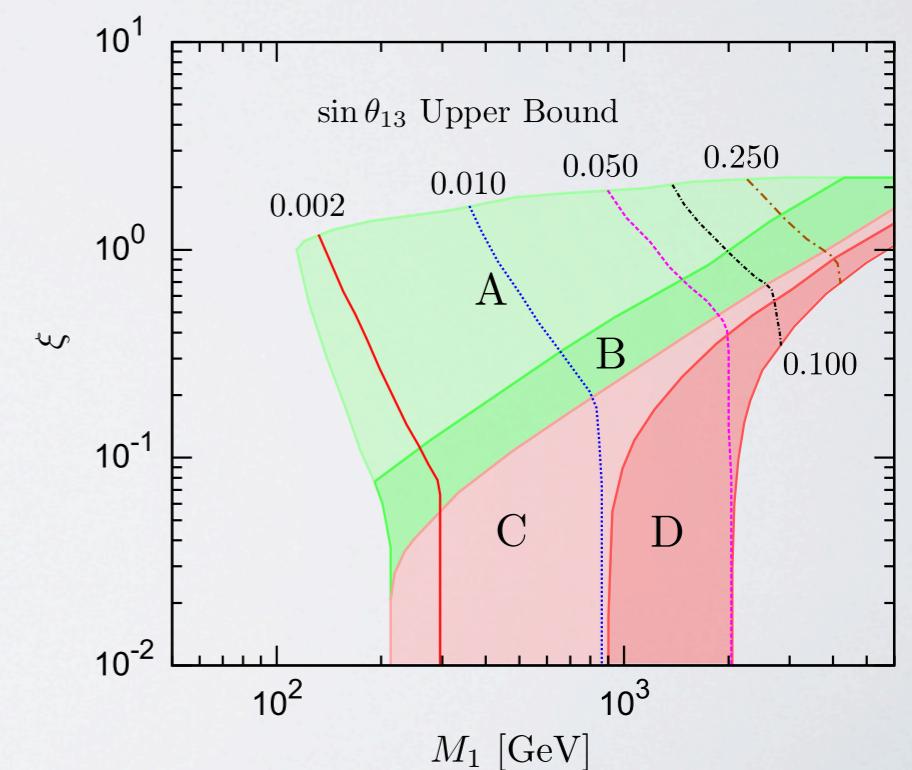
parameter regions:

- A : $2.0 < M_\eta/M_1 < 9.8$
- B : $1.2 < M_\eta/M_1 < 2.0$
- C : $1.05 < M_\eta/M_1 < 1.20$
- D : $1.0 < M_\eta/M_1 < 1.05$

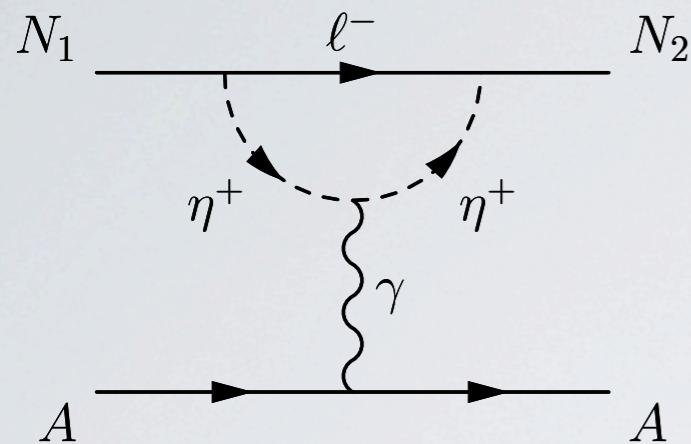
A,B,C, and D consistent with neutrino data,
LFV, perturbativity and DM relic density

$\sin^2 2\theta_{13} = 0.092 \pm 0.016 \pm 0.005$: additional contributions to $\mu \rightarrow e\gamma$

DAYA-BAY, Phys.Rev.Lett. 108(2012) 171803



DIRECT DETECTION: CROSS SECTIONS



charge-charge (CC)

dipole-charge (DC)

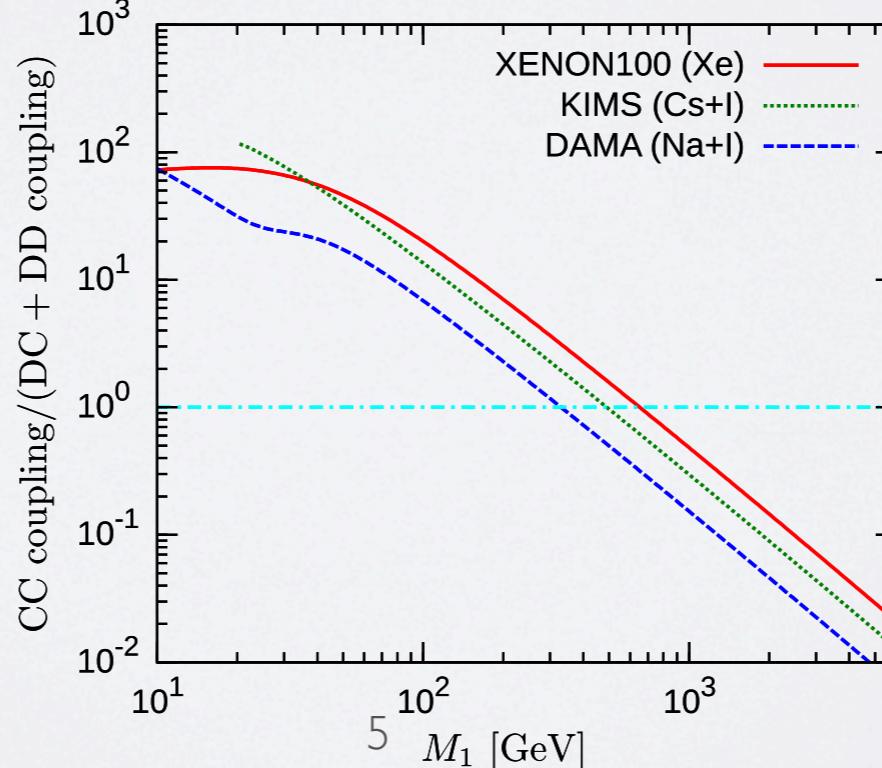
dipole-dipole (DD)

$$\mathcal{L}_{\text{eff}} = ia_{12}\overline{N_2}\gamma^\mu N_1 \partial^\nu F_{\mu\nu} + i\left(\frac{\mu_{12}}{2}\right)\overline{N_2}\sigma^{\mu\nu}N_1 F_{\mu\nu} + ic_{12}\overline{N_2}\gamma^\mu N_1 A_\mu$$

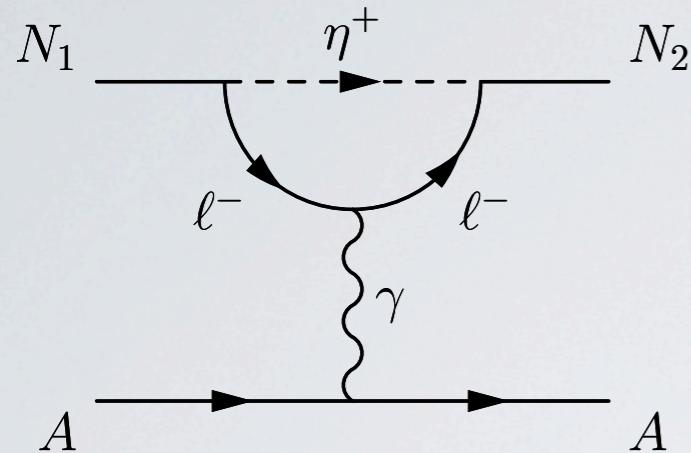
$$\frac{d\sigma_{\text{CC}}}{dE_R} = \frac{Z^2 b_{12}^2 m_A}{2\pi v^2} F^2(E_R) ; \quad b_{12} = (a_{12} + c_{12}/q^2)e$$

$$\frac{d\sigma_{\text{DC}}}{dE_R} = \frac{Z^2 \alpha_{\text{em}} \mu_{12}^2}{E_R} \left[1 - \frac{E_R}{v^2} \left(\frac{1}{2m_A} + \frac{1}{M_1} \right) - \frac{\delta}{v^2} \frac{1}{\mu_{\text{DM}}} - \frac{\delta^2}{v^2} \frac{1}{2m_A E_R} \right] F^2(E_R)$$

$$\frac{d\sigma_{\text{DD}}}{dE_R} = \frac{\mu_A^2 \mu_{12}^2 m_A}{\pi v^2} \left(\frac{J_A + 1}{3J_A} \right) F_D^2(E_R)$$



DIRECT DETECTION: CROSS SECTIONS



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dipole-charge (DC)

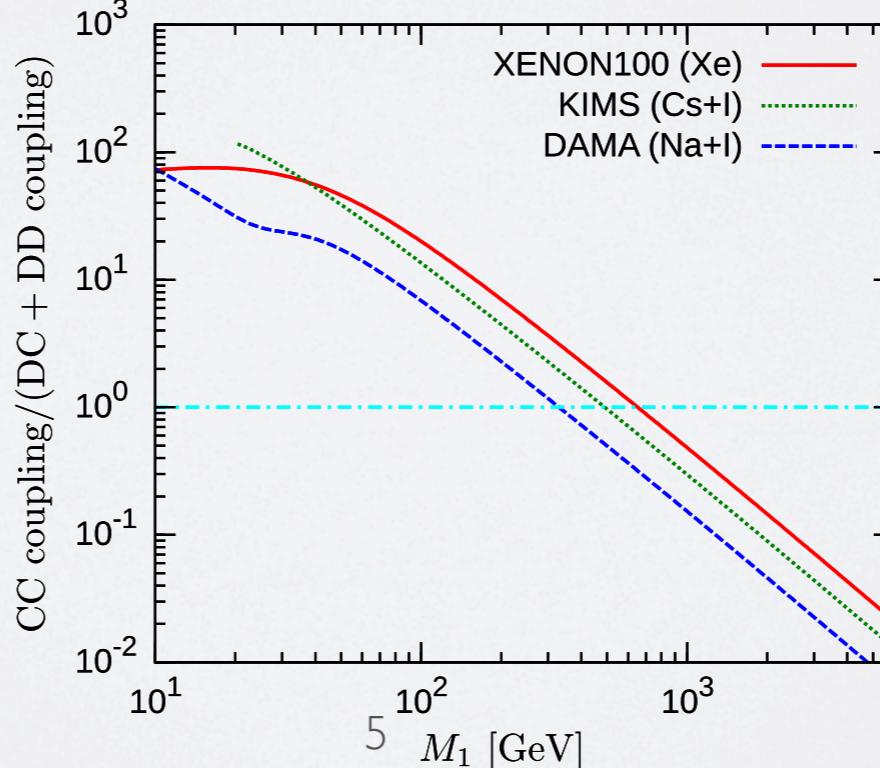
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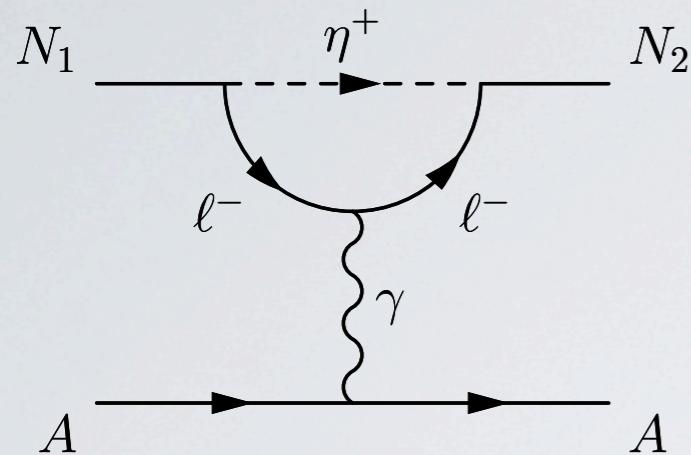
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DIRECT DETECTION: CROSS SECTIONS



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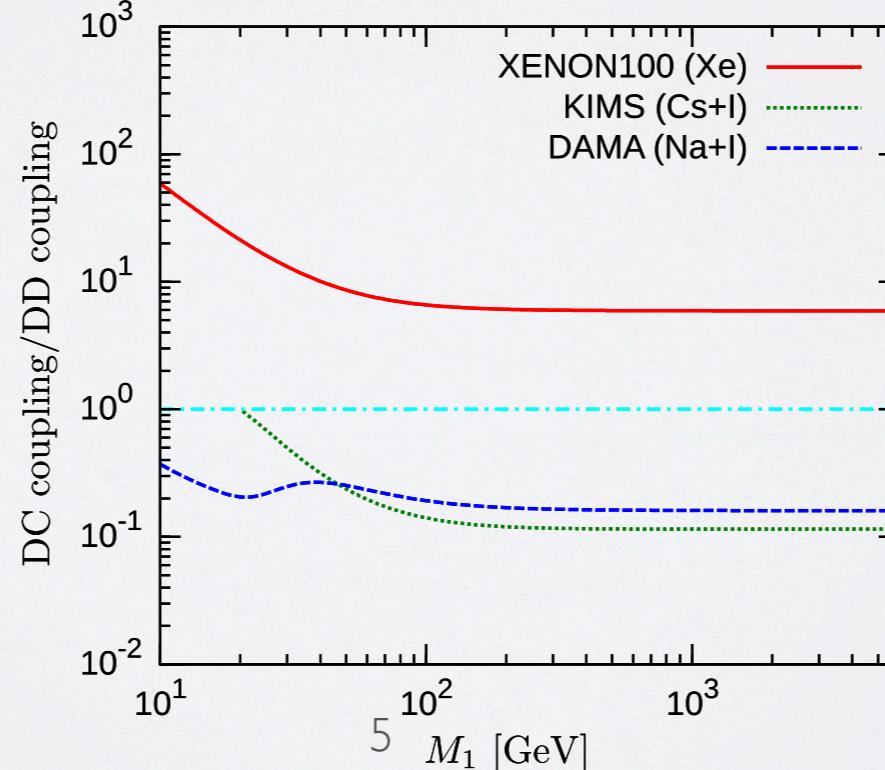
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$$\frac{d\sigma_{\text{CC}}}{dE_R} = \frac{Z^2 b_{12}^2 m_A}{2\pi v^2} F^2(E_R) ; \quad b_{12} = (a_{12} + c_{12}/q^2)e$$

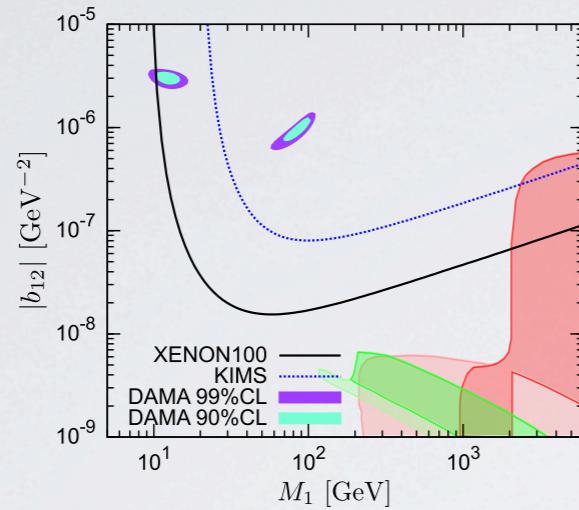
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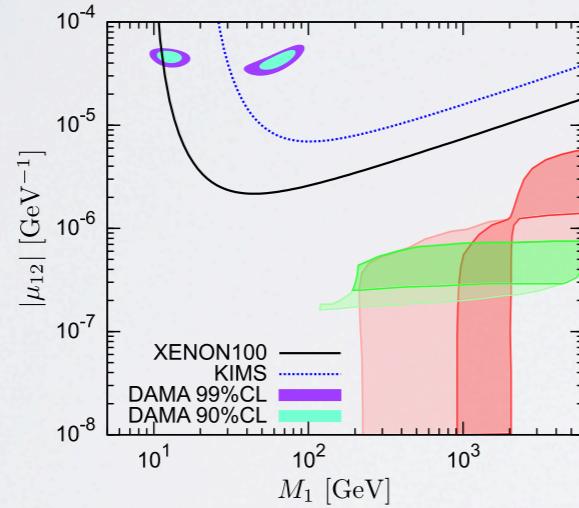
DIRECT DETECTION: COMPARISON WITH EXPERIMENTS

CC

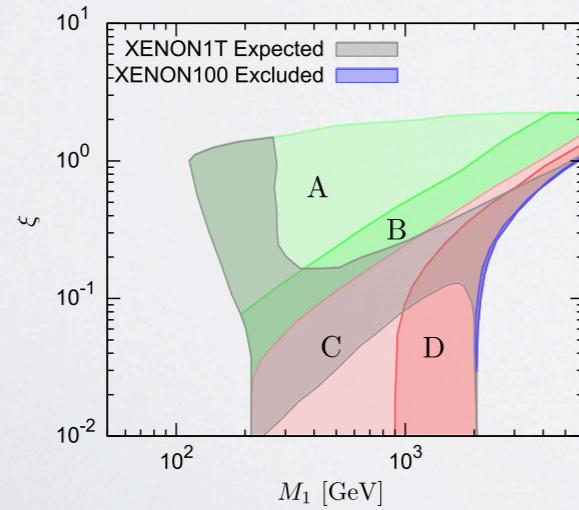


$\delta = 0 \text{ KeV}$

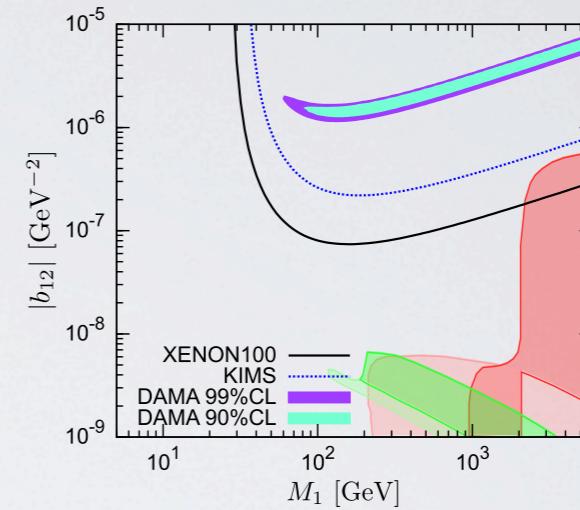
DC+ DD



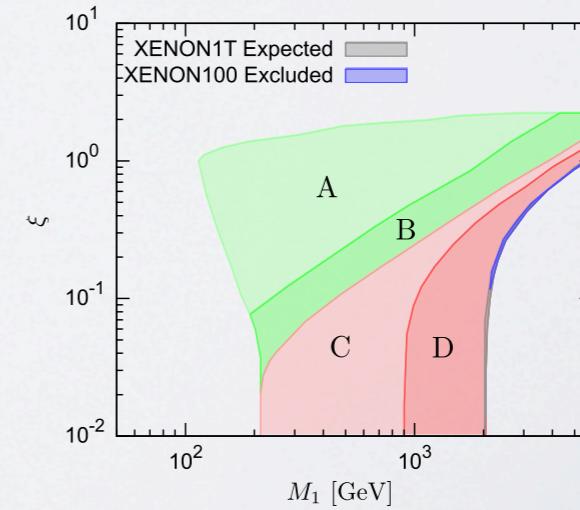
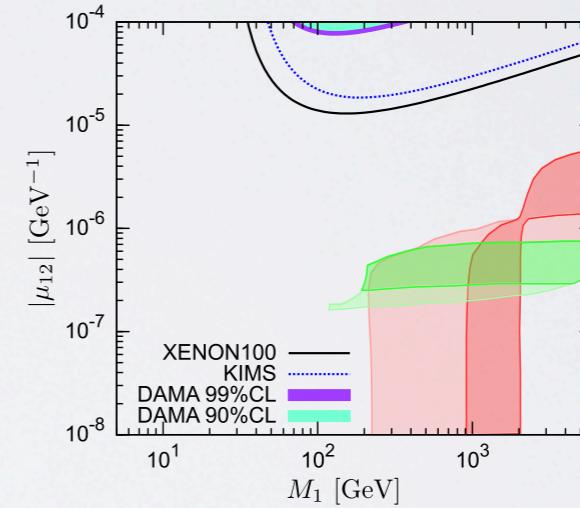
Yukawa



6



$\delta = 80 \text{ KeV}$



CONCLUSION

Ma model: **Neutrinos have mass and the Universe has Dark Matter.**

constraints from LFV processes

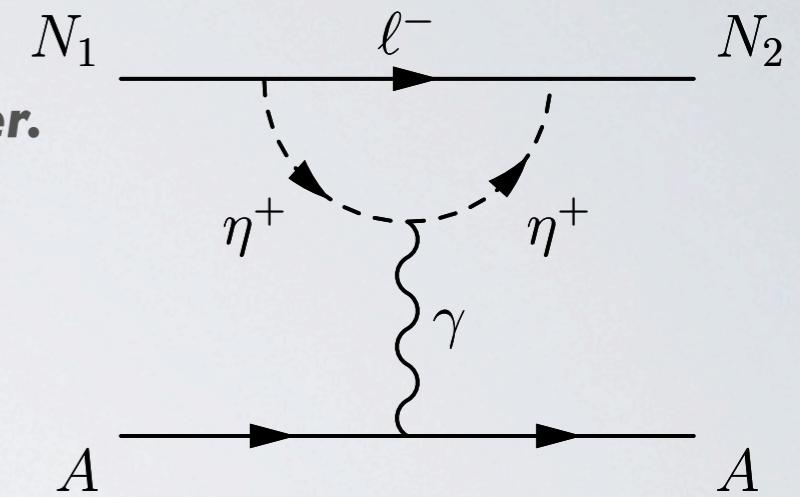
coannihilations important for DM relic density

direct detection of leptophilic DM possible by photon exchange

calculation of relevant loop processes for the first time

charge-charge interactions dominate XENON100 event rate

stringent test by XENON1T



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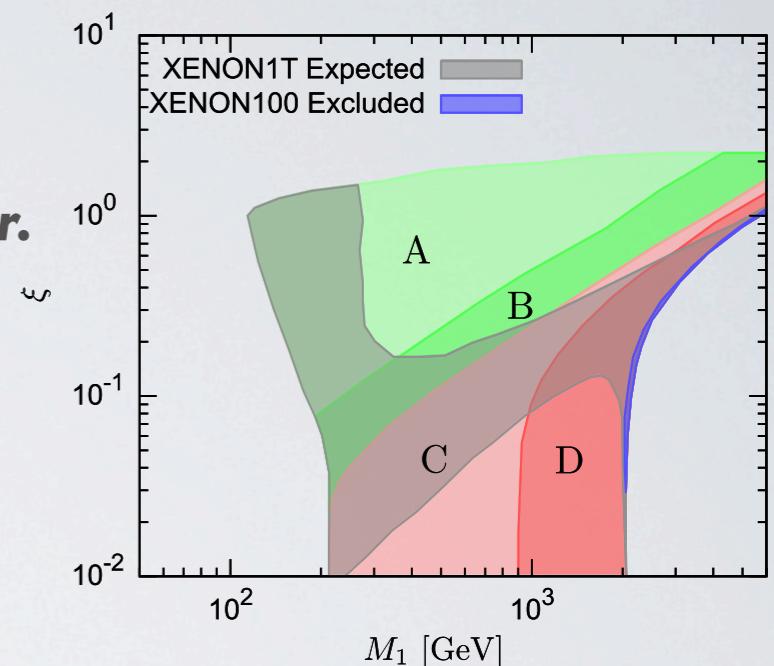


photo by M.Anzorg

THANKS



APPENDIX A: THE MODEL

$$\mathcal{L}_N = \overline{N_i} i\cancel{\partial} P_R N_i + (D_\mu \eta)^\dagger (D^\mu \eta) - \frac{M_i}{2} \overline{N_i^c} P_R N_i + h_{\alpha i} \overline{\ell_\alpha} \eta^\dagger P_R N_i + \text{h.c.}$$

$$\begin{aligned} \mathcal{V}(\phi, \eta) = & m_\phi^2 \phi^\dagger \phi + m_\eta^2 \eta^\dagger \eta + \frac{\lambda_1}{2} (\phi^\dagger \phi)^2 + \frac{\lambda_2}{2} (\eta^\dagger \eta)^2 \\ & + \lambda_3 (\phi^\dagger \phi) (\eta^\dagger \eta) + \lambda_4 (\phi^\dagger \eta) (\eta^\dagger \phi) + \frac{\lambda_5}{2} (\phi^\dagger \eta)^2 + \text{h.c.} \end{aligned}$$

$$(m_\nu)_{\alpha\beta} \simeq \sum_{i=1}^3 \frac{2\lambda_5 h_{\alpha i} h_{\beta i} \langle \phi^0 \rangle^2}{(4\pi)^2 M_i} I\left(\frac{M_i^2}{M_\eta^2}\right)$$

$$M_\eta^2 \simeq m_\eta^2 + (\lambda_3 + \lambda_4) \langle \phi^0 \rangle^2$$

$$I(x) = \frac{x}{1-x} \left(1 + \frac{x \log x}{1-x} \right)$$

APPENDIX B: FLAVOR STRUCTURE

perturbation:

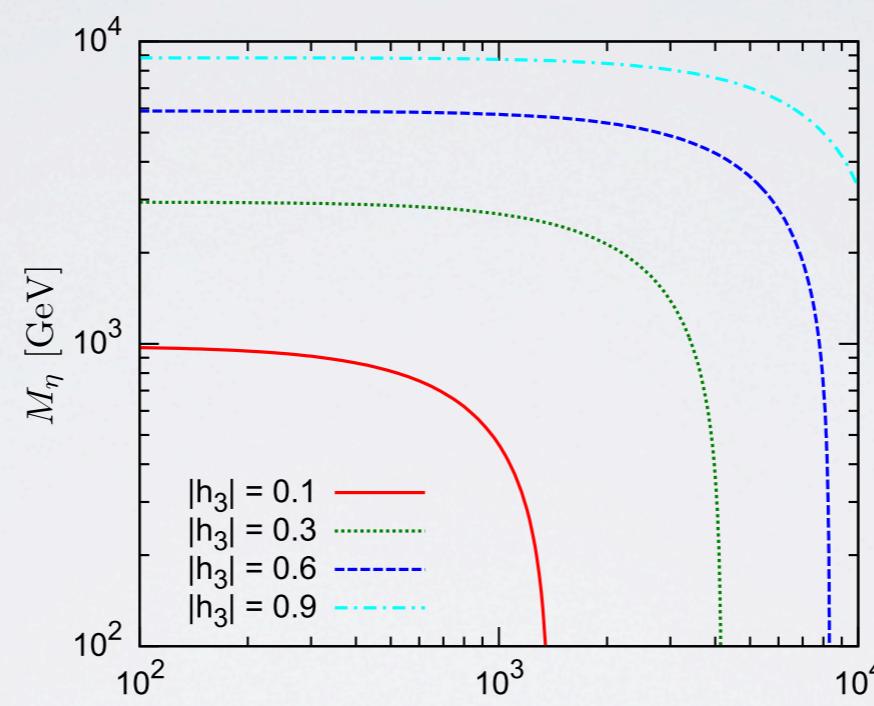
$$\sin \Theta_{13} = \epsilon_3$$

$$\cos \Theta_{13} = 1 + \mathcal{O}(\epsilon_3^2)$$

$$\sin \Theta_{23} = \frac{1}{\sqrt{2}} + \epsilon_4$$

$$\cos \Theta_{23} = \frac{1}{\sqrt{2}} - \epsilon_4$$

$$h_{\alpha i} = \begin{pmatrix} \epsilon_1 & \epsilon_2 & h'_3 \\ h_1 & h_2 & h_3 \\ h_1 & h_2 & -h_3 \end{pmatrix} + \mathcal{O}(\epsilon^2)$$



contours of $\text{Br}(\mu \rightarrow e\gamma) = 2.4 \times 10^{-12}$

loop function:

$$F_2(x) = \frac{1 - 6x + 3x^2 + 2x^3 - 6x^2 \log x}{6(1-x)^4}$$

$$\epsilon_4 = \frac{1}{\sqrt{2}} \frac{\tan \theta_{12} h_3^2 \Lambda_3}{(h_1^2 + h_2^2) \Lambda_1 - h_3^2 \Lambda_3} \epsilon_3$$

diagonalization of mass matrix: $\epsilon_1 h_1 + \epsilon_2 h_2 \approx \sqrt{2} (h_1^2 + h_2^2) \equiv P \epsilon_3$.

LFV: $\text{Br}(\mu \rightarrow e\gamma) = \frac{3\alpha_{\text{em}}}{64\pi G_F^2 M_\eta^4} \left| P \epsilon_3 F_2 \left(\frac{M_1^2}{M_\eta^2} \right) + \sqrt{2} \tan \theta_{12} |h_3|^2 F_2 \left(\frac{M_3^2}{M_\eta^2} \right) \right|^2$

APPENDIX C: DARK MATTER

$$\sigma_{\text{eff}} v = a_{\text{eff}} + b_{\text{eff}} v^2 + \mathcal{O}(v^4)$$

$$a_{\text{eff}}=\frac{\xi^2}{2\pi}\frac{M_1^2}{\left(M_\eta^2+M_1^2\right)^2}$$

$$b_{\text{eff}}=\frac{|h_1^2+h_2^2|^2}{24\pi}\frac{M_1^2\left(M_\eta^4+M_1^4\right)}{\left(M_\eta^2+M_1^2\right)^4}+\frac{\xi^2}{2\pi}\frac{M_1^2\left(M_\eta^4-3M_\eta^2M_1^2-M_1^4\right)}{\left(M_\eta^2+M_1^2\right)^4}$$

$$\Omega h^2\simeq\frac{1.07\times10^9x_f~[\mathrm{GeV}^{-1}]}{\sqrt{g_*}m_{\mathrm{pl}}\left(a_{\mathrm{eff}}+3b_{\mathrm{eff}}/x_f\right)}$$

$$x_f=\frac{M_1}{T}$$

APPENDIX D: EFFECTIVE INTERACTIONS

$$a_{12} = - \sum_{\alpha} \frac{\text{Im} (h_{\alpha 2}^* h_{\alpha 1}) e}{2(4\pi)^2 M_{\eta}^2} I_a \left(\frac{M_1^2}{M_{\eta}^2}, \frac{m_{\alpha}^2}{M_{\eta}^2} \right)$$

$$I_a(x, y) = \frac{1}{3} \int_0^1 \frac{3u^2 - 6u + 1}{xu^2 - (1+x-y)u + 1}$$

$$\mu_{12} = - \sum_{\alpha} \frac{\text{Im} (h_{\alpha 2}^* h_{\alpha 1}) e}{2(4\pi)^2 M_{\eta}^2} 2M_1 I_m \left(\frac{M_1^2}{M_{\eta}^2}, \frac{m_{\alpha}^2}{M_{\eta}^2} \right)$$

$$I_m(x, y) = - \int_0^1 \frac{u(1-u)}{xu^2 - (1+x-y)u + 1}$$

$$c_{12} = \sum_{\alpha} \frac{\text{Im} (h_{\alpha 2}^* h_{\alpha 1}) e}{2(4\pi)^2 M_{\eta}^2} q^2 I_c \left(\frac{M_1^2}{M_{\eta}^2}, \frac{m_{\alpha}^2}{M_{\eta}^2} \right)$$

$$I_c(x, y) = I_m(x, y)$$

APPENDIX E: DIRECT DETECTION

magnetic moments in units of nuclear magneton

	$^{19}_9\text{F}$	$^{23}_{11}\text{Na}$	$^{73}_{32}\text{Ge}$	$^{127}_{53}\text{I}$	$^{131}_{54}\text{Xe}$	$^{133}_{55}\text{Cs}$	$^{183}_{74}\text{W}$
J_A	1/2	3/2	9/2	5/2	3/2	7/2	1/2
μ_A/μ_N	2.629	2.218	-0.879	2.813	0.692	2.582	0.118

energy range and the quenching factor for XENON100, KIMS,DAMA

	Energy range	Quenching factor
XENON100	8.4 – 44.6 keV	–
KIMS	3.6 – 5.8 keVee	0.1 (Cs), 0.1 (I)
DAMA	2 – 8 keVee	0.3 (Na), 0.09 (I)