

# Enhanced 1–Loop Corrections to WIMP Annihilation

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# Introduction

Within standard cosmology: WIMP relic density can be computed from WIMP annihilation cross section!

$$\Omega_\chi h^2 \simeq \frac{2.09 \cdot 10^8 \text{ GeV}^{-1} \cdot x_F}{M_{\text{Pl}} \sqrt{g_*} (a + 3b/x_F)}$$

$x_F = m_\chi/T_F$ ;  $T_F$  : freeze-out temperature

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Universal cold DM density now quite well known (WMAP 9-year analysis) (arXiv:1212.5226)

$$\Omega_{\text{CDM}} h^2 = 0.1153 \pm 0.0019 \quad (1.6\% \text{ accuracy!})$$

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$\implies$  Need to include “large” radiative corrections! ( $> \alpha/\pi$ )

# Caveats

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- 1.6% error true for standard 6–parameter  $\Lambda$ CDM; allowing more parameters (in particular, finite curvature) increases the uncertainty
- Also need to know values of relevant parameters to similar accuracy! Very challenging in SUSY scenarios.

# **M**inimal **S**upersymmetric extension of the **S**tandard **M**odel

**WIMP candidate: Lightest neutralino  $\tilde{\chi}_1^0$ .**

$\tilde{\nu}$  excluded by direct searches.

Neutralinos are mixtures of  $\tilde{B}$ ,  $\tilde{W}_3$ ,  $\tilde{h}_1^0$ ,  $\tilde{h}_2^0$ .

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In order to compute loop corrections: need renormalization scheme!

# Improved on-shell renormalization of $\tilde{\chi}$ sector

Chatterjee, MD, Kulkarni, Xu, arXiv:1107.5218 [hep-ph]

$\tilde{\chi}_i^0$ ,  $i \in \{1, 2, 3, 4\}$ , have charged  $SU(2)$  partners: charginos

$\tilde{\chi}_a^\pm$ ,  $a \in \{1, 2\}$



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Mixing of current eigenstates described by mass matrices:

$$M^c = \begin{pmatrix} M_2 & \sqrt{2}M_W \sin \beta \\ \sqrt{2}M_W \cos \beta & \mu \end{pmatrix}.$$
$$M^n = \begin{pmatrix} M_1 & 0 & -M_Z s_W c_\beta & M_Z s_W s_\beta \\ 0 & M_2 & M_Z c_W c_\beta & -M_Z c_W s_\beta \\ -M_Z s_W c_\beta & M_Z c_W c_\beta & 0 & -\mu \\ M_Z s_W s_\beta & -M_Z c_W s_\beta & -\mu & 0 \end{pmatrix}$$

$M_1, M_2$ : gaugino masses;  $\mu$ : higgsino mass.

Other parameters ( $M_W, M_Z, \tan \beta$ ) are renormalized independently!

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In general, there are  $\binom{6}{3} = 20$  different choices of input states!

Selection criterion: Perturbative expansion should converge quickly, i.e. corrections to the remaining masses should be small!

# Reasons for perturbative instability

- $\exists$  parameter  $\in \{M_1, M_2, \mu\}$  that does not affect input masses significantly. E.g. no input state has sizable  $\tilde{B}$  component  $\implies$  finite part of  $\delta M_1$  can be very large.

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- Finite parts can have poles in parameter space. E.g. if  $\delta M_2, \delta\mu$  from  $m_{\tilde{\chi}_1^\pm}, m_{\tilde{\chi}_1^\pm}$ :  $\delta\mu \propto 1/(M_2^2 - \mu^2)$ !

# Solutions

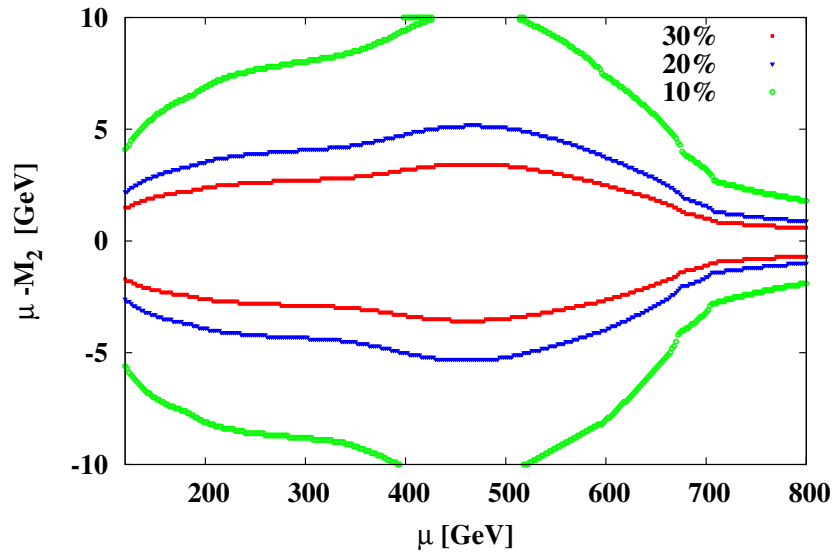
- First condition  $\implies$  must have one  $\tilde{B}$ -like, one  $\tilde{W}$ -like, one  $\tilde{h}$ -like input state. Choice of indices of input states depends on ordering of parameters! (Note: anyway need information on this ordering, since equations determining counterterms have multiple solutions.)



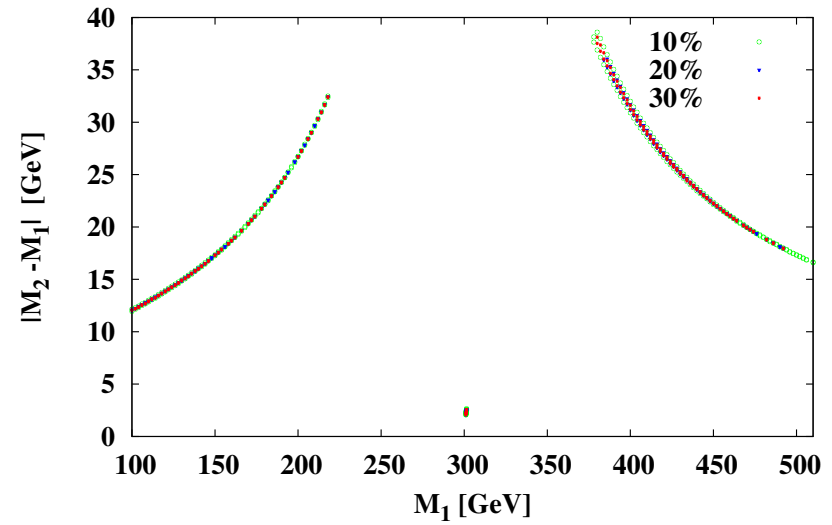
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- Second condition  $\implies$  best choice for input states: Wino-like chargino; higgsino- and bino-like neutralino! All other choices have significant regions with perturbative instability.

# Regions of instability



$\tilde{\chi}_1^\pm, \tilde{\chi}_2^\pm, \tilde{\chi}_{\tilde{B}}^0$  as input

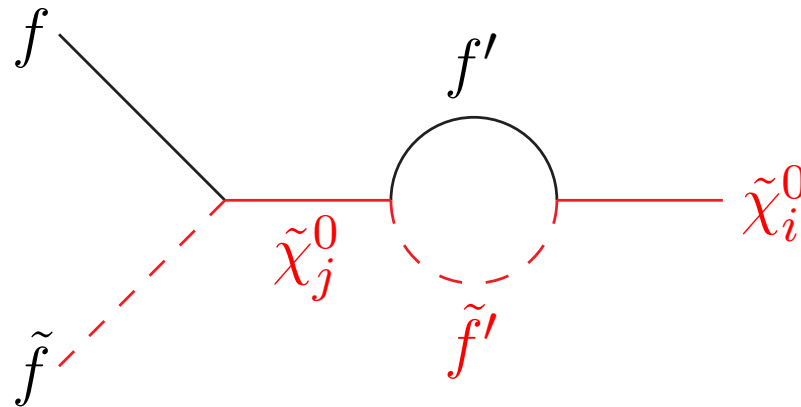


$\tilde{\chi}_{\tilde{B}}^0, \tilde{\chi}_{\tilde{W}}^0, \tilde{\chi}_{\tilde{h}}^0$  as input

# 1-loop corrections via effective $\tilde{\chi}$ couplings

Chatterjee, MD, Kulkarni, arXiv:1209.2328 [hep-ph]

Observation: matter (s)fermion correction to  $\tilde{\chi}$  two-point fcts., plus appropriate counterterms, form finite, gauge invariant subset of corrections! Guasch, Hollik & Sola 2002



Is the *only* diagram (apart from CTs) involving  $f' \neq f$ !

# Properties

- Corrections from “gauge” interactions enhanced by multiplicity factor  $N_F N_C!$

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- Correction from heavy sfermion  $\tilde{f}'$  is *enhanced* logarithmically: no decoupling!

# Pure gaugino sector

Consider spectrum  $m_{\tilde{q}} \gg m_{\tilde{\ell}} \simeq m_{\tilde{\chi}}$ . Effective  $\tilde{\chi}l\tilde{\ell}$  coupling after integrating out heavy squarks:

$$\tilde{g}(m_{\tilde{\chi}}) = \tilde{g}(m_{\tilde{q}}) - \beta_{\ell, \tilde{\ell}} \log \frac{m_{\tilde{q}}}{m_{\tilde{\chi}}}$$

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In effective theory approach: can also have non-decoupling corrections to off-diagonal entries of  $\tilde{\chi}$  mass matrices, i.e. to gaugino-higgsino mixing!

# Order of corrections

coupling	$\tilde{\chi} \simeq \text{higgsino}$		$\tilde{\chi} \simeq \text{gaugino}$	
	tree	one-loop	tree	one-loop
$\tilde{\chi}^{\ell_L} \tilde{\ell}_L, \tilde{\chi}^{\ell_R} \tilde{\ell}_R$	$\mathcal{O}(\epsilon g)$	$\mathcal{O}(\epsilon g^3, \epsilon g \lambda^2)$	$\mathcal{O}(g)$	$\mathcal{O}(g^3, \epsilon^2 g \lambda^2)$
$\tilde{\chi}^{\ell_L} \tilde{\ell}_R, \tilde{\chi}^{\ell_R} \tilde{\ell}_L$	$\mathcal{O}(\lambda_\ell)$	$\mathcal{O}(\lambda_\ell^3, \lambda_\ell \lambda_b^2, \epsilon^2 \lambda_\ell g^2, \epsilon^2 \lambda_\ell \lambda_t^2)$	$\mathcal{O}(\epsilon \lambda_\ell)$	$\mathcal{O}(\epsilon \lambda_\ell g^2, \epsilon \lambda_\ell \lambda^2)$

$g$ : generic electroweak gauge coupling;

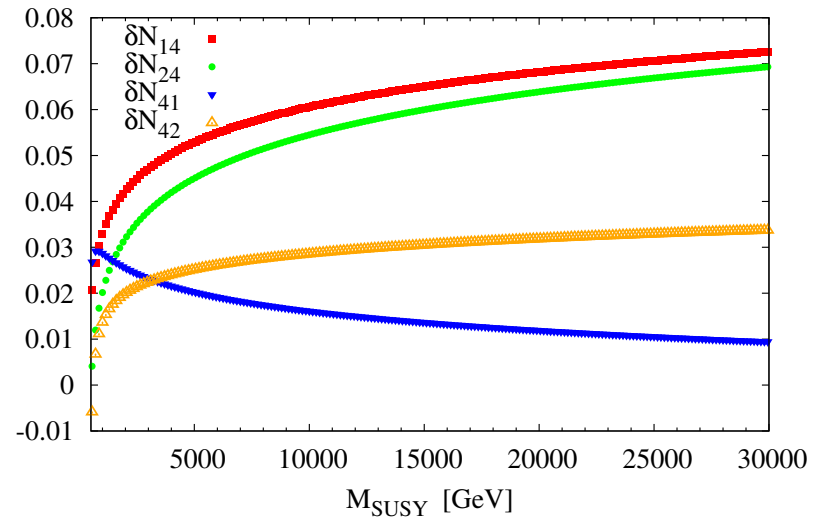
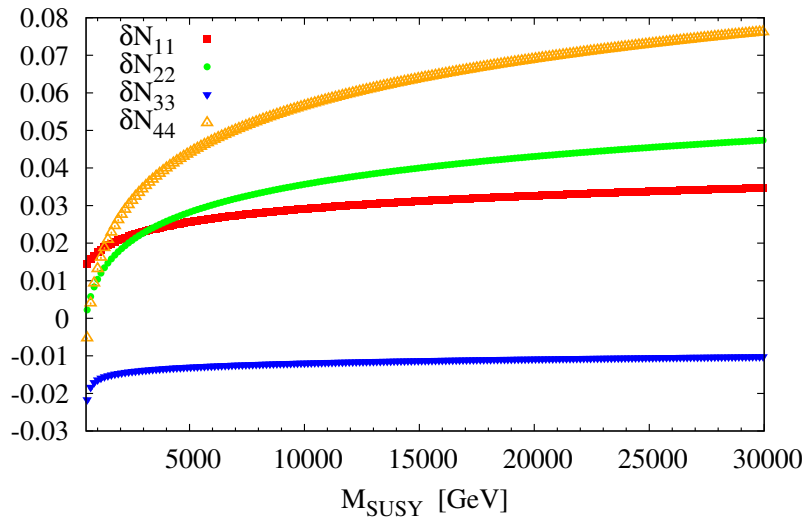
$\lambda_\ell$ : Yukawa coupling of lepton  $\ell$ ;

$\lambda$ : generic superpotential coupling;

$\epsilon$ : one factor of (small) gaugino–higgsino mixing.

**All corrections are non-decoupling!**

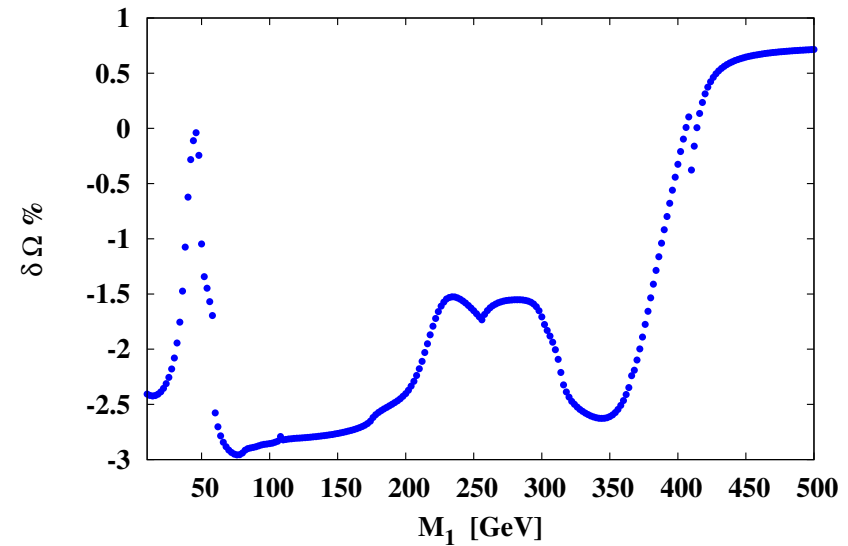
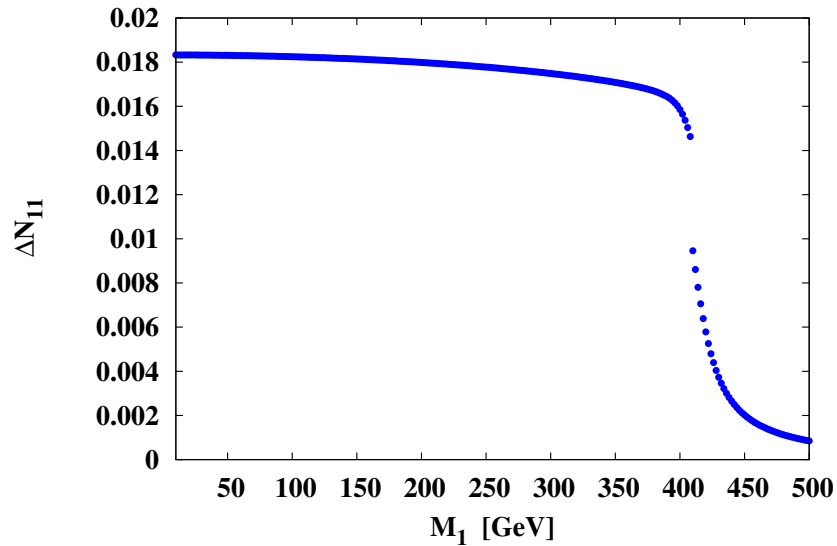
# Relative size of corrections to couplings



$M_2 = 3M_2 = 0.3 \text{ TeV}$ ,  $\mu = 0.6 \text{ TeV}$ ,  $\tan \beta = 10$ , common  $\tilde{f}$  masses

Note: Non-logarithmic non-decoupling corrections can be sizable, too!

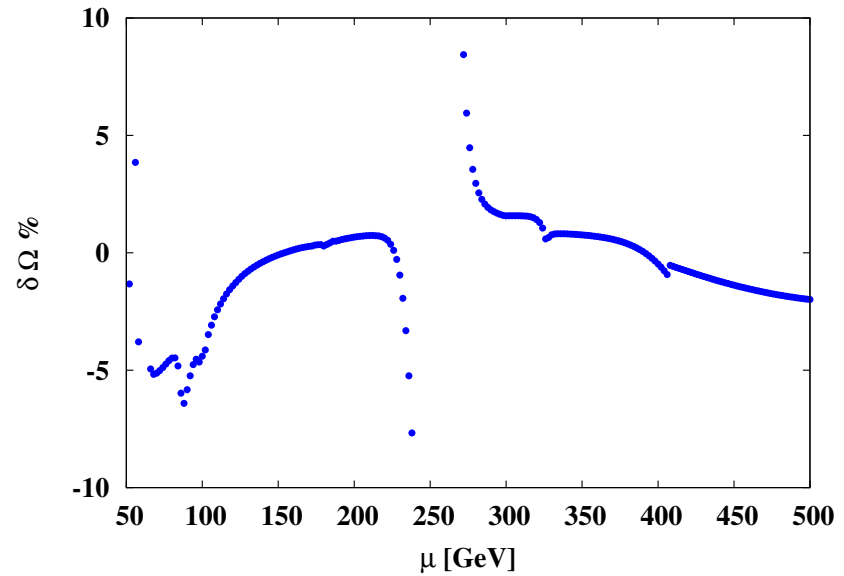
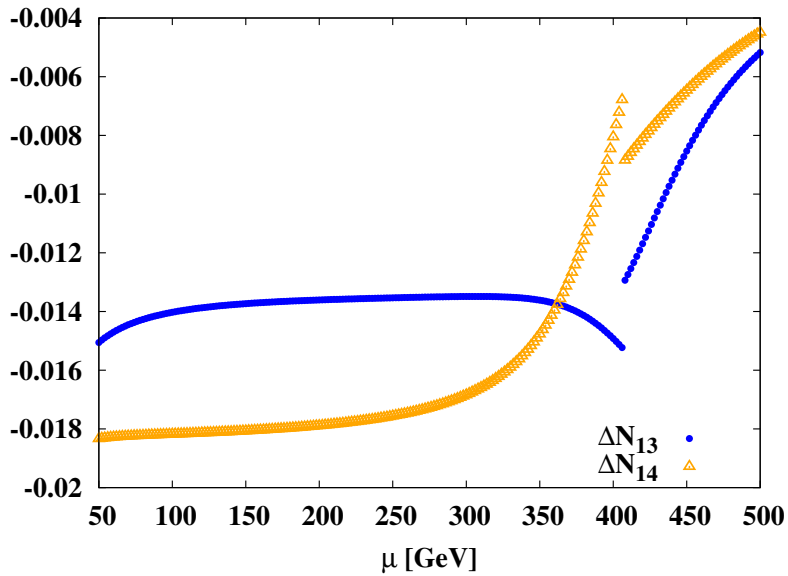
# Correction for (mostly) $\tilde{B}$ -like LSP



$m_{\tilde{q}} = 1.5, M_3 = 1.2, M_2 = 0.4, m_{\tilde{l}_L} = 0.55, m_{\tilde{l}_R} = 0.5,$   
 $\mu = 0.6, m_A = 0.5, A_3 = 1.0, \tan(\beta)(M_Z) = 10$   
(All dim.-ful parameters in TeV)

Obtained with modified micrOMEGAS.

# Correction for $\tilde{B}$ – or $\tilde{h}$ –like LSP



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# Bug in micrOMEGAs

Co-annihilation  $\tilde{\tau}_1 + \tilde{\chi}_1^0 \rightarrow A + \tau$  can diverge if  $m_A \gtrsim 2m_{\tilde{\chi}_1^0}$ !

Reason: Can be due to *on-shell*  $\tilde{\tau}_1 \rightarrow \tau + \tilde{\chi}_1^0$  decay,  
followed by  $\tilde{\chi}_1^0 + \tilde{\chi}_1^0 \rightarrow A$ !

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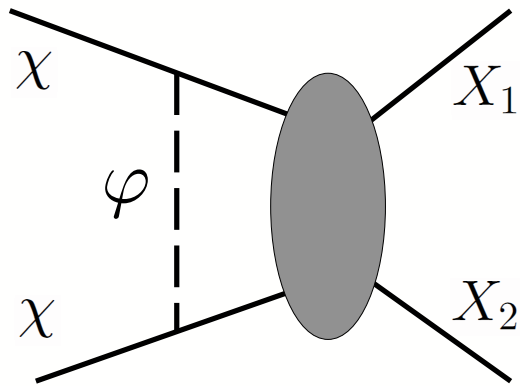
However,  $\tilde{\tau}_1 \leftrightarrow \tau + \tilde{\chi}_1^0$  is part of the (fast) processes maintaining relative equilibrium between  $\tilde{\chi}_1^0$  and  $\tilde{\tau}_1 \implies$  this kinematic configuration should *not* be included in the explicit co-annihilation cross section!



# Sommerfeld enhanced 1-loop corrections

MD, J.M. Kim, Nagao, arXiv:0911.3795 [hep-ph]

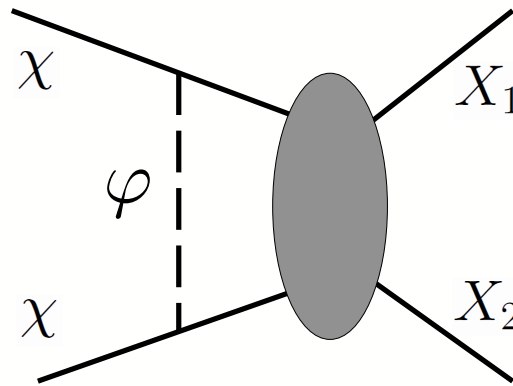
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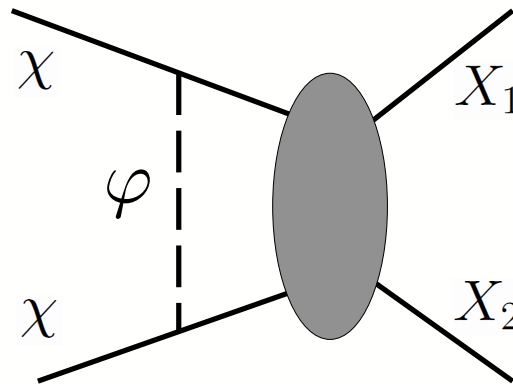


Are interested in case where 1-loop correction is sizable, but still a correction:  $\alpha m_\chi > \mu$ . ( $\alpha$  : relevant coupling;  $\mu$ : mass of exchanged boson.)

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Note:  $\mu < \alpha m_\chi$  may be technically unnatural: expect loop corrections  $\delta\mu^2 = \mathcal{O}(\alpha m_\chi^2/\pi)$ !

# Calculation of 1-loop Correction

Corrections are calculated in non-relativistic kinematics, with several approximations; **these are necessary to give a finite result!**

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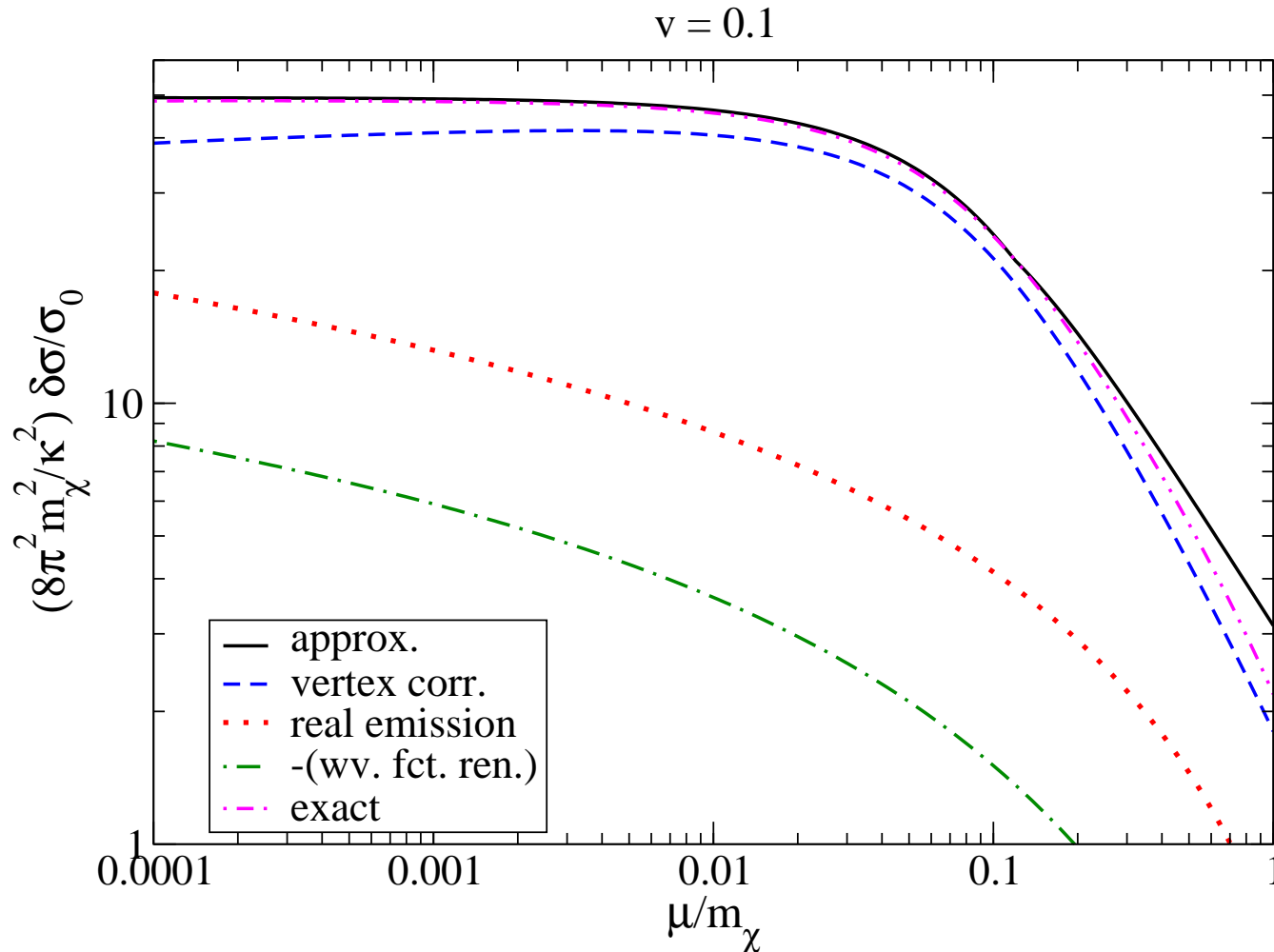
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- Ignore energy dependence of  $\varphi$  propagator

# Check against exact calculation



For scalar WIMP  $\chi$  coupling to (light) boson  $\varphi$  with (dimensionful) coupling  $\kappa$



# Result

- Correction factorizes!

$$\sigma_l^{1\text{-loop}}(\chi\chi \rightarrow \text{any}) = \sigma_l^{\text{tree}}(\chi\chi \rightarrow \text{any}) \cdot \left( 1 + \frac{g^2}{2\pi^2 v_\chi} I_l(\mu/|\tilde{p}_\chi|) \right)$$

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- Loop function:

$$I_S = I_P = \frac{\pi^2}{2} \quad \text{for } \mu \ll |\vec{p}_\chi|$$

$$I_\ell = \frac{2\pi}{(2\ell + 1) \sqrt{(\mu/|\vec{p}_\chi|) + 1}} \quad \text{for } \mu \gg |\vec{p}_\chi|$$

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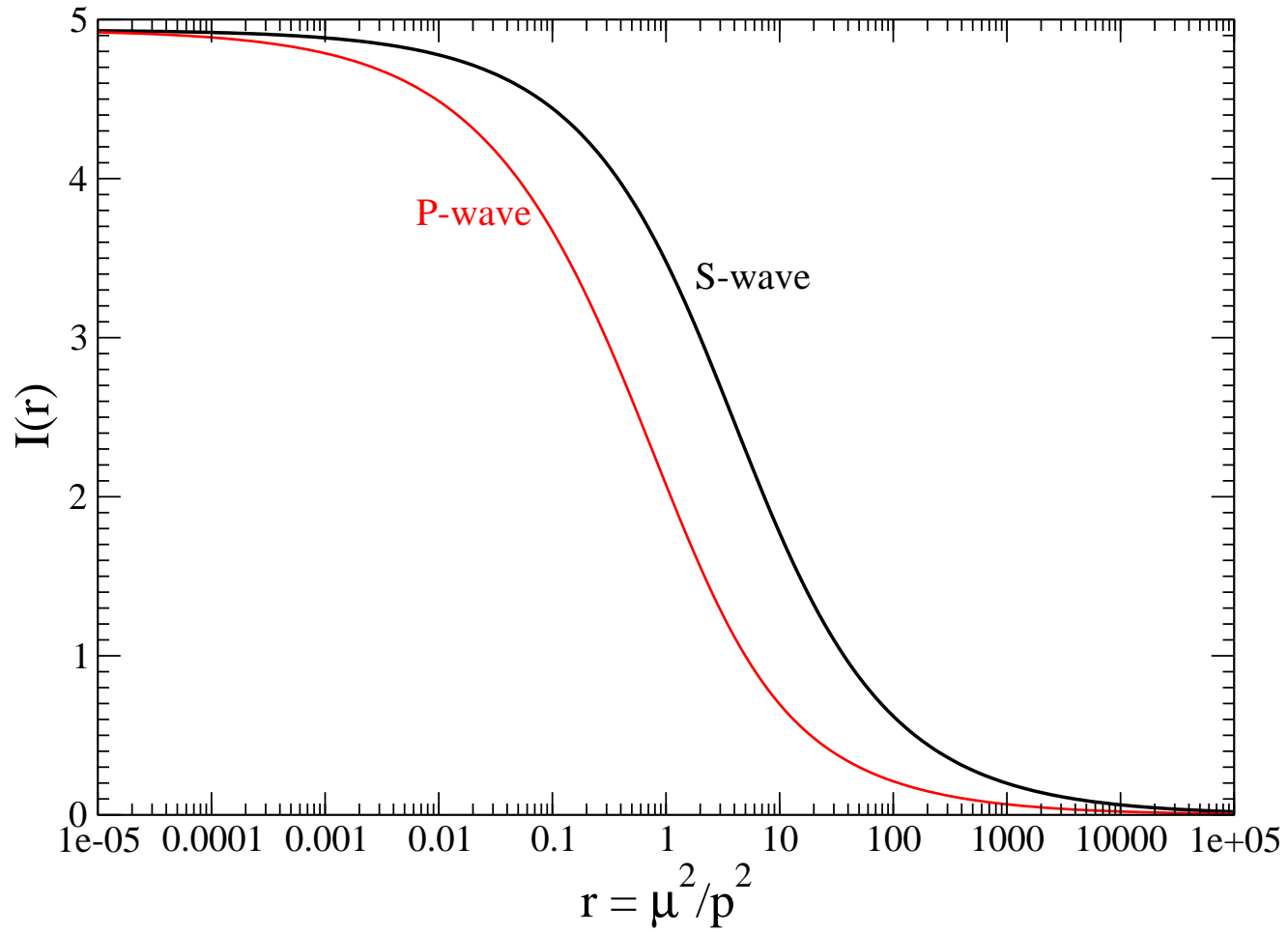
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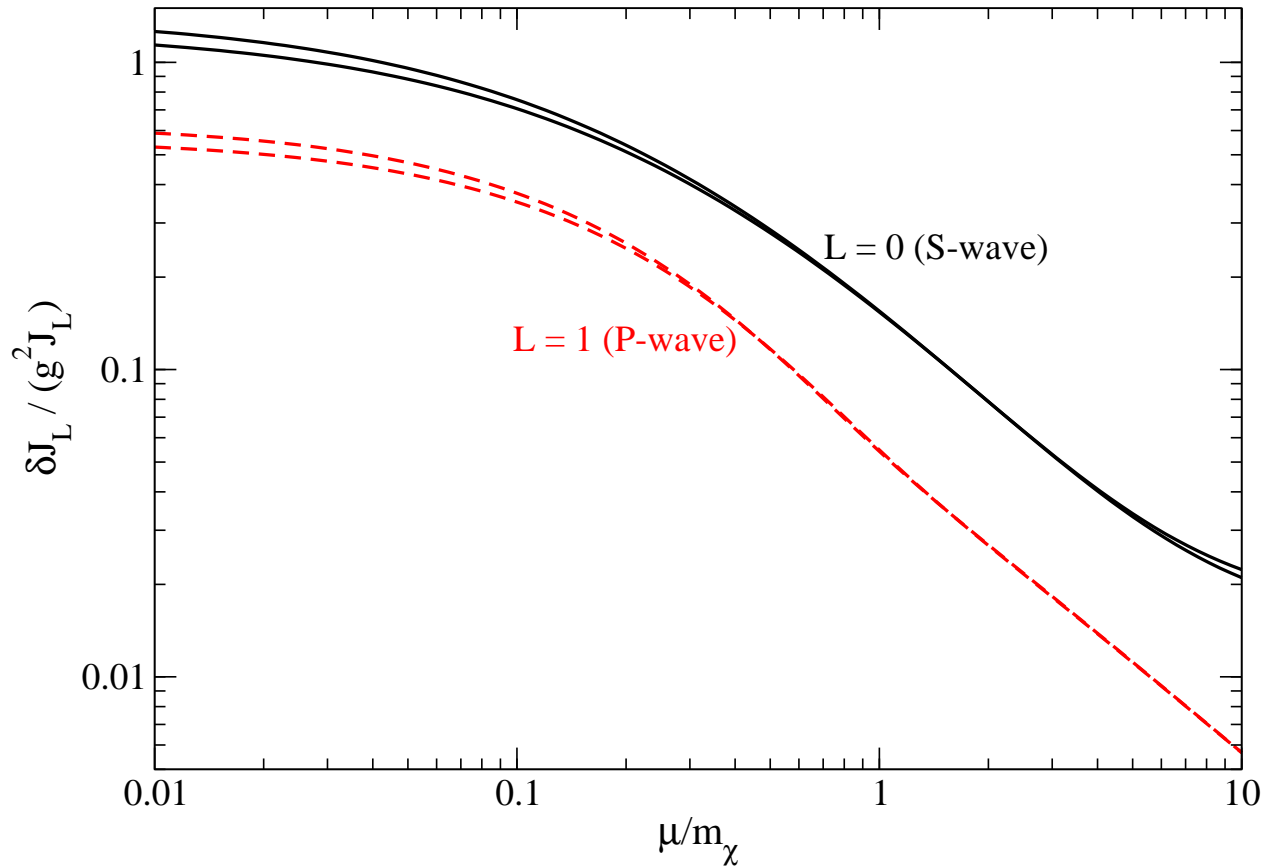
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- Correction depends only on partial wave, not on final state!

# Loop functions



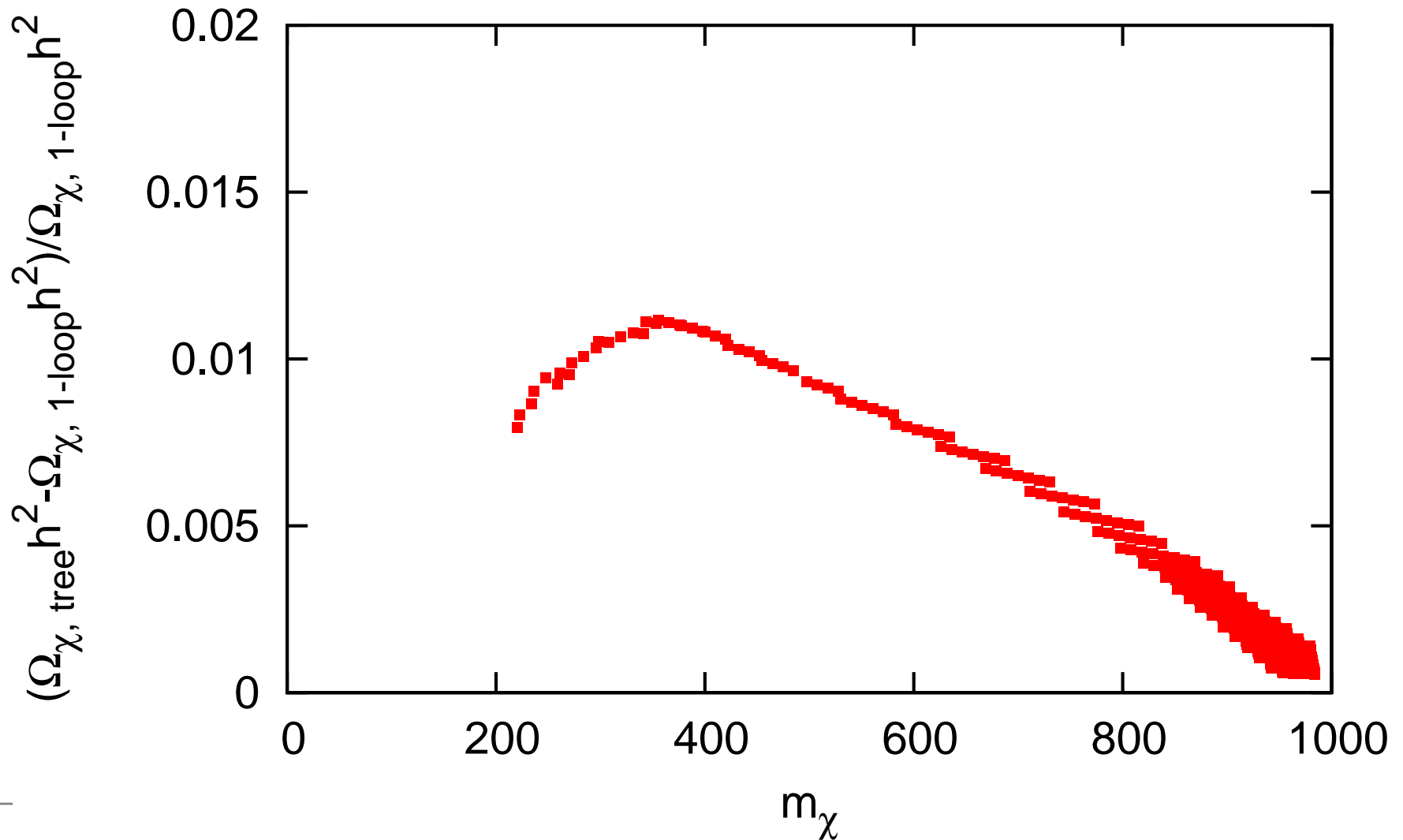
# Correction to annihilation integral



$$J(x_F) = \int_{x_F}^{\infty} \frac{\langle \sigma(\chi\chi \rightarrow \text{any}) v_\chi \rangle}{x^2} dx$$

# Corrections for “well-tempered neutralino” in MSSM

$\tan\beta=2.5, m_{sf}=5\text{TeV}, m_A=5\text{TeV}$



# Incorporating Co-Annihilation

MD, J. Gu, 2013

Complications:

- Masses in intermediate state differ (slightly) from those in the final state; can be bigger or smaller. Described by

$$\kappa = \frac{m_3 m_4}{m_1 m_2} \frac{m_1 + m_2}{m_3 + m_4} - \frac{2m_3 m_4}{m_3 + m_4} \frac{1}{p^2} (m_3 + m_4 - m_1 - m_2)$$

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- Correction now only factorizes on amplitude level:

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# Incorporating Co-Annihilation

MD, J. Gu, 2013

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- Masses in intermediate state differ (slightly) from those in the final state; can be bigger or smaller. Described by

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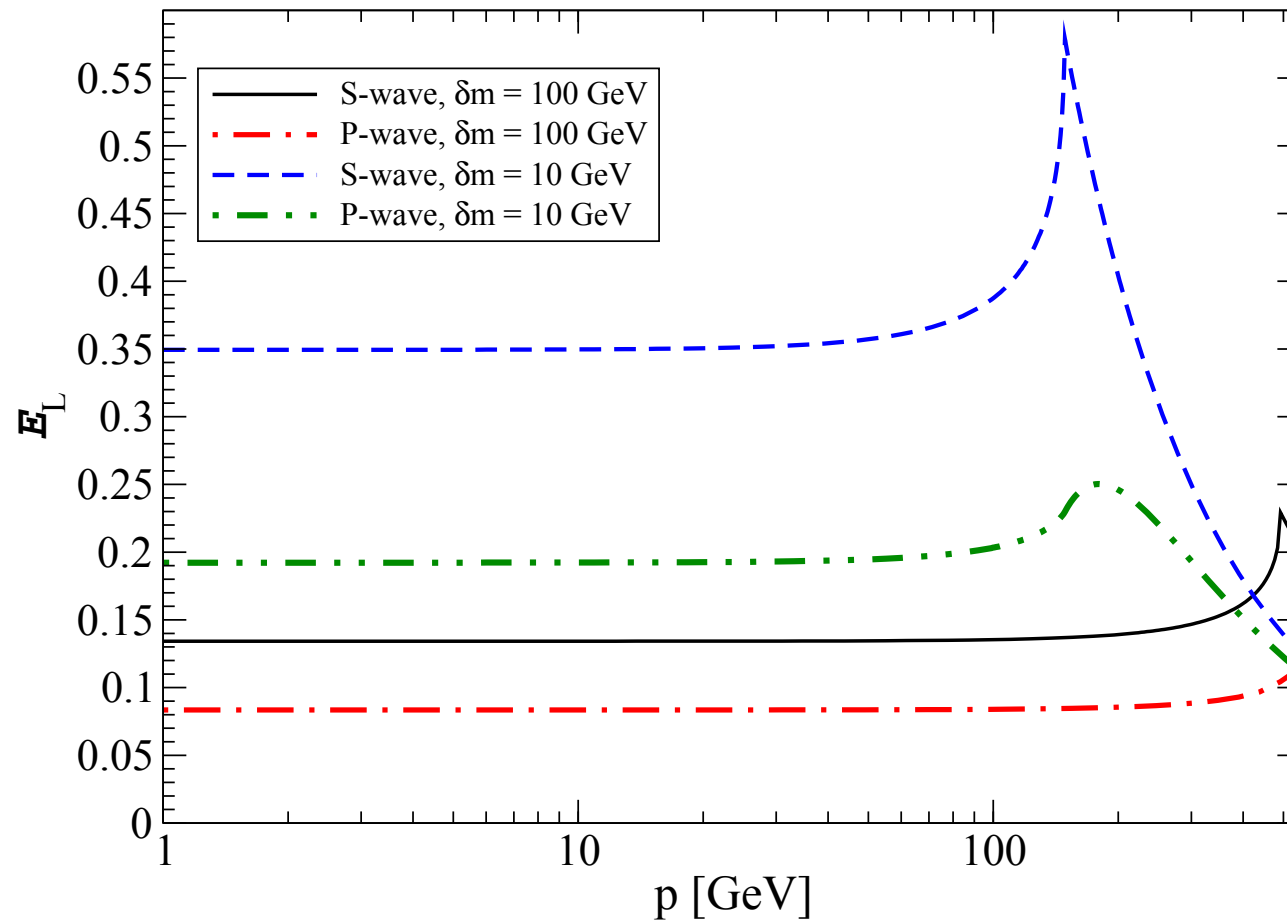
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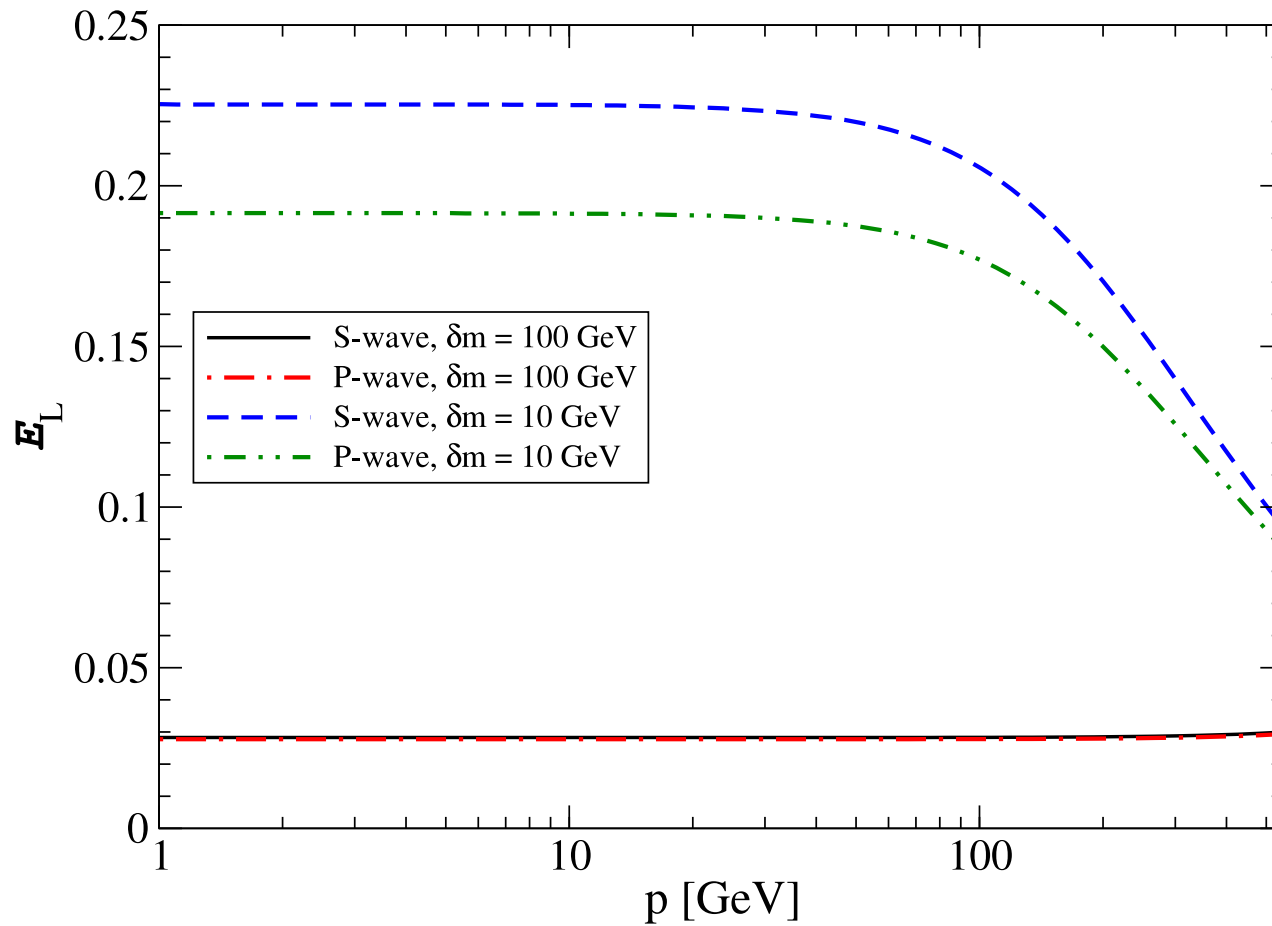
- In general, have to sum over several intermediate states  $\chi_3 \chi_4$ !
- Need to worry about signs (or phases)!

# Loop functions: heavier intermediate state



Correction suppressed, except near threshold for on-shell production of intermediate state

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Correction always suppressed.

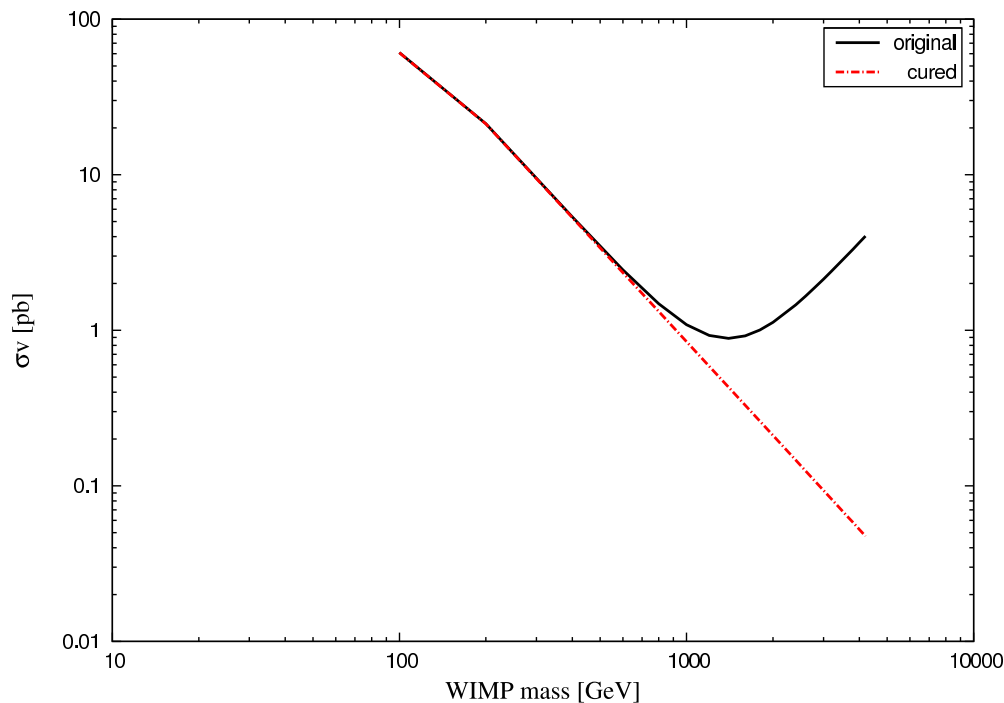
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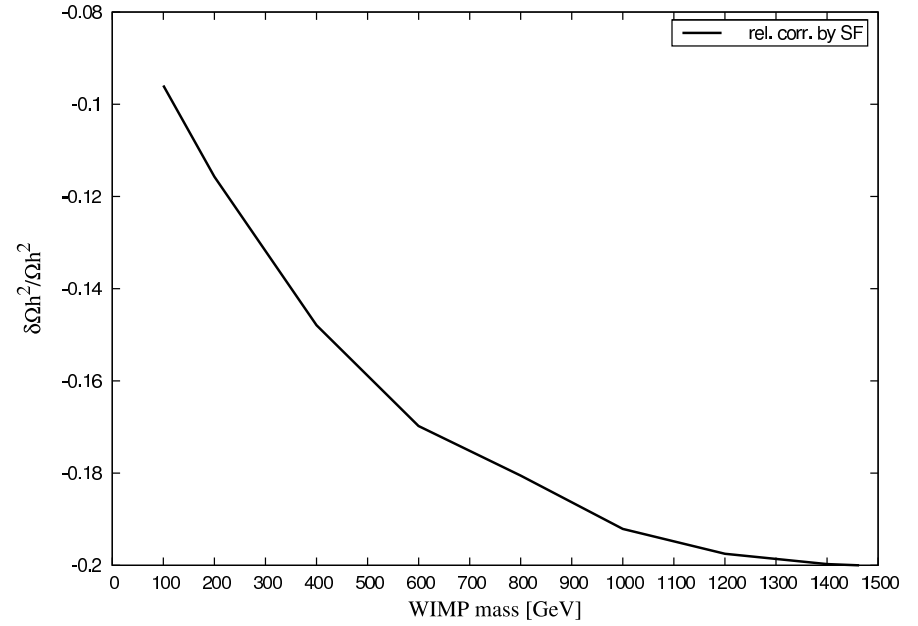
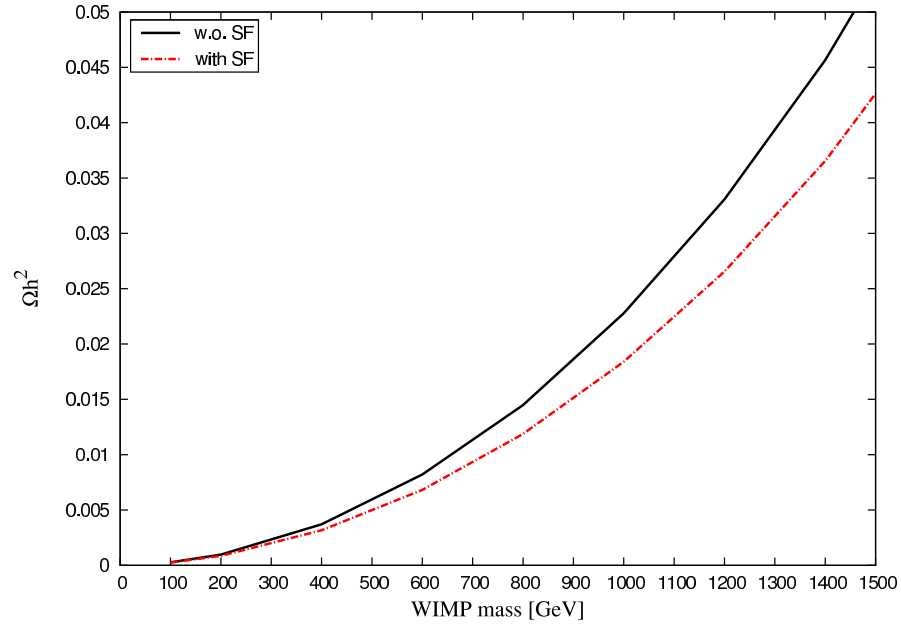
Unfortunately, original (summer 2012) version of DarkSUSY violated unitarity badly:



$v_\chi = 0.001$   
Wino-like LSP

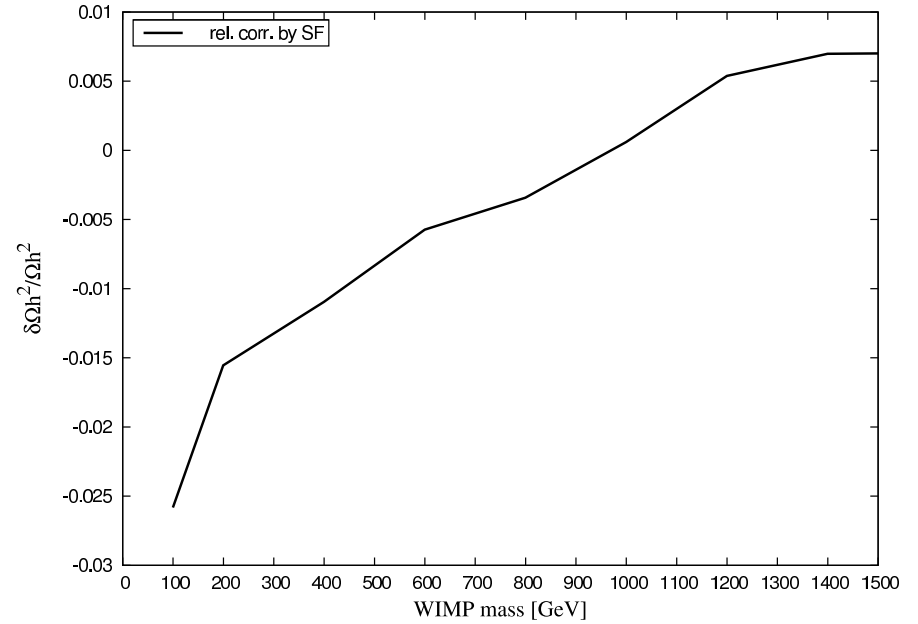
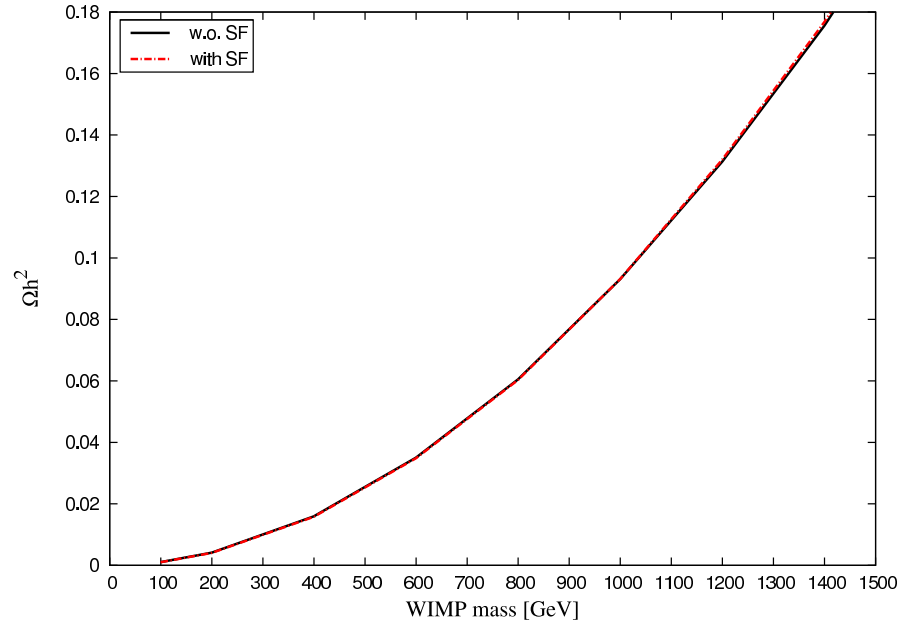
Problem due to use of finite widths for  $t$ -,  $u$ -channel propagators.

# Results for wino-like LSP



Size of effect increases with LSP mass

# Results for higgsino-like LSP



Strong cancellations, e.g. between two neutral higgsinos in loop!



# Difference between higgsinos and winos

Wino-like states form (approximate)  $SU(2)$  triplet of Majorana fermions: no cancellations.

Higgsino-like states form (approximate)  $SU(2)$  doublet of Dirac fermions: strong cancellations.

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- Developed model-independent method to compute one-loop Sommerfeld enhanced corrections; small for  $\tilde{\chi}_1^0$  annihilation in MSSM
- Extended this method to include co-annihilation: large effects for wino-like neutralino, small for higgsino-like neutralino.