

# The Scale-Invariant NMSSM and the 126 GeV Higgs Boson

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in collaboration with

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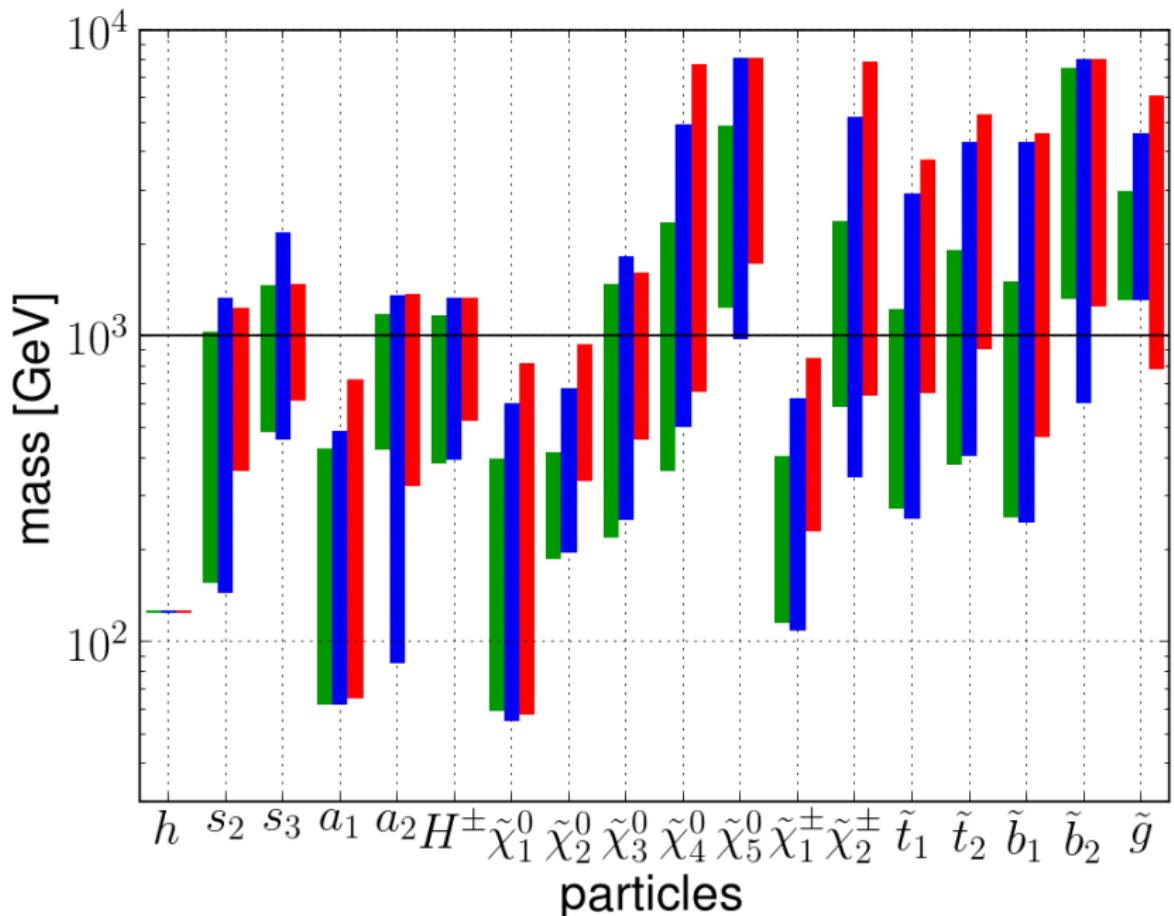


THE UNIVERSITY OF  
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**CoEPP**

ARC Centre of Excellence for  
Particle Physics at the Terascale



# Outline

- ① Introduction
- ② Naturalness in the Scale-Invariant NMSSM
- ③ Phenomenology
- ④ Conclusions

# Outline

1 Introduction

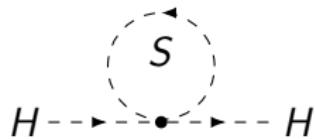
2 Naturalness in the Scale-Invariant NMSSM

3 Phenomenology

4 Conclusions

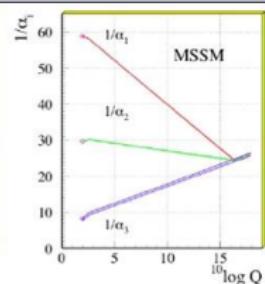
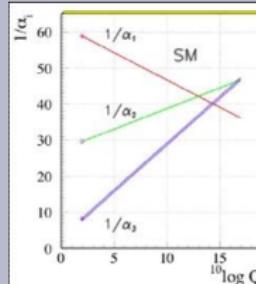
# Low-energy Supersymmetry

## Solution to hierarchy problem

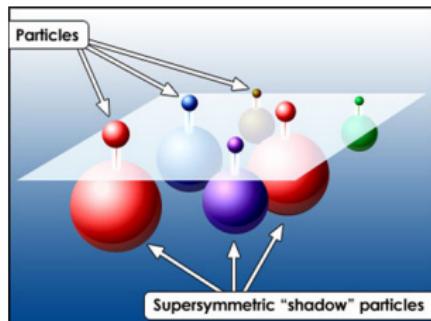
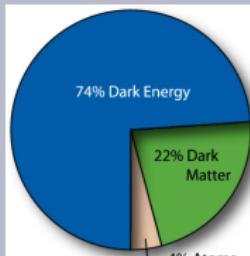


$$\Delta m_H^2 \propto \frac{\lambda_S}{16\pi^2} m_S^2$$

## Gauge coupling unification



## DM candidate



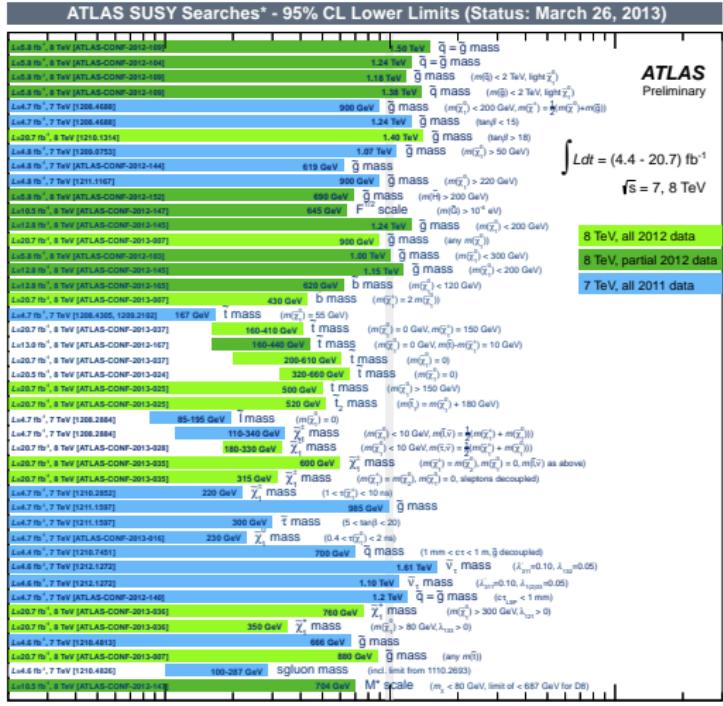
## Higgs mass prediction

At tree level

$$m_{h^0} < m_Z |\cos 2\beta|$$

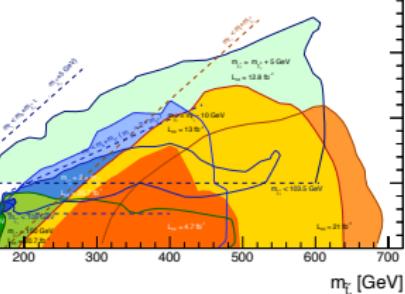
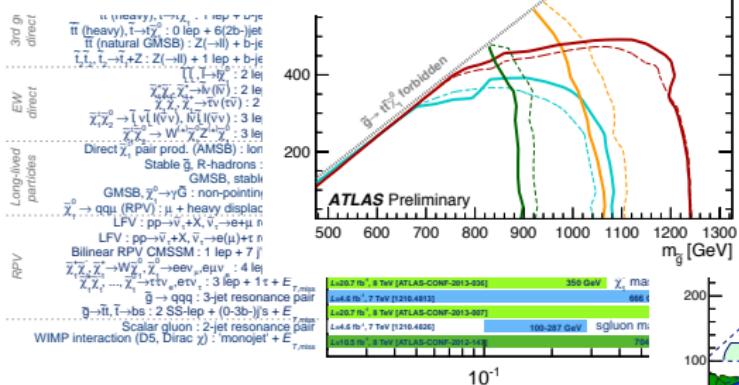
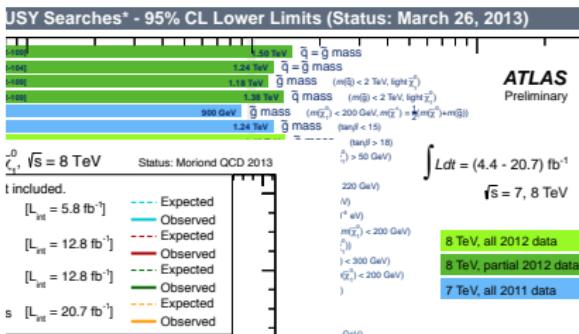
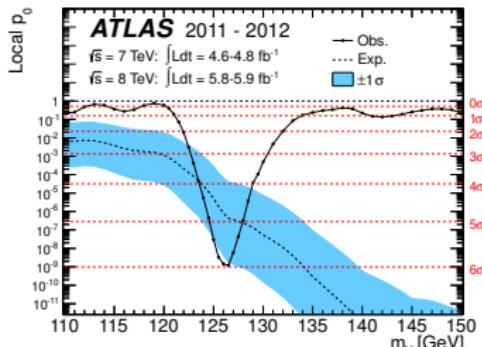
# After 8TeV LHC: (C)MSSM and Co.

Inclusive searches  
3rd gen. squarks  
3rd gen. gluino  
mediated  
EW direct production  
Long-lived particles  
RPV



\*Only a selection of the available mass limits on new states or phenomena shown.  
All limits quoted are observed minus 1 $\sigma$  theoretical signal cross section uncertainty.

# After 8TeV LHC: (C)MSSM and Co.



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# Fine-tuning in the MSSM

- rely on loop corrections to Higgs mass from (s)top sector

$$m_h^2 = m_Z^2 \cos^2 2\beta \left( 1 - \frac{3}{8\pi^2} y_t^2 t \right) + \frac{3}{4\pi^2} y_t^2 \left[ \frac{1}{2} X_t + t + \dots \right]$$

with  $t = \ln \frac{m_{\text{soft}}^2}{M_t^2}$ ,  $X_t = \frac{2(A_t - \mu \cot \beta)^2}{m_{\text{soft}}^2} \left( 1 - \frac{(A_t - \mu \cot \beta)^2}{12m_{\text{soft}}^2} \right)$ ,  $m_{\text{soft}} = \sqrt{m_{t_1} m_{\bar{t}_2}}$

- Large contribution to  $m_{H_u}^2$  from RG evolution in the supersymmetric theory (finite corrections are small compared to RG effects for  $\Lambda \gg m_{\text{soft}}$ )

$$m_{H_u}^2(m_{\text{soft}}) = m_{H_u}^2(\Lambda) - \frac{3y_t^2}{8\pi^2} \left[ m_{Q_3}^2(\Lambda) + m_{u_3}^2(\Lambda) + A_t^2(\Lambda) \right] \ln \left[ \frac{\Lambda}{m_{\text{soft}}} \right] + \dots$$

## Tree-level minimization conditions

$$m_Z^2 = \frac{m_{H_u}^2 - m_{H_d}^2}{\cos 2\beta} - m_{H_u}^2 - m_{H_d}^2 - 2\mu^2$$

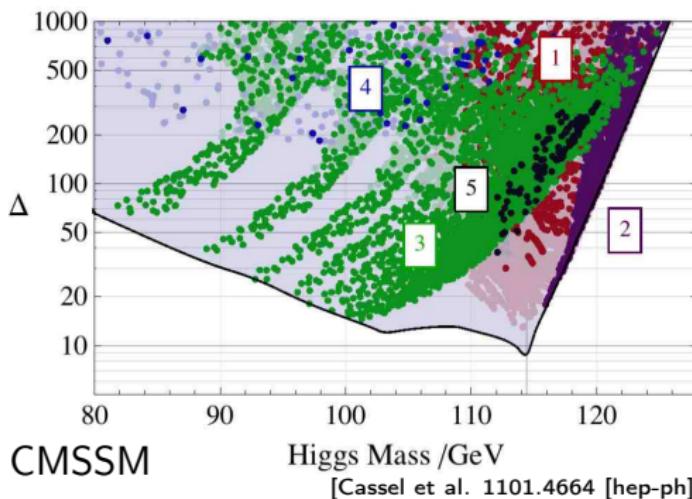
$$0 = \frac{2b}{\sin 2\beta} - m_{H_u}^2 - m_{H_d}^2 - 2\mu^2$$

# Fine-tuning in the MSSM 2

## Finetuning measure [Barbieri, Giudice (1988)]

$$\Sigma^\nu = \Delta \equiv \max_{\xi} \left| \frac{\partial \ln \nu^2}{\partial \ln \xi(\Lambda)} \right|$$

with  $\xi = (m_{H_u}^2, m_{H_d}^2, m_{Q_3}^2, m_{u_3}^2, m_{d_3}^2, A_t, A_b, M_1, M_2, M_3)$



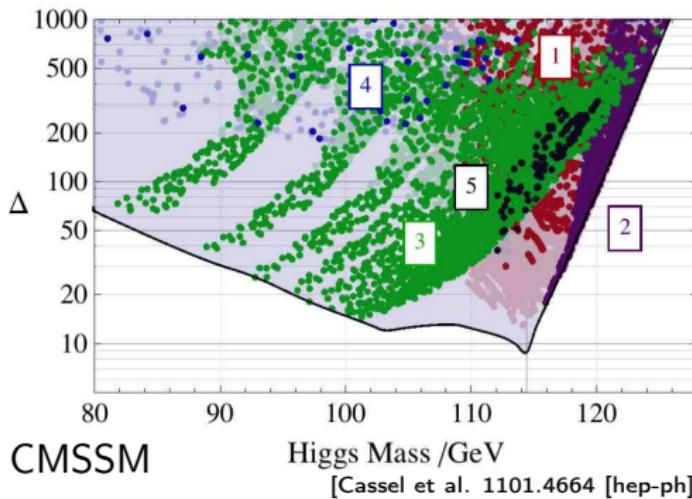
- after LEP:  $\Sigma^\nu \gtrsim 10$ ,  
i.e. 10% fine tuning
- after 8TeV LHC:  $\Sigma^\nu \gtrsim 100$ ,  
i.e. worse than 1% fine tuning

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- after 8TeV LHC:  $\Sigma^\nu \gtrsim 100$ ,  
i.e. worse than 1% fine tuning



NEVER, NEVER,  
NEVER GIVE UP!





### New tree-level contributions to the Higgs mass

- additional  $D$ -term contributions
- additional  $F$ -term contributions

→ e.g. in the NMSSM: additional gauge singlet  $S$

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# NMSSM

## Superpotential

$$W_{NMSSM} = \lambda S H_d H_u + \frac{\kappa}{3} S^3$$

## Soft breaking terms

$$V_{soft} = m_{H_d}^2 |H_d|^2 + m_{H_u}^2 |H_u|^2 + m_s^2 |S|^2 - (a_\lambda S H_d H_u + \frac{a_\kappa}{3} S^3 + h.c.)$$

## Bound on lightest Higgs mass $m_h$

$$m_h^2 \leq m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta \rightarrow \begin{cases} m_Z^2 \cos^2 2\beta & \text{large } \tan \beta \\ \lambda^2 v^2 \sin^2 2\beta & \text{small } \tan \beta \end{cases}$$

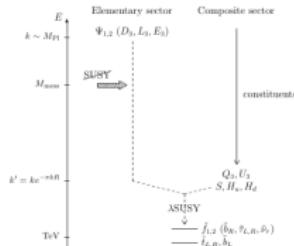
- No gain for large  $\tan \beta$  compared to MSSM → small  $\tan \beta$   
→ relatively large  $\lambda \gtrsim 0.7$  or additional stop loop contribution  
→ Landau pole below GUT scale

# Large $\lambda$ : Landau pole below GUT scale

- Additional vector-like  $5 + \bar{5}$  SU(5) multiplets delay Landau pole  
[Mazip, Munoz-Tapia, Pomarol (1998); Barbieri et al. (2007); Barbieri et al. (2013)]
- Strong-coupling effects might even correct for 3% discrepancy in unification [Hardy, March-Russell, Unwin (1207.1435)]
- New physics at  $\Lambda \sim 10$  TeV ( $\lambda$ SUSY [Barbieri, Hall, Nomura, Rychkov (2006)])  
 → e.g. Fat Higgs [Harnik, Kribs, Larson, Murayama (2003)]  
 [Chang, Kilic, Mahbubani (2004); Delgado, Tait; Birkedal, Chacko, Nomura (2004)]  
 → e.g. within extra dimensional models



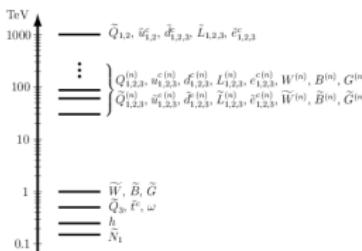
[Gherghetta, v.Harling, Setzer (2011)]



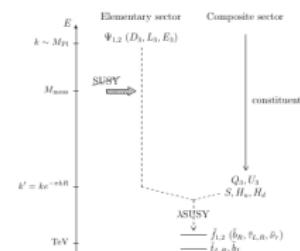
[Larsen, Nomura, Roberts (2012)]

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 → e.g. within extra dimensional models



[Gherghetta, v.Harling, Setzer (2011)]



[Larsen, Nomura, Roberts (2012)]

- Our assumptions: decouple first two generations of sparticles as well as third generation of sleptons

# EW symmetry breaking

## MSSM minimization conditions

$$m_Z^2 = \frac{m_{H_u}^2 - m_{H_d}^2}{\cos 2\beta} - m_{H_u}^2 - m_{H_d}^2 - 2\mu^2$$

$$0 = m_A^2 - m_{H_u}^2 - m_{H_d}^2 - 2\mu^2$$

We obtain the following minimization conditions in the NMSSM [ $\mu \equiv \lambda v_S$  and  $m_A^2 \equiv \frac{2(a_\lambda + \lambda \kappa v_S)v_S}{\sin 2\beta}$ ]

$$m_Z^2 = \frac{m_{H_u}^2 - m_{H_d}^2}{\cos 2\beta} - m_{H_u}^2 - m_{H_d}^2 - 2\lambda^2 v_S^2$$

$$\lambda^2 v^2 = \frac{2(a_\lambda + \lambda \kappa v_S)v_S}{\sin 2\beta} - m_{H_u}^2 - m_{H_d}^2 - 2\lambda^2 v_S^2$$

$$m_S^2 = \lambda \kappa v^2 \sin 2\beta - 2\kappa^2 v_S^2 - \lambda^2 v^2 - \frac{a_\lambda v^2}{2v_S} \sin 2\beta - a_\kappa v_S$$

# Beyond tree-level

$$\mathcal{S}_{\text{eff}} = \int d^4x \left\{ \sum_{n=0}^{\infty} Z_i^n \partial_\mu \phi_i^\dagger \partial^\mu \phi_i - \sum_{n=0}^{\infty} V_n \right\}$$

The NLO corrections to the potential are given by the Coleman-Weinberg potential

$$V_1 = \frac{1}{64\pi^2} \text{STr } M^4 \left[ \ln \left( \frac{M^2}{\mu_r^2} \right) - \frac{3}{2} \right]$$

Defining

$$\hat{m}_{H_u}^2 \equiv m_{H_u}^2 + \frac{d}{dv_u^2} V_1, \quad \hat{m}_{H_d}^2 \equiv m_{H_d}^2 + \frac{d}{dv_d^2} V_1, \quad \hat{m}_S^2 \equiv m_S^2 + \frac{d}{dv_S^2} V_1$$

the minimization conditions at NLO can be written as

$$m_Z^2 = \frac{\hat{m}_{H_u}^2 - \hat{m}_{H_d}^2}{\cos 2\beta} - \hat{m}_{H_u}^2 - \hat{m}_{H_d}^2 - 2\lambda^2 v_S^2.$$

$$\lambda^2 v^2 = 2 \frac{(a_\lambda v_S + \lambda \kappa v_S^2)}{\sin 2\beta} - \hat{m}_{H_u}^2 - \hat{m}_{H_d}^2 - 2\lambda^2 v_S^2$$

$$\hat{m}_S^2 = \lambda \kappa v^2 \sin 2\beta - 2\kappa^2 v_S^2 - \lambda^2 v^2 - \frac{a_\lambda v^2}{2v_S} \sin 2\beta - a_\kappa v_S$$

# Fine-tuning in the NMSSM

## Finetuning measure [Barbieri, Giudice (1988)]

$$\Sigma^\nu = \max_i \left| \frac{\partial \ln \nu^2}{\partial \ln \xi_i(\Lambda)} \right|$$

with  $\xi_i = (m_{H_u}^2, m_{H_d}^2, m_S^2, \lambda, \kappa, a_\lambda, a_\kappa, m_{Q_3}^2, m_{u_3}^2, m_{d_3}^2, A_t, A_b, M_1, M_2, M_3)$

Finite one-loop corrections alleviate tuning by 10-20% compared to tree-level (cf. [Cassel, Ghilencia, Ross (2010)])

$$\Sigma^\nu = \max_i \left| \sum_j \frac{d \ln \nu^2}{d \ln \xi_j(m_{\text{soft}})} \frac{d \ln \xi_j(m_{\text{soft}})}{d \ln \xi_i(\Lambda)} \right|$$

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For example in the case of  $m_{Q_3}^2$  neglecting finite one-loop correction:

$$\left| \frac{d \ln \nu^2}{d \ln m_{Q_3}^2(\Lambda)} \right| \approx \left| \frac{3y_t^2}{8\pi^2} \frac{m_{Q_3}^2(\Lambda)}{\nu^2} \ln \left[ \frac{\Lambda}{m_{\text{soft}}} \right] \times \frac{d\nu^2}{dm_{H_u}^2(m_{\text{soft}})} \right| \lesssim \Sigma^\nu$$

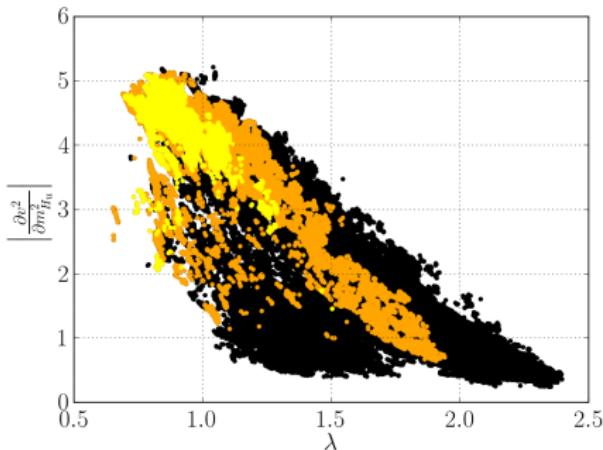
### MSSM

$$\frac{d\nu^2}{dm_{H_u}^2(m_{\text{soft}})} = -2 \frac{\nu^2}{m_Z^2} + \mathcal{O}\left(\frac{1}{\tan \beta}\right)$$

### NMSSM

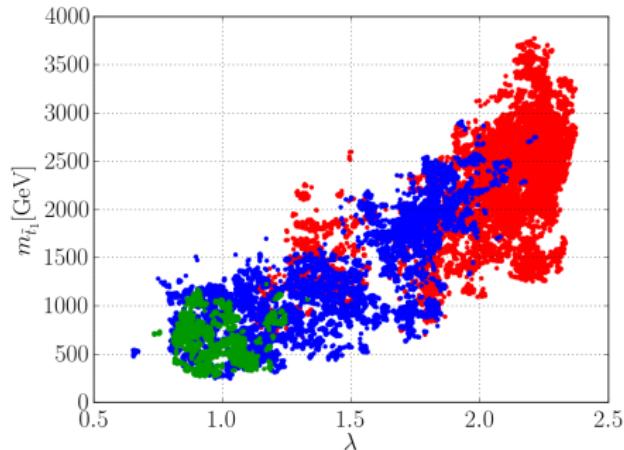
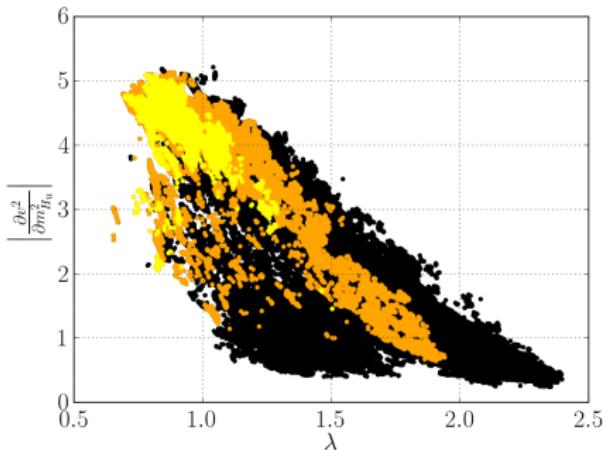
$$\frac{d\nu^2}{dm_{H_u}^2(m_{\text{soft}})} = \frac{\kappa}{\lambda^3} \cot 2\beta + \mathcal{O}\left(\frac{1}{\lambda^4}\right)$$

## Large $\lambda$ helps



- $\frac{dv^2}{dm_{H_u}^2(m_{\text{soft}})} \sim \frac{\kappa}{\lambda^3} \cot 2\beta$
- large  $\lambda$  reduces finetuning
- black  $\Lambda = 20$  TeV, orange  $\Lambda = 100$  TeV, yellow  $\Lambda = 1000$  TeV

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- black  $\Lambda = 20$  TeV, orange  $\Lambda = 100$  TeV, yellow  $\Lambda = 1000$  TeV
- right:  $\Sigma^\nu < 20$ : better than 5% VEV tuning
- large  $\lambda$  allows larger stop masses

## Large $\lambda$ hurts

$$m_h^2 \simeq m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta + \delta m_{h,\text{stop}}^2 + \delta m_{h,S}^2$$

- For large  $\lambda$ :  $\lambda v \sin 2\beta > m_h$ : cancellation required
  - generally  $\delta m_{h,\text{stop}}^2 > 0$
  - partly cancelled by mixing with singlet (cf. [Agashe, Cui, Franceschini (2012)])
  - In our scan, it amounts to at most 40% of  $m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta$
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## Higgs mass tuning

$$\Sigma^h \equiv \max_{\xi_i} \left| \frac{d \ln m_h^2}{d \ln \xi_i} \right| \approx \max_{\xi_i} \left| \frac{\xi_i}{m_h^2} \frac{dm_{h,\text{tree}}^2}{d\xi_i} \right|$$

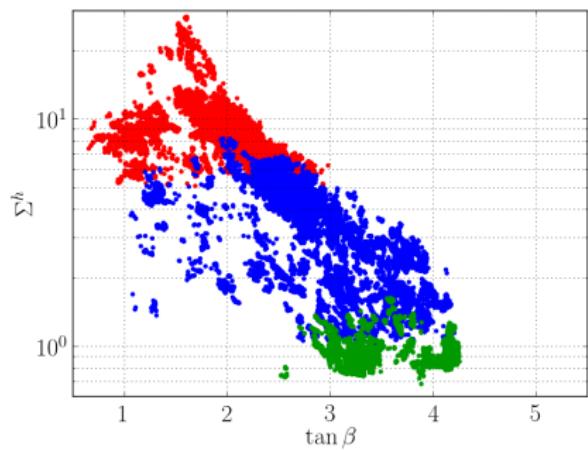
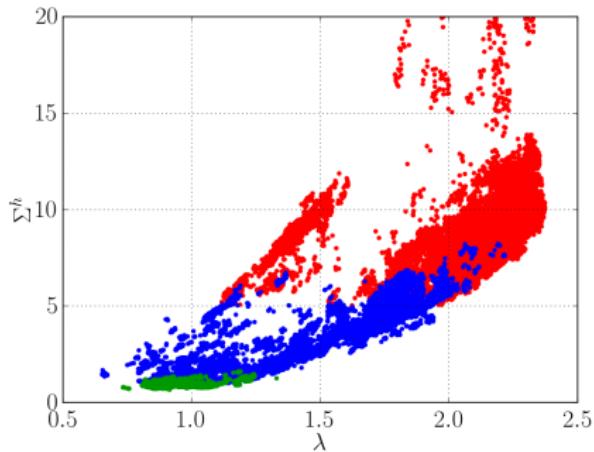
with  $\xi_i = (\lambda, \kappa, a_\lambda, a_\kappa, m_{Q_3}^2, m_{u_3}^2, m_{d_3}^2, A_t, A_b, M_1, M_2, M_3)$  and fix  $v_{u,d,S}$

# Large $\lambda$ hurts 2

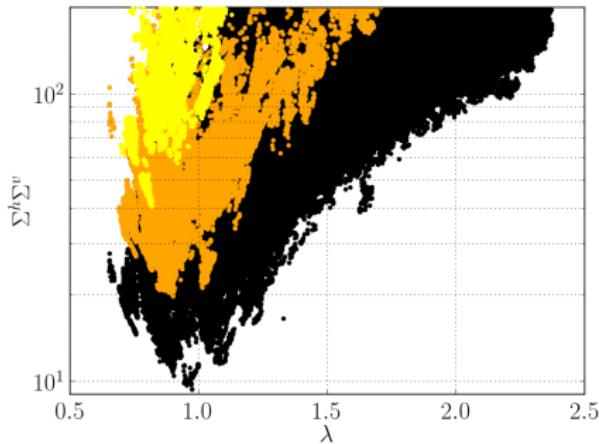
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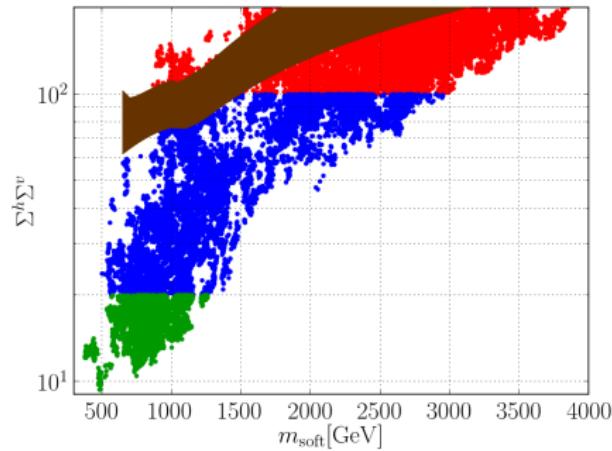
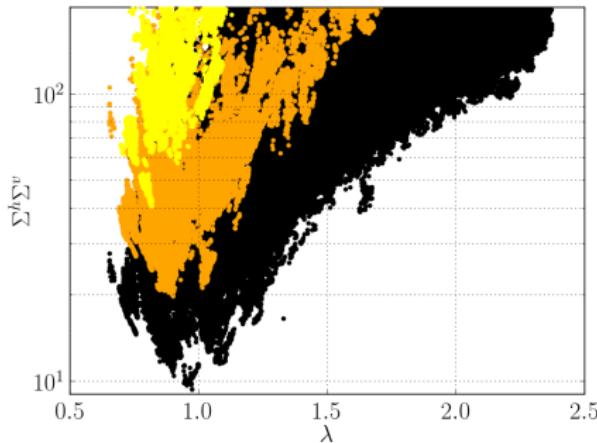


## Combined fine-tuning



- In a statistical sense, if two quantities are not correlated, the probability involving both is  $P(A \cap B) = P(A) * P(B)$ .
- black  $\Lambda = 20$  TeV, orange  $\Lambda = 100$  TeV, yellow  $\Lambda = 1000$  TeV

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- black  $\Lambda = 20$  TeV, orange  $\Lambda = 100$  TeV, yellow  $\Lambda = 1000$  TeV
- right:  $\Sigma^\nu < 20$ : better than 5% VEV tuning
- green  $\Sigma^h \Sigma^\nu < 20$ , blue  $20 < \Sigma^h \Sigma^\nu < 100$ , red  $\Sigma^h \Sigma^\nu > 100$
- brown band: MSSM with  $\mu = 200$  GeV,  $\tan \beta = 20$  and  $m_A = 1$  TeV  
[generated using FeynHiggs]

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# Numerical Scan

- Markov Chain Monte Carlo
- We used a modified version of NMHDECAY, which is part of NMSSMTools 3.2.1, as well as MicrOMEGAs 2.4.5 for DM
- We scanned linearly in 14-dimensional parameter space

$\tan \beta$	$\tan \beta > 0.08$	$m_{Q_3}$	$\Delta_{\tilde{g}} m_{Q_3} < m_{Q_3} < 5 \text{ TeV}$	$M_1$	$0 < M_1 < 8 \text{ TeV}$
$\mu$	$ \mu  < 1 \text{ TeV}$	$m_{u_3}$	$\Delta_{\tilde{g}} m_{u_3} < m_{u_3} < 5 \text{ TeV}$	$M_2$	$0 < M_2 < 8 \text{ TeV}$
$\lambda$	$0 < \lambda < 3$	$m_{d_3}$	$0 < m_{d_3} < 8 \text{ TeV}$	$M_3$	$0.5 \text{ TeV} < M_3 < 8 \text{ TeV}$
$\kappa$	$ \kappa  < 2.75$	$A_t$	$ \Delta_{\tilde{g}} A_t  <  A_t  < 5 \text{ TeV}$	$v$	$174 \text{ GeV}$
$A_\lambda$	$ A_\lambda  < 2 \text{ TeV}$	$A_b$	$ A_b  < 8 \text{ TeV}$	$\Lambda$	$20, 100, 1000 \text{ TeV}$
$A_\kappa$	$ A_\kappa  < 1 \text{ TeV}$				

⇒ We did not sample the full parameter space. Therefore, **there is no statistical interpretation of the scatter plots.**

- Likelihood function: product of Gaussians for the Higgs mass centered at 126 GeV and for the VEV finetuning centered at 0.
- We impose the hard cuts summarised in the table as well as

$$|\xi(\Lambda) - \xi(m_{\text{soft}})| < |\xi(\Lambda)| \quad \text{for } \xi = m_{Q_3}^2, m_{u_3}^2, m_{d_3}^2, A_t, A_b$$

similar to "gluino sucks the stop mass up" [Arvanitaki, Craig, Dimopoulos, Villadoro(2012)]

- We conservatively impose the collider, EW precision and flavour constraints.<sup>20</sup>

## Electroweak Precision Tests (EWPT)

- impose EWPT at  $2\sigma$  with  $m_{h,\text{ref}} = 117 \text{ GeV}$  [PDG(2012)].

$S_0 = -0.04 \pm 0.09$ ,  $T_0 = 0.07 \pm 0.08$  and correlation of 88% at 95% C.L.

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 $S_0 = -0.04 \pm 0.09$ ,  $T_0 = 0.07 \pm 0.08$  and correlation of 88% at 95% C.L.
- Consistent with previous analyses[Barbieri et. al.(2006); Franceschini, Gori (2010)], we find
- Higgs and singlet scalars do not contribute much to S,T
- (3rd gen) squarks:  $\tan \beta$  can not be too large unless cancellation  
 $\tan \beta \neq 1$  breaks custodial SU(2)  $\rightarrow$  contribution to  $T$

# Electroweak Precision Tests (EWPT)

- impose EWPT at  $2\sigma$  with  $m_{h,\text{ref}} = 117 \text{ GeV}$  [PDG(2012)].  
 $S_0 = -0.04 \pm 0.09$ ,  $T_0 = 0.07 \pm 0.08$  and correlation of 88% at 95% C.L.
- Consistent with previous analyses[Barbieri et. al.(2006); Franceschini, Gori (2010)], we find
- Higgs and singlet scalars do not contribute much to  $S, T$
- (3rd gen) squarks:  $\tan \beta$  can not be too large unless cancellation  
 $\tan \beta \neq 1$  breaks custodial  $SU(2) \rightarrow$  contribution to  $T$
- Neutralino/chargino sector imposes strong constraint on  $\lambda$  as well as  $\tan \beta$

$$M_{\psi^0} = \begin{pmatrix} M_1 & 0 & -\cos \beta \sin \theta_W m_Z & \sin \beta \sin \theta_W m_Z & 0 \\ . & M_2 & \cos \beta \cos \theta_W m_Z & -\sin \beta \cos \theta_W m_Z & 0 \\ . & . & 0 & -\mu & -\lambda v \sin \beta \\ . & . & -\mu & 0 & -\lambda v \cos \beta \\ . & . & . & . & -2 \frac{\kappa}{\lambda} \mu \end{pmatrix}$$

$$M_{\psi^\pm} = \begin{pmatrix} 0 & X^T \\ X & 0 \end{pmatrix} \quad \text{with} \quad X = \begin{pmatrix} M_2 & \sqrt{2} m_W \sin \beta \\ \sqrt{2} m_W \cos \beta & \mu \end{pmatrix}$$

in gauge-basis  $\psi^0 = (\tilde{B}, \tilde{W}^3, \tilde{H}_d^0, \tilde{H}_u^0, \tilde{S})$  and  $\psi^\pm = (\tilde{W}^+, \tilde{H}_u^+, \tilde{W}^-, \tilde{H}_d^-)$

- We find  $\tan \beta \lesssim 5$  depending on  $\lambda$

## SUSY searches

Assuming a neutralino LSP, we **conservatively** exclude the following regions:

- gluino search:  $\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$  [ATLAS-CONF-2012-151]

$$m_{\tilde{g}} < 1310 \text{ GeV} \quad \text{if } m_{\tilde{\chi}_1^0} < 650 \text{ GeV}$$

- sbottom search [PDG (2012); CMS-PAS-SUS-12-028; ATLAS-CONF-2012-106]

$$m_{\tilde{b}} < 89 \text{ GeV}$$

$$150 \text{ GeV} < m_{\tilde{b}} < 650 \text{ GeV} \quad \text{if } m_{\tilde{\chi}_1^0} < 230 \text{ GeV}$$

- stop search [PDG(2012); ATLAS (1208.1447, 1208.2590)]

$$m_{\tilde{t}} < 95.7 \text{ GeV}$$

$$220 \text{ GeV} < m_{\tilde{t}} < 500 \text{ GeV} \quad \text{if } m_{\tilde{\chi}_1^0} < 160 \text{ GeV}$$

- chargino search [PDG (2012)]

$$m_{\tilde{\chi}^\pm} < 94 \text{ GeV}$$

The neutralino/chargino searches by ATLAS/CMS did not lead to further constraints.

# Higgs searches

$$R_X \equiv \frac{\sigma(h) \times BR(h \rightarrow X)}{\sigma(h_{\text{SM}}) \times BR(h_{\text{SM}} \rightarrow X)}$$

- Higgs resonance at 126 GeV

[ATLAS (1207.7214), ATLAS-CONF-2012-162; ATLAS-CONF-2012-170; CMS (1207.7235), CMS-HIG-12-045]

$$\begin{aligned} 0.81 < R_{ZZ} < 1.32, \quad & 0.74 < R_{WW} < 1.40, \\ 0 < R_{b\bar{b}} < 1.10, \quad & 0.27 < R_{\tau\tau} < 1.15, \end{aligned}$$

- Heavy Higgs searches [CMS-PAS-Higgs-11-024, CMS-PAS-Higgs-11-041]

$$\begin{aligned} \frac{\sigma(s_i) \times BR(s_i \rightarrow ZZ)}{\sigma(h_{\text{SM}}) \times BR(h_{\text{SM}} \rightarrow ZZ)} &< 0.09 \\ \frac{\sigma(s_i) \times BR(s_i \rightarrow WW)}{\sigma(h_{\text{SM}}) \times BR(h_{\text{SM}} \rightarrow WW)} &< 0.2 \end{aligned}$$

- Charged Higgs [PDG(2012)]

$$m_{H^\pm} \gtrsim 79.3 \text{ GeV}$$

# Flavour Constraints

We use the following flavour physics constraints

- $B$ -meson mixing [HFAG (1207.1158)]

$$\Delta M_s = (17.719 \pm 0.086) \text{ ps}^{-1}$$

$$\Delta M_d = (0.507 \pm 0.008) \text{ ps}^{-1}$$

- rare  $B$ -decays [HFAG (1207.1158)]

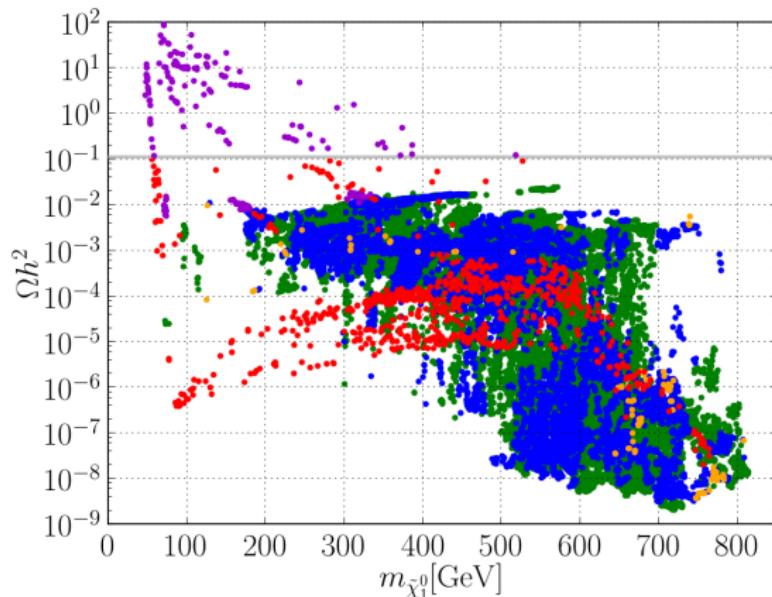
$$\text{Br}(B^+ \rightarrow \tau^+ \nu_\tau) = (1.67 \pm 0.60) \times 10^{-4}$$

$$\text{Br}(B \rightarrow X_s \gamma) = (3.55 \pm 0.48 \pm 0.18) \times 10^{-4}$$

- recently measured rare decay  $B_s^0 \rightarrow \mu^+ \mu^-$  [LHCb (1211.2674)]

$$\text{Br}(B_s^0 \rightarrow \mu^+ \mu^-) = 3.2_{-2.4-0.6}^{+3.0+1.0} \times 10^{-9}$$

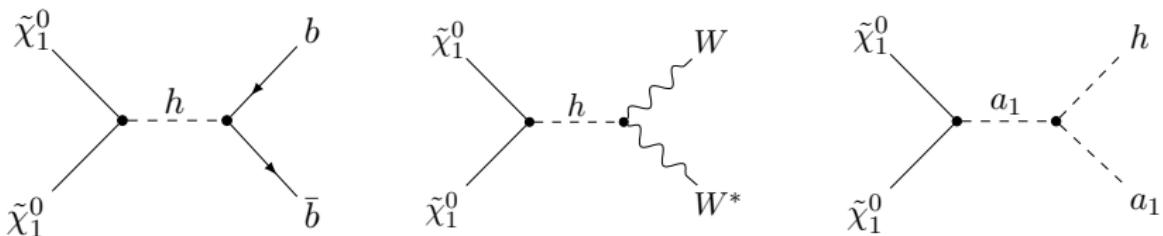
# Dark Matter



- WMAP-7 constraint:  $\Omega_{LSP} h^2 < \Omega_{WMAP-7} h^2 = 0.1120 \pm 0.0056$
- green (singlino), blue (Higgsino), orange (wino), red (bino) DM
- Purple points overclose the Universe or are excluded by XENON100  
⇒ but they are viable if the gravitino is the LSP.

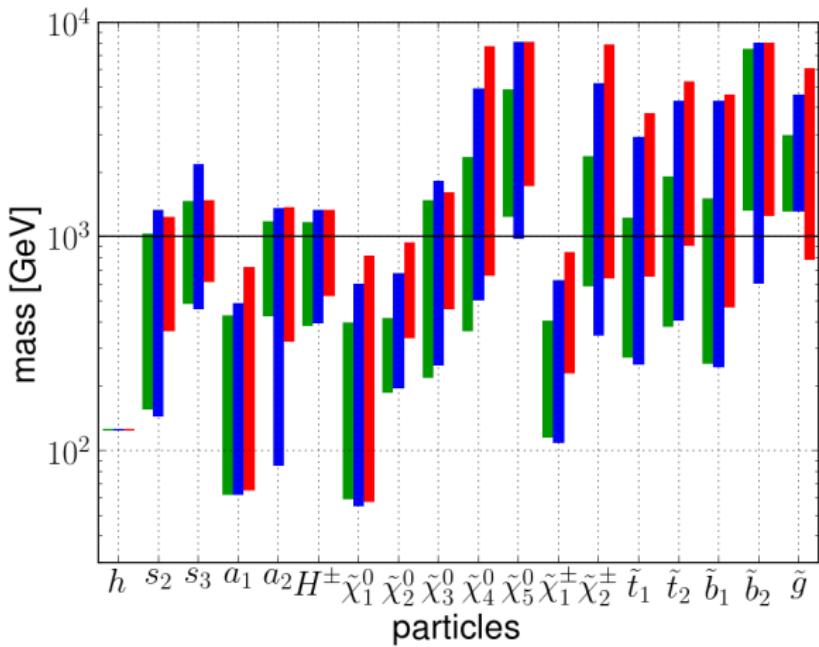
## Dark Matter 2

- A mostly Bino with Higgsino admixture is viable DM candidate
- We find:
  - lightest chargino mostly Higgsino with  $m_{\chi^\pm} \gtrsim 200$  GeV
  - charged Higgs  $m_{H^+} \gtrsim 750$  GeV
- ⇒  $t$ -channel annihilation via chargino and LSP-chargino coannihilation via charged Higgs are suppressed
- Dominant annihilation channels:



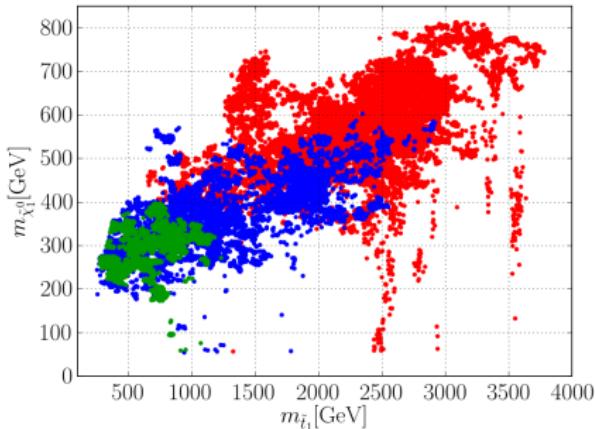
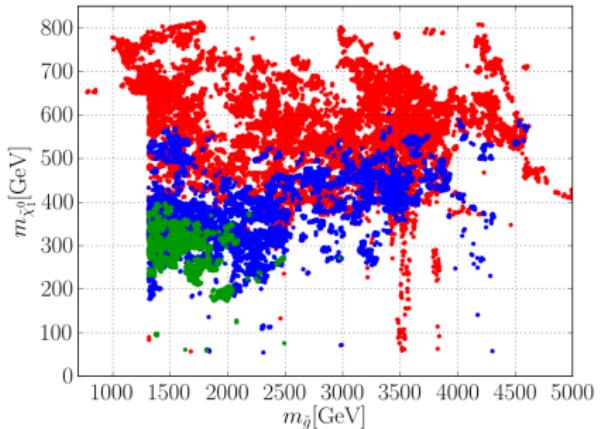
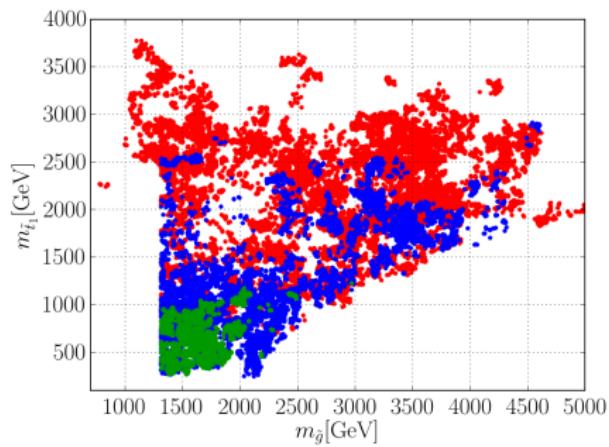
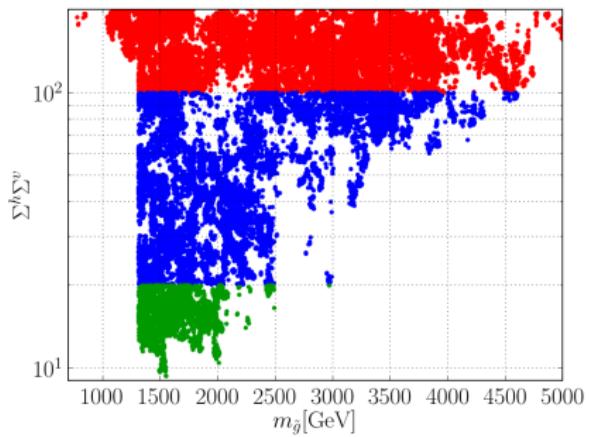
- on resonance for CP even Higgs-mediated annihilation ( $p$ -wave)
- $s$ -wave CP odd Higgs-mediated annihilation
- Why did we not find more points, which fit WMAP-7?  
→ scan linearly in parameters → large values are preferred,  $M_1$  generally large

# Particle spectrum

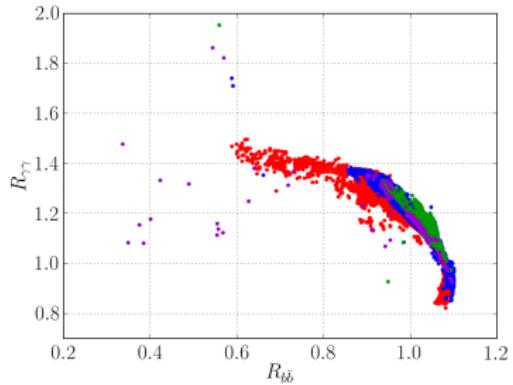
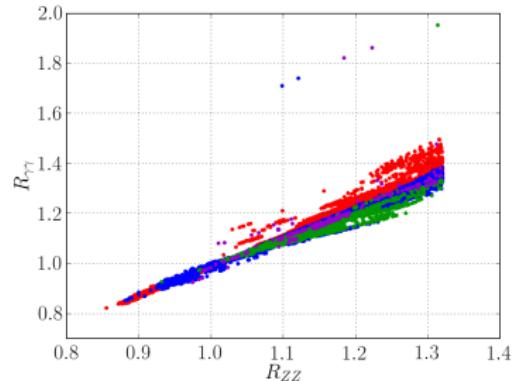


- all colored particles can be above 1 TeV with tuning better than 5%
- but light EW sector required
- due to limits on invisible widths:  $m_{a_1} \gtrsim \frac{m_h}{2}$  and  $m_{\tilde{\chi}_1^0} \gtrsim \frac{m_Z}{2}$  (or  $\frac{m_h}{2}$ )

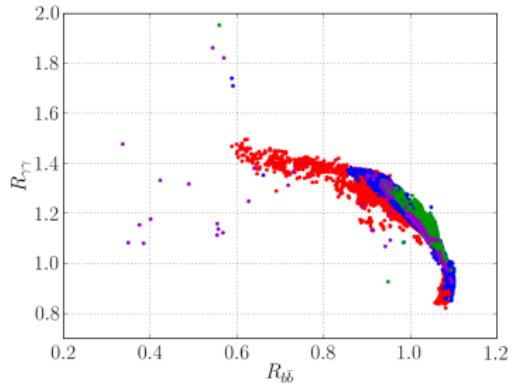
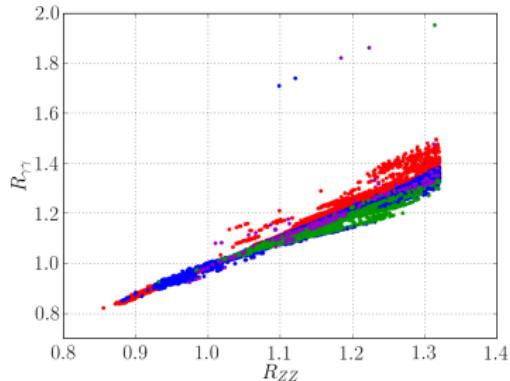
# SUSY searches



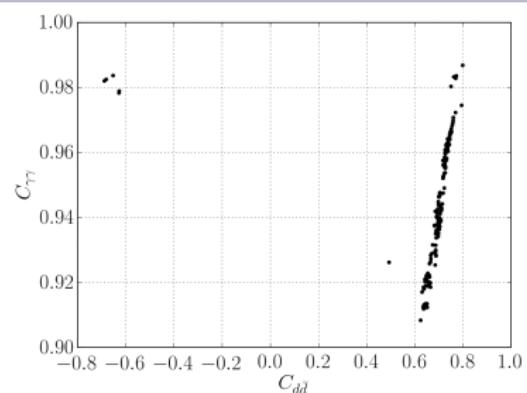
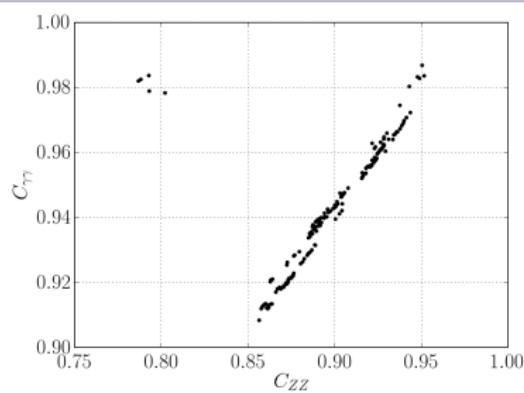
# Higgs diphoton rate



# Higgs diphoton rate



Require  $h \rightarrow \gamma\gamma$  to be within  $1\sigma$ :  $1.40 < R_{\gamma\gamma} < 1.99$



# Outline

1 Introduction

2 Naturalness in the Scale-Invariant NMSSM

3 Phenomenology

4 Conclusions

## Conclusions

- natural SUSY is not excluded (yet)
  - TeV-scale stop masses allowed with 5% tuning for a cutoff scale of 20 TeV
  - a large value of  $\lambda$  reduces tuning in Higgs VEV,  $\Sigma^v$
  - but introduces an additional tuning in the Higgs mass,  $\Sigma^h$
- ⇒ consider combined tuning  $\Sigma^v \Sigma^h$

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## Particle spectrum

We find for the particle spectrum with a combined tuning,  $\Sigma^v \Sigma^h$ , better than 5(1)% with  $\Lambda = 20$  TeV

- lightest stop  $m_{\tilde{t}_1} \lesssim 1.2(2.6)$  TeV
  - gluino masses  $m_{\tilde{g}} \lesssim 3.0(4.6)$  TeV
  - lightest EW sparticles (charginos/neutralinos)  $m_{\tilde{\chi}} \lesssim 400(600)$  GeV
- We have to wait for the 14 TeV LHC

