Towards true frequentist p-values for global SUSY fits with Fittino



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FITTINO

- C++ program for SUSY model testing and SUSY parameter analysis
- Currently supported SUSY models:
 CMSSM, GMSB, AMSB, MSSM24, NMSSM, NUHM1, NUHM2
- Measurements from low/high energy experiments, direct SUSY search LEP/SLC, Tevatron, cosmology, LHC and LC, (g-2), B, K...
- Use puplic theory codes: SPheno, SuperIso, Micromegas, FeynHiggs, HDecay
- Parameter analysis using full correlation information: Auto-adaptive Markov Chain Monte Carlo (MCMC)

Previous publications:

<u>arXiv:0412012</u> [hep-ph] (2004), <u>arXiv:0511006</u> [hep-ph] (2005), <u>arXiv:0907.2589</u> [hep-ph] (2009), <u>arXiv:0909.1820</u> [hep-ph] (2009), <u>arXiv:1105.5398</u> [hep-ph] (2011), <u>arXiv:1102.4693</u> [hep-ph] (2011), <u>arXiv:1204.4199</u> [hep-ph] (2012),



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Non-minimal model: additional Higgs mass parameter: NUHM1, NUHM2



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NEW Observables



- <u>Higgs limits</u> via *HiggsBounds* or requiring $m(h^0) = (126 \pm 2 \text{ (exp)} \pm 3(\text{theo}))$ GeV

Higgs discovery via HiggsSignals (fit of m(h⁰) and μ (h⁰))

- LEP chargino limit (implications for neutralino limits)
- LHC exclusion from 'jets+MET' analysis L=4.7/fb

Fine-tuning of the simulation to match the ATLAS exclusion

- Direct detection of DM via AstroFit



generated by HiggsSignals

Higgs session: Tuesday 11:00 in SR 4a

Full 5D fit (M_0 , M_{12} , tan β , A_0 , m(top)) including all observables: low energy, astrophysics, direct searches, Higgs combined constraints (m(h^0), μ)



Pretty good fit, with tendency to higher masses and tan β (coupling throug (g-2)_u)



mininal Chi2/ndf = 33.5/31





All Observables, no combined Higgs constraints, Only asking for $m(h^0)=(125 \pm 2 \pm 3)$ GeV

First measurement of $B_s \rightarrow \mu^+ \mu^-$ by LHCb !

arxiv:1211.2674 [hep-ex]



| M ₀ : | 477.1 (730.7) | [-396.1 <i>,</i> +700.5] GeV |
|-------------------|--------------------|-------------------------------|
| M ₁₂ : | 861.3 (880.5) | [-158.0 <i>,</i> +395.0] GeV |
| A ₀ : | -2332.6 (-2225.83) | [-1975.2 <i>,</i> +768.4] GeV |
| Tanβ: | 20.5 (32.7) | [-18.0, +9.1] |
| Chi2: | 33.4 (34.0) | |

Minor impact on the best fit point, increase of M_0 and tan β





Well-spread thought:

"The fit is good if this number is roughly one"

 \rightarrow Because the mean value of a $\chi 2$ distribution is the number of degrees of freedom

"our" **x2** :
$$\chi^2 = (M - O(P))^T \operatorname{cov}_M^{-1}(M - O(P))$$

M: measurements O(P): predictions at point P covM: covariance matrix

This assumption that our $\chi 2$ is distributed according to a $\chi 2$ distribution is correct only:

- if the Oⁱ are Gaussian-distributed and if the dependency O(P) is linear.
- or if a large number of observables is used

Therefore TMath::Prob may not give the right value !!

→ Risk of drawing a wrong conclusion on the validity of a model

What is the p-value?

Assuming the best fit point found is the real one, if measurements are repeated, what is the probability to get an agreement (i.e. χ 2) at least as worse as the one observed?

Computation of the p-value of the best fit point (lower $\chi 2$) with toys:

- take the observables values at this point
- smear the observables values
- calculate the $\chi 2$ for these new pseudo-measurements
- spot the new best fit point
- repeat that procedure many times
- integrate the distribution for $\chi 2 \ge \chi 2$ (real fit)

Statistical uncertainty (binomial rule: $N_p = #(\chi 2 > \chi 2 \text{ (real fit)}), N_g = #(\text{total})$ $\frac{\sigma(X)}{\langle X \rangle} = \sqrt{\frac{1}{N_p} - \frac{1}{N_q}}$

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Results of the toys: 1st frequentist p-values for the CMSSM !

Total toy statistics per model: 1000

| Model | Toys | χ2 approx. |
|-------|-------------------|------------|
| CMSSM | 29.6 ± 1.4 | 41.7 |
| NUHM1 | 31.7 ± 1.5 | 41.4 |
| NUHM2 | 31.2 ± 1.5 | 34.7 |



Parameters distributions for the CMSSM (more in backup)



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A few observables distributions for the CMSSM (more in backup)



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Conclusion & Plans

The first real toy-based frequentist p-values for the CMSSM!

SM-like Higgs sector easily achievable within the CMSSM, NUHM

- → good agreement of the fit with the SM-like Higgs rate (Tevatron, LHC)
- → improvement in Fit Quality (45%)
- First measurement of $B_s \rightarrow \mu^+ \mu^-$ in agreement with the SM **perfectly expected within constraint SUSY**

Still room for M_{12} , $M_0 < 1$ TeV



Higgs mass and rate measurements push the CMSSM into a parameter space where the **differences to the SM in all observables (apart** Ω_{DM} **) is small** \rightarrow negligible improvement in (g-2)_u wrt the SM





 χ^2

Where is the focus point gone?



Searching for SUSY at the TeV scale ... and ?

2011: long LHC run, center-of-mass energy 7 TeV, luminosity ~5/fb.

- Direct step into Terascale
- No significant excess seen

https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/PAPERS/SUSY-2011-19/



Short reminder of the previous Fittino results

Main Conclusions

- **Current LHC exclusion leads to tension** within CMSSM
- Accommodate Higgs mass ≥ 125 GeV very hard in mSUGRA

Measurement of Higgs branching ratios can discriminate SM/MSSM



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And now what ...?



Improvements

Observables

| $\mathcal{B}(b ightarrow s \gamma)$ | $(3.55\pm0.34)	imes10^{-4}$ |
|--|------------------------------|
| $\mathcal{B}(B_s \to \mu \mu)$ | $< 4.5 	imes 10^{-9}$ |
| $\mathcal{B}(B \to \tau \nu)$ | $(1.67\pm0.39)	imes10^{-4}$ |
| Δm_{B_s} | $17.78\pm5.2{\rm ps}^{-1}$ |
| $a_{\mu}^{\mathrm{exp}}-a_{\mu}^{\mathrm{SM}}$ | $(28.7\pm8.2)	imes10^{-10}$ |
| m_W | $(80.385 \pm 0.015){ m GeV}$ |
| $\sin^2	heta_{ m eff}$ | 0.23113 ± 0.00021 |
| $\Omega_{ m CDM} h^2$ | 0.1123 ± 0.0118 |
| m _t | $(173.2\pm1.34)~{ m GeV}$ |

- <u>Higgs limits</u> via *HiggsBounds* or requiring $m(h^0) = (126 \pm 2 \text{ (exp)} \pm 3(\text{theo}))$ GeV

- LEP chargino limit (implications for neutralino limits)
- LHC exclusion from 'jets+MET' analysis L=4.7/fb
- Direct and indirect detection of DM via AstroFit

Comparison of models performed using p-values

 \rightarrow Computed assuming the observables likelihood to be Gaussian

 \rightarrow integration of a $\chi 2$ distribution function

What is the p-value?

Assuming the best fit point found is the real one, if measurements are repeated, what is the probability to get an agreement (i.e. χ 2) at least as worse as the one observed?





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Toy fits (χ^2 distribution)



Mass spectra from the previous Fittino results: JHEP06(2012)098



Figure 9. Predicted distribution of sparticle and Higgs boson masses from (a) the LEO fit and (b) the LHC fit. The full uncertainty band gives the 1-dimensional 2σ uncertainty of each mass defined by the region $\Delta\chi^2 < 4$ after profiling over all hidden dimensions.

Mass spectra from the previous Fittino results: JHEP06(2012)098



Figure 13. Predicted distribution of sparticle and Higgs boson masses from the CMSSM fit with $m_h = (126 \pm 2 \pm 3) \text{ GeV}$ in (a) and the NUHM1 fit to the same observable set in (b).



Figure 16. Predicted distribution of sparticle and Higgs boson masses from (a) the LHC and (b) the LHC+XENON1T fit. The full uncertainty band gives the 1-dimensional 2σ uncertainty of each mass defined by the region $\Delta \chi^2 < 4$ after profiling over all hidden dimensions. Note the different scales on the ordinate (mass) axis compared to figure 9.