

Prospects of antideuteron detection from Dark Matter annihilations or decays at AMS-02 and GAPS

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Based on [arXiv:1209.5539](#) in collaboration with Alejandro Ibarra

Outline

- 1 Indirect DM detection with antideuteron
- 2 Constraining \bar{d} prospects with PAMELA \bar{p}/p data
- 3 Coalescence model: enhancement of \bar{d} yield?
- 4 Conclusion

Indirect DM detection with antideuteron

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Search for annihilation/decay products of Dark Matter in cosmic rays

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- Motivation for using antideuterons ($\bar{d} = [\bar{p}\bar{n}]$):

\bar{d} background flux is **suppressed** for $T_{\bar{d}} \lesssim 1 \text{ GeV/n}$

\hookrightarrow kinematical suppression of $p + H \rightarrow \bar{d} (T_{\bar{d}} \lesssim 1 \text{ GeV/n}) + X$

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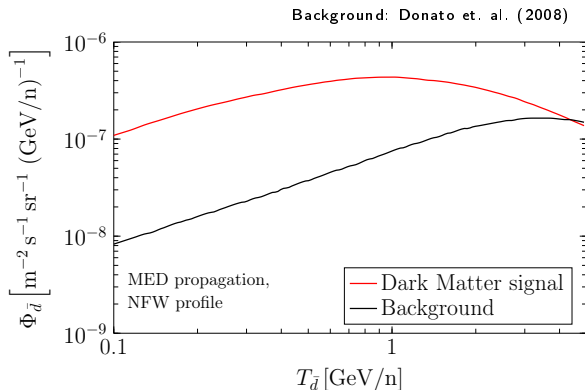
\hookrightarrow kinematical suppression of $p + H \rightarrow \bar{d} (T_{\bar{d}} \lesssim 1 \text{ GeV/n}) + X$



The detection of **only a few** low-energetic antideuterons therefore can be a **smoking-gun signal** for DM annihilations or decays

Indirect DM detection with antideuteron

- **Example:** $\chi\chi \rightarrow b\bar{b}$, $m_\chi = 100 \text{ GeV}$, $\langle\sigma v\rangle = 3 \cdot 10^{-26} \text{ cm}^3/\text{s}$



Current experimental situation

AMS-02 is currently taking data

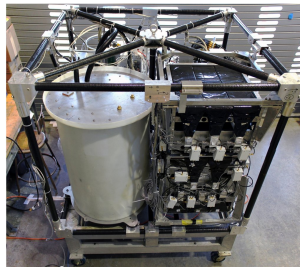


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GAPS is a balloon-borne experiment
↪ scheduled for 2016

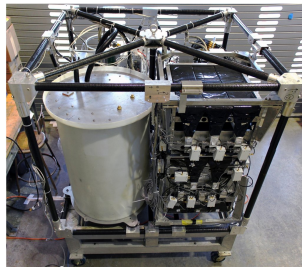


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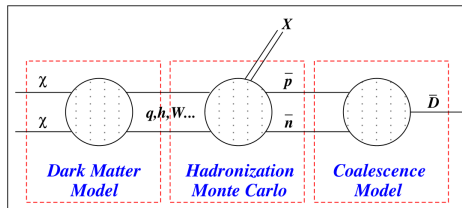


Main issue of this talk:

Prospects of DM discovery with antideuteron at AMS-02 / GAPS,
taking into account \bar{p}/p constraints from PAMELA

Production of antideuterons in DM annihilations/decays

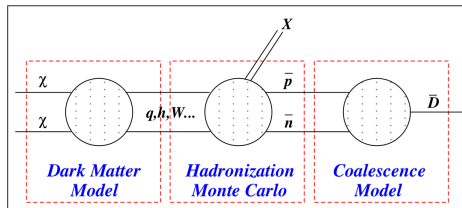
Figure from Baer, Profumo (2005)



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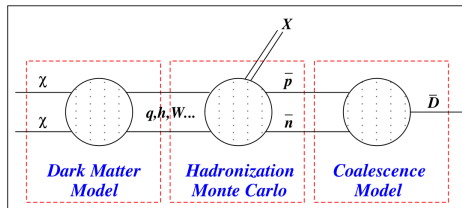
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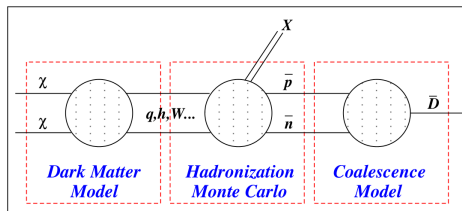
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- Considered annihilation/decay channels: $\chi\chi \rightarrow W^+ W^-$ and $\chi\chi \rightarrow b\bar{b}$
- Hadronization simulated with PYTHIA 8
- We use three different **halo profiles**: NFW, Einasto and Isothermal
 \hookrightarrow we investigate the astrophysical uncertainties regarding our results

The coalescence model

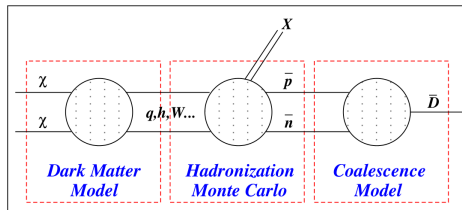
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Formation of an antideuteron: **Coalescence model**

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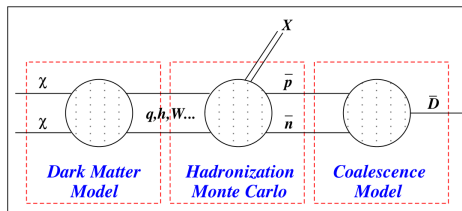
Formation of an antideuteron: **Coalescence model**

- \bar{d} forms if $|\vec{k}_{\bar{p}} - \vec{k}_{\bar{n}}| \leq p_0 = 192 \text{ MeV}$

\hookrightarrow **Coalescence momentum** p_0 is tuned against measured \bar{d} yield per Z decay

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 \hookrightarrow **Coalescence momentum** p_0 is tuned against measured \bar{d} yield per Z decay
- We search on an event-by-event basis for \bar{p} - \bar{n} pairs satisfying this condition

Propagation of antideuterons in the galaxy

- We employ the usual two-zone diffusion model
 \hookrightarrow diffusion, convection, annihilation
- We use three sample sets of propagation parameters, covering the allowed parameter space

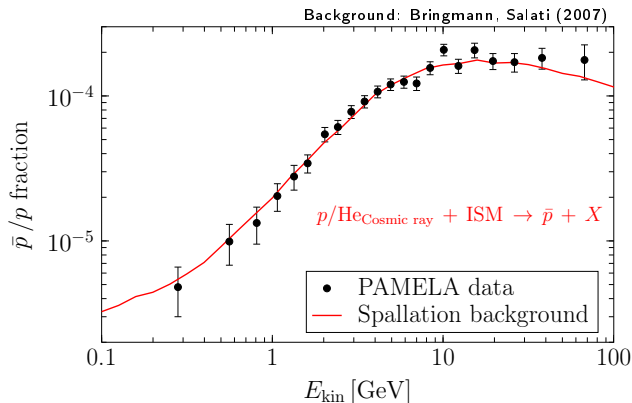
MINimal
MEDium
MAXimal } number of antideuterons

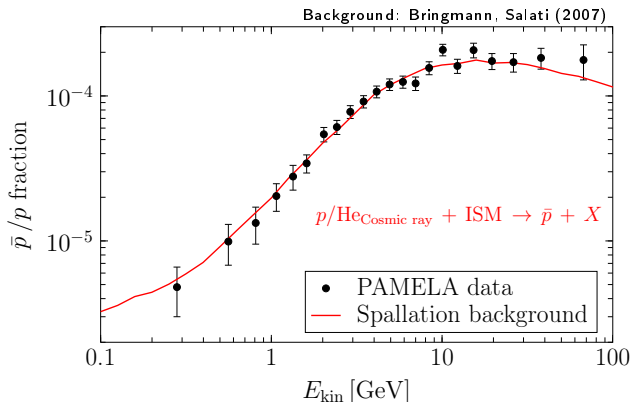
Let's ask:

How many \bar{d} events can at most be expected at AMS-02 and GAPS?



Severe constraints from
PAMELA measurement of **antiprotons**

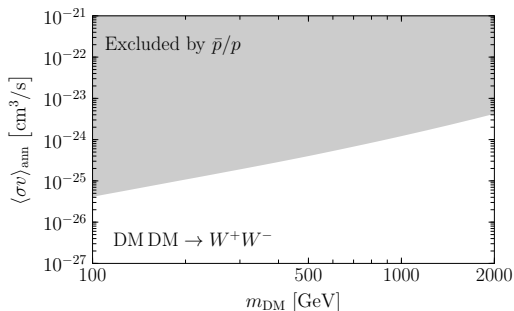
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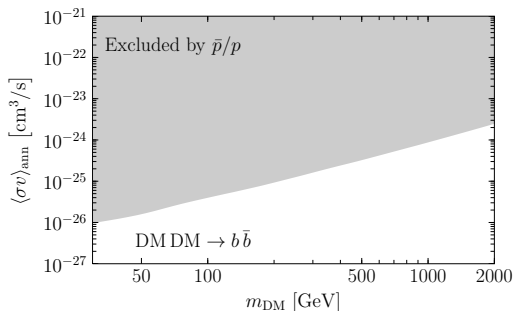
\Rightarrow No need for an exotic component

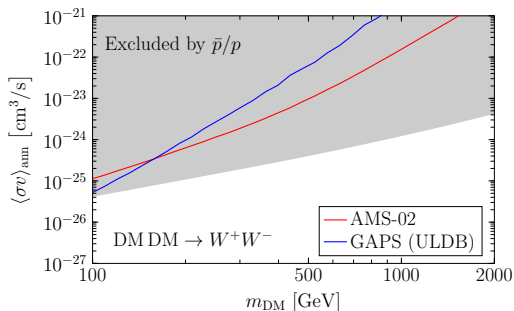
\Rightarrow **Antiproton constraints** on Dark Matter models:

Spallation background + DM induced flux \leq PAMELA data

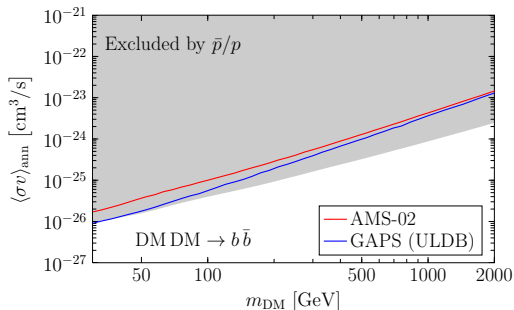


- Shaded regions: 95% C.L. exclusion from PAMELA \bar{p}/p
 \hookrightarrow using NFW profile, MED propagation parameters



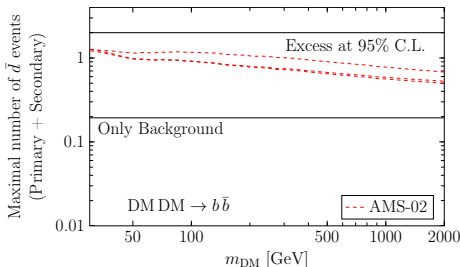
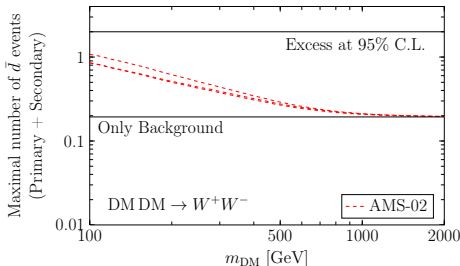


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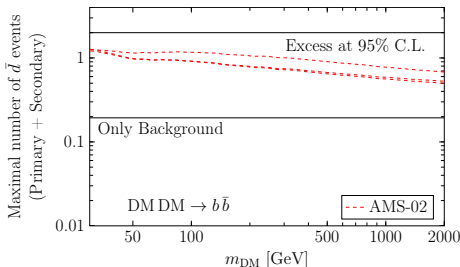
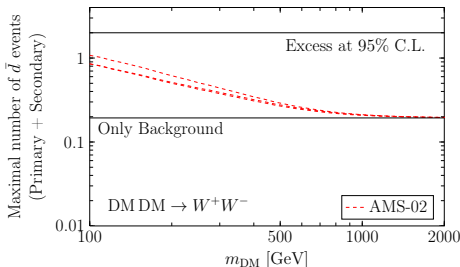


- Red and blue: cross sections necessary for an expectation of a primary \bar{d} signal at 95% C.L.

Maximal number of \bar{d} events at AMS-02

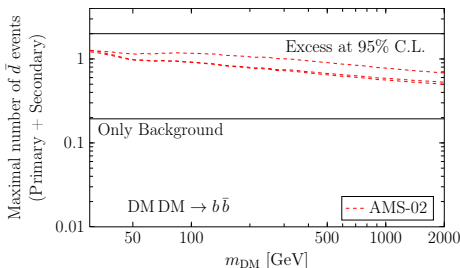
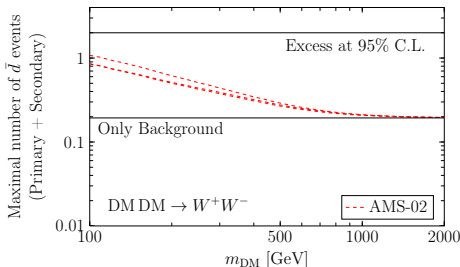


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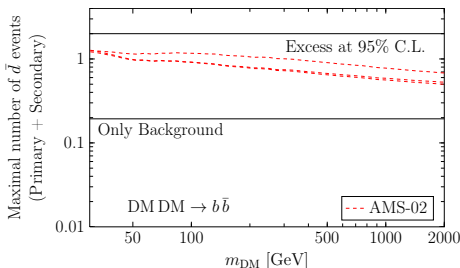
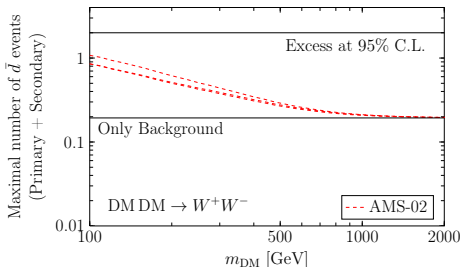
- Red curves:
Maximal number of \bar{d} at AMS-02 compatible with \bar{p}/p constraints (MIN, MED, MAX)

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- **Red curves:**
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- Propagation uncertainties largely cancel out (similar for halo profile uncertainties)

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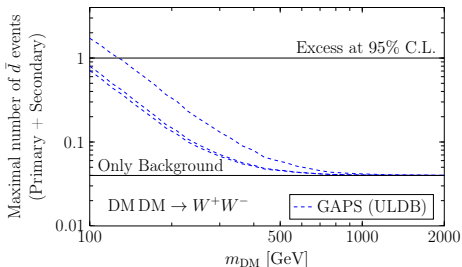


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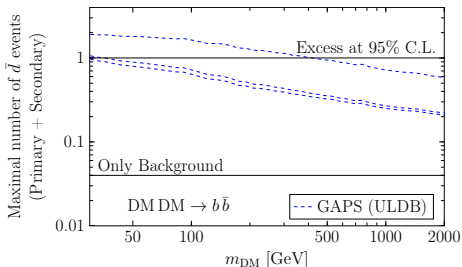


Excess at 95% C.L. (= 2 events) is in **strong tension** with \bar{p}/p constraints!

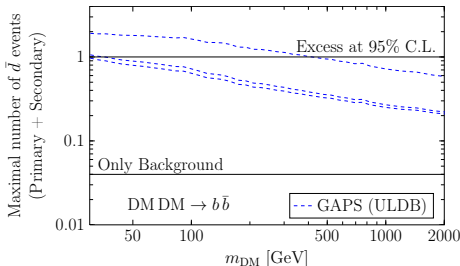
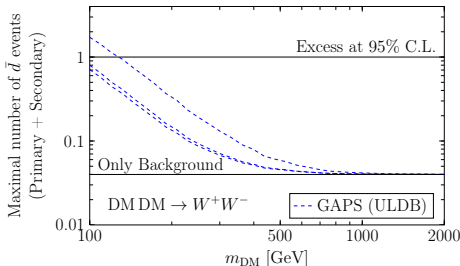
Maximal number of \bar{d} events at GAPS (ULDB)



- Blue curves:
Maximal number of \bar{d} at GAPS (ULDB) compatible with \bar{p}/p constraints (MIN, MED, MAX)



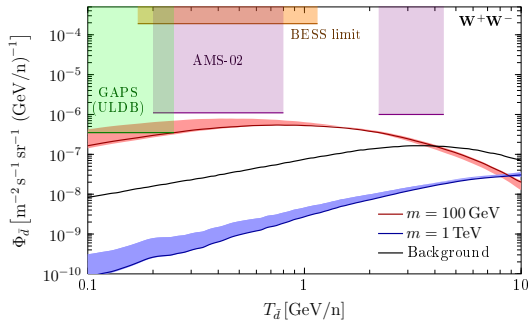
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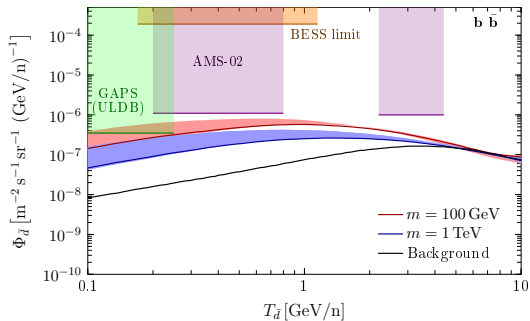
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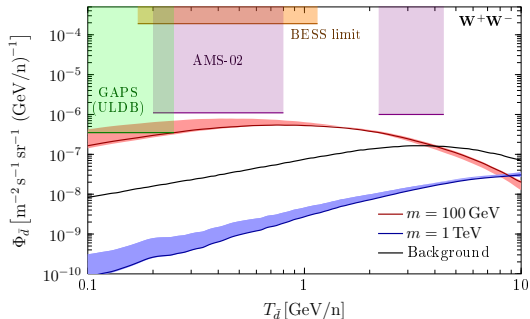


Excess at 95% C.L. (= 1 event)
only possible for MAX
propagation and
 $m_{DM} < 125$ GeV (W^+W^-)
 $m_{DM} < 400$ GeV ($b\bar{b}$)



- Red and blue curves:
Maximally allowed \bar{d} fluxes
($m_{\text{DM}} = 0.1/1 \text{ TeV}$)

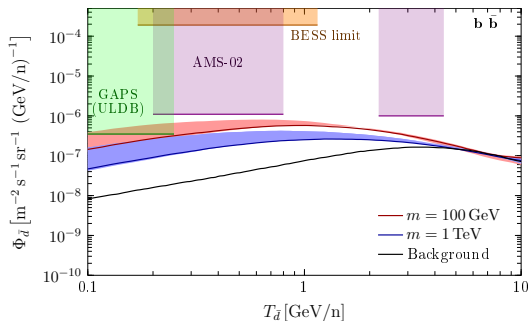




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More sensitive experiments
could detect antideuterons!



Coalescence model: enhancement of \bar{d} yield?

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- However:** By calibrating p_0 against several other experiments ($p p$, Υ decay, $e^- p$) we find a **significant dependence of p_0 on the process and the energy**: $133 \text{ MeV} \leq p_0 \leq 236 \text{ MeV}$
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Coalescence model needs further investigation!
 More lab experiments with antideuteron are needed for
 a better understanding of the coalescence process!

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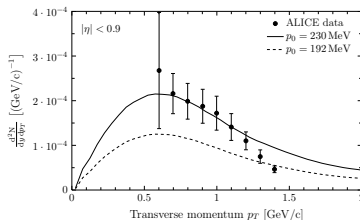
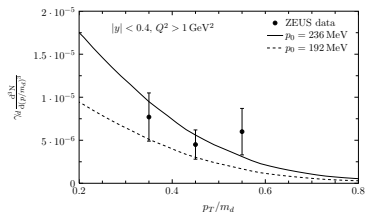
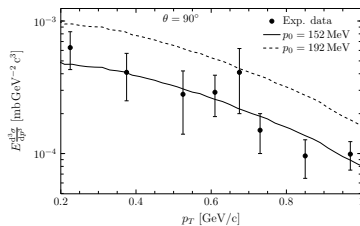
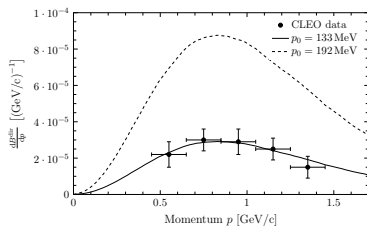
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- We found an energy and process dependent coalescence momentum p_0
 - ↪ Coalescence model seems not to be fully understood!
 - ↪ Further investigations necessary!

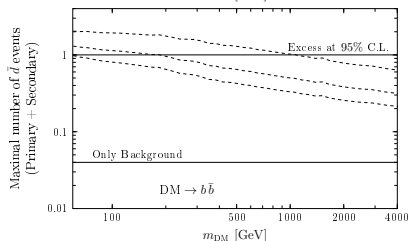
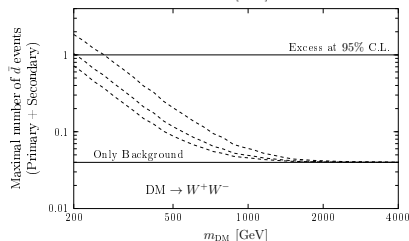
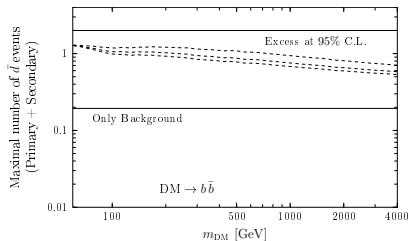
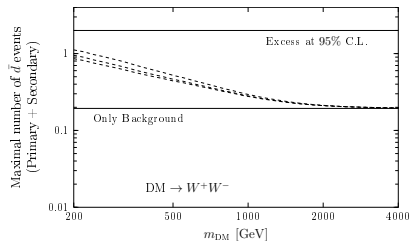
Backup slides

Determination of the coalescence momentum p_0



Top left: Υ decay, top right: pp collisions at $\sqrt{s} = 53$ GeV (CERN ISR), bottom left: e^-p collisions at $\sqrt{s} = 318$ GeV (ZEUS), bottom right: pp collisions at $\sqrt{s} = 7$ TeV (ALICE, deuteron spectrum)

Maximal number of events for decaying Dark Matter



Upper panel: AMS-02, lower panel: GAPS (ULDB)