The disk entanglement entropy of free Maxwell

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Motivation: 3d F-theorem

Free 3d Maxwell is not a CFT

Calculation using replica trick

Hilbert space interpretation

F-theorem

Finite term in the disk entanglement entropy:

- works away from conformality
- proof of monotonicity
- hard to calculate and even define

[Myers Sinha]

■ Log of the S³ partition function:

- inequality of values at UV and IR conformal fixed points
- much easier to calculate

[DLJ Klebanov Pufu Safdi]

c-theorems in various dimensions

- A measure of the number of degrees of freedom in interacting field theories. It should decrease along rg flow.
- Most obvious conjecture is the thermal free energy. Not constant along conformal manifolds. Also, in 3d, the critical O(N) model is a counter-example.

c-theorems in various dimensions

- In 2d, the coefficient of the trace anomaly has this property. RG flow is the gradient flow for this quantity.
 Zamolodchikov
- In 4d, $16\pi^2 \langle T^{\mu}{}_{\mu} \rangle = c(\text{Weyl})^2 2a(\text{Euler})$, and it is *a* that plays this role.

Komargodski Schwimmer

In odd dimensions, there are no anomalies, so this has long been an open problem.

Sphere partition function

Any conformal field theory can be put on the sphere – it is conformal to flat space.

IR finite observable in odd dim, but non-local.

Weyl invariance is maintained, hence 1-point functions vanish, and Z is constant along conformal manifolds.

S³ partition function well-defined

In general, a calculation in an effective theory with a lower cutoff Λ' < Λ differs by a local effective action for the background fields.

 $\int \sqrt{g}, \quad \int \sqrt{g} \mathcal{R}$

- In odd dimensions, no such counter term can integrate to a number, so the finite term is well-defined.
- Would be interesting to develop the analogous theory to deal with divergences in entanglement entropy.





Shown that $-S + r\partial_r S$ is smaller for the IR fixed point than in the UV, *when* the Hilbert factorizes on the lattice.

[Casini Huerta]

Equal to the sphere partition function for conformal field theories

[Casini Huerta Myers]

Counts topological d.o.f

- The original motivation for studying entanglement entropy in QFT was in condensed matter physics as a way to characterize topological order in gapped systems.
- For example, WZW edge states contribute in pure Chern-Simons.
- Therefore, F cannot be determined only from local correlation functions (likewise for S³ free energy).

Replica trick

- There is a path integral description of the Renyi entropies, $S_n(r) = -\frac{\log(\mathrm{tr}\rho^n)}{n-1}$
- Q is given in the semi-classical wavefunctional of fields basis as the path integral on R³ with boundary conditions on across the disk.

One obtains an n-sheeted branched cover.

 $\mathrm{tr}\rho^n = Z_n/Z_1^n$

From entanglement to spheres

[Casini Huerta Myers]

- The ball in flat space is conformal to the hemisphere in S^{d-1} × R. Its casual diamond can be mapped by a time dependent Weyl rescaling to the static patch in de Sitter. The CFT vacuum maps to the euclidean vacuum.
- The reduced density matrix of any QFT in the static patch is thermal at the dS temperature.
- Analytic continuation of the static patch is S^d. $\rho = \frac{e^{-\beta H}}{\operatorname{tr}(e^{-\beta H})} \qquad -S = \operatorname{tr}(\rho \log \rho) = \operatorname{tr}(\rho(-\beta H - \log Z)) = -\log Z$

Sketch of the proof of the monotonity of entenglement entropy

Strong sub-additivity

 $S(A) + S(B) \ge S(A \cup B) + S(A \cap B)$



 $2S(\sqrt{rR}) \ge S(R) + S(r)$

[Casini Huerta]

 $rS(r) + S(r) \le 0$

Known values

For massless fields, can compute via the sphere partition function. Massive fields known in an expansion in 1/(r m).



$$F_{\rm CS} = \frac{1}{2}\log k$$

Exotic rg flows?

- Consider a weakly coupled abelian Chern-Simons-matter CFT. $F \approx \frac{1}{2} \log k + N_f^{(UV)}$
- The F-theorem permits flows to IR theories with $N_f^{(IR)} > N_f^{(UV)}$.

No examples are known. Maybe there a stronger constraint? Or we should find such flows.

Non-conformality of Maxwell

The stress tensor is $T_{\mu\nu} = F_{\nu\rho}F^{\rho}_{\nu} - \frac{1}{4}\eta_{\mu\nu}F^2$

• This is not traceless $T^{\mu}_{\mu} = \frac{4-d}{4}F^2$

Nor is the virial current $V_{\mu} = \frac{d-4}{2} A^{\nu} F_{\mu\nu}$ gauge invariant.

But charge associated to $\Delta_{\mu} = x^{\nu}T_{\mu\nu} + V_{\mu}$ exists.

[EI-Showk Nakayama Rychkov]

Dual photon description

- It's useful to dualize to a free compact scalar with radius $2\pi g$. $\partial_{\mu} \phi = \epsilon_{\mu\nu\rho} \partial^{\nu} A^{\rho}$
- This is gauging the shift symmetry of the scalar with a Z-gauge field.

 $\int (\partial_{\mu}\phi + 2\pi g A_{\mu})(\partial^{\mu}\phi + 2\pi g A^{\mu}) + 4\pi i \ \epsilon^{\mu\nu\rho}A_{\mu}\partial_{\nu}B_{\rho}$

• The monopole operator is $e^{i\phi/g}$

Spontaneous symmetry breaking

- In d > 2 dimensions, the shift symmetry of the compact scalar theory is spontaneously broken.
 The field takes a definite value, say 0, at infinity in the superselection sector of the vacuum.
- Otherwise cluster decomposition would be violated for correlation functions of $e^{i\phi/g}$

Corresponds to monopole condensation.

S³ partition function

- The theory cannot be conformally coupled because the scalar is compact (no improvement term, it's not a CFT, nor does virial current exist).
- There is a zero mode integral, proportional to the circle size, g. Together with the one loop determinant, one finds a logarithmic dependence

$$F = -\frac{1}{2}\log(g^2 r) + \frac{\zeta(3)}{4\pi^2}$$

Can't see the symmetry breaking on the sphere.

[Klebanov Pufu Sachdev Safdi]

Maxwell-Chern-Simons rg flow

- Consider Maxwell-Chern-Simons at level k. This represents an rg flow from free Maxwell to the gapped Chern-Simons theory.
- The crossover scale is $g^2 k$, so $F_{\text{Maxwell}}(r = \frac{1}{g^2 k}) \sim \frac{1}{2} \log k$
- This is valid for any large k (weak coupling), so for large r, one expects $F_{\text{Maxwell}}(r) \sim -\frac{1}{2}\log(rg^2)$

Instanton sum

- In the n-fold branched cover of R³, there are n asymptotic regions. The field must go to 0 mod 2πg there.
- This results in a sum over winding sectors
 (0, w₁, , w_{n-1}). The fluctuation determinants are all equal.

$$Z_n = Z_{\text{scalar}}^{1-\text{loop}} \sum_w e^{-S(w)}$$

$$S(w) = (2\pi g)^2 r \sum_{jk} (M_n)_{jk} w^j w^k$$

UV limit

At short distances, the sum becomes an integral

$$Z_n^{\text{inst}}(r) = \sum_w e^{-4\pi^2 g^2 r(M_n)_{jk} w^j w^k} \sim \left(\det(\frac{4\pi^2 g^2 r M_n}{\pi})\right)^{-1/2}$$

The Renyi entropies become

$$\Delta S_n(r) = \frac{1}{2} \log(g^2 r) + \frac{1}{2(n-1)} \log \det M_n$$

Thus, $F \approx \frac{\log 2}{8} - \frac{3\zeta(3)}{16\pi^2} + \frac{1}{2} + \frac{\gamma}{2} - \frac{1}{2}\log(rg^2)$

Finding the instanton solutions

- It's convenient to regard the field on the nsheeted cover as an n-vector with monodromy.
- One can then go to a diagonal complex basis. $(M_n)_{jk} = \frac{1}{n} \sum_{m=1}^{n-1} e^{2\pi i (j-k)m/n} J(k/n)$

where $J(\beta) = \frac{1}{2} \int d^3x \partial_\mu \phi^*_\beta \partial^\mu \phi_\beta$ is the action for a solution with multiplicative cut $e^{2\pi i\beta}$

On-shell action

- The on-shell action is equivalent to a boundary integral, which equals the flux since the field approaches 1 at infinity.
- Actually found the solutions in a convenient basis of harmonic functions numerically.

 $J(\beta) = 2\pi(1 - 2\beta)\tan\pi\beta$

Numerical extension

• Unknown how to continue Θ function in n.

 Fitting with rational functions works much better than polynomials.

Check that convergence is good by approximating the known S₂(r) from higher S_n(r), and checking at small r limit.



Hilbert space structure

- The Hilbert space of the compact scalar is the same as the ordinary scalar in flat space, due to spontaneous symmetry breaking by the vacuum.
- However, the Hilbert spaces in the disk (really a sum of selection sectors with different boundary conditions) differ between the two theories.

Disk Hilbert space

- In continuum QFT, one has continuity across the disk, moreover the field value there defines boundary conditions, $\mathcal{H}_{physical} = \bigoplus_f \mathcal{H}_f^{in} \otimes \mathcal{H}_f^{out}$
- Suggests that the reduced density matrix is block diagonal in the boundary field basis (superselection).
- Can check this using the fact that $\rho = e^{-K}$ for conformal field theories. K leaves scalar primaries at the boundary invariant.

$$[\rho,\Phi|_{S^1}]=0$$

Averaging versus interfering

In the compact scalar, the wavefunctionals must be invariant under the shift, in other words one first sums the states and then forms the density matrix.

This gives a purer density matrix than summing afterward, hence the EE deficit, F, is larger.

$$\langle \phi_+ | \rho_{\rm c} | \phi_- \rangle = \sum_n \langle \phi_+ + 2\pi g n | \rho_{\rm nc} | \phi_- + 2\pi g n \rangle$$

Why does F depend on such subtleties?

- It might seem strange that to obtain an rg monotonic quantity one needs something that depends on more data than the Hilbert space and Hamiltonian in flat space.
- However, in principle more rg flows (in the space of local QFTs) are allowed if fewer operators in the UV are considered to be local.

Punchlines

- The 3d compact Maxwell theory flows from the noncompact R gauge theory to the free noncompact scalar.
- Found the interpolating F(r). It diverges in the UV.
- F depends on more data than H and the Hilbert space in Minkowski space.