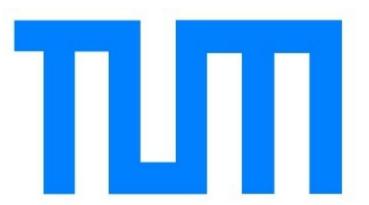
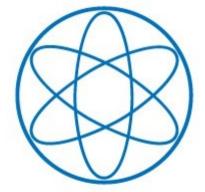
Study of the Internal Bremsstrahlung in the Inert Doublet Model

Camilo A. Garcia Cely Technische Universität München

> Planck Conference 22 May 2013





Based on work in progress done in collaboration with Pr. Alejandro Ibarra

Outline

- Inert doublet model and dark matter
- Indirect searches and spectral features
- Benchmark points and effect of the model parameters on the internal Bremsstrahlung
- H.E.S.S. Upper limits
- Conclusions

The inert doublet model

Let
$$\eta = \begin{pmatrix} H^+ \\ \frac{1}{\sqrt{2}} (H + iA) \end{pmatrix}$$
 be an extra doublet, and Φ the SM doublet
 $\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{\eta}$ $\mathcal{L}_{SM} \supset -\mu_1^2 \Phi^{\dagger} \Phi - \lambda_1 (\Phi^{\dagger} \Phi)^2$
 $\mathcal{L}_{\eta} = (D_{\mu}\eta)^{\dagger} (D^{\mu}\eta) - \mu_2^2 \eta^{\dagger} \eta - \lambda_2 (\eta^{\dagger}\eta)^2 - \lambda_3 (\Phi^{\dagger} \Phi) (\eta^{\dagger} \eta)$ Invariant
 $-\lambda_4 (\Phi^{\dagger}\eta) (\eta^{\dagger} \Phi) - \frac{1}{2} \left(\lambda_5 (\Phi^{\dagger}\eta) (\Phi^{\dagger}\eta) + \text{h.c.} \right)$.
 $(Z_2 \ symmetry)$

Electroweak symmetry breaking

$$\langle \Phi \rangle = \begin{pmatrix} 0 \\ \frac{v}{\sqrt{2}} \end{pmatrix}$$
, $\langle \eta \rangle = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$ \checkmark Z_2 is not spontaneously broken

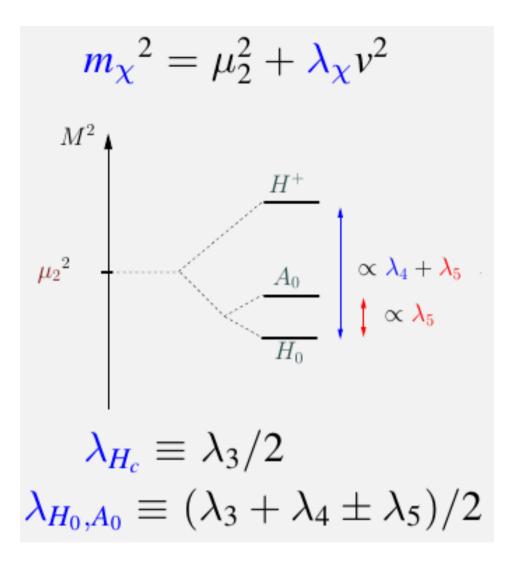
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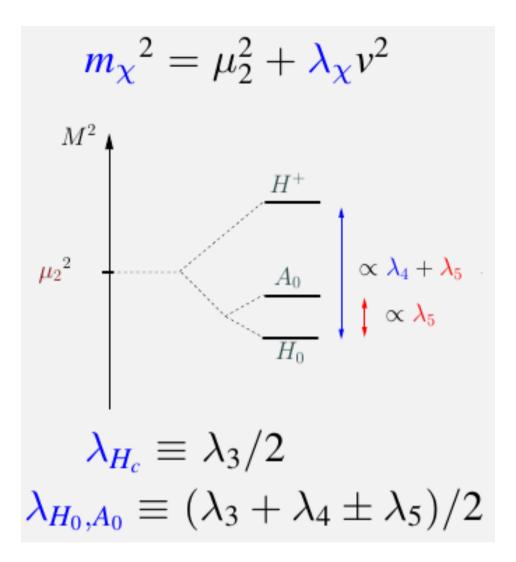
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Electroweak symmetry breaking

$$\langle \Phi \rangle = \begin{pmatrix} 0 \\ \frac{v}{\sqrt{2}} \end{pmatrix} , \qquad \langle \eta \rangle = \begin{pmatrix} 0 \\ 0 \end{pmatrix} \qquad \checkmark \qquad Z_2 \text{ is not spontaneously broken}$$

If the lightest particle that is charged under Z_2 is neutral : we have a **dark matter** candidate!!!





For a heavy dark matter candidate $(M_{H^0} \gg M_W)$ the splitting is relatively small and we expect the particles belonging to the extra doublet to have nearly degenerate masses .

 $m_{H_0} \lesssim m_W$: GeV range $H_0 H_0 \to h^* \to \bar{f}f$ and $H_0 A_0 \to Z^* \to \bar{f}f$

Barbieri PRD06, LLH JCAP06, Gustafsson PRL07, Cao PRD07, Andreas JCAP08,...

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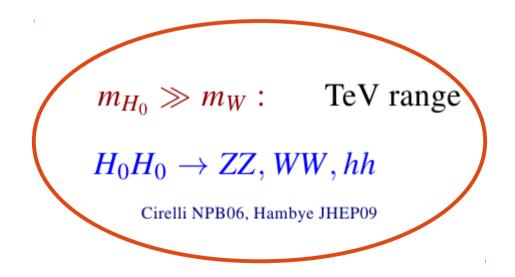
 $m_{H_0} \gg m_W$: TeV range

 $H_0H_0 \rightarrow ZZ, WW, hh$

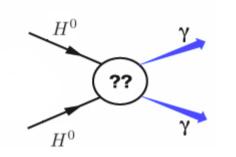
Cirelli NPB06, Hambye JHEP09

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Indirect Searches



No astrophysical uncertainties "Smoking gun" Potentially low statistics.

Gustaffson et al. 2007

Indirect Searches

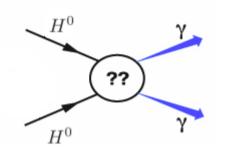


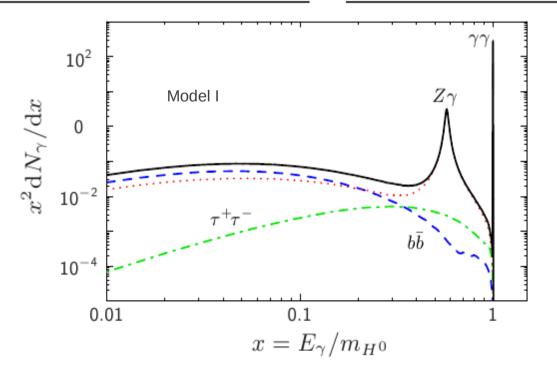
TABLE I: IDM benchmark models. (In units of GeV.)

Model	m_h	m_{H^0}	m_{A^0}	$m_{H^{\pm}}$	μ_2	$\lambda_2 \times 1 \text{ GeV}$
Ι	500	70	76	190	120	0.1
II	500	50	58.5	170	120	0.1
III	200	70	80	120	125	0.1
IV	120	70	80	120	95	0.1

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TABLE II: IDM	benchmark	model	results.
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Model	$v\sigma_{tot}^{v\to 0}$	Branching ratios [%]:	$\Omega_{\rm CDM} h^2$
	$[{\rm cm}^3 {\rm s}^{-1}]$	$\gamma\gamma$ $Z\gamma$ $bar{b}$ $car{c}$ $ au^+ au^-$	
Ι	1.6×10^{-28}	36 33 26 2 3	0.10
II	8.2×10^{-29}	29 0.6 60 4 7	0.10
III	8.7×10^{-27}	2 2 81 5 9	0.12
IV	1.9×10^{-26}	$0.04 \ 0.1 \ 85 \ 5 \ 10$	0.11



Gustaffson et al. 2007

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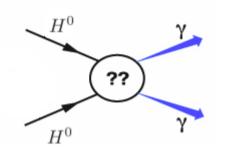


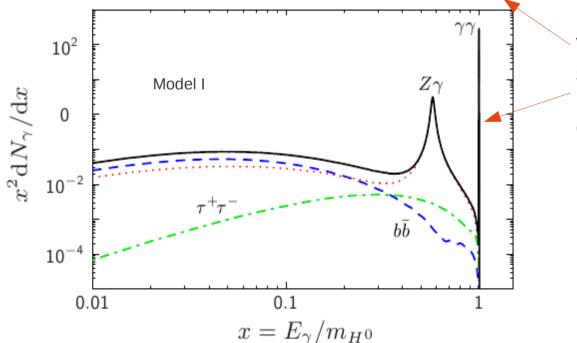
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Very prominent spectral features, but very small cross sections (loop suppressed)

Gustaffson et al. 2007

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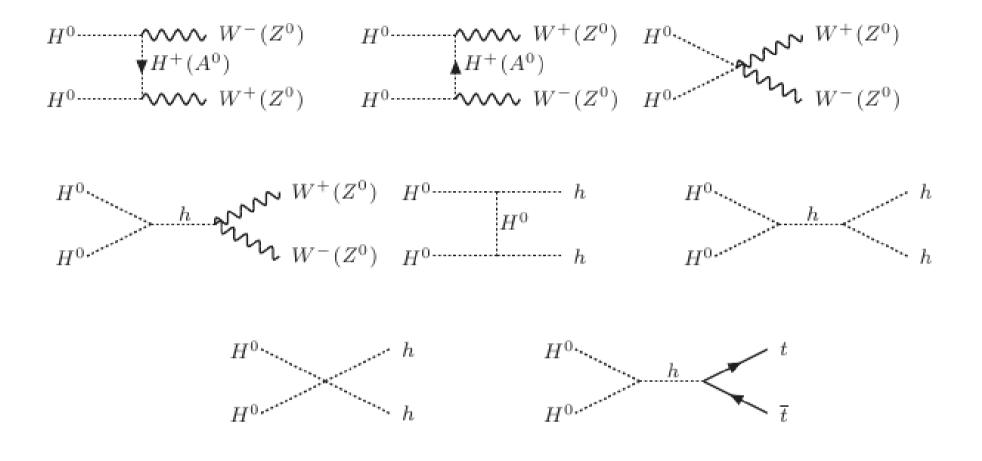
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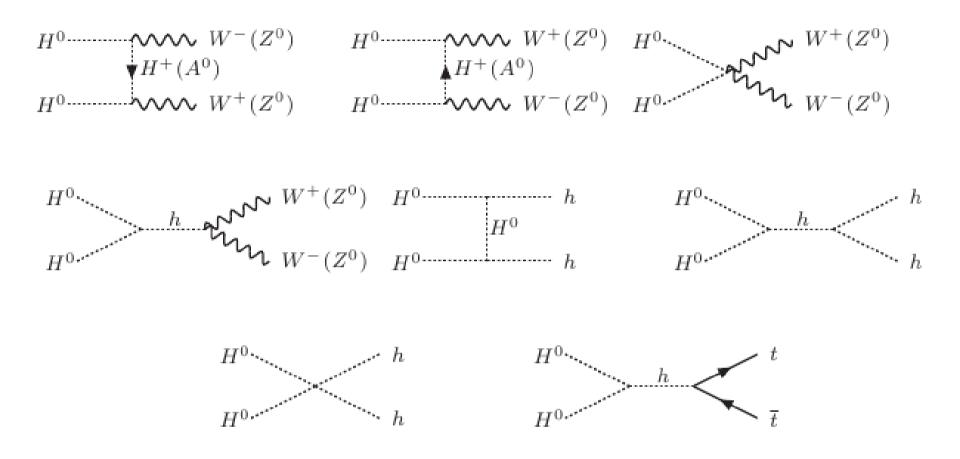
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That is the case for the inert doublet model in the high mass regime if X is a W boson!

Annihilation diagrams



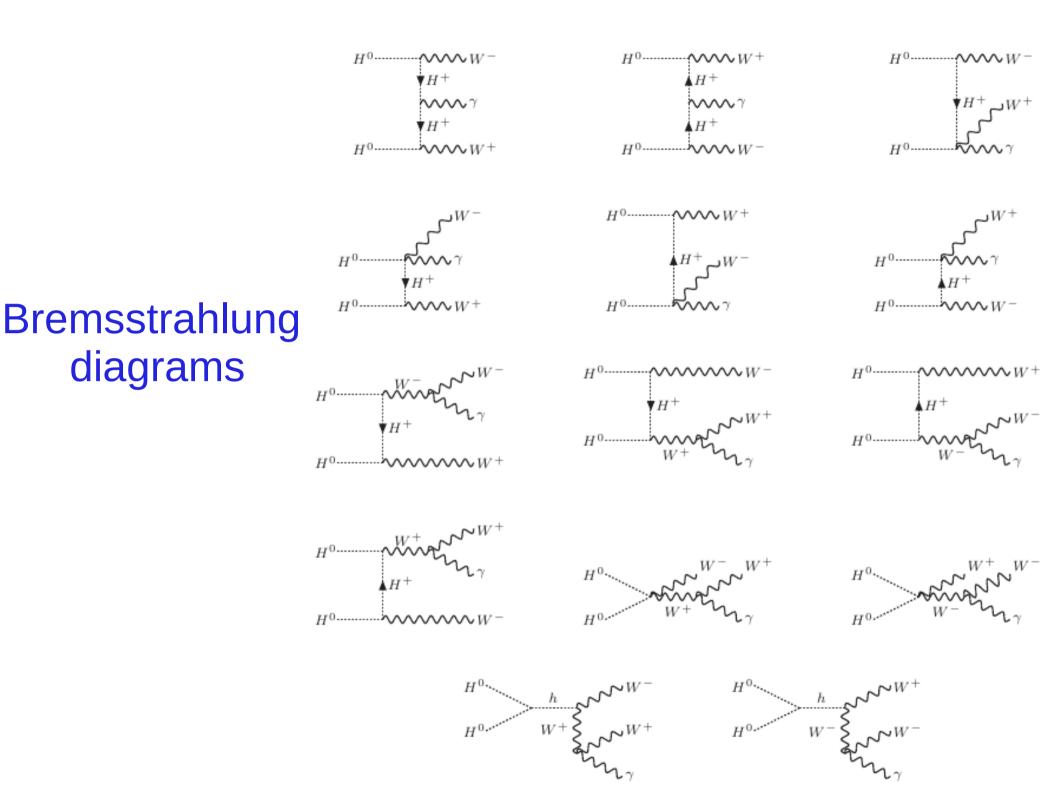
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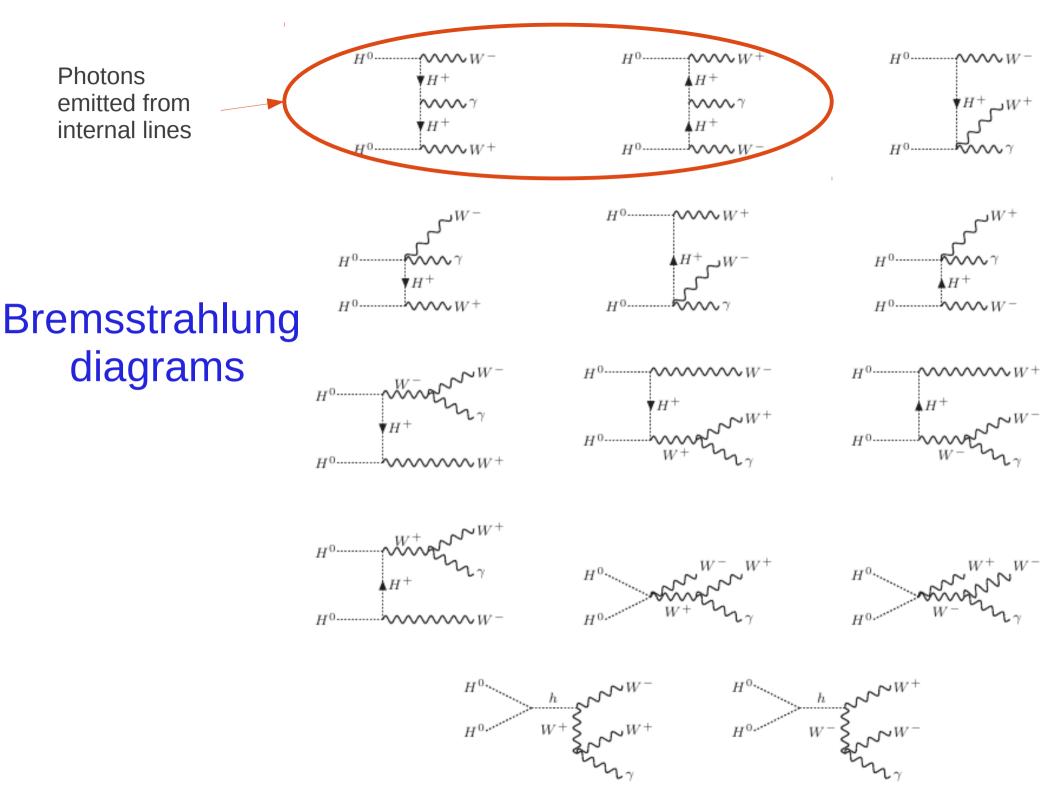


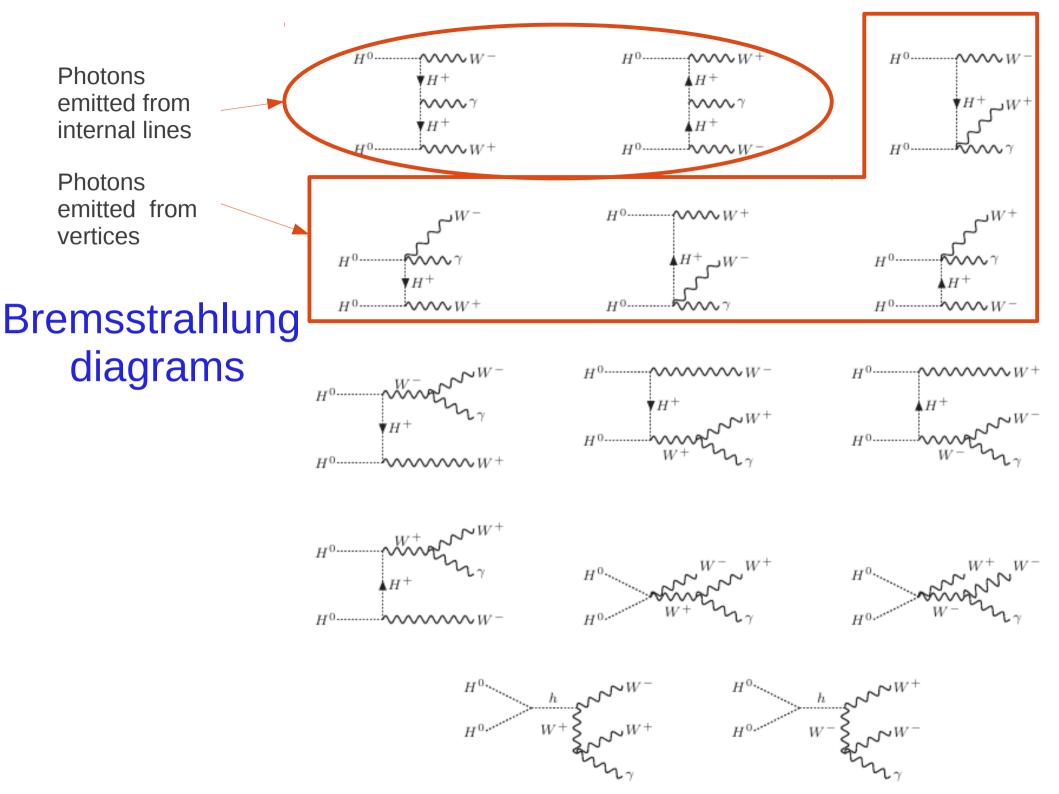
Why the t-channel?

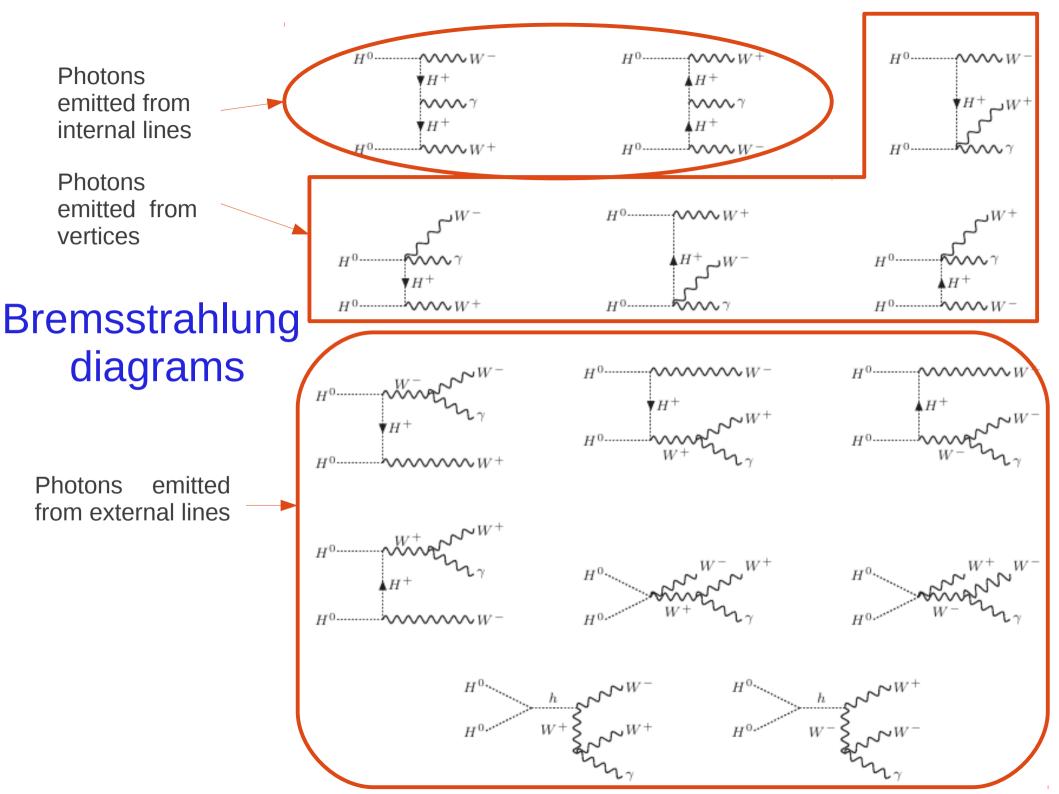
$$D_t(p_W) \propto \left((p_{H^0} - p_W)^2 - M_{H^+}^2 \right)^{-1} \\\approx \left(M_{H^0}^2 + M_W^2 - M_{H^+}^2 - 2M_{H^0} E_W \right)^{-1}$$

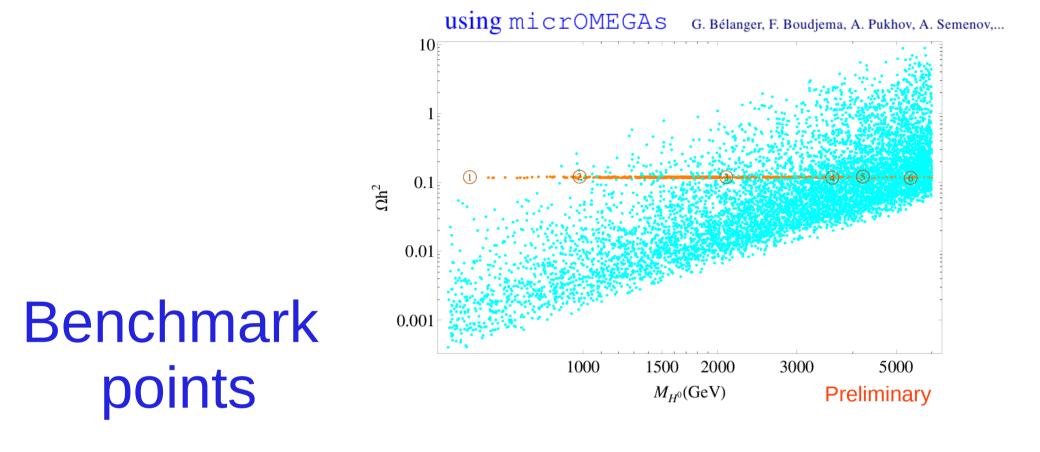
If H^0 and H^+ are almost degenerate in mass, one thus finds an enhancement for small E_W .



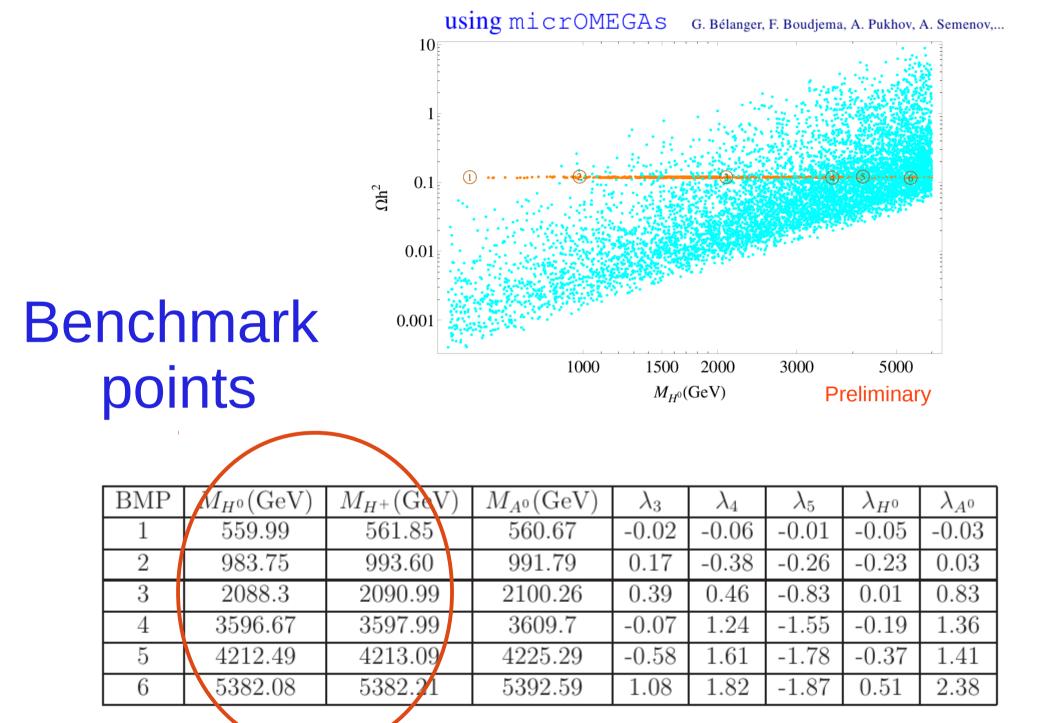








BMP	$M_{H^0}(\text{GeV})$	$M_{H^+}(\text{GeV})$	$M_{A^0}(\text{GeV})$	λ_3	λ_4	λ_5	λ_{H^0}	λ_{A^0}
1	559.99	561.85	560.67	-0.02	-0.06	-0.01	-0.05	-0.03
2	983.75	993.60	991.79	0.17	-0.38	-0.26	-0.23	0.03
3	2088.3	2090.99	2100.26	0.39	0.46	-0.83	0.01	0.83
4	3596.67	3597.99	3609.7	-0.07	1.24	-1.55	-0.19	1.36
5	4212.49	4213.09	4225.29	-0.58	1.61	-1.78	-0.37	1.41
6	5382.08	5382.21	5392.59	1.08	1.82	-1.87	0.51	2.38



I.

Cross Sections and Spectra

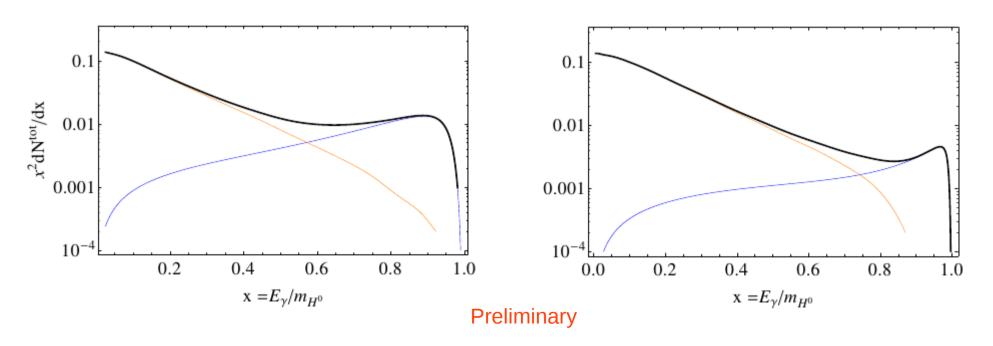
BMP	$\sigma v (10^{-26} {\rm cm}^3/{\rm s})$		Ωh^2				
DMI	<i>00</i> (10 Cm /S)	W^+W^-	ZZ	hh	$t\bar{t}$	$W^+W^-\gamma$	2211
1	6.42	50.25	43.54	2.57	0.60	3.03	0.119
2	3.95	34.91	22.87	36.02	3.03	3.17	0.121
3	4.58	13.80	84.35	0.01	0.00	1.84	0.119
4	3.52	2.36	95.29	1.83	0.01	0.51	0.117
5	3.14	8.75	83.82	5.80	0.03	1.60	0.121
6	5.38	9.29	84.92	3.92	0.01	1.86	0.116

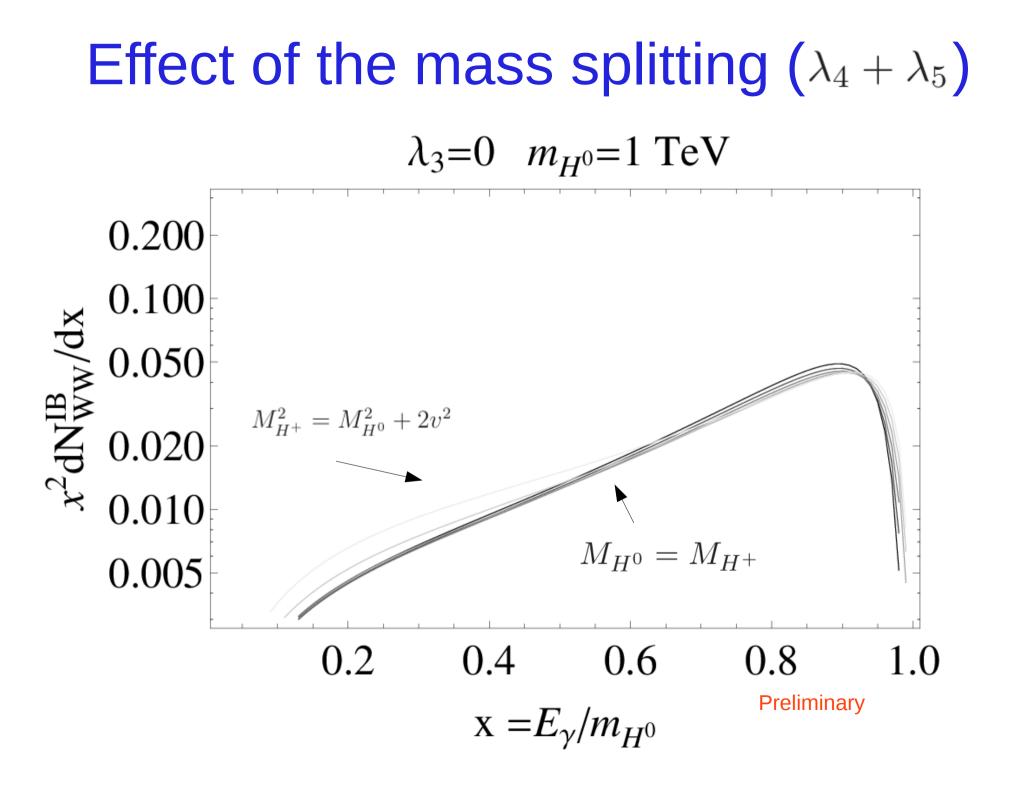
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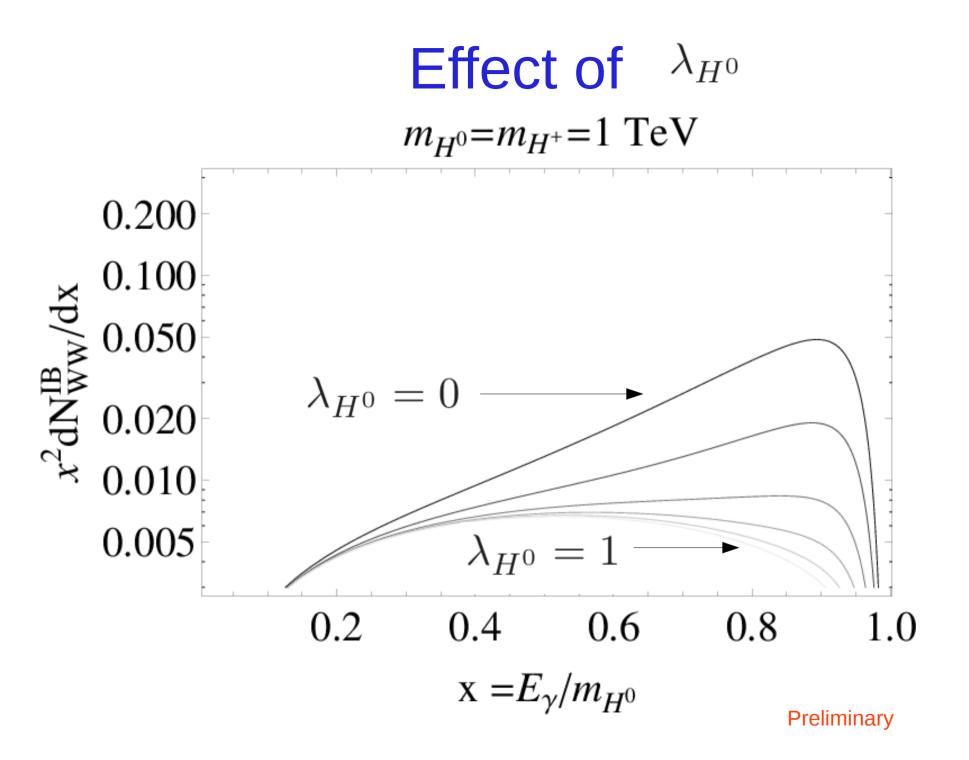
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BMP 2

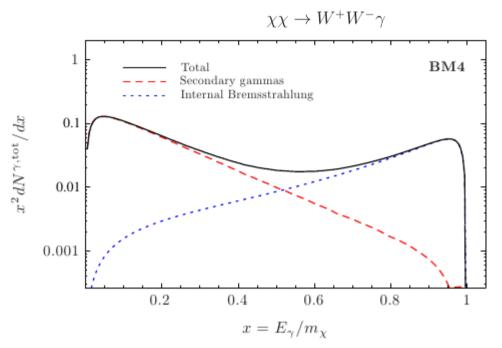






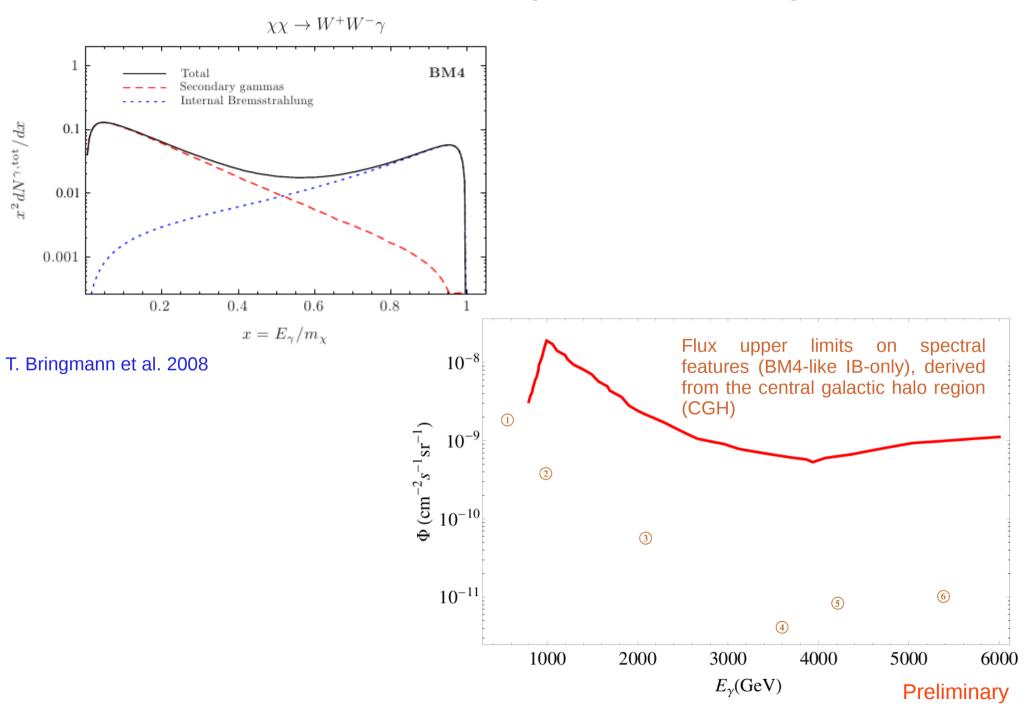


H.E.S.S. searches for photon-like signatures



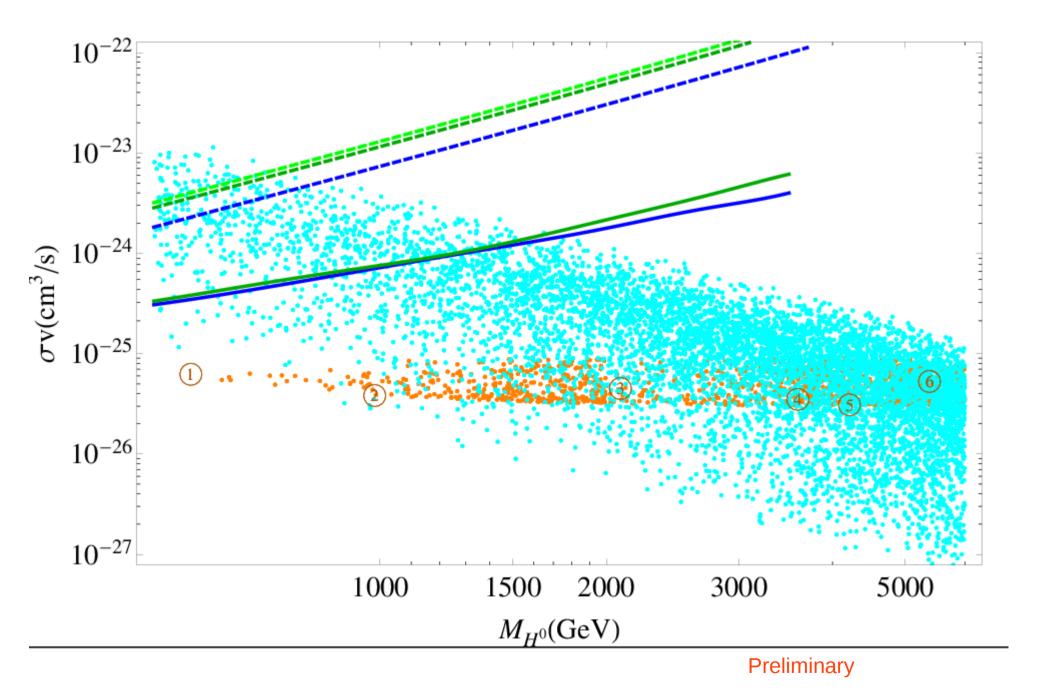
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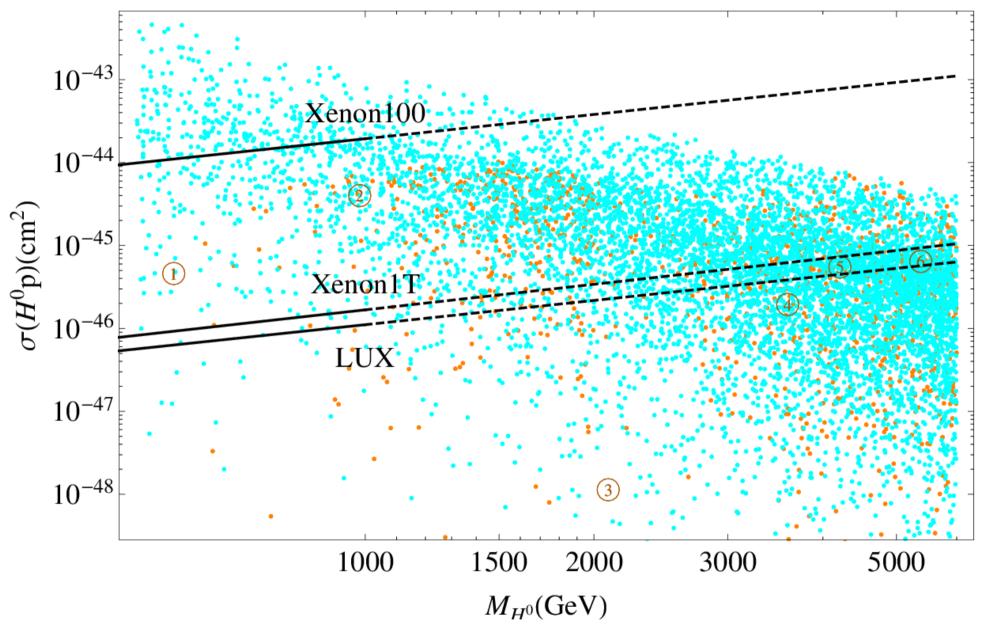
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Conclusions

- Internal Bremsstrahlung signatures are present in the high-mass regime of the inert doublet model.
- In the case of small quartic couplings the feature is more prominent.
- In the high mass regime of the inert doublet model, the internal bremsstrahlung can lead to observable signatures in gamma-ray telescopes





Preliminary