Evolution and Dynamics of Cusped Light-Like Wilson Loops in Loop Space

3rd Workshop on the QCD Structure of the Nucleon Bilbao, Spain

Frederik F. Van der Veken † Igor O. Cherednikov $^{\dagger\,\gamma}$ Tom Mertens †

[†]Universiteit Antwerpen $^{\gamma}$ JINR Dubna

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2 Wilson Loops and their Evolution

3 Geometrical Approach



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			TMDs	
Transverse M	omentum Depend	lent PDFs		
$f(x,\mathbf{k}_{\perp})=\frac{1}{2}$	$\int \frac{\mathrm{d}z^- \mathrm{d}^2 \mathbf{z}_\perp}{2\pi (2\pi)^2} e^{ik \cdot z}$	$z \langle P, S \overline{\psi}(z) \mathcal{U}^{\dagger}_{(z;\infty)} \mathcal{U}_{(\infty;z)}$	$_{0)}\psi(0)\left P,S\right\rangle\Big _{z^{+}=0}$	
Gauge link:	$\mathcal{U}(z;\infty) = e^{-\mathrm{i}g}$	$\int_{\infty}^{z} \mathrm{d}x^{\mu} A_{\mu}(x)$		I

Singularities

- UV poles $\sim rac{1}{\epsilon}$
- Rapidity divergencies $\, \sim \ln heta \,$
- ullet Overlapping divergencies $\,\sim rac{1}{\epsilon} \ln heta\,$ generalised renormalisatio
- Self-energy divergencies

removed by standard renormalisation resummed by use of Collins-Soper

treated by modification of soft factors

See talk by I.O. Cherednikov on friday.

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Transverse M	omentum Depenc	lent PDFs	
$f(x,\mathbf{k}_{\perp}) = \frac{1}{2}$	$\int \frac{\mathrm{d}z^- \mathrm{d}^2 \mathbf{z}_\perp}{2\pi (2\pi)^2} e^{ik \cdot z}$	$\mathbb{P}\left\langle P,S \overline{\psi}(z) \mathcal{U}_{(z;\infty)}^{\dagger} \mathcal{U}_{(\infty)} \right\rangle$	$_{0)}\psi(0)\left P,S\right\rangle\Big _{z^{+}=0}$
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Singularities

- UV poles $\sim \frac{1}{\epsilon}$
- Rapidity divergencies $\sim \ln \theta$
- Overlapping divergencies $\sim \frac{1}{\epsilon} \ln \theta$
- Self-energy divergencies

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- resummed by use of Collins-Soper
- generalised renormalisation

treated by modification of soft factors

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			TMDs
Cusp Anomalous	s Dimension		
(0 ⁺ ,0 ⁻ , 0 _⊥)	$(0^+, x^-, \mathbf{x}_\perp)$	• "Hidden cusp" con • $\Gamma_{cusp} = \frac{\alpha_s C_F}{\pi} (\chi$ • $\cosh \chi = \frac{p \cdot k}{ p k } \stackrel{\text{on}}{\rightarrow}$ • $\Gamma_{cusp} \stackrel{\text{on-LC}}{\rightarrow} \frac{\alpha_s C_F}{\pi}$	tribution to Γ_{cusp} $\cosh \chi - 1$) $\stackrel{\text{LC}}{\rightarrow} \infty$

Nucl. Phys. B283 (1987) 342-364, G.P. Korchemsky and A.V. Radyushkin Phys. Rev. D86 (2012) 085035, I.O. Cherednikov, T. Mertens and F.F. VdV.

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		Definition	f Wilson Loona	

Definition of Wilson Loops

Definition

$$\mathcal{W}[C] = \frac{1}{N_c} \operatorname{tr} \langle 0 | \mathcal{P}e^{ig \oint_C dz^{\mu} A^a_{\mu}(z)t_a} | 0 \rangle$$

$$\mathcal{C}: \quad z^{\mu}(s) \quad s = 0 \dots 1 \quad \text{where} \quad z^{\mu}(0) \equiv z^{\mu}(1)$$

Wilson Loops...

- are characterised by their geometry
- can be used as elementary objects defining QCD (in loop space)
- can exhibit dualities to objects in standard QCD, depending on the structure of the loop

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- Duality between a planar Wilson loop made up from N light-like segments and the N-gluon planar scattering amplitude in $\mathcal{N}=4$ SYM.
- Momenta p_i of external gluons in SYM are equal to light-like segment lengths $p_i \equiv x_i x_{i+1}$ of the loop



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- The UV singularities of the Wilson loop are related to the IR singularities of the ${\cal N}=4$ SYM scattering amplitude.
- The evolution equation for the amplitude as function of IR-cutoff is governed by the cusp anomalous dimension of the Wilson loop.



- It is thus instructive to investigate further how the Wilson loop behaves in function of (changes of) its geometry and derive its evolution equations.
- These might lead to evolution equations for TMDs, as the singular parts of the latter are related to those of Wilson loops.

Leading Order Calculation

Expansion and Calculation



- Mandelstam Energy/rapidity variables: $s = (p_1 + p_2)^2, t = (p_2 + p_3)^2$
- s = -t, but watch out for pole structure: $(-s + i\varepsilon)^{\epsilon} + (-t + i\varepsilon)^{\epsilon}$
- Absorb π and expansion of $\Gamma(1-\epsilon)$ in μ : $\overline{\mu}^2 = \pi \mu^2 e^{\gamma_E}$

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One-Loop Evolution

$$\mathcal{W}_{LO} = 1 - \frac{\alpha_s C_F}{\pi} \left[\frac{1}{\epsilon^2} \left((-\overline{s} + i\varepsilon)^\epsilon + (\overline{s} + i\varepsilon)^\epsilon \right) + \frac{\pi^2}{2} - 2\zeta(2) + \mathcal{O}(\epsilon) \right]$$

Not multiplicatively renormalisable due to extra light-cone divergencies

Dual to the TMD case

Energy logarithmic derivative gives an evolution equation
$$(\overline{s} = \overline{\mu}^2 s)$$
:

$$\frac{\mathrm{dln} \mathcal{W}}{\mathrm{dln} s} = -\frac{\alpha_s C_F}{\pi} \frac{1}{\epsilon} \left[(\overline{s} + \mathrm{i}\varepsilon)^\epsilon - (-\overline{s} + \mathrm{i}\varepsilon)^\epsilon \right]$$

$$\frac{\mathrm{d}}{\mathrm{dln} \mu} \frac{\mathrm{dln} \mathcal{W}}{\mathrm{dln} s} = -2 \frac{\alpha_s C_F}{\pi} \left[(\overline{s} + \mathrm{i}\varepsilon)^\epsilon + (-\overline{s} + \mathrm{i}\varepsilon)^\epsilon \right] \rightarrow -2 \frac{\alpha_s C_F}{\pi}$$

$$= -2\Gamma_{cusp}$$

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Conclusion

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 $\begin{array}{c|c} \hline \mbox{Mison Loops} & \hline \mbox{Geometrical Approach} & \hline \mbox{Conclusion} & \hline \mbox{Conclusion} & \hline \mbox{Makeenko-Migdal Approach} & \hline \mbox{Makeenko-Migdal Approach} & \hline \mbox{Mison Loops as Fundamental Gauge-Invariant Degrees of Freedom} & \hline \mbox{Wilson Loops as Fundamental Gauge-Invariant Degrees of Freedom} & \hline \mbox{W}_n(\Gamma_1,\ldots,\Gamma_n) = \mathrm{tr}\,\langle 0 | \, \mathcal{T} \frac{1}{N_c^n} \Phi(\Gamma_1) \cdots \Phi(\Gamma_n) \, | 0 \rangle \\ & \Phi(\Gamma_i) = \mathcal{P} e^{\mathrm{i}g \int_{\Gamma_i} \mathrm{d} z^\mu A_\mu(z)} & \hline \end{array}$

FormalismArea derivative: $\frac{\delta}{\delta\sigma_{\mu\nu}(x)} \Phi(\Gamma) = \lim_{|\delta\sigma|} \frac{\Phi(\Gamma\delta\Gamma_x) - \Phi(\Gamma)}{|\delta\sigma_{\mu\nu}(x)|}$ $\frac{\delta}{\delta\sigma_{\mu\nu}(x)} \Phi(\Gamma)$ Path derivative: $\partial_\mu \Phi(\Gamma) = \lim_{|\delta x_\mu|} \frac{\Phi(\delta x_\mu^{-1} \Gamma \delta x_\mu) - \Phi(\Gamma)}{|\delta x_\mu|}$ $\frac{\delta}{\delta\sigma_{\mu\nu}(x)}$

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$$\Phi(\Gamma_i) = \mathcal{P}e^{\mathrm{i}g\int_{\Gamma_i} \mathrm{d}z^{\mu}A_{\mu}(z)}$$

Formalism

Area derivative:
$$\frac{\delta}{\delta\sigma_{\mu\nu}(x)}\Phi(\Gamma) = \lim_{|\delta\sigma|} \frac{\Phi(\Gamma\delta\Gamma_x) - \Phi(\Gamma)}{|\delta\sigma_{\mu\nu}(x)|} \qquad \bigcirc$$
Path derivative:
$$\partial_{\mu}\Phi(\Gamma) = \lim_{|\delta x_{\mu}|} \frac{\Phi(\delta x_{\mu}^{-1}\Gamma\delta x_{\mu}) - \Phi(\Gamma)}{|\delta x_{\mu}|} \qquad \bigcirc$$

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Evolution of Cusped Loops

δΓ



$$\frac{\delta}{\delta\sigma_{\mu\nu}(x)}\operatorname{tr}\Phi(\Gamma) = \operatorname{i}g\operatorname{tr}\left\{F^{\mu\nu}(x)\Phi(\Gamma)\right\}$$

- Relates geometry evolution of a loop with its gauge content
- But area functional derivative is not well-defined for arbitrary contours..

 \Rightarrow no information about cusps, divergencies, etc...!

Schwinger approach

- Fundamental quantum dynamical principle: $\delta \langle | ' \rangle = \frac{i}{\hbar} \langle | \delta S | '' \rangle$
- Applied to \mathcal{W}_1 , this gives the Makeenko-Migdal equation:

$$\partial^{\nu} \frac{\delta}{\delta \sigma_{\mu\nu}(x)} \mathcal{W}_1(\Gamma) = g^2 N_c \oint_{\Gamma} \mathrm{d}z^{\mu} \, \delta^{(4)}(x-z) \mathcal{W}_2(\Gamma_{xz} \Gamma_{zx})$$

Migdal (1980); Makeenko, Migdal (1981); Brandt et al. (1982)

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Migdal (1980); Makeenko, Migdal (1981); Brandt et al. (1982)

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Remarks:

- MM equation is exact and non-perturbative, but not closed and difficult.
- Most interesting loops are divergent and have obstructions. Of particular interest are cusped loops, but in that case we lack a renormalised version of the MM equation.
- The area functional derivative is not a well-defined operation. In particular, the area differentiation of cusped loops is (at least) not straightforward.
- Problems when doing continuous deformations of loops in Minkowski space: a consistent definition of the derivatives is missing.
- Connecting loop functionals to observables is highly non-trivial.
- No known solution to the MM equations in four-dimensional Mikowskian space-time.

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Makeenko-Migdal Approach

Simplifications:

- Large- N_c limit: factorisation property $\mathcal{W}_2(\Gamma_1, \Gamma_2) \approx \mathcal{W}_1(\Gamma_1)\mathcal{W}_1(\Gamma_2)$.
- Null-plane light-cone rectangular contours are essentially two-dimensional.
- Light-like polygons with conserved angles: no angle-dependent contributions (which might break the MM equation).
- By using area differentiation the power of divergencies decreases.

Can we combine the geometric approach of the MM equations with the evolution equations governed by the cusp anomalous dimension?

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Angle-Conserving Deformations of the Null Plane

Deformations



 $\bullet~$ Null-plane defined by $\mathbf{z}_{\perp}\equiv 0$

• Light-like lines:
$$p_i^2 = (x_i - x_{i+1})^2 \equiv 0$$

•
$$\delta \sigma^{+-} = \oint \mathrm{d} z^- z^+ = p^+ \delta p^-$$

$$\delta\sigma^{-+} = \oint \mathrm{d}z^+ \, z^- = p^- \delta p^+$$

• Area variable:
$$\Sigma = p_1 \cdot p_2 = rac{1}{2}s = p^+p^-$$

Use energy evolution to find the geometrical evolution:

$$\frac{\mathrm{d}}{\mathrm{d}\ln\mu}\frac{\mathrm{d}\ln\mathcal{W}}{\mathrm{d}\ln s} = -2\Gamma_{cusp}$$

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Angle-Conserving Deformations of the Null Plane

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- Null-plane defined by $\mathbf{z}_{\perp}\equiv 0$
- Light-like lines: $p_i^2 = (x_i x_{i+1})^2 \equiv 0$
- $\delta \sigma^{+-} = \oint \mathrm{d} z^- z^+ = p^+ \delta p^-$

$$\partial \partial = \mathcal{Y} dz^2 z = p^2 \partial p^2$$

• Area variable:
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Use energy evolution to find the geometrical evolution:

$$\frac{\mathrm{d}}{\mathrm{dln}\,\mu}\frac{\mathrm{dln}\,\mathcal{W}}{\mathrm{dln}\,\boldsymbol{s}} = -2\Gamma_{cusp}$$

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• Area variable:
$$\Sigma = p_1 \cdot p_2 = rac{1}{2}s = p^+p^-$$

Use energy evolution to find the geometrical evolution:

$$\frac{\mathrm{d}}{\mathrm{dln}\,\mu}\frac{\mathrm{dln}\,\mathcal{W}}{\mathrm{dln}\,\Sigma} = -4\Gamma_{cusp}$$

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Angle-Conserving Deformations of the Null Plane

Deformations



- Null-plane defined by $\mathbf{z}_{\perp} \equiv 0$
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$$\delta \sigma^{+-} = \oint dz^- z^+ = p^+ \delta p^-$$

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• Area variable:
$$\Sigma = p_1 \cdot p_2 = \frac{1}{2}s = p^+p^-$$

Motivated by this, we conjecture a more general evolution equation:

$$\frac{\mathrm{d}}{\mathrm{dln}\,\mu} \left[\sigma_{\mu\nu} \frac{\delta}{\delta \sigma_{\mu\nu}} \ln \mathcal{W} \right] = -\sum_{i} \Gamma_{cusp}$$

Phys. Rev. D86 (2012) 085035, I.O. Cherednikov, T. Mertens and F.F. VdV.

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			Relation to TMDs
Relation to TN	ИDs		

- Related by overlapping divergencies ($\sim \frac{1}{\epsilon} \ln \theta$).
- Those are the only divergencies that contribute after $\frac{d}{d \ln \mu} \frac{d}{d \ln \theta}$.
- $\theta = \frac{\eta}{p \cdot N^{-}}$ is the rapidity cut-off

•
$$p \sim N^+$$
, so $\theta \sim (N^+N^-)^{-1}$

Cherednikov, Stefanis (2008)

Evolution Equation for TMD On the Light-Cone

$$\frac{\mathrm{d}}{\mathrm{dln}\,\mu}\left[\frac{\mathrm{d}}{\mathrm{dln}\,\theta}\ln f(x,\mathbf{k}_{\perp})\right] = 2\Gamma_{cusp}$$

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			Conclusion

- Loop space:
 - Wilson loops as fundamental degrees of freedom
 - Makeenko-Migdal approach gives good description of geometrical properties
 - But system of MM eqs is not closed and cannot be applied trivially...
- For \square : $\delta_{\sigma_{\mu\nu}} \sim \partial_s$ and MM eqs \sim energy/rapidity evolution eqs.
- Geometrical properties of loop space \sim dynamics in cusps
- Conjecture: MM approach can be applied to construct energy/rapidity evolution equations in many interesting situations. Specific properties are determined by contours with cusps.