Comments on Racetrack Kahler Uplift

Sumitomo, Tye
 arXiv:1204.5177, JCAP 1208 (2012) 032
 arXiv:1209.5086, JCAP 1302 (2013) 006
 arXiv:1211.6856, PLB 723 (2013) 406-410

Sumitomo, Tye, Wong
 arXiv:1305.0753, JHEP 07 (2013) 052

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Just moved! (June 17)







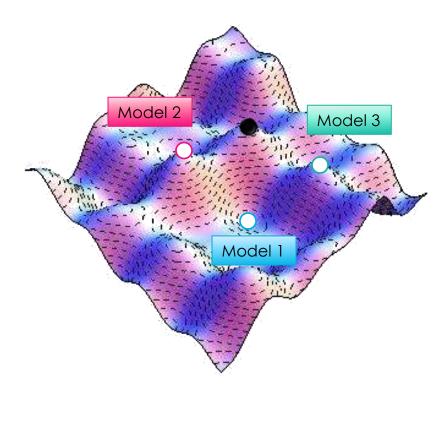


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Landscape

Metastable vacua in moduli space



Inflation I rolling down (& tunneling)

dS vacua

Low energy

Initial conditions?

On which directions (models), can we achieve the tiny cosmological constant?

 $\Lambda \sim 10^{-123} M_P^4$

Stringy Landscape

There are many types of vacua in string theory, as a result of a variety of (Calabi-Yau) compactification.

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 $ds_{10}^2 = ds_4^2 + ds_6^2$

A class of Calabi-Yau gives Swiss-cheese type of volume.

$$\mathcal{V}_6 = \gamma_1 (T_1 + \bar{T}_1)^{3/2} - \sum_{i=2} \gamma_i (T_i + \bar{T}_i)^{3/2}$$
,

Examples:

[Denef, Doualas, Florea, 04]

• $\mathbb{P}^4_{[1,1,1,6,9]}$: $h^{1,1} = 2$, $h^{2,1} = 272$ All can be stabilized

•
$$\mathcal{F}_{11}$$
: $h^{1,1} = 3$, $h^{2,1} = 111$

•
$$\mathcal{F}_{18}$$
: $h^{1,1} = 5$, $h^{2,1} = 89$

(a la KKLT),

with a variety of fluxes.

 $(h^{1,1}: # \text{ of Kahler}, h^{2,1}: # \text{ of c.s. moduli})$

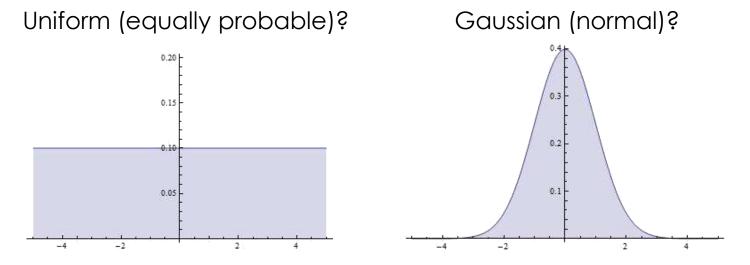
More recently, for $2 \le h^{1,1} \le 4$, 418 manifolds

Rich vacuum structures!

[Gray, He, Jejjala, Jurke, Nelson, Simon, 12]

Distribution

What distribution do you imagine for physical quantities?



Yes, if the system is quite simple enough.

However, complicated system + moduli stabilization

somewhat non-trivial distributions

Statistical Approach in Stringy Models

Kahler Uplift

[Balasubramanian, Berglund, 04], [Westphal, 06], [Rummel, Westphal, 11], [de Alwis, Givens, 11]

Similar setup as that of Large Volume Scenario

$$K = -2 \ln \left(\mathcal{V}_6 + \frac{\xi}{2} \right) + \cdots, \qquad \mathcal{V}_6 = \gamma_1 (T_1 + \overline{T}_1)^{3/2} - \sum_{i=2} \gamma_i (T_i + \overline{T}_i)^{3/2},$$

$$W = W_0 + A_1 e^{-a_1 T_1} + \sum_{i=2} A_i e^{-a_i T_i}$$
Swiss-cheese
This term plays a roll. (cf. negligible in LVS)

less large volume than LVS, but still $|W_0| \gg |W_{NP}|, \mathcal{V} \gg \xi$ E.g. single modulus [Rummel, Westphal, 11],.010 $\Lambda \equiv V \Big|_{\min} \sim \frac{1}{9} \left(\frac{2}{5}\right)^{\frac{9}{2}} \frac{-W_0 a_1^3 A_1}{\gamma_1^2} (C - 3.65) \overset{0.004}{0.002}$ $C = \frac{-27W_0 \xi a_1^{\frac{3}{2}}}{64\sqrt{2}\gamma_1^2 A_1}$

Distribution of Kahler Uplift

Starting with the simplified potential:

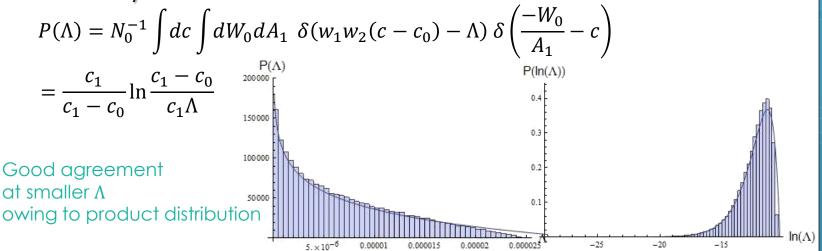
[YS, Tye, 12 (Apr)]

$$\Lambda \propto -W_0 A_1 (c - c_0), \qquad c_0 \le c = \frac{-W_0}{A_1} < c_1$$

Many ways to fix W_0 , A_1 due to varieties of fluxes, and that each c.s. moduli (with $h^{2,1}$) stabilization gives different values.

 $\Rightarrow W_0, A_1 \in [-1, 1],$ uniform distribution (for simplicity)

Probability distribution function



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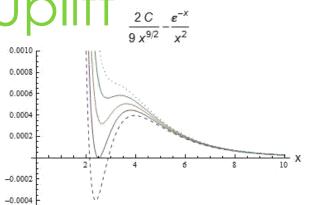


Kahler Uplift has a potential problem:

 $x_1 = a_1 Re T_1 \leq 3.11$

by tadpole condition.

Volume moduli is restricted from above.



 $a_1 = \frac{2\pi}{N}$ (for SU(N)) would be bounded below [Cicoli, Mayrhofer, Valandro, 11] [Louis, Rummel, Valandro, Westphal, 12]

e.g. [Pedro, Rummel, Westphal, 13] concerns about α' -corrections

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Let's introduce racetrack superpotential:

$$W = W_0 + A_1 e^{-a_1 T_1} + B_1 e^{-b_1 T_1}$$



[YS, Tye, Wong, 13] There opens up a new region of solutions.

(See also [Westphal, 05], [de Alwis, Givens, 11])

Solutions in Racetrack Kahler Uplift

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dS vacua at large x_1 (> 3.11): [YS, Tye, Wong, 13]

$$x_1 = a_1 \operatorname{Re} T_1 \sim \frac{1}{\beta - 1} \ln \left[\frac{\beta^3}{-z} \right]$$

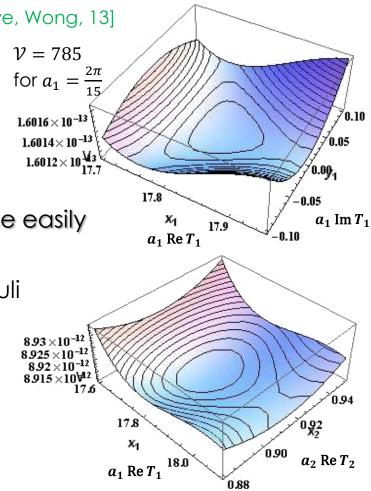
when
$$\beta = b_1/a_1 \approx 1$$
, $|z| = |A_1/B_1| \ll 1$

No upper bound 📃 Larger volume easily

Applicable even for multi Kahler moduli

$$\mathcal{V}_6 = \gamma_1 (T_1 + \bar{T}_1)^{3/2} - \sum_{i=2} \gamma_i (T_i + \bar{T}_i)^{3/2}$$
$$W = W_0 + A_1 e^{-a_1 T_1} + B_1 e^{-b_1 T_1} + \sum_{i=2} A_i e^{-a_i T_i}$$

Large volume holds.



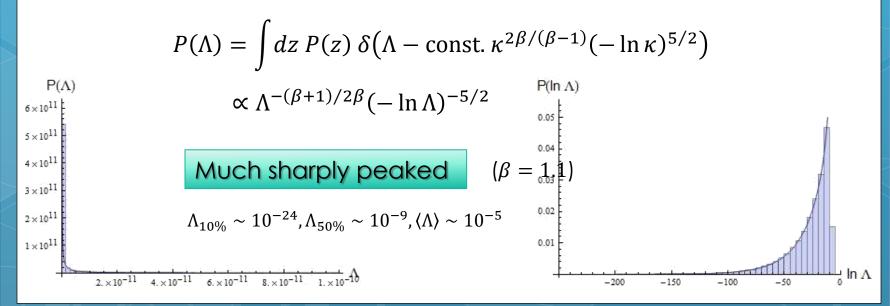
 $W = W_0 + A_1 e^{-a_1 T_1} + B_1 e^{-b_1 T_1}$

Distributions of Racetrack Kahler Uplift

As $\beta \approx 1$, exponentially suppressed CC is realized.

Cosmological constant

Introduce uniformly distributed $-1 \le z = A_1/B_1 \le 0$.



 $\Lambda \propto \kappa^{2\beta/(\beta-1)} (-\ln \kappa)^{5/2}, \qquad \kappa = \frac{-z}{\beta^3} \qquad \qquad \beta = b_1/a_1 \approx 1, \\ |z| = |A_1/B_1| < 1$

[YS, Tye, 12 (Nov)]

Summary & Discussion

Random mini-Landscape in String Theory

We may expect that stringy motivated models have the following properties:

- Non-trivial function (Racetrack Kahler Uplift, SUSY KKLT)
- Product of parameters (Kahler Uplift)
- Correlation of each term via dynamics (multi-moduli)
 [YS, Tye, 12 (Sep)]

All seem to work for diverging peaked distributions that prefer smaller physical quantities.

Complex sector stabilization may make W_0 smaller. Smallness of W_0 is good for the hierarchical structure.

Kahler stabilization with NP-effect is more reliable in the presence of large $h^{2,1}$. [YS, Tye, 12 (Nov)] (Probability of positive mass matrix increases as $h^{2,1}$ increases)

Detail of Racetrack Kahler Uplift

$$\frac{1}{z} \equiv \frac{B_1}{A_1} \sim -\frac{1}{\beta^3} e^{(\beta-1)x_1}, \qquad \hat{C} \equiv -\frac{3a_1^{3/2}W_0\xi}{32\sqrt{2}A_1} \sim \frac{2(\beta-1)}{9\beta} e^{-x_1}x_1^{7/2}$$

or

$$x_1 \sim \frac{1}{\beta - 1} \ln\left(\frac{\beta^3}{-z}\right), \qquad \hat{C} \sim \frac{2}{9\beta(\beta - 1)^{5/2}} \left(\frac{\beta^3}{-z}\right)^{-1/(\beta - 1)} \left(\ln\frac{\beta^3}{-z}\right)^{7/2}$$

Solution:

$$W_0 = -0.223, A_1 = 1.65, B_1 = -4.77, a_1 = \frac{2\pi}{15}, b_1 = \frac{2\pi}{14}, \xi = 3.41 \times 10^{-3}$$

or

$$z = -0.346, \hat{C} = 8.28 \times 10^{-6}, \beta = 1.07, \mathcal{V} = 785$$