## Dominikus Hellgartner

## Technische Universität München Physik Department, E15

522. Wilhelm and Else Heraeus-Seminar

## $2^{\text {nd }}$ January 2013

## Contents

Motivation
Long baseline neutrino oscillations
Performance of LENA without tracking
Backtracking
Algorithm
First results
Wonsak's tracking
Application to LENA events
Application to CNGS- $\mu^{-}$in Borexino
Likelihood-fit
The PDF
Application to simple events
Conclusion

## Long baseline neutrino oscillations

Physics goals

- Determination of the mass hierarchy
- Measurement of the CP-violating phase $\delta_{\mathrm{CP}}$


## Detector requirements

- Good energy resolution
- Large target mass and cheap target material
- Capability to discriminate between $\nu_{\mu}$-CC, $\nu_{e}$-CC and $\nu_{x}$-NC interactions


## Current situation for LENA

NC background discrimination

- Based on multivariate analysis (boosted decision trees)
- Input parameters from overall photon pulse shape
- Only first photon time and total charge on each PMT used
- Time of flight correction w.r.t. charge barycenter

Sensitivity ( $2300 \mathrm{~km}, 10^{20} \mathrm{POT} / \mathrm{a}$ )

Mass Hierarchy



## How to improve the situation



- The arrival times of the photons on the individual PMTs contains important information
$\Rightarrow$ Go beyond the overall pulse shape
$\Rightarrow$ Use the hit pattern to reconstruct general event structure
$\Rightarrow$ Additional input parameters for multivariate analysis


## Basic considerations for LSc detectors

Problem: Scintillation photons are emitted isotropically
$\Rightarrow$ No directional information from the charge distribution
$\Rightarrow$ Use photons' arrival times for track reconstruction
General idea

- Isotropic emission over total track length
- Superposition of spherical "waves" leads to first photon cone
- The shape of the cone contains information about the track direction



## Contents

Motivation
Long baseline neutrino oscillations
Performance of LENA without tracking
Backtracking
Algorithm
First results

```
Wonsak's tracking
    Application to LENA events
    Application to CNGS- }\mp@subsup{\mu}{}{-}\mathrm{ in Borexino
Likelihood-fit
    The PDF
    Application to simple events
```

Conclusion

## Basics

## Requirements

- Obtain bubble chamber like images
- Do not require any input knowledge
$\Rightarrow$ Goal: Get basic picture of an event


## General idea

- Use only PMTs with a high charge
$\Rightarrow$ The first detected photon on each PMT is emitted instantly
$\Rightarrow$ The first detected photon on each PMT is not scattered
$\Rightarrow$ Time resolution dominated by PMTs
- Photons from a point source in the detector cluster in time after taking the photon TOF into account

This algorithm is currently developed by Kai Loo (University of Jyväskylä)

## Algorithm I

1. Choose a point to qualify whether a track was there: $\boldsymbol{x}_{g}$
2. Create a vector with the TOF-corrected hit times w.r.t. $\boldsymbol{x}_{g}$ :

$$
\boldsymbol{t}^{\mathrm{TOF}}=\left(t_{i}^{\mathrm{TOF}}\right)=\left(t_{i}^{\mathrm{hit}}-\frac{n}{c}\left|\boldsymbol{x}_{g}-\boldsymbol{x}_{i}^{\mathrm{PMT}}\right|\right)
$$



## Algorithm II

3. Calculate:

$$
h\left(\boldsymbol{x}_{g}, t\right)=\sum_{i=1}^{N_{\text {PMT }}}\left(t-t_{i}^{\text {tof_corr }}\right) \cdot \exp \left[-\frac{\left(t_{i}^{\text {tof_corr }}-t\right)^{2}}{2 \sigma_{\text {tts }}^{2}}\right]
$$



## Algorithm III

4. Calculate: $\left|h\left(\boldsymbol{x}_{g}, t\right)\right|^{2}$

5. The figure of merit is:

$$
f_{\mathrm{FCN}}\left(\boldsymbol{x}_{g}\right)=\int_{-\infty}^{\infty}\left|h\left(\boldsymbol{x}_{g}, t\right)\right|^{2} d t
$$

## Capabilities




- $\mu^{-}$
- 1 GeV
- Origin (0,0,0)
- Direction (-1,0,0)
- Only first hit information used
- $2 \mu^{-}$
- 1 GeV each
- Origin ( $0,0,0$ )
- Enclosed angle $45^{\circ}$
- Multi-hit information used


## Contents

Motivation
Long baseline neutrino oscillations
Performance of LENA without tracking
Backtracking
Algorithm
First results
Wonsak's tracking
Application to LENA events
Application to CNGS- $\mu^{-}$in Borexino
Likelihood-fit
The PDF
Application to simple events
Conclusion

## Basics

## Assumption

- One reference point on/near track known $\left(\boldsymbol{r}_{0}, t_{0}\right)$.

Allowed photon emission points

- $\left(t-t_{0}\right) c=d_{T}+n \cdot d_{P}$
- All allowed photon emission points $\boldsymbol{r}_{e}$ are
- on an ellipsoid ( $n=1$ )
- on a drop-like surface $(n \neq 1)$
- Drop surfaces are smeared due to PMT-resolution and scintillator decay time
- Superimposing all smeared drop surfaces reveals the track
Invented and developed by Björn Wonsak (Universität Hamburg)


## Results for LENA

- $2 \mu^{-}, 750 \mathrm{MeV}$ each, enclosed angle $90^{\circ}$.
- Full hit information used



## Ridge line analysis

- Tracks should show up as a ridge
$\Rightarrow$ Take only bins with more than 17 smaller neighbors



## Application to CNGS- $\mu^{-}$in Borexino



## Application to CNGS- $\mu^{-}$in Borexino

## Contents

Motivation
Long baseline neutrino oscillations
Performance of LENA without tracking
Backtracking
Algorithm
First results
Wonsak's tracking
Application to LENA events
Application to CNGS- $\mu^{-}$in Borexino
Likelihood-fit
The PDF
Application to simple events
Conclusion

## Basics

## Assumption

- There is at least one model for the event (1 track, 2 tracks, particle types ...)

Likelihood fitting

- Determine free parameters of model: $\boldsymbol{x}$.
- Obtain a set of seed parameters for the fit ( $\rightarrow$ Input from other tracking mechanisms)
- Calculate PDF for photon arrival times and charge: $P(\boldsymbol{t}, \boldsymbol{q} \mid \boldsymbol{x})$
- Maximize likelihood $L(\boldsymbol{x} \mid \boldsymbol{t}, \boldsymbol{q})=P(\boldsymbol{t}, \boldsymbol{q} \mid \boldsymbol{x})$ w.r.t. $\boldsymbol{x} \Rightarrow \hat{\boldsymbol{x}}$.
- Distinguish different models by taking the model with the highest $L(\hat{\boldsymbol{x}} \mid \boldsymbol{t})$.


## Considerations for calculating the PDF

- Mean number of photons emitted per unit track length
- Particle/shower propagation in time
- Time resolution of the PMTs
- Finite dimensions of the PMTs
- Winston Cone acceptance function
- Decay time distribution of the scintillator
- Absorption/scattering of photons in the scintillator (changes both the number of detected photons as well as their arrival times, have to use MC input)




## Comparison of calculated time PDF with Geant4




- Good agreement for total spectrum as well as for first-hit spectrum
- Agreement gets worse for lower energies as tracks deviate from straight line.


## 500 MeV muons

Sample

- Simulation: $1000 \mu^{-}$from $(0,0,0)$ along negative $x$-axis
- Require an identified muon decay with decay time $>500 \mathrm{~ns}$

Results







## Multi track fitting

$2 \mu^{-}$with 500 MeV each, enclosed angle $45^{\circ}$

- Currently uses MC-truth as seed.
- Vertex resolution (in 1 direction) $\approx 4 \mathrm{~cm}$
- Energy resolution of each track $\approx 4 \%$
- Charge no longer strongly constrains energy
- Energy of track determined from its length



## Conclusion

## Current status

- Three approaches to reconstruct spatially extended events in LENA
- All algorithms work on muon events
- Two algorithms are currently tested on muons in Borexino.


## Outlook

- Connect the algorithms: Use the result of one algorithm as input to the next algorithm.
- Extend the algorithms to be applied to full neutrino events
- Use the gathered information to improve the LENA beam performance



## LENA Migration matrices

## Simulation

- Neutrino events created with GENIE
- Detector simulation using GEANT4
- Simplified set-up with $\sim 10000$ PMTs
- Analysis uses only total charge and position of barycenter




## Resolution Start Point Wonsack's tracking



Distributions for distance of reference point to the track

- Mean distance of constructed reference point to the track: ~ 60 cm


## Application CNGS- $\mu^{-}$in Borexino

Complications

- No flash ADCs $\Rightarrow$ Only first hit information
- PMTs/Electronics in saturation $\Rightarrow$ No charge information
$\Rightarrow$ No charge barycenter available
$\Rightarrow$ Getting reference point challenging


## Strategy

- Construct reference point

1. For each spatial point, calculate required time correction
2. Histogram time corrections
3. Take point with overall highest bin

- Construct last point on reconstructed track
- Use this point to track backwards in time
- Use the first point on the backwards track as reference point for final fit


## Resolution w.r.t. Borexino tracking



Note: Resolution of Borexino tracking $\approx 0.09$ rad

## Single muon tracking

- Divide event in time snapshots and propagate the PMT pulses backwards to this time.
- Follow maximum







# Beam performance input parameters 

Sebastian Lorenz, Universität Hamburg

## General setup

- GLoBES simulation
- 50 kt LAB LSc detector
- X-secs simulated with GENIE for carbon and hydrogen
- 5\% systematic error for signal and background
- Migration matrices for $\nu_{e} C C, \nu_{\mu} C C$ and NC events
- $\nu_{\tau} \mathrm{CC}$ events not included
$\rightarrow$ migration matrices in preparation
- $90 \%$ efficiency for $\nu_{\mu}$ CC + efficiency decreases linearly between 3 and 7 GeV


## Oscillation parameters and priors

| Parameter | Value | Error |
| :---: | :---: | :---: |
| $\theta_{12}$ | $33.8^{\circ}$ | $5 \%$ |
| $\theta_{13}$ | $9.1^{\circ}$ | $7 \%$ |
| $\theta_{23}$ | $45^{\circ}$ | $10 \%$ |
| $\Delta m^{\circ}{ }_{12}\left[\mathrm{eV}^{\circ}\right]$ | $7.5 \times 10^{-5}$ | $3 \%$ |
| $\Delta \mathrm{~m}_{23}[\mathrm{eV} 2]$ | $2.5 \times 10^{-3}$ | $5 \%$ |
| $\delta_{\text {cP }}$ | changed | free |
| $\rho_{\text {mass }}$ | PREM | $5 \%$ |

## Neutrino beams

CERN to Pyhäsalmi (CN2PY)

- Based on LBNO EOI
- 400 GeV protons
- 1x1020 POT / yr (shared mode)
- Baseline: 2288 km
- Used „official" 50 GeV spectrum and rescaled



## Protvino to Pyhäsalmi (P2PY)

- 70 GeV protons
- $4.01 \times 10^{20} \mathrm{POT} / \mathrm{yr}$
$\rightarrow 450 \mathrm{~kW}$
- Baseline: 1160 km



## Neutrino mass hierarchy I

Simulated: normal mass hierarchy
Fitted: inverted mass hierarchy

## CERN to Pyhäsalmi (CN2PY)



## Protvino to Pyhäsalmi (P2PY)



Different scales!

## Neutrino mass hierarchy II

Simulated: normal mass hierarchy
Fitted: inverted mass hierarchy

## CERN to Pyhäsalmi (CN2PY)



## Protvino to Pyhäsalmi (P2PY)



Different scales!

## CP phase I

Simulated: different true values of $\delta_{\mathrm{CP}}$
Fitted:
$\delta_{\mathrm{CP}}=0^{\circ}$ and $\delta_{\mathrm{CP}}=180^{\circ}$

## CERN to Pyhäsalmi (CN2PY)



## Protvino to Pyhäsalmi (P2PY)



## CP phase II

Simulated: different true values of $\delta_{\mathrm{CP}}$
Fitted:
$\delta_{\mathrm{CP}}=0^{\circ}$ and $\delta_{\mathrm{CP}}=180^{\circ}$

## CERN to Pyhäsalmi (CN2PY)



## Protvino to Pyhäsalmi (P2PY)



