

FastLim: a fast LHC limit calculator

Kazuki Sakurai



In collaboration with:

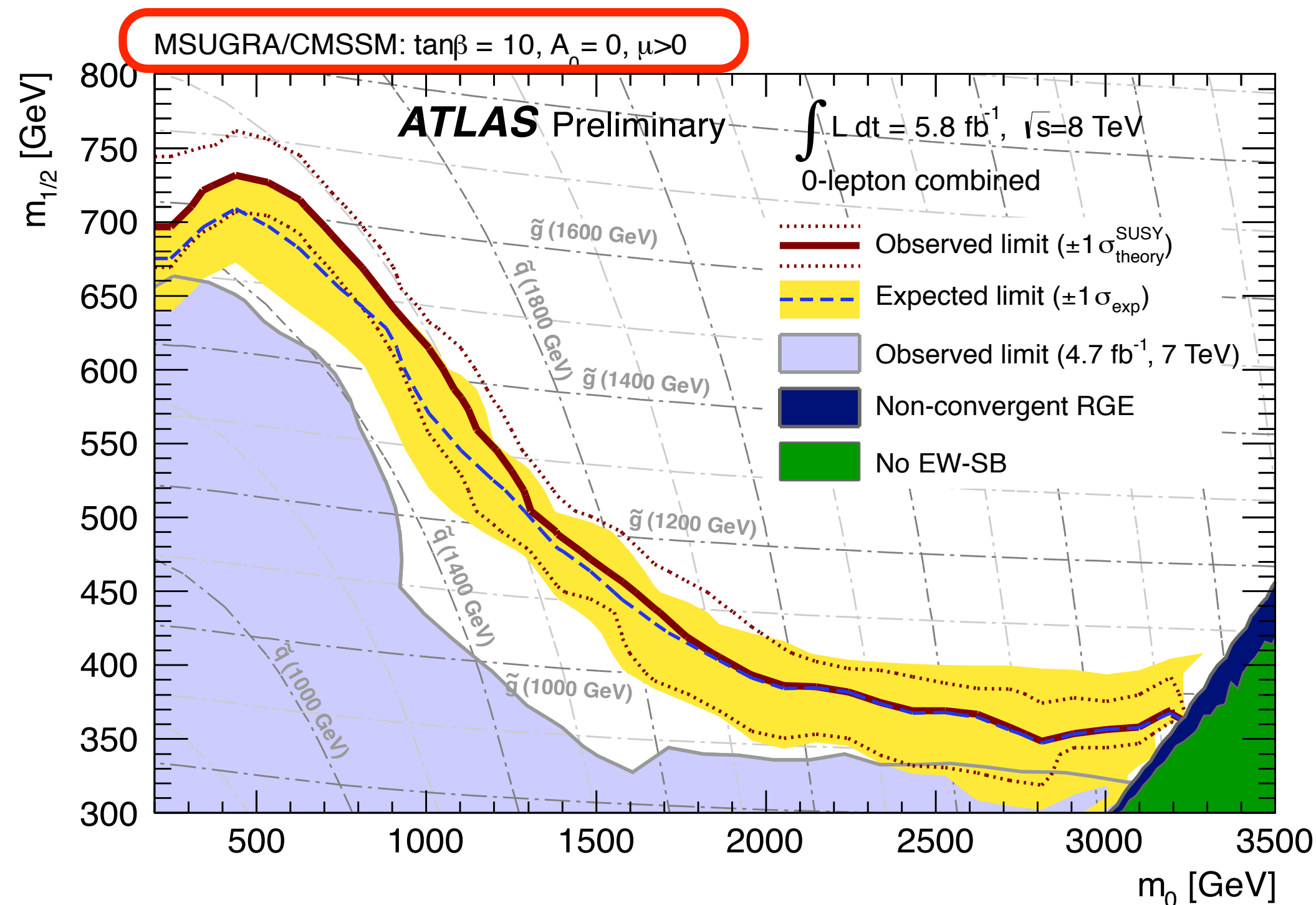
Michele Papucci (LBNL), Andreas Weiler, Lisa Zeune

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- Introduction
- A fast LHC limit calculation and **FastLim**
- FastLim demonstration
- Application
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Introduction

- LHC worked well and produced a lot of SUSY limits
- The limits are not generic and employ assumptions → cannot be applied to your model
- MSSM has more than 100 parameters → no way to calculate/visualise the generic limit



MSUGRA/CMSSM

MSUGRA/CMSSM

$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$

$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$

$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_1^\pm \rightarrow qqW^\pm\tilde{\chi}_1^0$

$\tilde{g}\tilde{g} \rightarrow qqql\ell(\ell\ell)\tilde{\chi}_1^0\tilde{\chi}_1^0$

GMSB ($\tilde{\ell}$ NLSP)

GMSB ($\tilde{\ell}$ NLSP)

GGM (bino NLSP)

GGM (wino NLSP)

GGM (higgsino-bino NLSP)

GGM (higgsino NLSP)

ATLAS Preliminary

$\tilde{g}\tilde{g}$	1.2 TeV	any $m(\tilde{q})$
$\tilde{g}\tilde{g}$	1.1 TeV	any $m(\tilde{q})$
$\tilde{q}\tilde{q}$	740 GeV	$m(\tilde{\chi}_1^0) = 0 \text{ GeV}$
$\tilde{g}\tilde{g}$	1.3 TeV	$m(\tilde{\chi}_1^0) = 0 \text{ GeV}$
$\tilde{g}\tilde{g}$	1.18 TeV	$m(\tilde{\chi}_1^0) < 200 \text{ GeV}$
$\tilde{g}\tilde{g}$	1.1 TeV	$m(\tilde{\chi}_1^0) < 650 \text{ GeV}$
$\tilde{g}\tilde{g}$	1.24 TeV	$\tan\beta < 15$
$\tilde{g}\tilde{g}$	1.4 TeV	$\tan\beta > 18$
$\tilde{g}\tilde{g}$	1.07 TeV	$m(\tilde{\chi}_1^0) > 50 \text{ GeV}$
$\tilde{g}\tilde{g}$	619 GeV	$m(\tilde{\chi}_1^0) > 50 \text{ GeV}$
$\tilde{g}\tilde{g}$	900 GeV	$m(\tilde{\chi}_1^0) > 220 \text{ GeV}$
$\tilde{g}\tilde{g}$	690 GeV	$m(\tilde{H}) > 200 \text{ GeV}$

How to find the limit on your model?

ATLAS-CONF-2011-086

Signal Region	≥ 2 jets	≥ 3 jets	≥ 4 jets
$E_{\text{T}}^{\text{miss}}$ [GeV]	> 130	> 130	> 130
Leading jet p_{T} [GeV]	> 130	> 130	> 130
Second jet p_{T} [GeV]	> 40	> 40	> 40
Third jet p_{T} [GeV]	–	> 40	> 40
Fourth jet p_{T} [GeV]	–	–	> 40
$\Delta\phi(\text{jet}_i, E_{\text{T}}^{\text{miss}})_{\text{min}}$ ($i = 1, 2, 3$)	> 0.4	> 0.4	> 0.4
$E_{\text{T}}^{\text{miss}}/m_{\text{eff}}$	> 0.3	> 0.25	> 0.25
m_{eff} [GeV]	> 1000	> 1000	> 1000

Process	Signal Region		
	≥ 2 jets	≥ 3 jets	≥ 4 jets
Prediction	12.1 ± 2.8	10.1 ± 2.3	7.3 ± 1.7
Observed	10	8	7

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ATLAS-CONF-2011-086

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statistically consistent 

How to find the limit on your model?

additional contribution from SUSY events

$N_{\text{SUSY}} = \dots, 2, \dots, 10, \dots ?$

statistically consistent 

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How to find the limit on your model?

$$\frac{N_{\text{SUSY}}^{(i)}}{N_{\text{UL}}^{(i)}} \equiv R^{(i)} \begin{cases} > 1 & \text{: excluded} \\ \leq 1 & \text{: allowed} \end{cases}$$

95% CL upper limit

N_{UL}

allowed

excluded

additional contribution from SUSY events

$N_{\text{SUSY}} = \dots, 2, \dots, 10, \dots ?$

statistically consistent

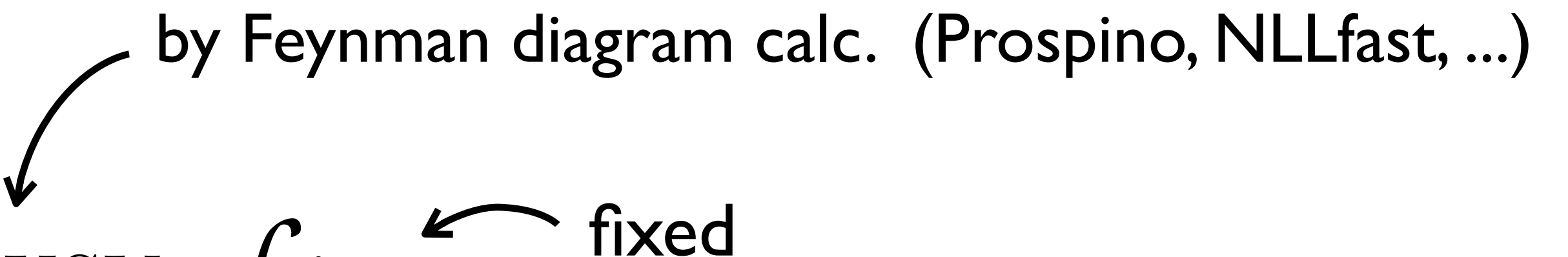
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How to evaluate N_{SUSY} ?

$$N_{\text{SUSY}}^{(i)} = \epsilon_{\text{SUSY}}^{(i)} \cdot \sigma_{\text{SUSY}} \cdot \mathcal{L}_{\text{int}}$$

by Feynman diagram calc. (Prospino, NLLfast, ...)

fixed



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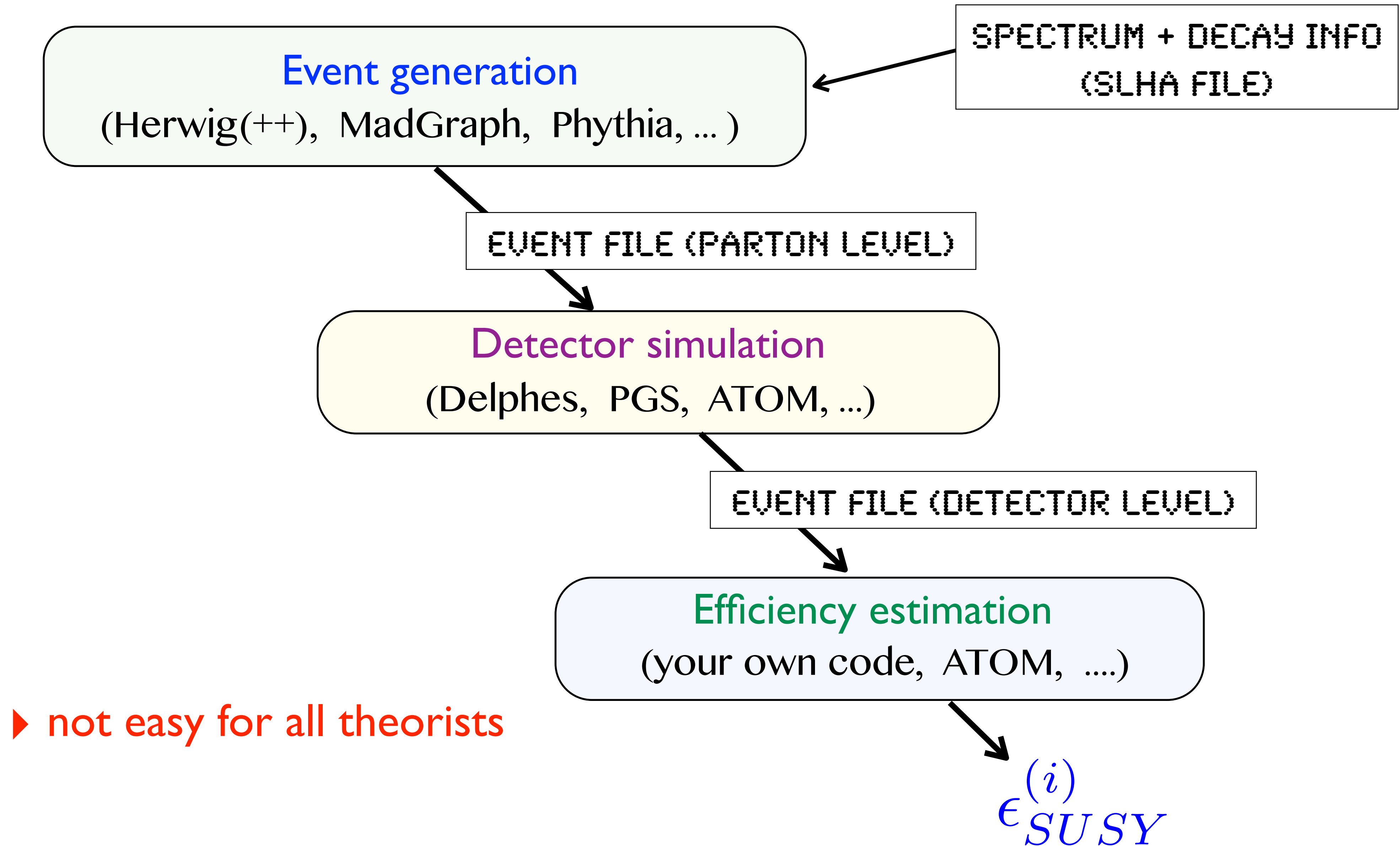
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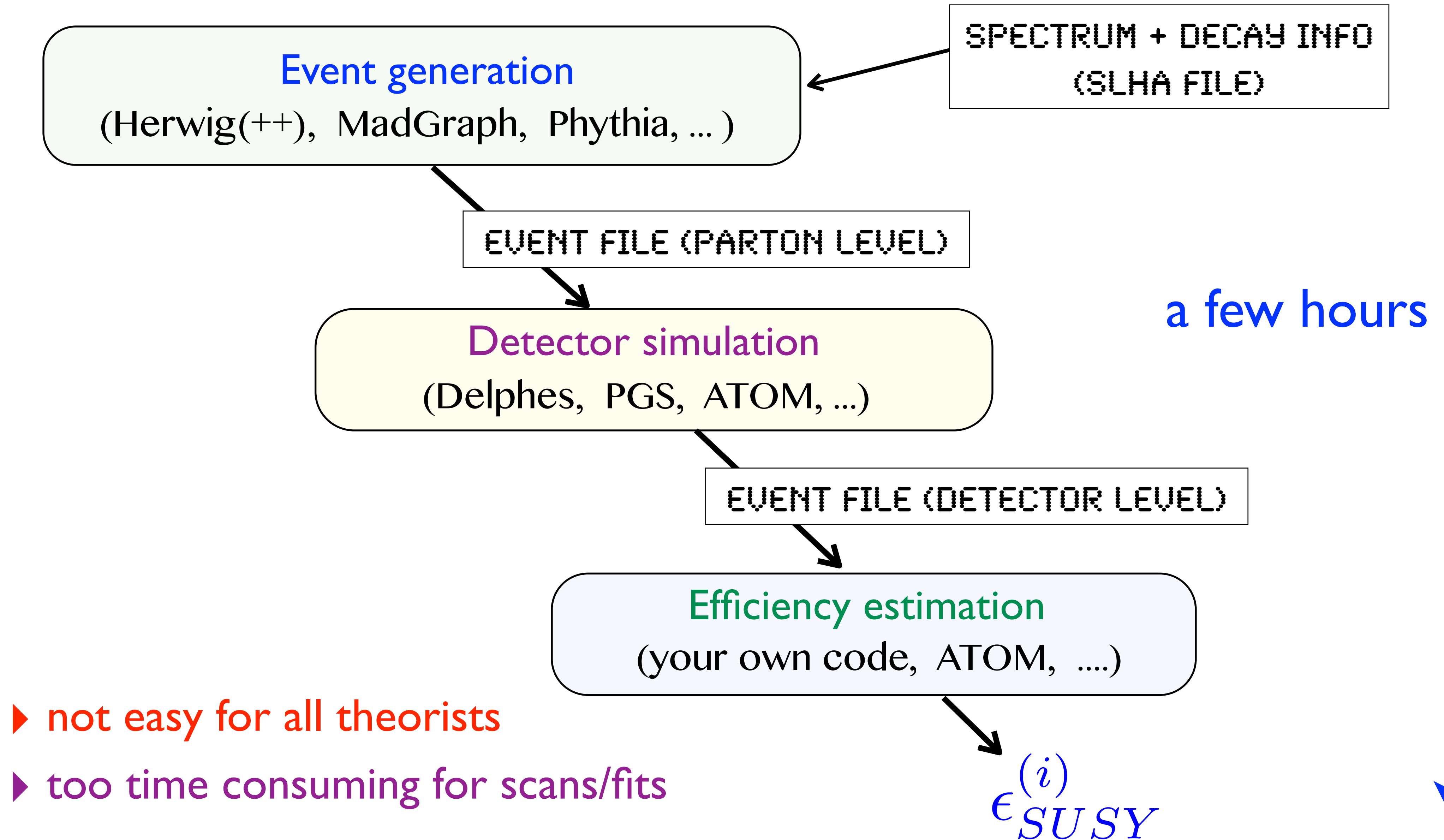
requires a chain of MC simulations

$$\epsilon_{\text{SUSY}}^{(i)} = \lim_{N_{\text{MC}}^{\text{gen.}} \rightarrow \infty} \frac{N_{SR}^{(i)} \left(\begin{array}{l} \text{Events fall into} \\ \text{Signal Region (i)} \end{array} \right)}{N_{MC}^{\text{gen.}}}$$

Chain of simulations



Chain of simulations



A fast evaluation of N_{SUSY}

- We propose a new approach to estimate N_{SUSY}

Key Idea: to reconstruct N_{SUSY} using simplified model processes

$$Q = \tilde{q}$$

$$G = \tilde{g}$$

$$N1 = \tilde{\chi}_1^0$$

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$$Q = \tilde{q}$$

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$$N1 = \tilde{\chi}_1^0$$

$$N_{\text{SUSY}}^{(i)} = \left\{ \begin{array}{l} N_{QqN1:QqN1} \\ + \\ N_{GqqN1:GqqN1} \\ + \\ N_{GqqN1:QqN1} \\ + \\ \vdots \end{array} \right.$$

A fast evaluation of N_{SUSY}

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$$N_{\text{SUSY}}^{(i)} = \left\{ \begin{array}{l} N_{QqN1:QqN1} = \mathfrak{E}_{QqN1:QqN1}(m_Q, m_{N1}) \cdot \sigma_{QQ} \cdot \text{BR}_{QqN1:QqN1} \cdot L_{\text{int}} \\ + \\ N_{GqqN1:GqqN1} = \mathfrak{E}_{GqqN1:GqqN1}(m_G, m_{N1}) \cdot \sigma_{GG} \cdot \text{BR}_{GqqN1:GqqN1} \cdot L_{\text{int}} \\ + \\ N_{GqqN1:QqN1} = \mathfrak{E}_{GqqN1:QqN1}(m_Q, m_G, m_{N1}) \cdot \sigma_{GQ} \cdot \text{BR}_{GqqN1:QqN1} \cdot L_{\text{int}} \\ + \\ \vdots \end{array} \right.$$

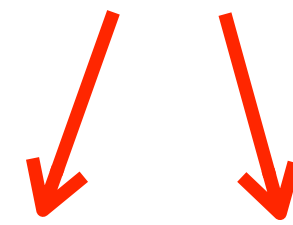
A fast evaluation of N_{SUSY}

$$Q = \tilde{q}$$

$$G = \tilde{g}$$

$$N1 = \tilde{\chi}_1^0$$

Efficiencies for simplified processes
depend only on a few mass parameters



$$N_{\text{SUSY}}^{(i)} = \left\{ \begin{array}{l} N_{QqN1:QqN1} = \epsilon_{QqN1:QqN1}(mQ, mN1) \cdot \sigma_{QQ} \cdot \text{BR}_{QqN1:QqN1} \cdot L_{\text{int}} \\ + \\ N_{GqqN1:GqqN1} = \epsilon_{GqqN1:GqqN1}(mG, mN1) \cdot \sigma_{GG} \cdot \text{BR}_{GqqN1:GqqN1} \cdot L_{\text{int}} \\ + \\ N_{GqqN1:QqN1} = \epsilon_{GqqN1:QqN1}(mQ, mG, mN1) \cdot \sigma_{GQ} \cdot \text{BR}_{GqqN1:QqN1} \cdot L_{\text{int}} \\ + \\ \vdots \end{array} \right.$$

A fast evaluation of N_{SUSY}

$$Q = \tilde{q}$$

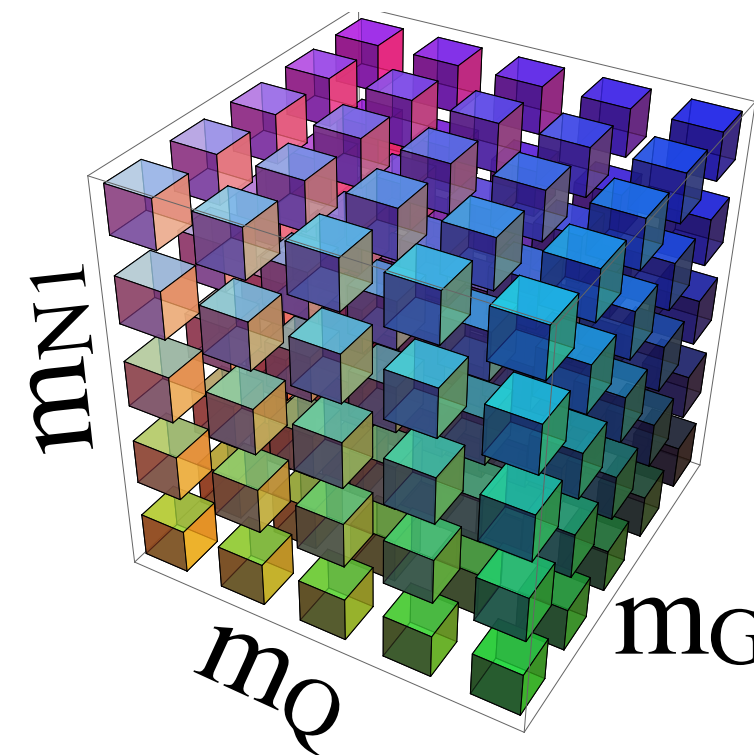
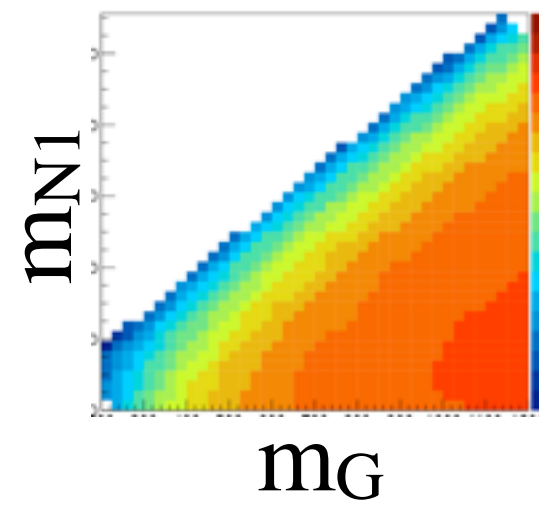
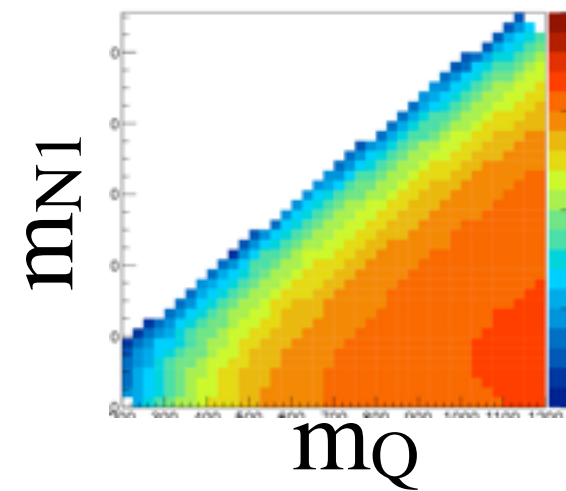
$$G = \tilde{g}$$

$$N1 = \tilde{\chi}_1^0$$

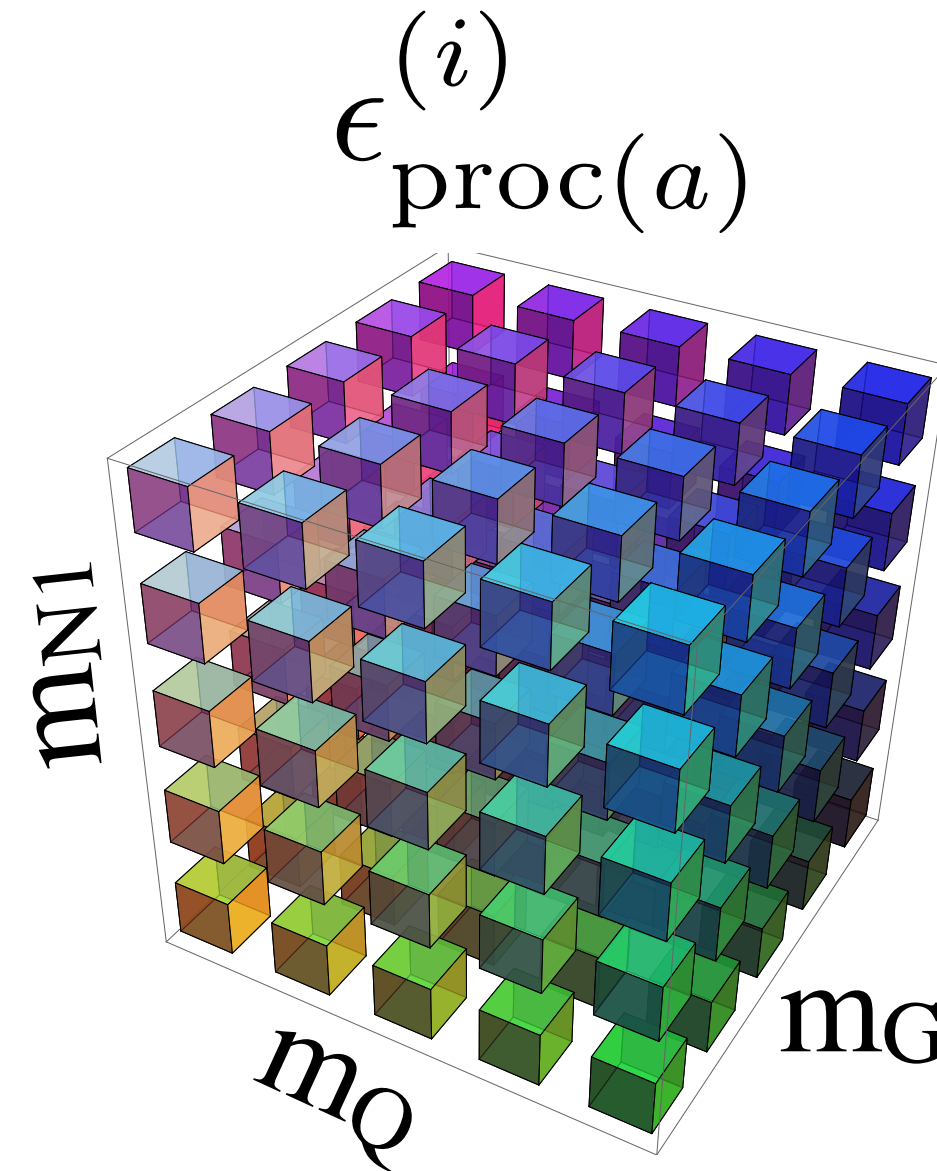
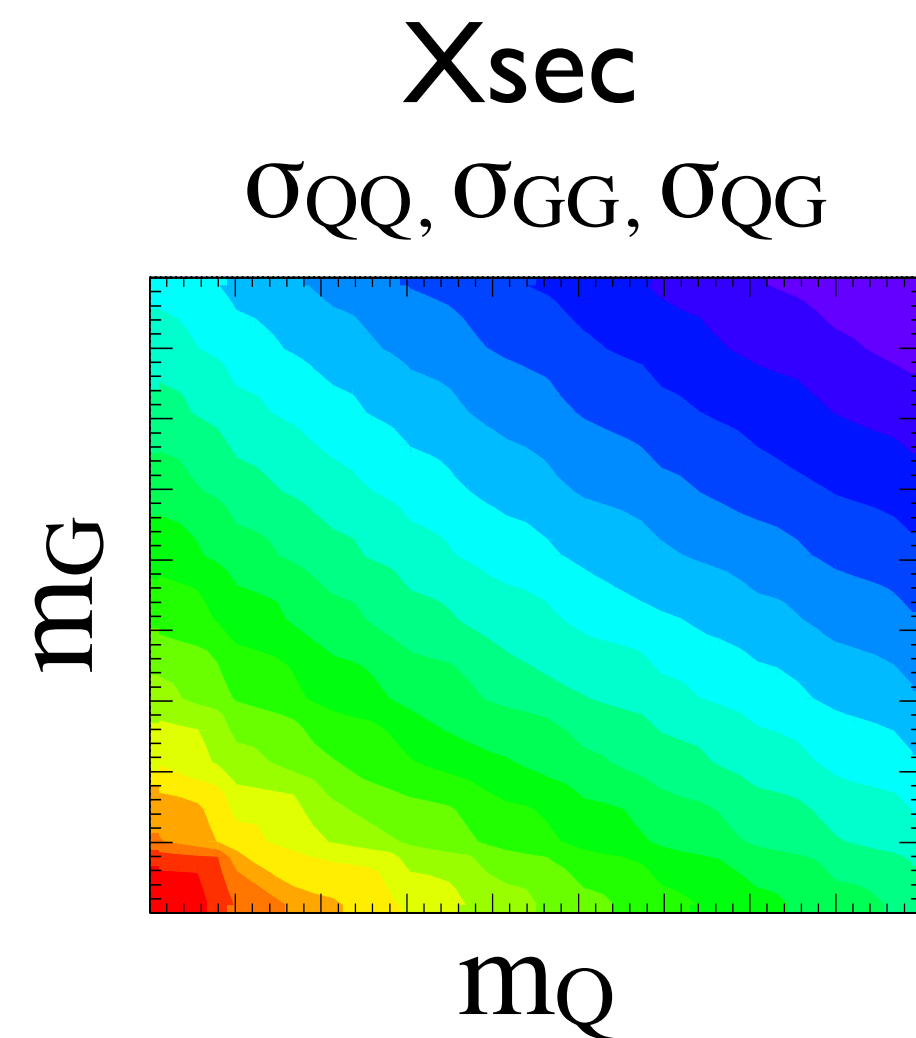
- Once one has the efficiency tables for the simplified model processes, one can read off the efficiencies and re-assemble N_{SUSY} of your model.

no MC simulation is required !

$$N_{\text{SUSY}}^{(i)} = \left\{ \begin{array}{l} N_{QqN1:QqN1} = \\ \quad + \\ N_{GqqN1:GqqN1} = \\ \quad + \\ N_{GqqN1:QqN1} = \\ \quad + \\ \vdots \end{array} \right. \cdot \sigma_{QQ} \cdot \text{BR}_{QqN1:QqN1} \cdot L_{\text{int}} \\ \cdot \sigma_{GG} \cdot \text{BR}_{GqqN1:GqqN1} \cdot L_{\text{int}} \\ \cdot \sigma_{GQ} \cdot \text{BR}_{GqqN1:QqN1} \cdot L_{\text{int}}$$



FastLim



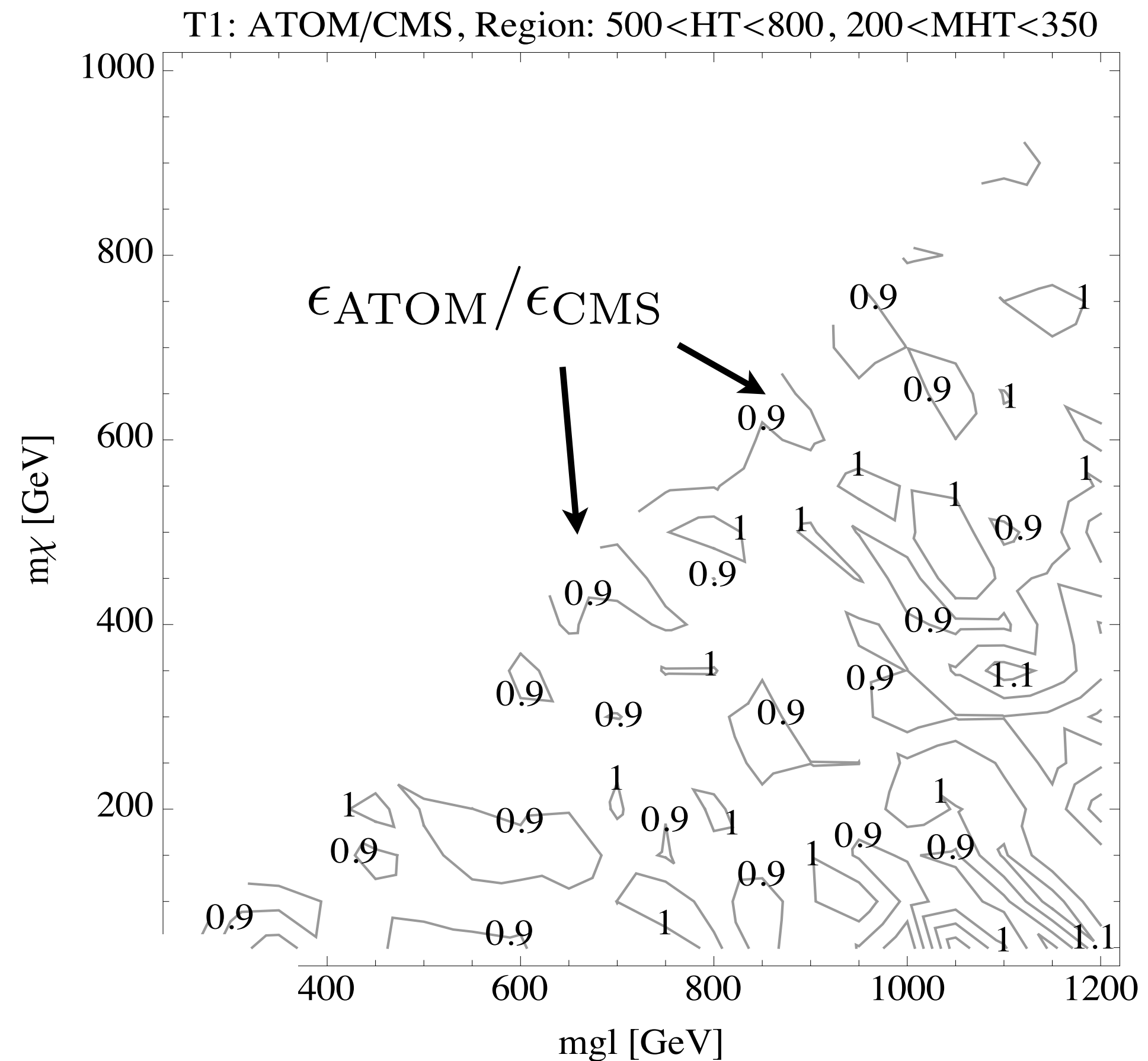
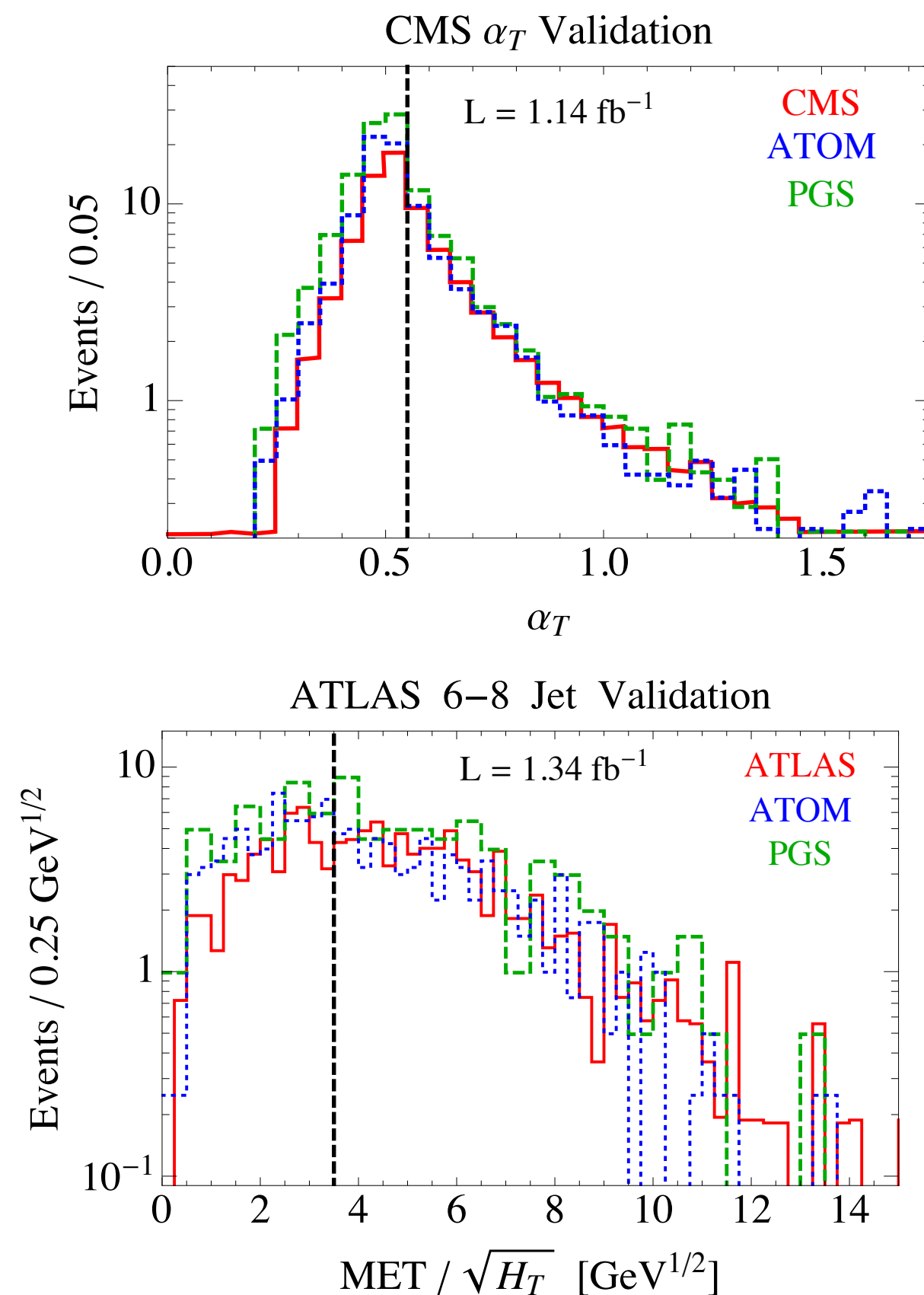
pre-calculated **cross section tables**
calculated by **NLLfast**

Prospino2.1 is interfaced
for the **EWkino cross section**

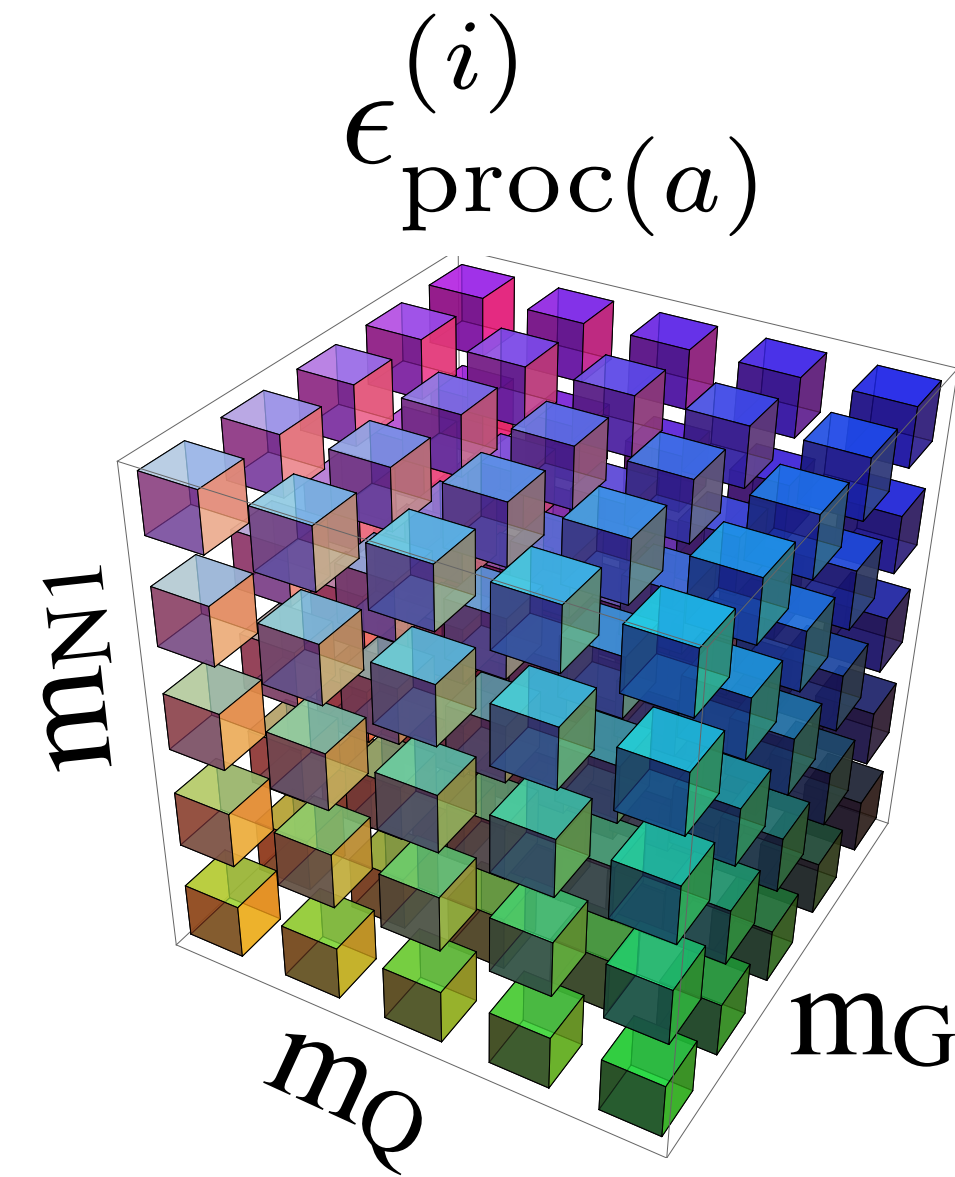
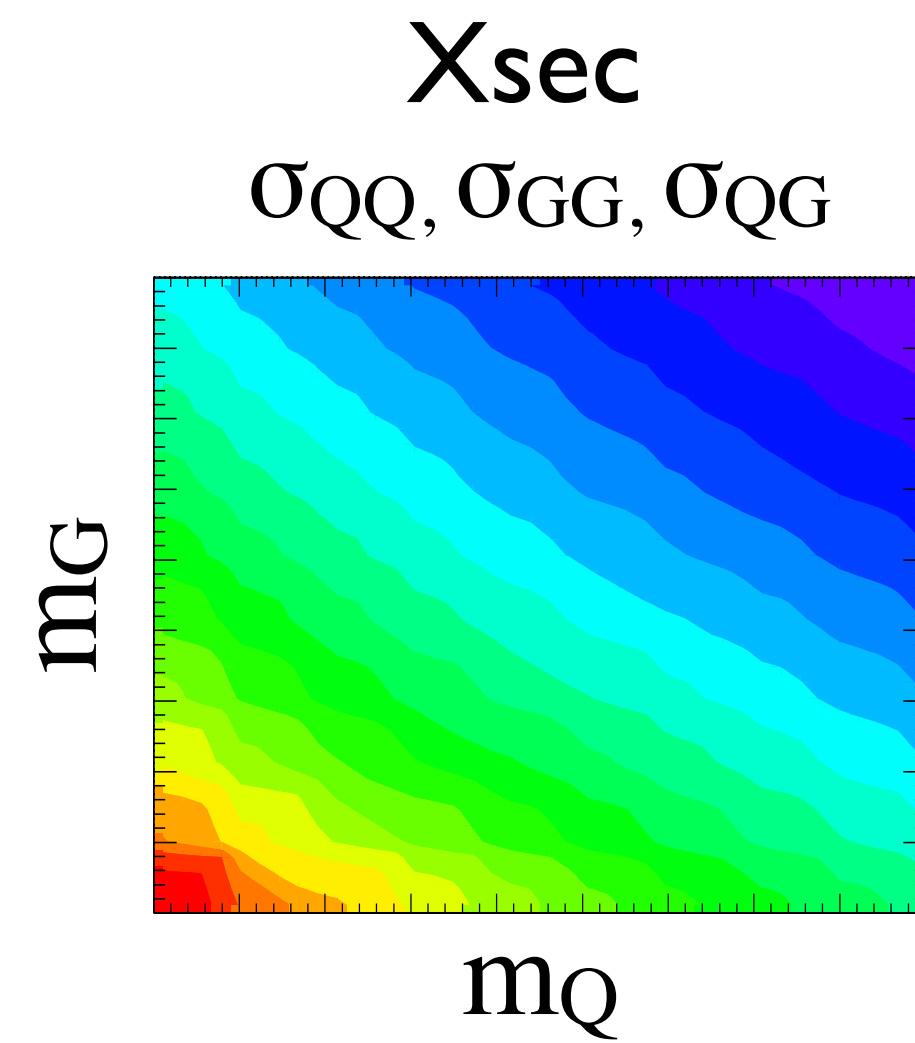
efficiency tables for the simplified processes
taken from ATLAS/CMS if available
otherwise calculated by **MadGraph + ATOM**

(Automated Testing Of Models)

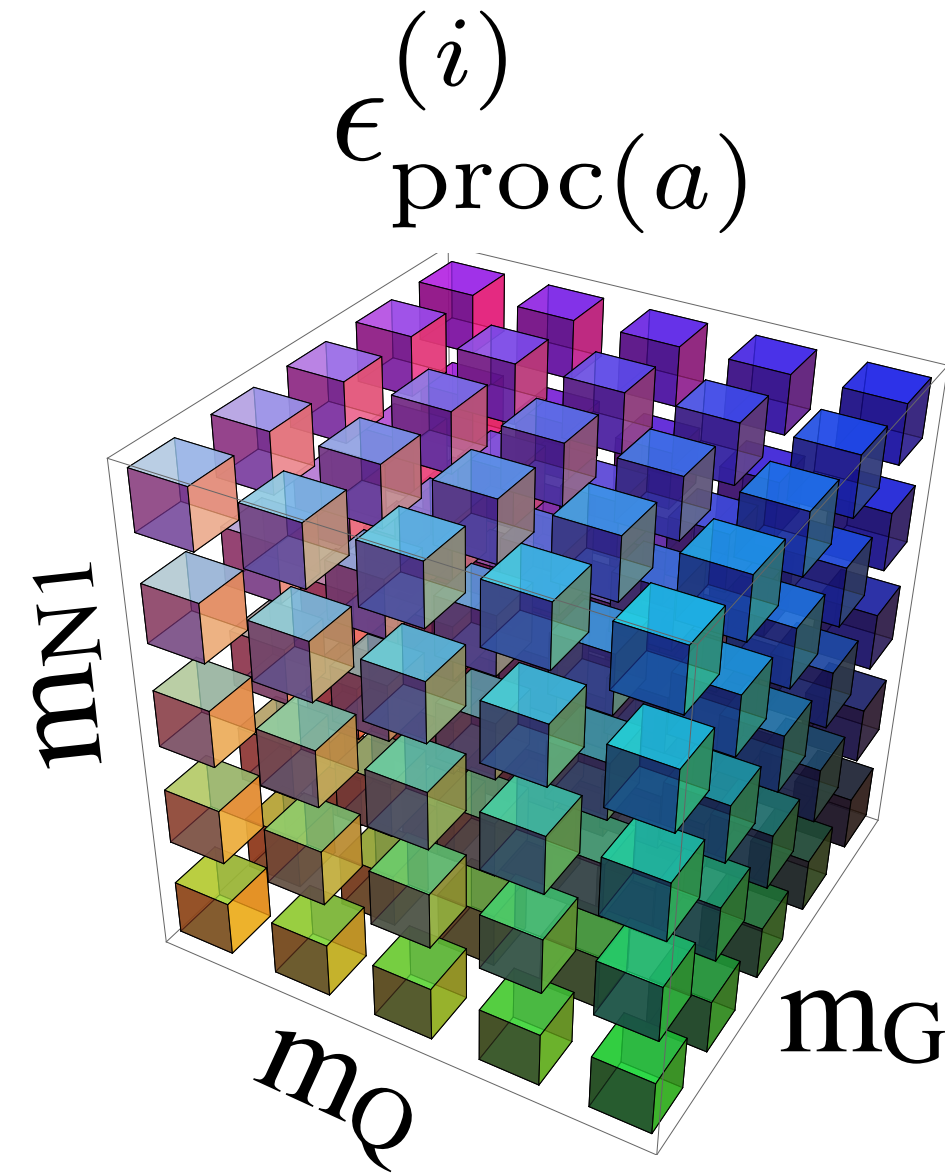
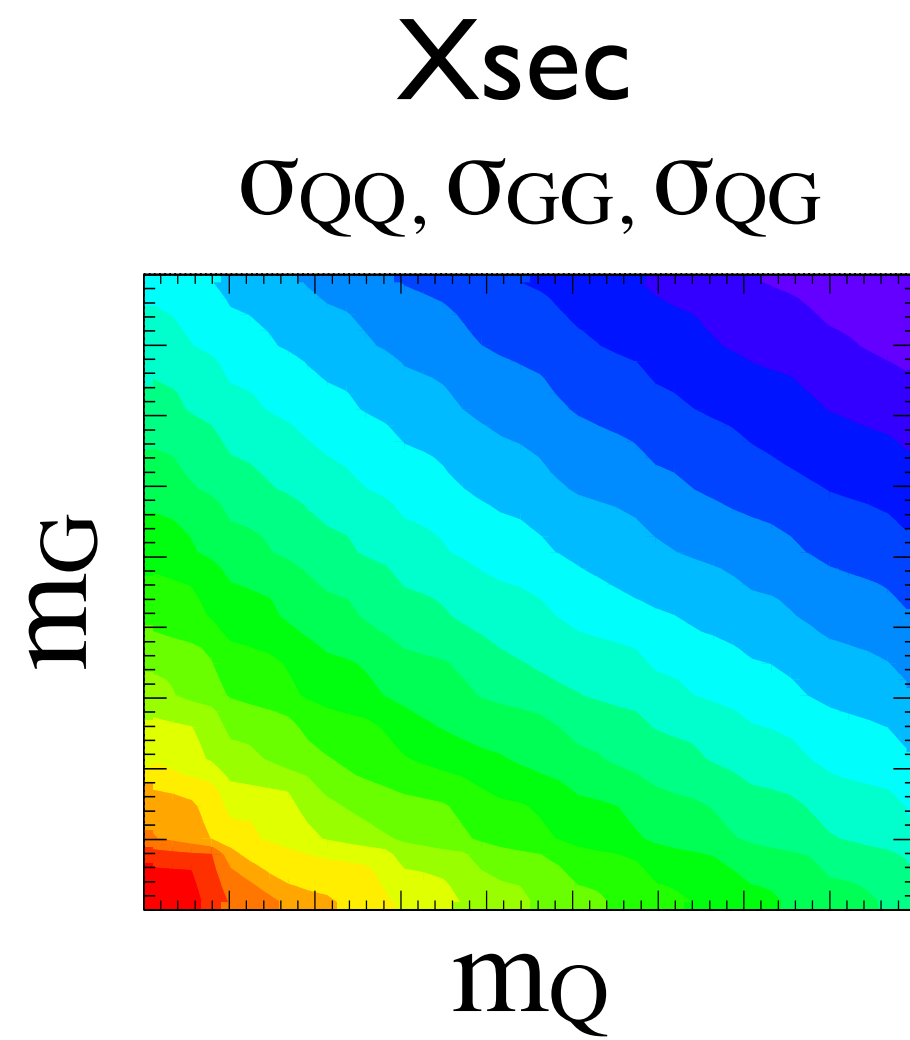
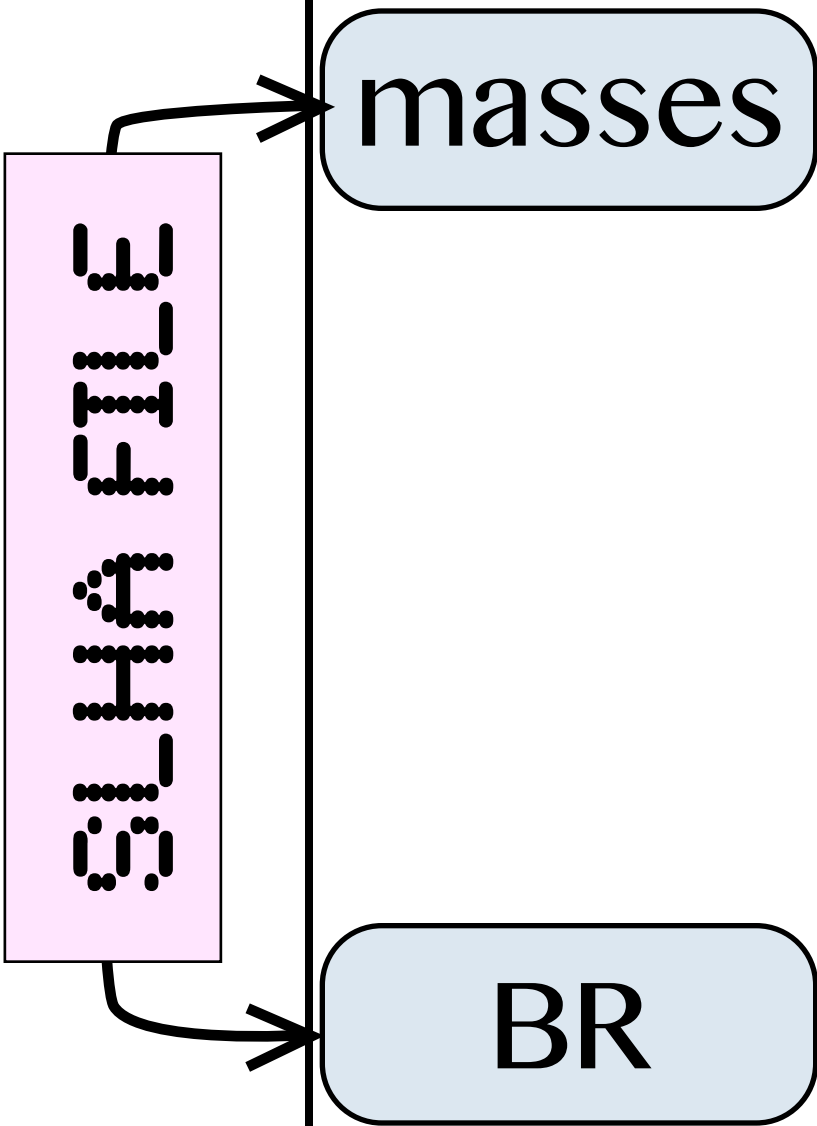
- ATOM: a program to calculate the efficiencies including detector effects
- well validated and reliable



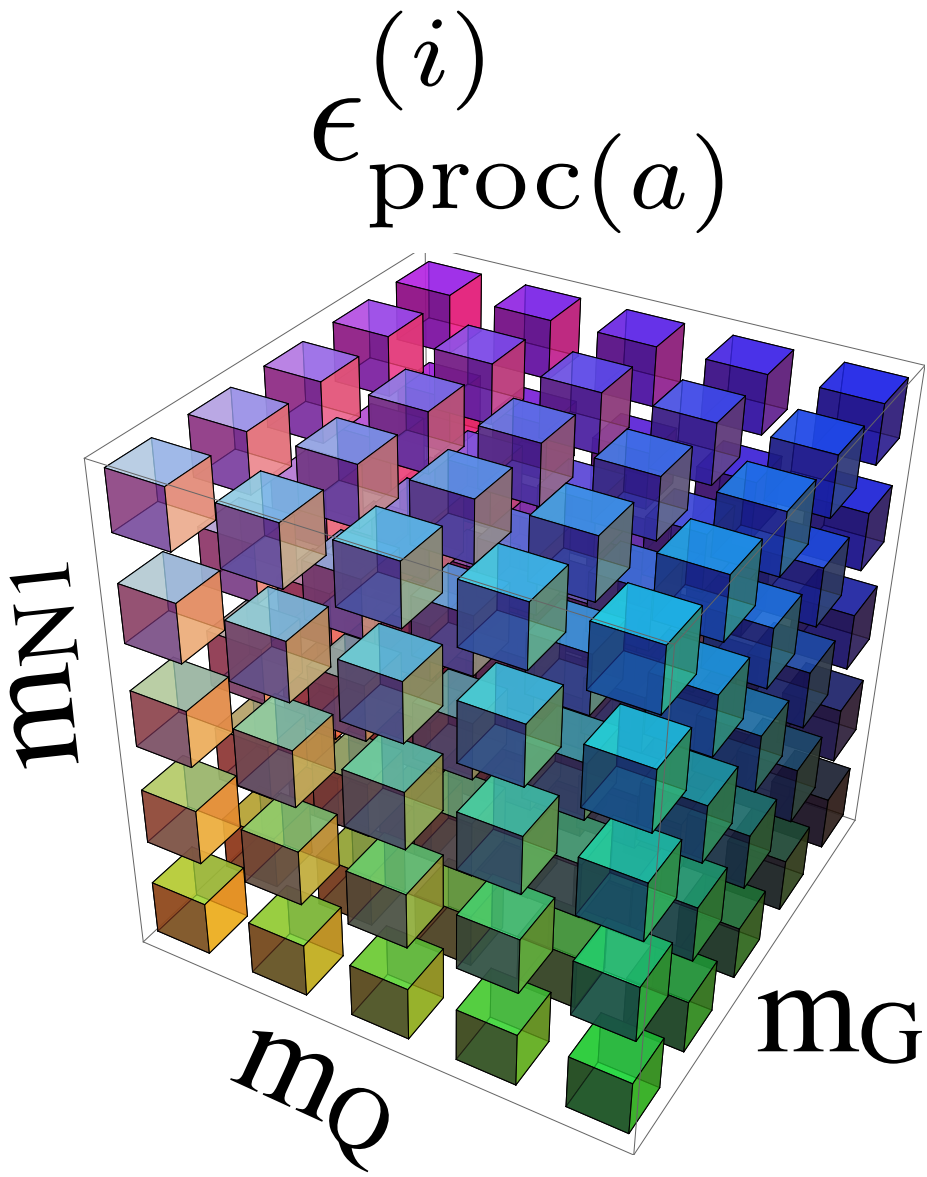
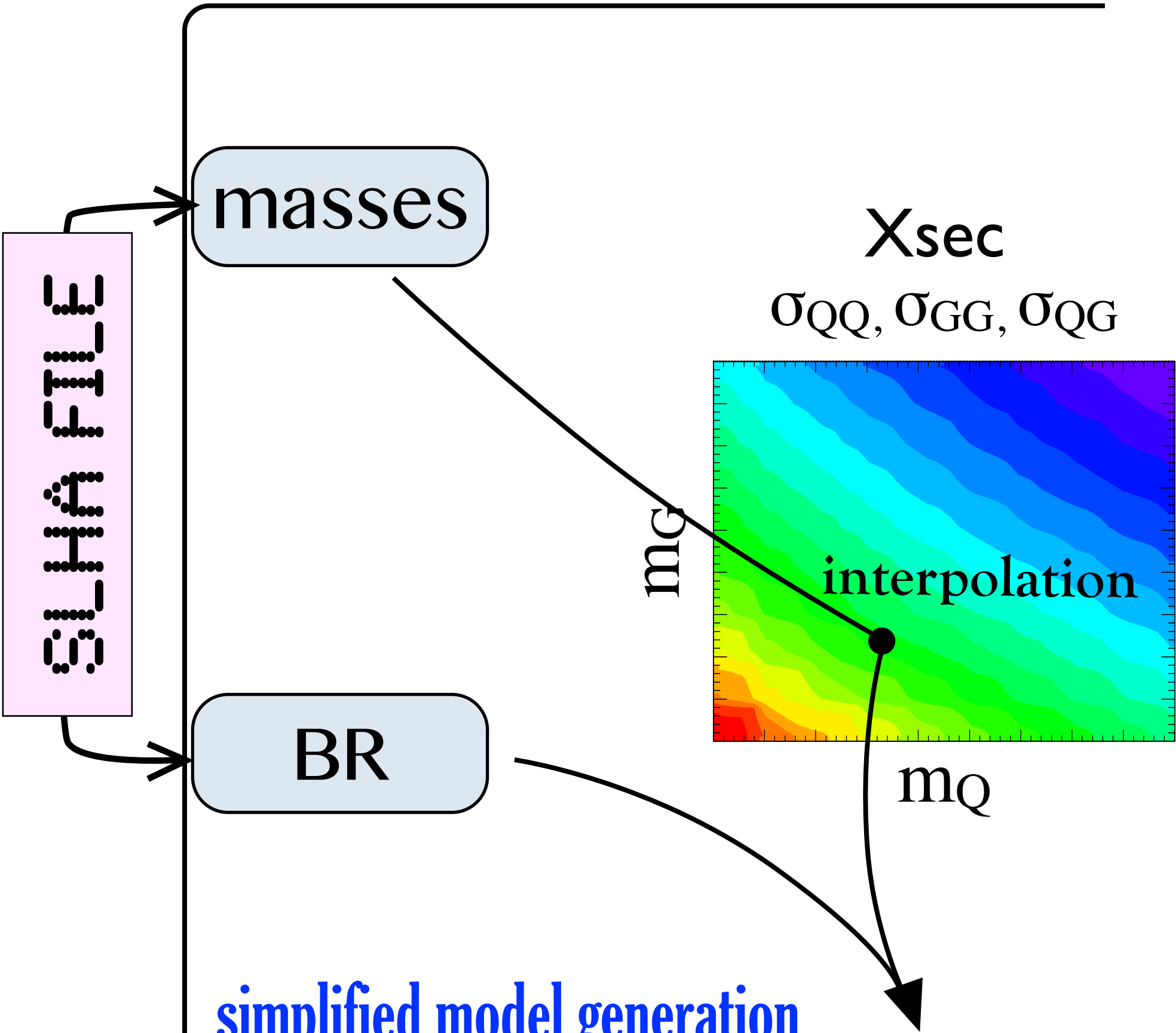
FastLim



FastLim

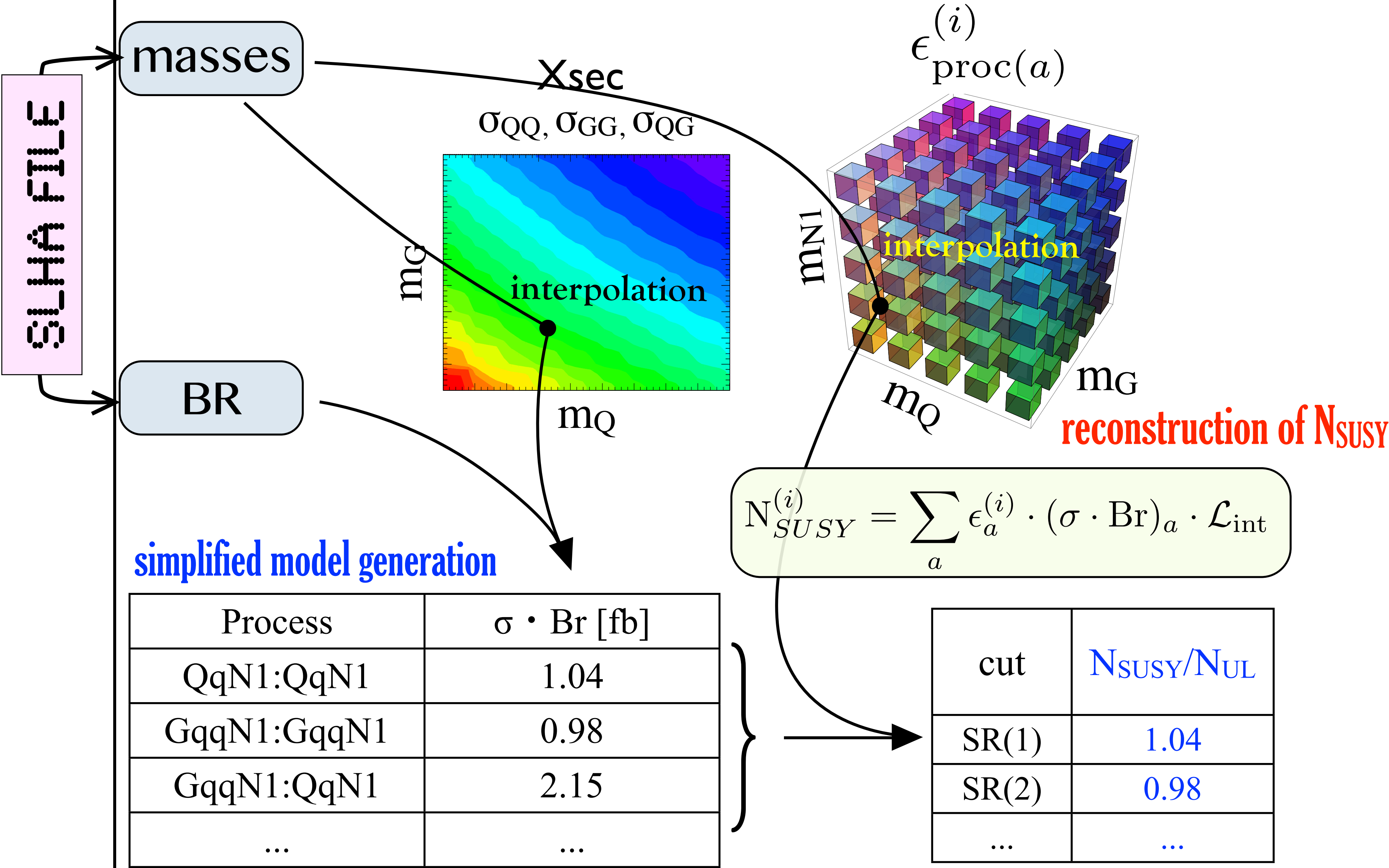


FastLim

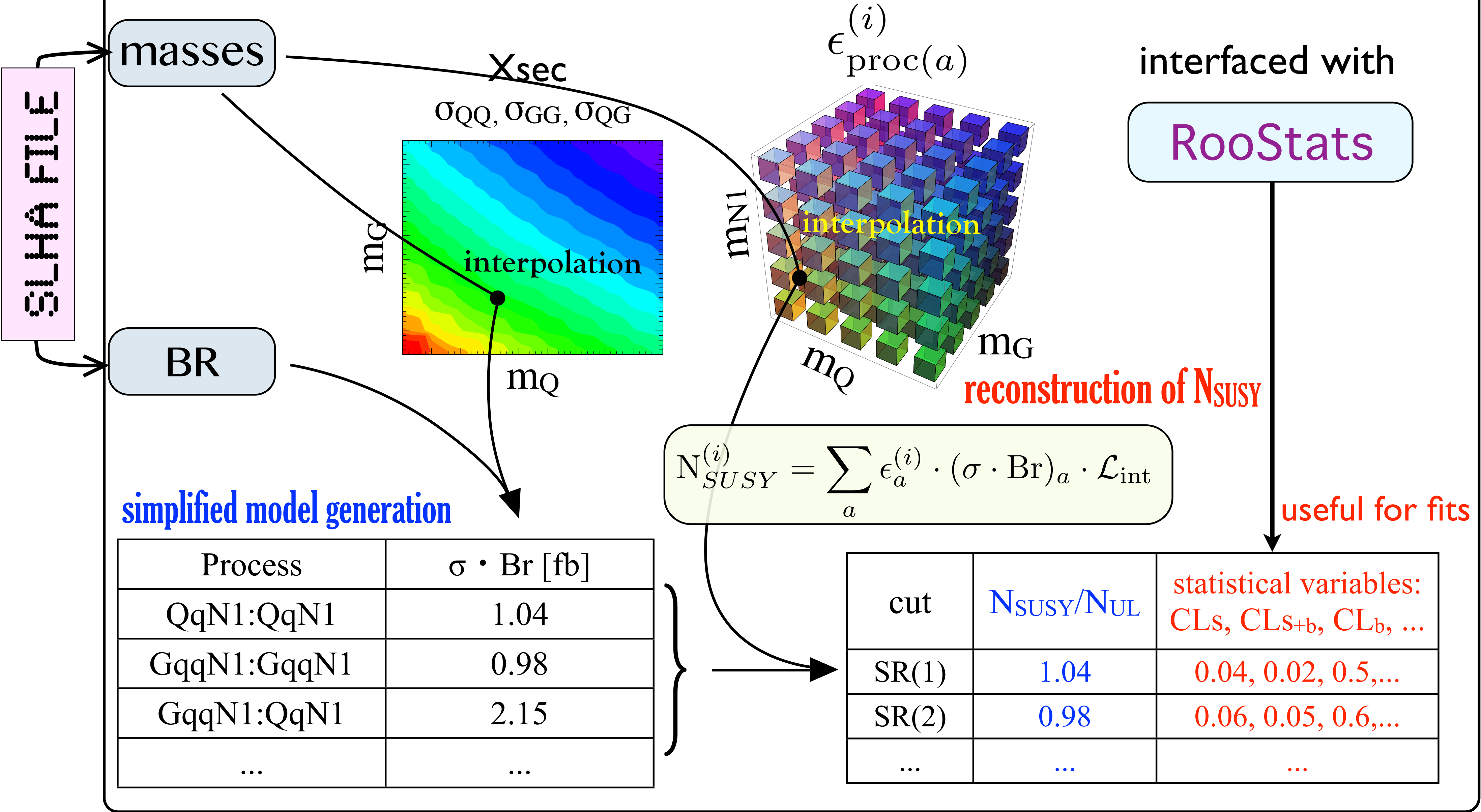


Process	$\sigma \cdot \text{Br} \text{ [fb]}$
QqN1:QqN1	1.04
GqqN1:GqqN1	0.98
GqqN1:QqN1	2.15
...	...

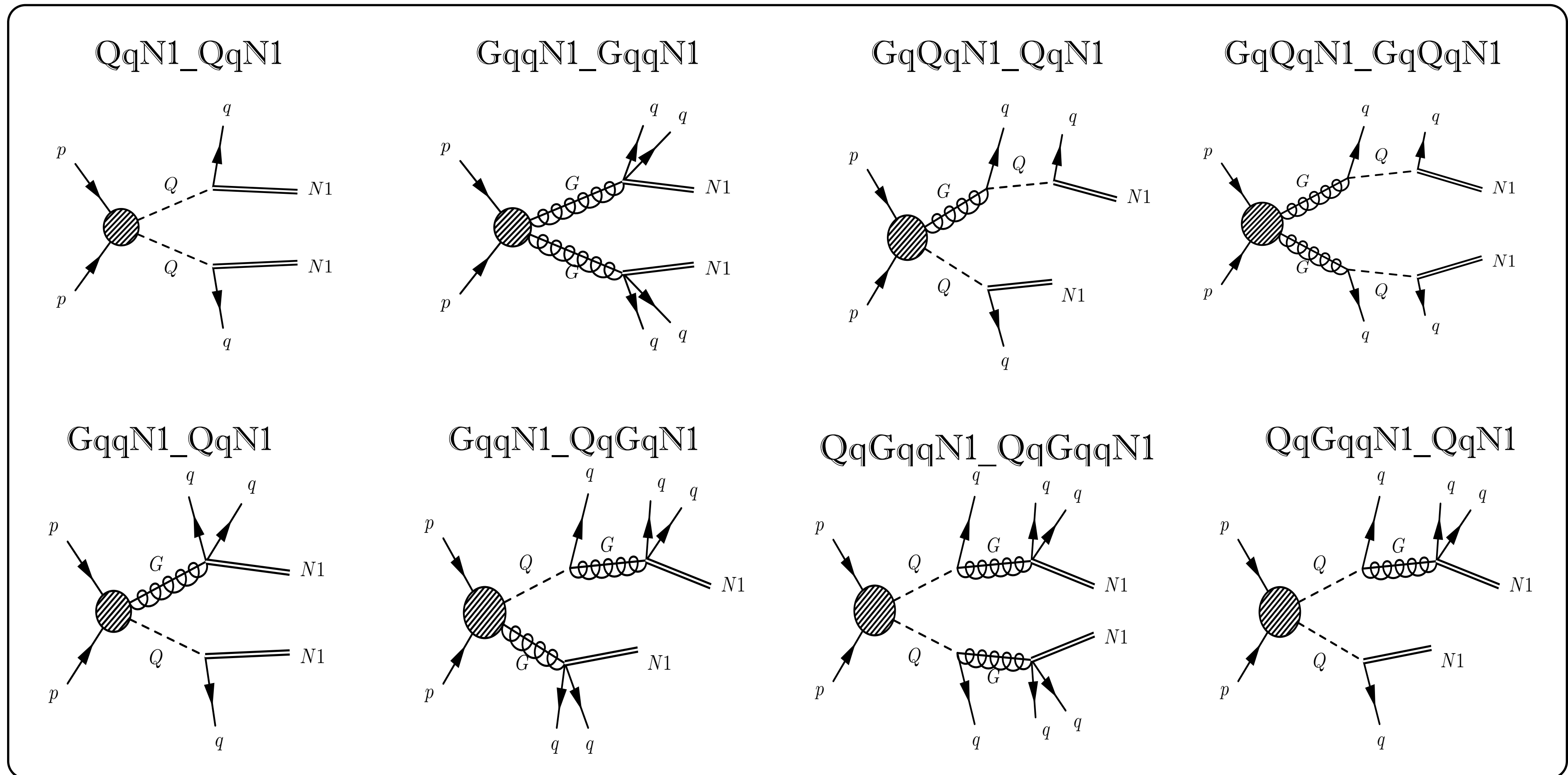
FastLim



FastLim



- A complete set of Q-G-N1 model (w/o top) has been implemented



The other processes are missing ➡ Reconstructed N_{SUSY} is conservative

- Many analyses have been implemented (thanks to ATOM)

Name	E _{cm}	Short description
ATLAS_2011_CONF_2011_086	7	Jets+MET at 7 TeV with 165pb ⁻¹ .
ATLAS_2011_CONF_2011_090	7	1 lepton+jets+MET at 7 TeV with 165pb ⁻¹ .
ATLAS_2011_CONF_2011_098	7	bjets+MET+0L at 7 TeV with 830pb ⁻¹ .
ATLAS_2011_CONF_2011_126	7	Search for Anomalous Production of Prompt Like-sign Muon Pairs with 1.6 fb ⁻¹ .
ATLAS_2011_CONF_2011_130	7	bjets+1lept+jets+MET SUSY search at 7TeV with 1fb ⁻¹
ATLAS_2011_S8970084	7	1 lepton+jets+MET at 7 TeV with 35pb ⁻¹ .
ATLAS_2011_S8983313	7	Jets+MET at 7 TeV with 35pb ⁻¹ .
ATLAS_2011_S9011218	7	bjets+MET at 7 TeV with 35pb ⁻¹ .
ATLAS_2011_S9019553	7	SF lepton pairs SUSY search at 7 TeV with 35pb ⁻¹ .
ATLAS_2011_S9019561	7	2leptons+MET at 7TeV with 35pb ⁻¹ .
ATLAS_2011_S9225137	7	multijet SUSY search at 7TeV
ATLAS_2012_CONF_2012_033	7	2-6 jets + MET SUSY search at 7TeV
CMS_2011_S8932190	7	Jets+MET with alpha _T variable with 35pb ⁻¹
CMS_2011_S8991847	7	OS dileptons at 7TeV with 35pb ⁻¹
CMS_2011_S9036504	7	Same Sign dileptons at 7TeV in 35pb ⁻¹
CMS_PAS_SUS_10_005	7	HT,MHT susy search in jets+MET at 7 TeV with 35pb ⁻¹ .
CMS_PAS_SUS_10_009	7	razor analysis on jets+MET and 1lepton+jets+MET at 7 TeV with 35pb ⁻¹ .
CMS_PAS_SUS_10_011	7	alpha _T analysis on b jets+MET at 7 TeV with 35pb ⁻¹ .
CMS_PAS_SUS_11_003	7	Jets+MET with alpha _T variable with 1.1 fb ⁻¹
CMS_PAS_SUS_11_017	7	Search for New Physics in Events with a Z Boson and Missing Transverse Energy
ATLAS_2012_CONF_2012_109	8	2-6 jets + MET SUSY search at 8TeV
CMS_PAS_SUS_12_028	8	CMS 8 TeV analysis

FastLim demonstration

Applications

Because FastLim contains a finite set of simplified processes, we would like to check how well the code covers the interesting models:

- CMSSM
- NUHM
- natural SUSY
- spread SUSY

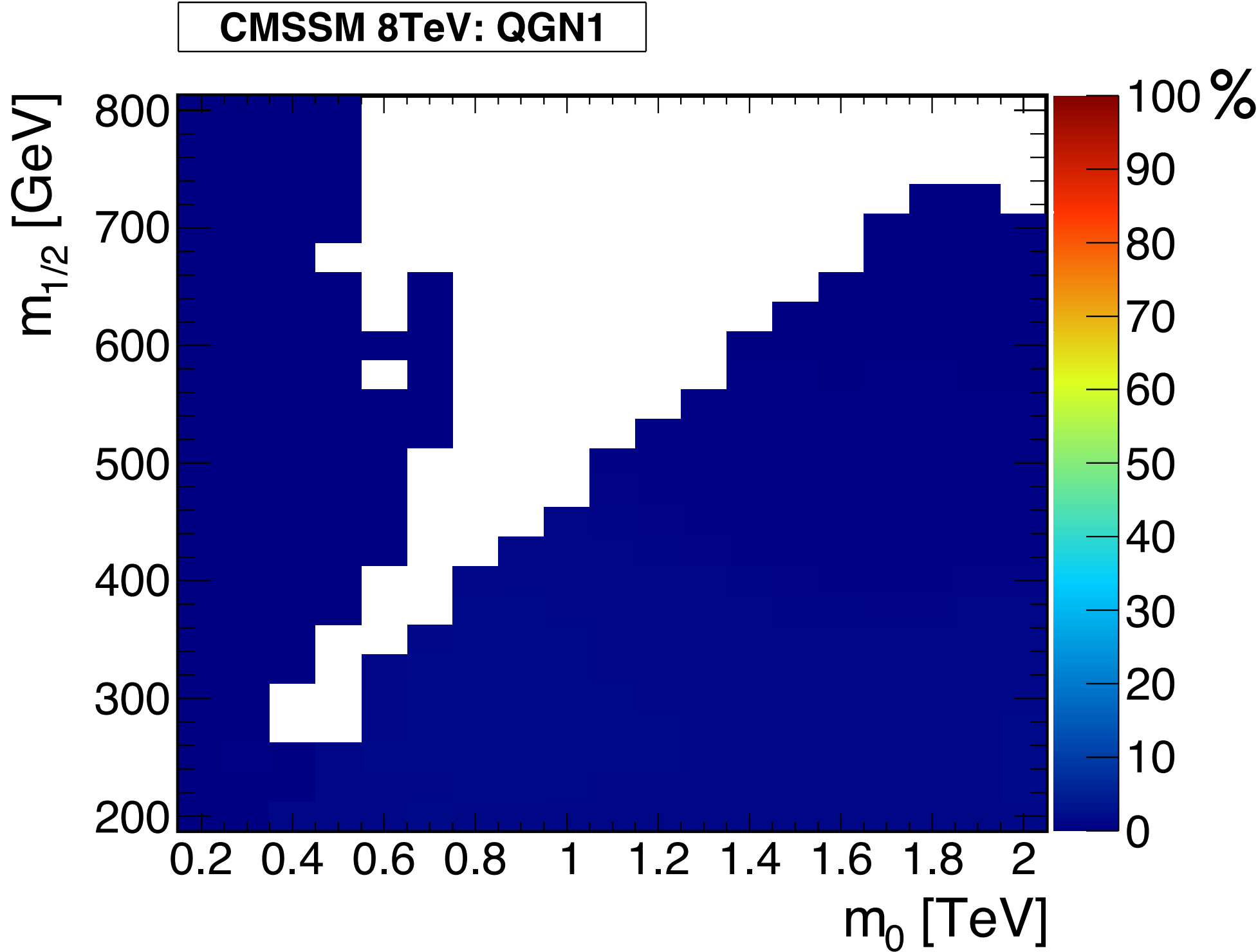
CMSSM coverage

$$\begin{aligned} Q &= \tilde{q} \\ G &= \tilde{g} \\ C1 &= \tilde{\chi}_1^\pm \\ N2 &= \tilde{\chi}_2^0 \\ N1 &= \tilde{\chi}_1^0 \end{aligned}$$

$$\text{Coverage} = \frac{\sum_i \sigma_i \left(\begin{smallmatrix} \text{implemented} \\ \text{processes} \end{smallmatrix} \right)}{\sigma_{\text{tot}}}$$

(only Q, G and N1 can appear in the process)

Q-G-N1 model

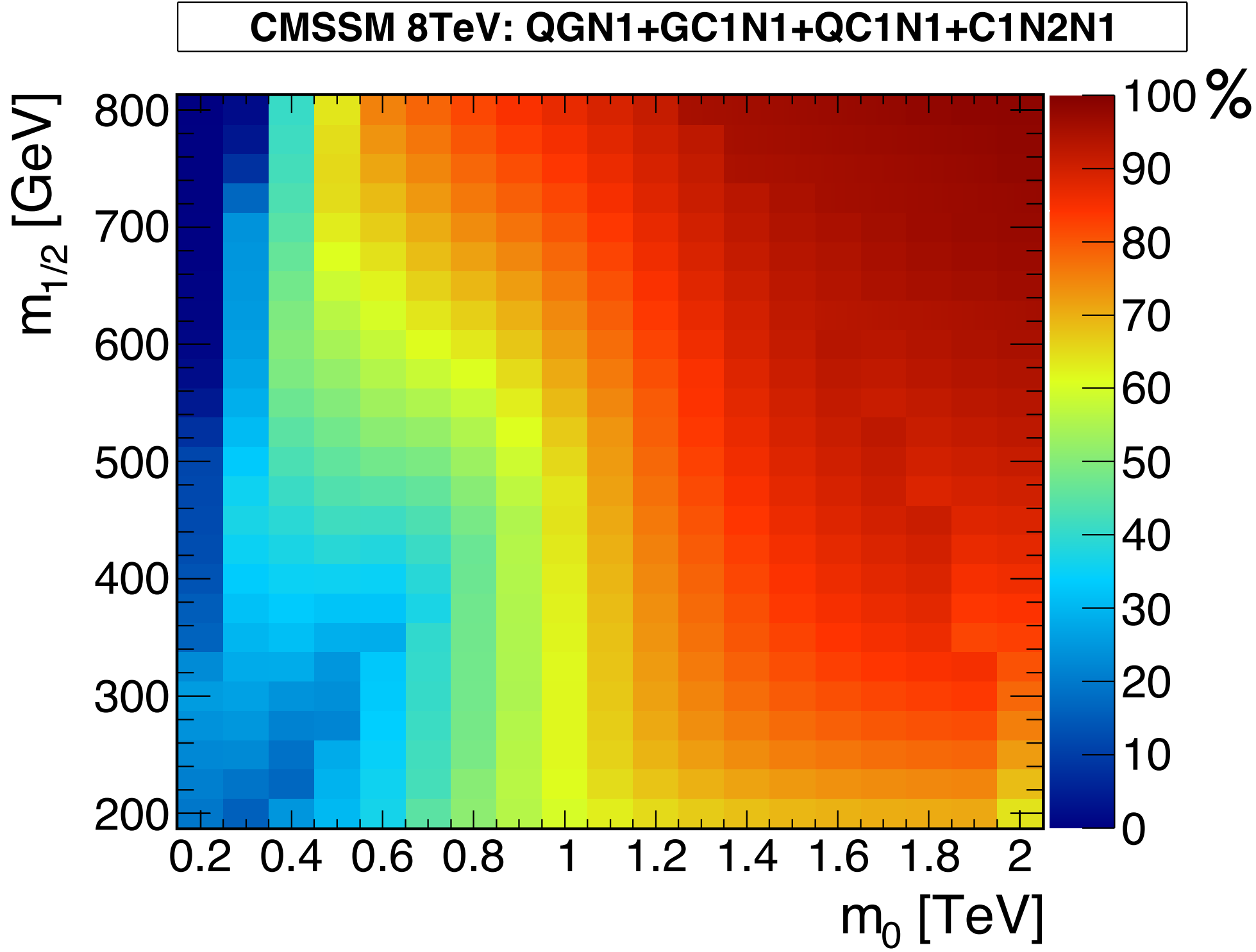


GttNI_QqGttNI	GbbNI_QqGqqNI
GqqNI_QqGqqNI	QqGbbNI_QqGbbNI
GttNI_QqNI	QqGttNI_QqGttNI
GbbNI_GbbNI	GbbNI_QqNI
GbbNI_QqGttNI	GttNI_GttNI
QqGttNI_QqNI	QqGbbNI_QqGttNI
GbbNI_GttNI	QqGbbNI_QqGqqNI
QqGqqNI_QqGqqNI	GqQqNI_GqQqNI
GbbNI_GqqNI	GqqNI_QqGttNI
GttNI_QqGqqNI	GttNI_QqGbbNI
GqQqNI_QqNI	QqNI_QqNI
GqqNI_GttNI	QqGqqNI_QqNI
GqqNI_QqGbbNI	GbbNI_QqGbbNI
QqGbbNI_QqNI	QqGqqNI_QqGttNI
GqqNI_GqqNI	GqqNI_QqNI

CMSSM coverage

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Q-G-N1 + G(Q)-C1-N1 + C1-N2-N1 model

GttNI_QqGttNI	GbbNI_QqGqqNI	GqqClwNI_GttNI
GqqNI_QqGqqNI	QqGbbNI_QqGbbNI	GbtClwNI_GqqClwNI
GttNI_QqNI	QqGttNI_QqGttNI	GbbNI_GqqClwNI
GbbNI_GbbNI	GbbNI_QqNI	GqqClwNI_GqqNI
GbbNI_QqGttNI	GttNI_GttNI	GbtClwNI_GttNI
QqGttNI_QqNI	QqGbbNI_QqGttNI	GbbNI_GbtClwNI
GbbNI_GttNI	QqGbbNI_QqGqqNI	GqqClwNI_GqqClwNI
QqGqqNI_QqGqqNI	GqQqNI_GqQqNI	GbtClwNI_GqqNI
GbbNI_GqqNI	GqqNI_QqGttNI	GbtClwNI_GbtClwNI
GttNI_QqGqqNI	GttNI_QqGbbNI	QqClwNI_QqNI
GqQqNI_QqNI	QqNI_QqNI	QqClwNI_QqClwNI
GqqNI_GttNI	QqGqqNI_QqNI	ClwNI_N2zNI
GqqNI_QqGbbNI	GbbNI_QqGbbNI	ClwNI_ClwNI
QqGbbNI_QqNI	QqGqqNI_QqGttNI	ClwNI_N2h0NI
GqqNI_GqqNI	GqqNI_QqNI	

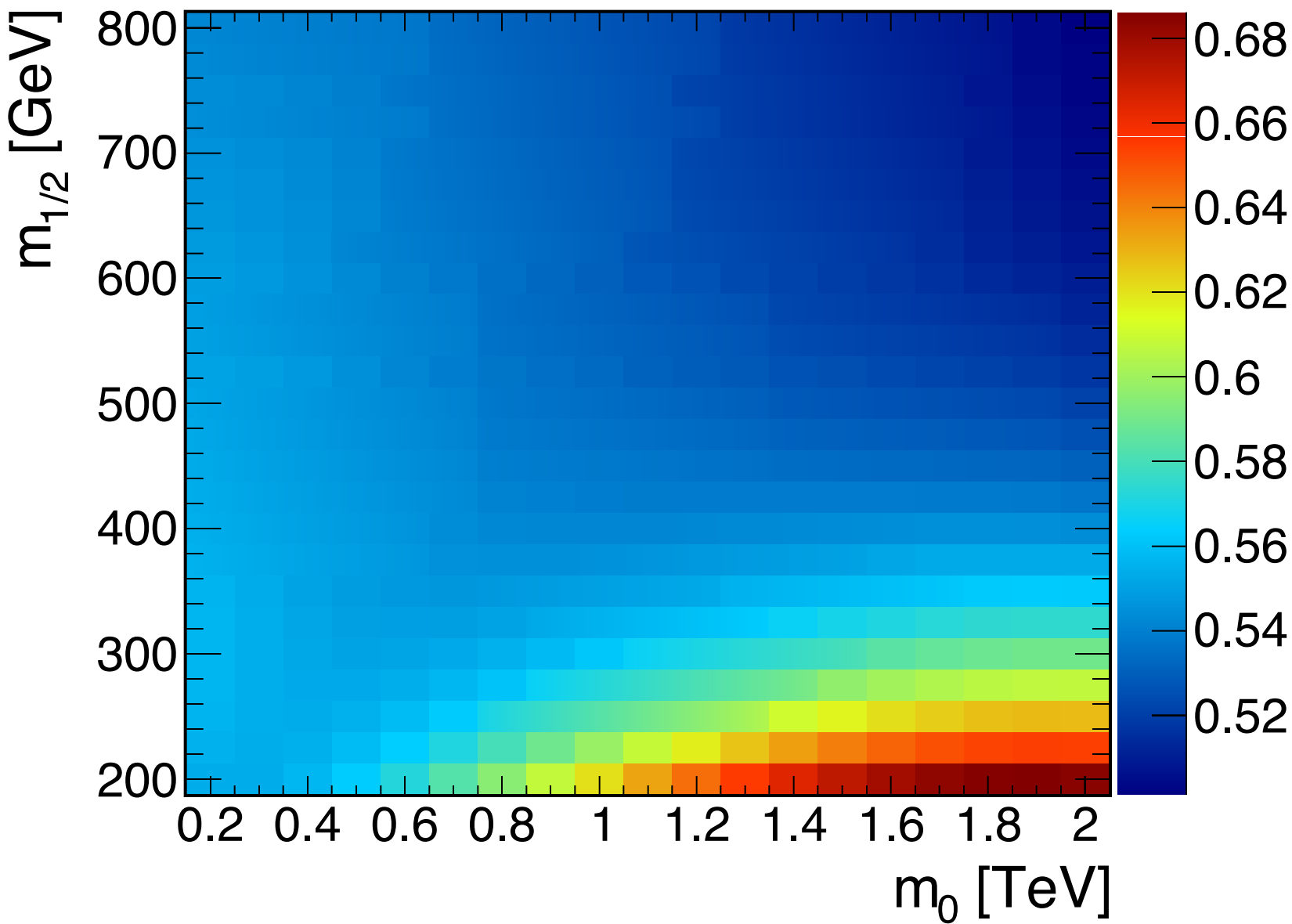
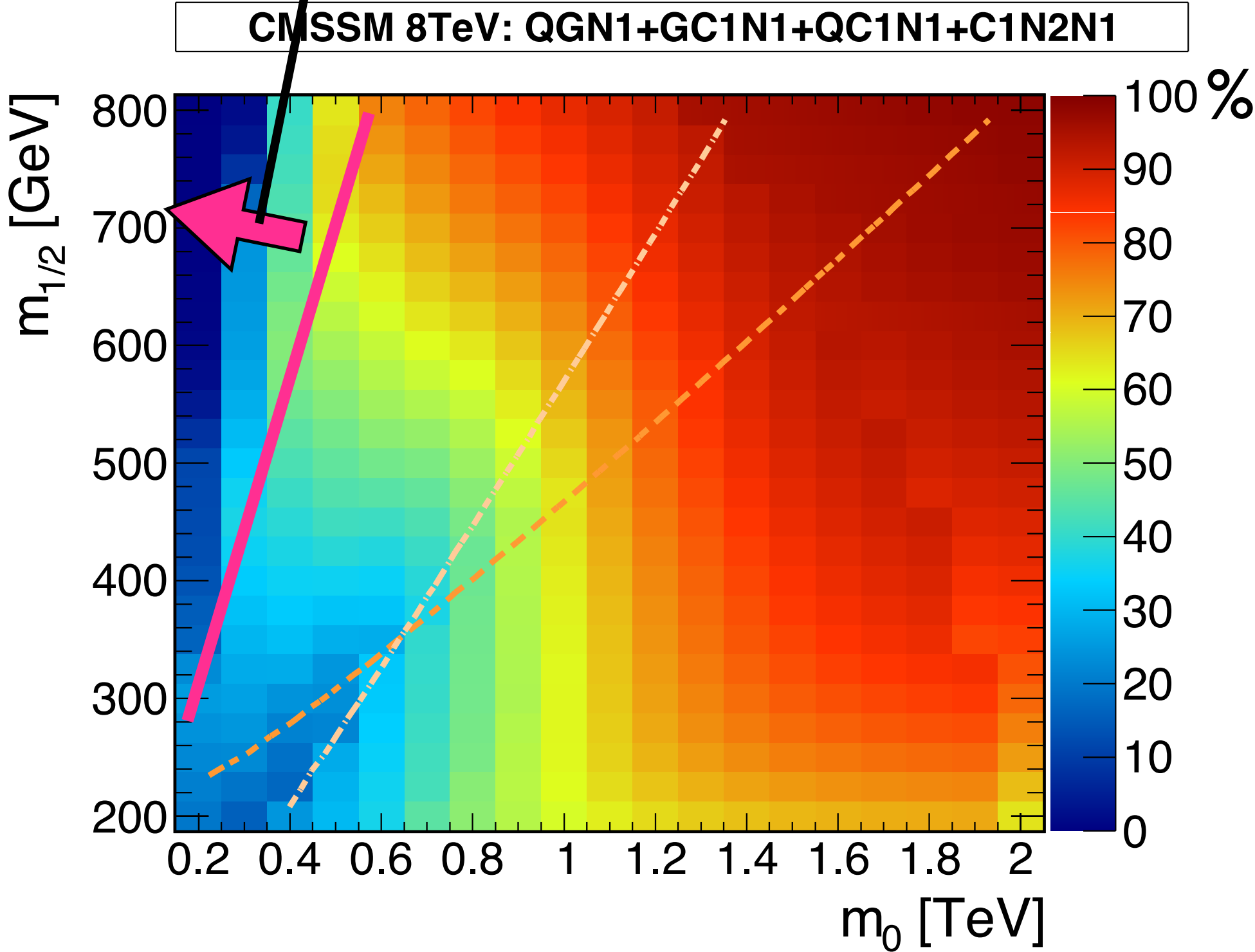
CMSSM coverage

$Q = \tilde{q}$
 $G = \tilde{g}$
 $C1 = \tilde{\chi}_1^\pm$
 $N2 = \tilde{\chi}_2^0$
 $N1 = \tilde{\chi}_1^0$

gluino decay modes to the higgsino states
dominate due to the top Yukawa coupling
and that makes decay chains longer

$\tilde{\chi}_1^+ \rightarrow \nu \tilde{\tau}^+$ open
 $\tilde{\chi}_1^+ \rightarrow W^+ \tilde{\chi}_1^0$ suppressed

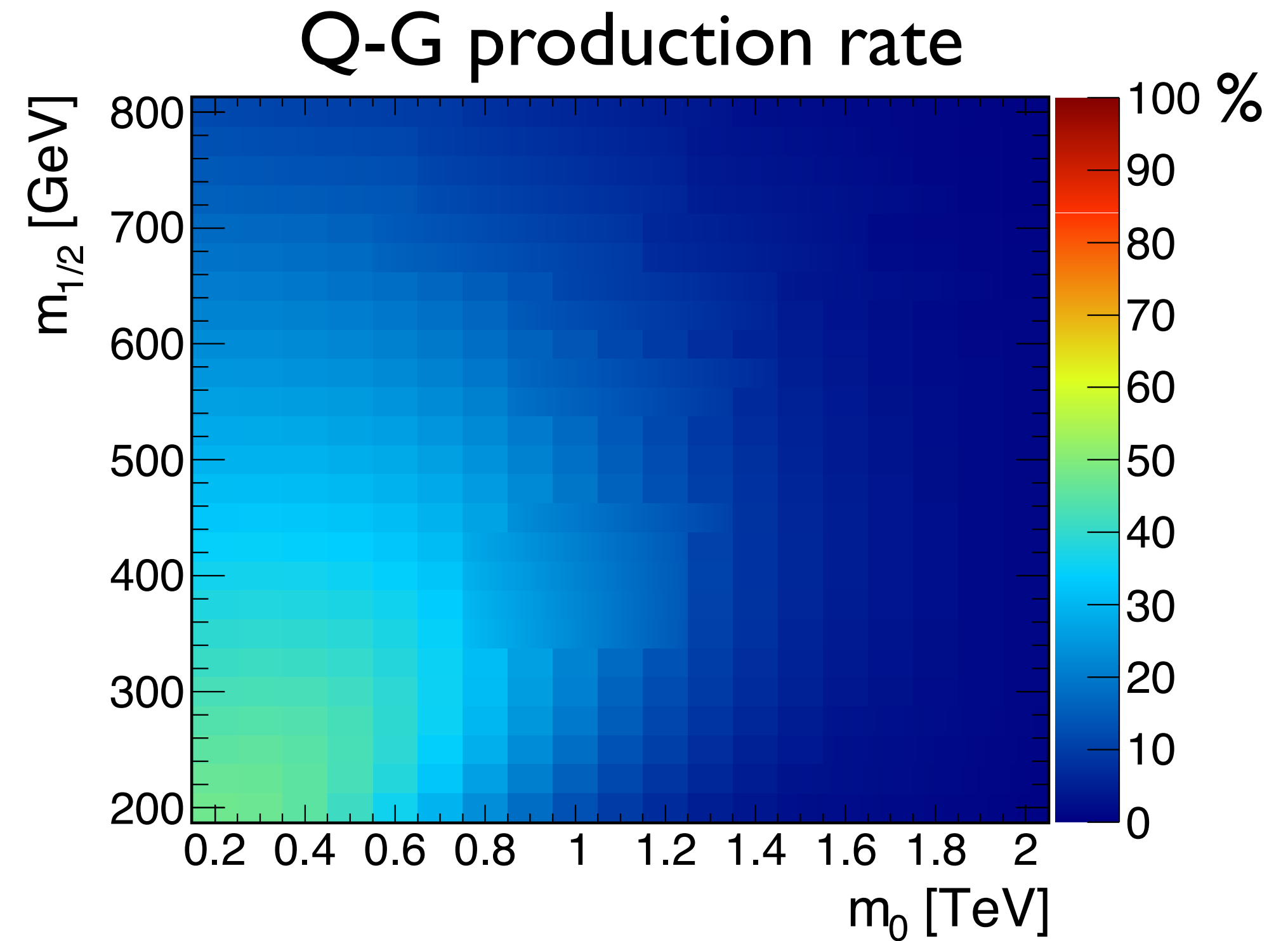
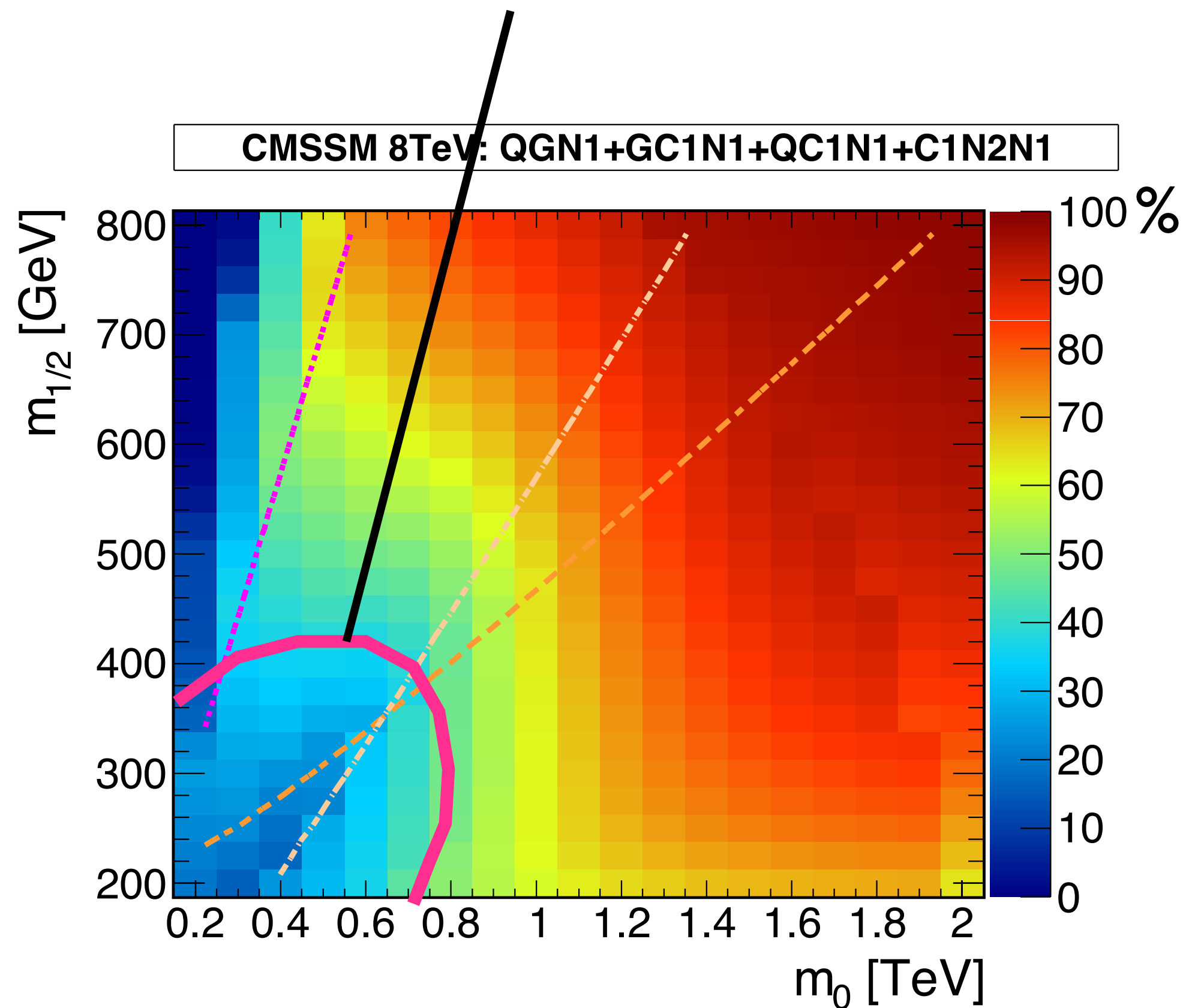
$\tilde{g} \rightarrow t \bar{t} \chi_3^0 \rightarrow \dots$
 $\tilde{g} \rightarrow b \bar{t} \chi_2^+ \rightarrow \dots$
 $\mu/m_{\tilde{g}}$



CMSSM coverage

$$\begin{aligned} Q &= \tilde{q} \\ G &= \tilde{g} \\ C1 &= \tilde{\chi}_1^\pm \\ N2 &= \tilde{\chi}_2^0 \\ N1 &= \tilde{\chi}_1^0 \end{aligned}$$

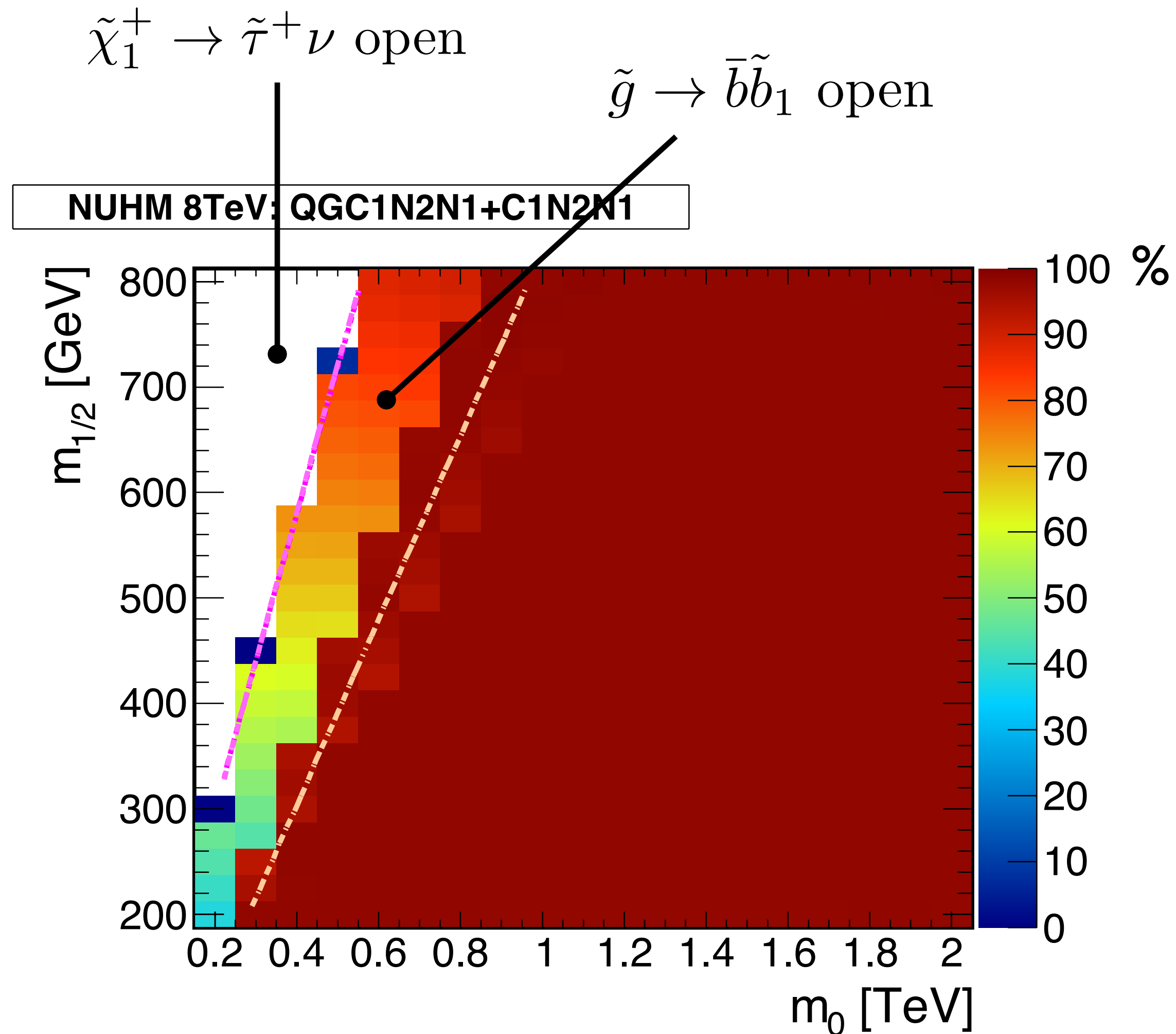
Q-G-CI-NI model is necessary => 4 dimensional



NUHM coverage

Non Universal Higgs Mass (NUHM) model

$$\mu = m_A = 3 m_{1/2}$$



- $\tilde{g} \rightarrow t\bar{t}\tilde{\chi}_3^0, b\bar{t}\tilde{\chi}_2^+$ not open due to large μ
- 4D Q-G-C1(N2)-N1 model is assumed

- After the null SUSY search results and $\sim 125\text{GeV}$ Higgs discovery at the LHC, **natural SUSY** and **spread SUSY** models become attractive

natural SUSY

- tension between fine tuning and null SUSY search results is optimised
- the Higgs sector should be extended to explain 125GeV Higgs mass, but such extension may not alter the LHC signatures

R.Kitano, Y.Nomura, 0602096

M.Papucci, J.T.Ruderman, A.Weiler, 1110.6926

...

spread SUSY

- EW naturalness is removed from the SUSY motivation
- no flavour/CP problem
- 125GeV Higgs mass is realised
- model building is simple

L.J.Hall, Y.Nomura, 1111.4519

M.Ibe, S.Matsumoto, T.T.Yanagida, 1202.2253

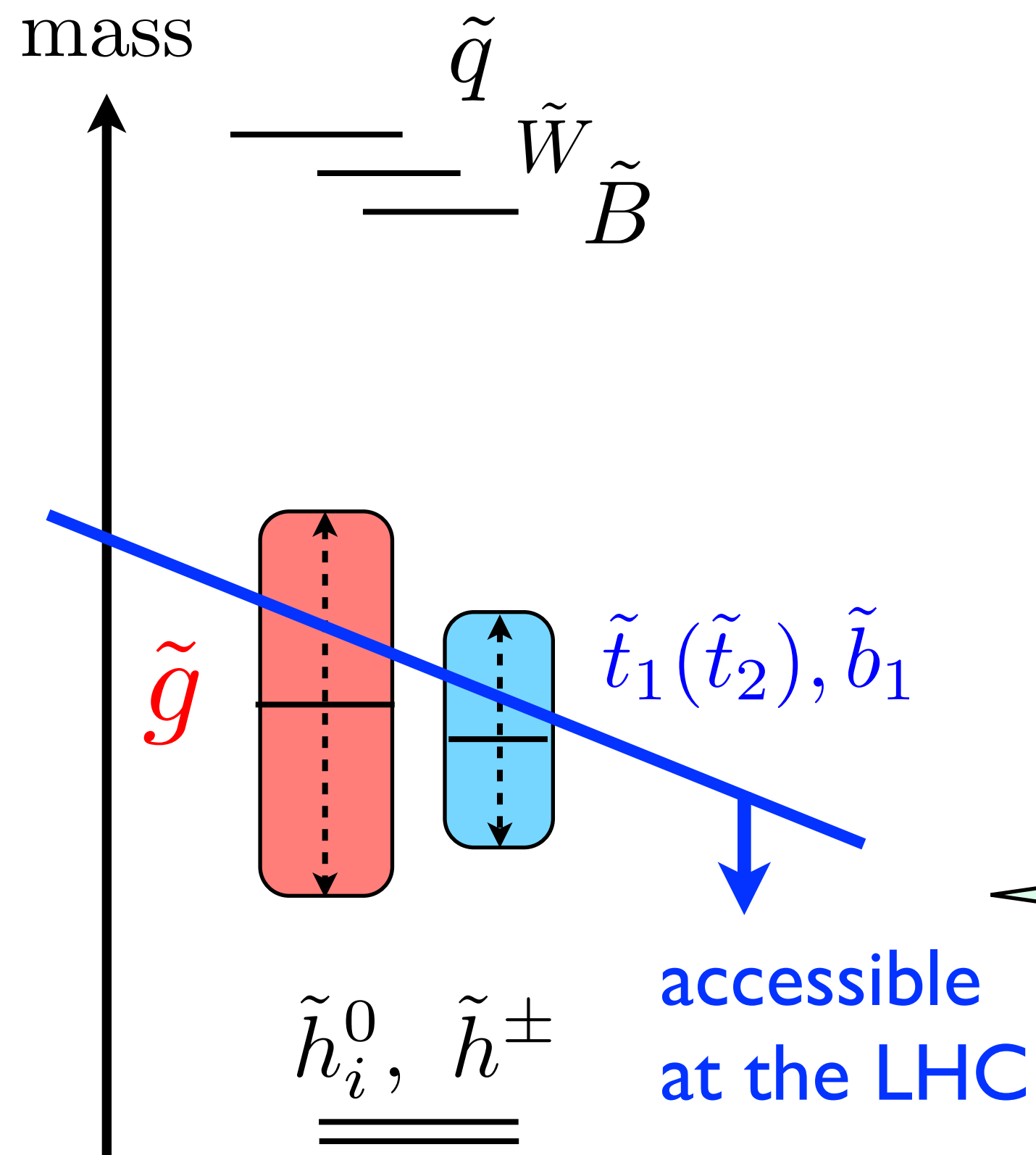
A.Arvanitakia, N.Craigb, S.Dimopoulousa, G.Villadoroc, 1210.0555

N.Arkani-Hamed, A.Gupta, D.E.Kaplan, N.Weiner, T.Zorawski, 1212.6971

...

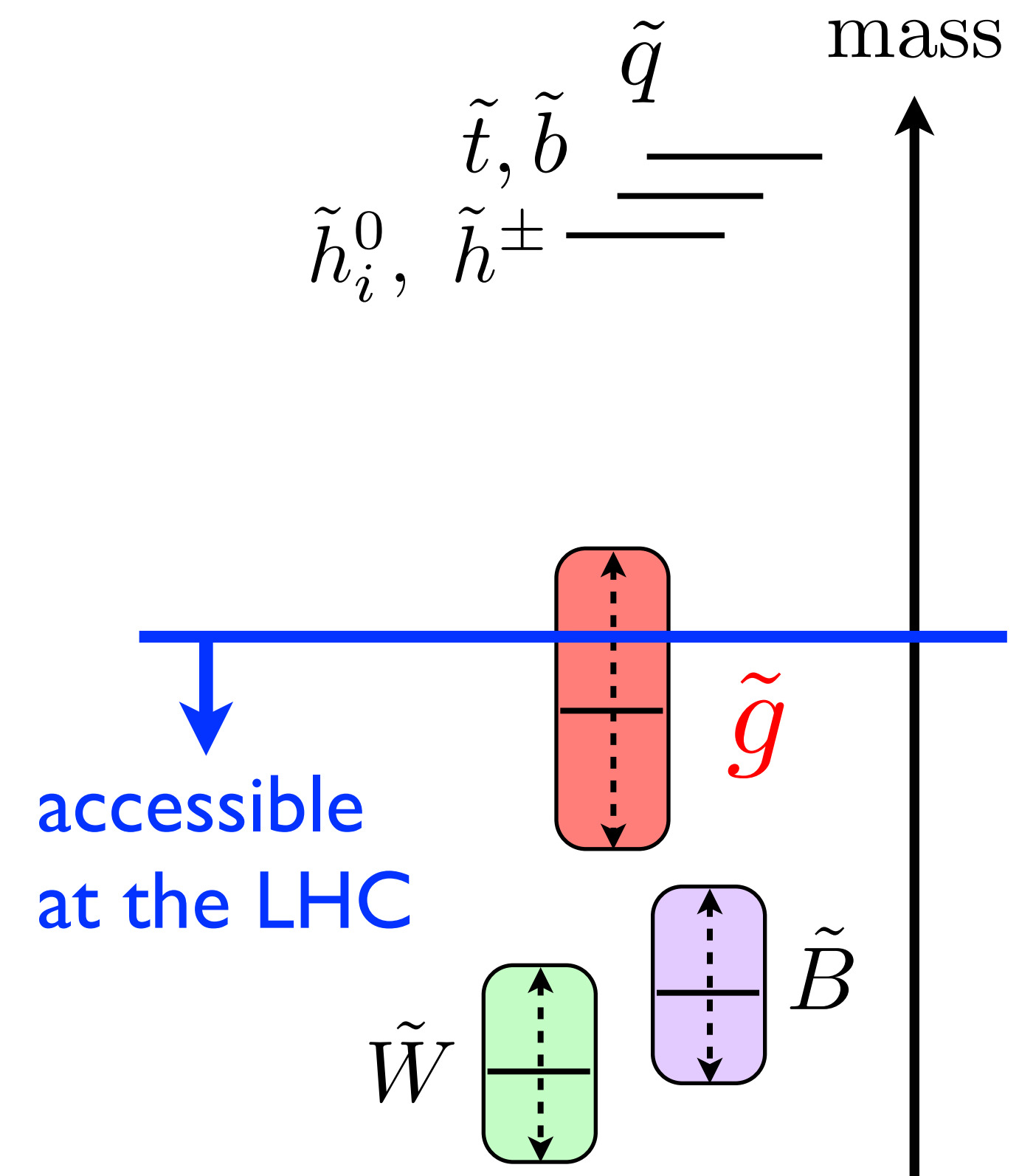
- After the null SUSY search results and $\sim 125\text{GeV}$ Higgs discovery at the LHC, **natural SUSY** and **spread SUSY** models become attractive

natural SUSY



A few particles participate the LHC signature

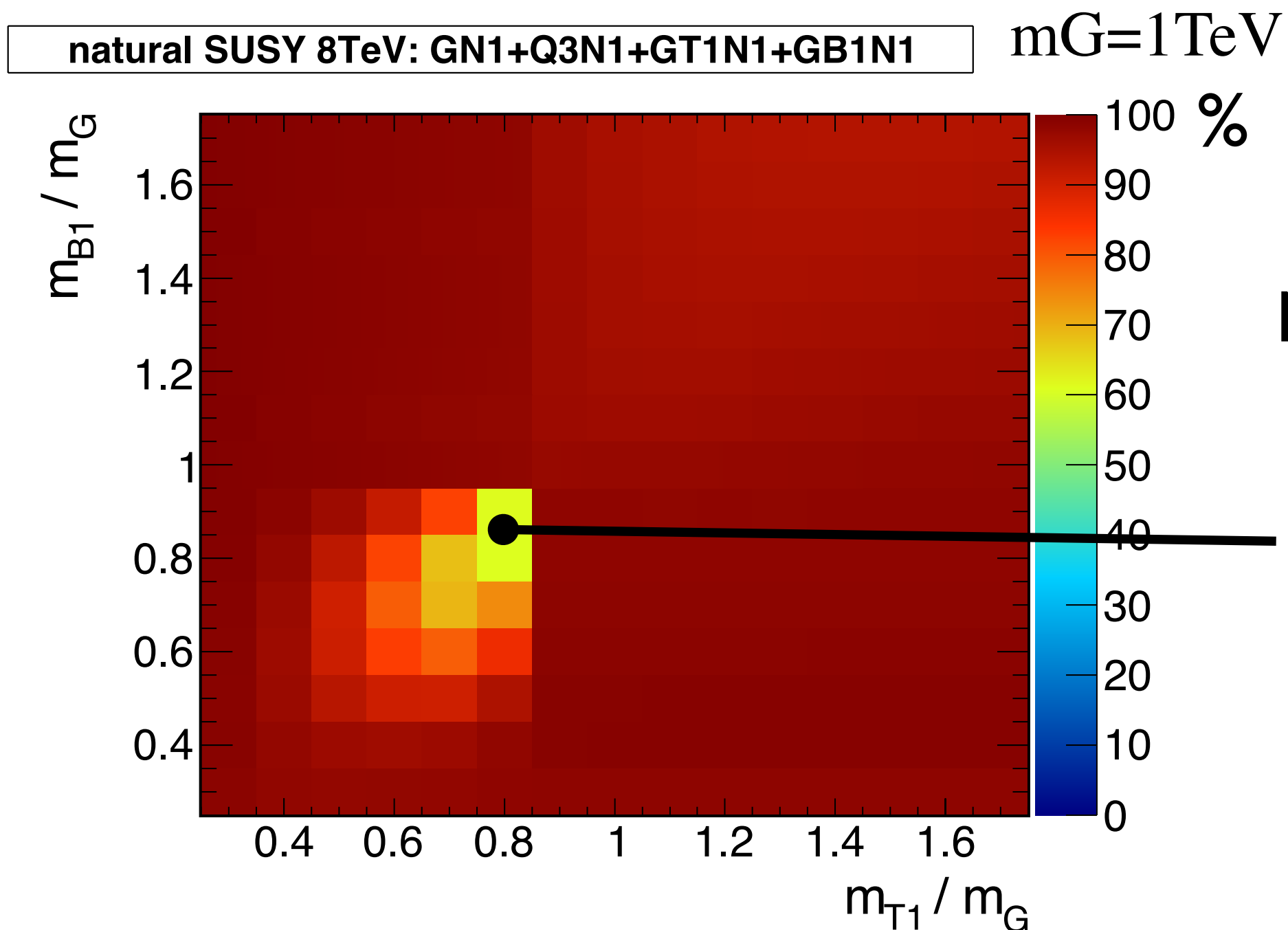
spread SUSY



- After the null SUSY search results and $\sim 125\text{GeV}$ Higgs discovery at the LHC, **natural SUSY** and **spread SUSY** models become attractive

natural SUSY

G-N1 + G-T1(B1)-N1 model

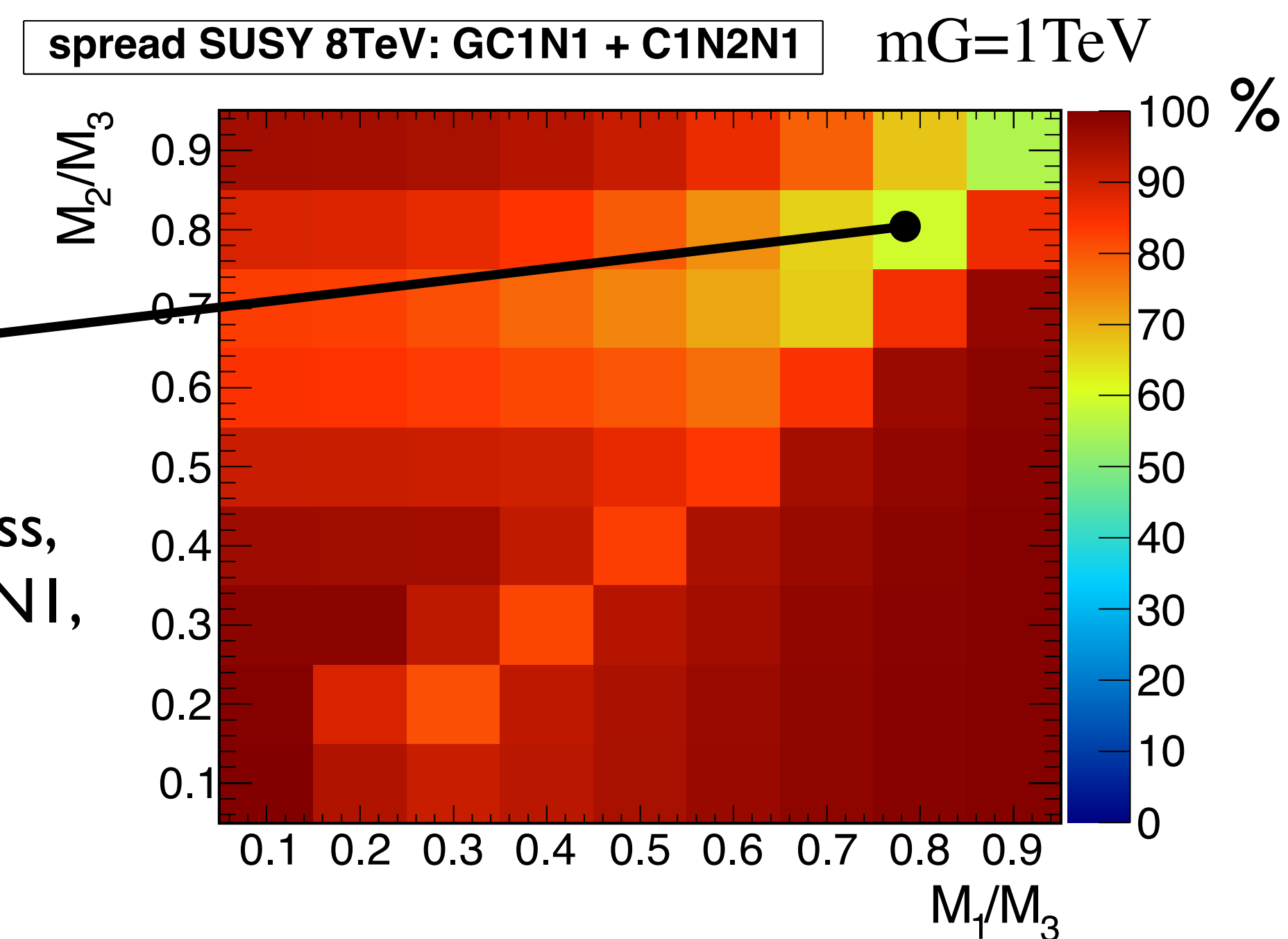


N2 is required

asymmetric process,
GtT1tN1_GbB1bN1,
required (4D)

spread SUSY

G-C1-N1 + C1-N2-N1 model



good coverage can be achieved by a few sets of simplified models

Summary

- FastLim reconstructs N_{SUSY} of a given model from the efficiency tables for the simplified processes and calculate the LHC constraints.
- A finite set of simplified processes are implemented in FastLim, but it can provide good coverage for interesting SUSY models: CMSSM, NUHM, natural SUSY, spread SUSY

simplified processes



your favourite models

Slim-Fast

Creamy Milk Chocolate

ARTIFICIALLY FLAVORED

11 FL OZ (325mL)

220 calories | 24 vitamins & minerals | high in calcium | 99% fat free

Healthy Ready To Drink Meal

Nutrition Facts

Serving Size 1 Can (325 mL)

Amount Per Serving

Calories (Energy)	220	Fat Cal	25
% Daily Value*			
Total Fat	3g		5%
Saturated Fat	1g		5%
Trans Fat	0g		
Polyunsaturated Fat	0.5g		
Monounsaturated Fat	1.5g		
Cholesterol	5mg		2%
Sodium	220mg		9%
Potassium	600mg		17%
Total Carbohydrate	40g		13%
Dietary Fiber	5g		20%
Sugars	34g		
Protein	10g		20%
Vitamin A	35%	Vitamin C	100%
Calcium	40%	Iron	15%
Vitamin D	35%	Vitamin E	100%
Vitamin K	25%	Thiamin	35%
Riboflavin	35%	Niacin	35%
Vitamin B6	35%	Folate	30%
Vitamin B12	35%	Biotin	35%
Pantothenic Acid	35%	Phosphorus	40%
Iodine	35%	Magnesium	35%
Zinc	15%	Selenium	25%
Manganese	35%	Chromium	35%
Molybdenum	35%		

*Percent Daily Values are based on a 2,000 calorie diet.

Thank you for listening

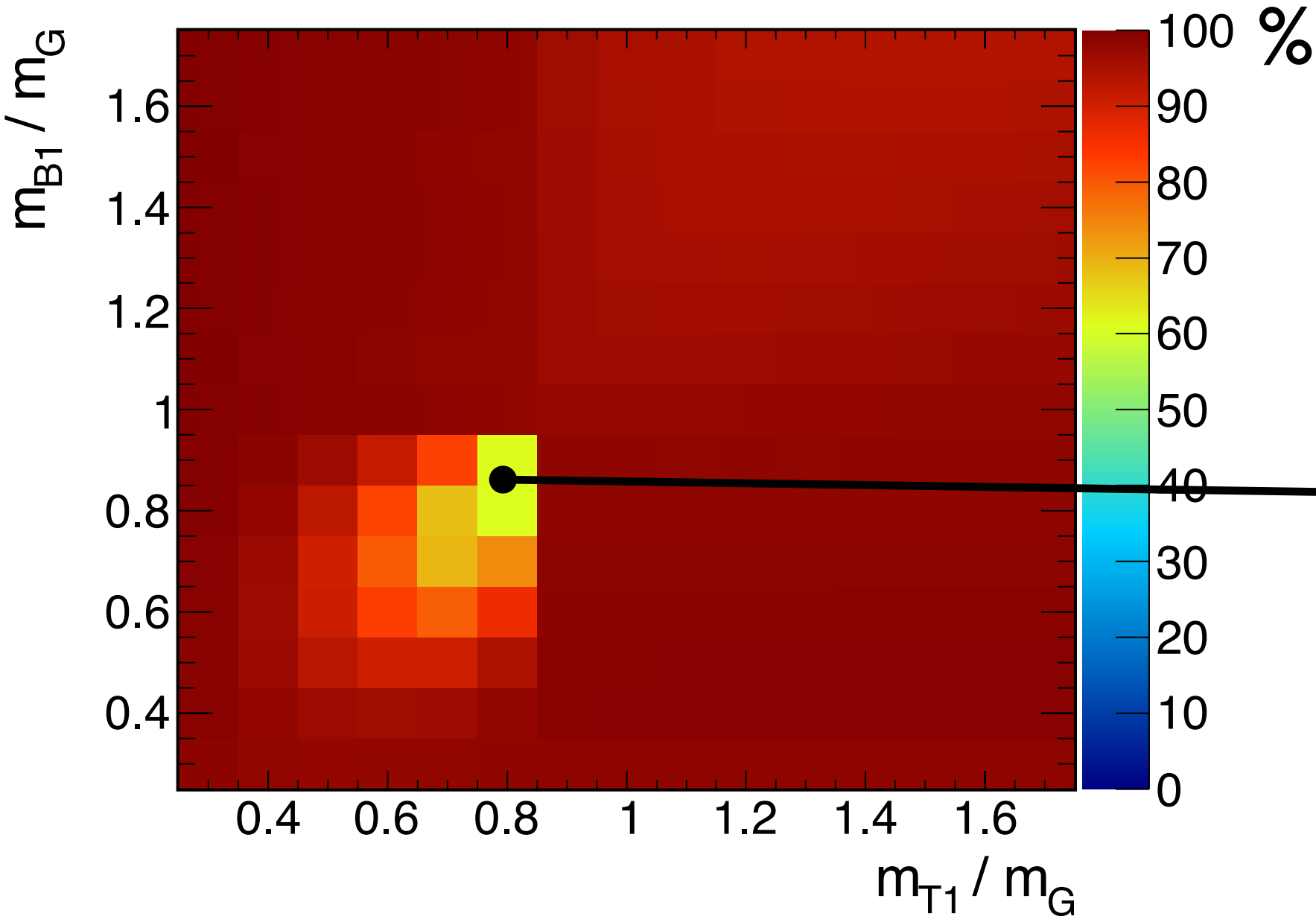
G-N1 + G-T1(B1)-N1 model

GbbNI_GbtNI	GbbNI_GqqNI	BibNI_BltNI
GbtNI_GqqNI	GqqNI_GqqNI	BibNI_BibNI
GqqNI_GttNI	GtTibNI_GtTltNI	TibNI_TltNI
GttNI_GttNI	GtTltNI_GtTltNI	TltNI_TltNI
GbbNI_GbbNI	GtTibNI_GtTibNI	BltNI_BltNI
GbtNI_GbtNI	GbBibNI_GbBltNI	TibNI_TibNI
GbbNI_GttNI	GbBltNI_GbBltNI	
GbtNI_GttNI	GbBibNI_GbBibNI	

G-C1-N1 + C1-N2-N1 model

GbbNI_GbtNI	GbbNI_GbtClwNI	GttNI_GttNI
GqqNI_GttNI	GqqClwNI_GqqNI	GbtNI_GttNI
GbbNI_GbbNI	GbtNI_GqqClwNI	GbtClwNI_GttNI
GbbNI_GttNI	GbtClwNI_GqqNI	GbbNI_GqqClwNI
GqqNI_GqqNI	GbtClwNI_GbtNI	GqqClwNI_GttNI
GbtNI_GqqNI	GqqClwNI_GqqClwNI	ClwNI_ClwNI
GbbNI_GqqNI	GbtClwNI_GbtClwNI	ClwNI_N2zNI
GbtNI_GbtNI	GbtClwNI_GqqClwNI	ClwNI_N2h0NI

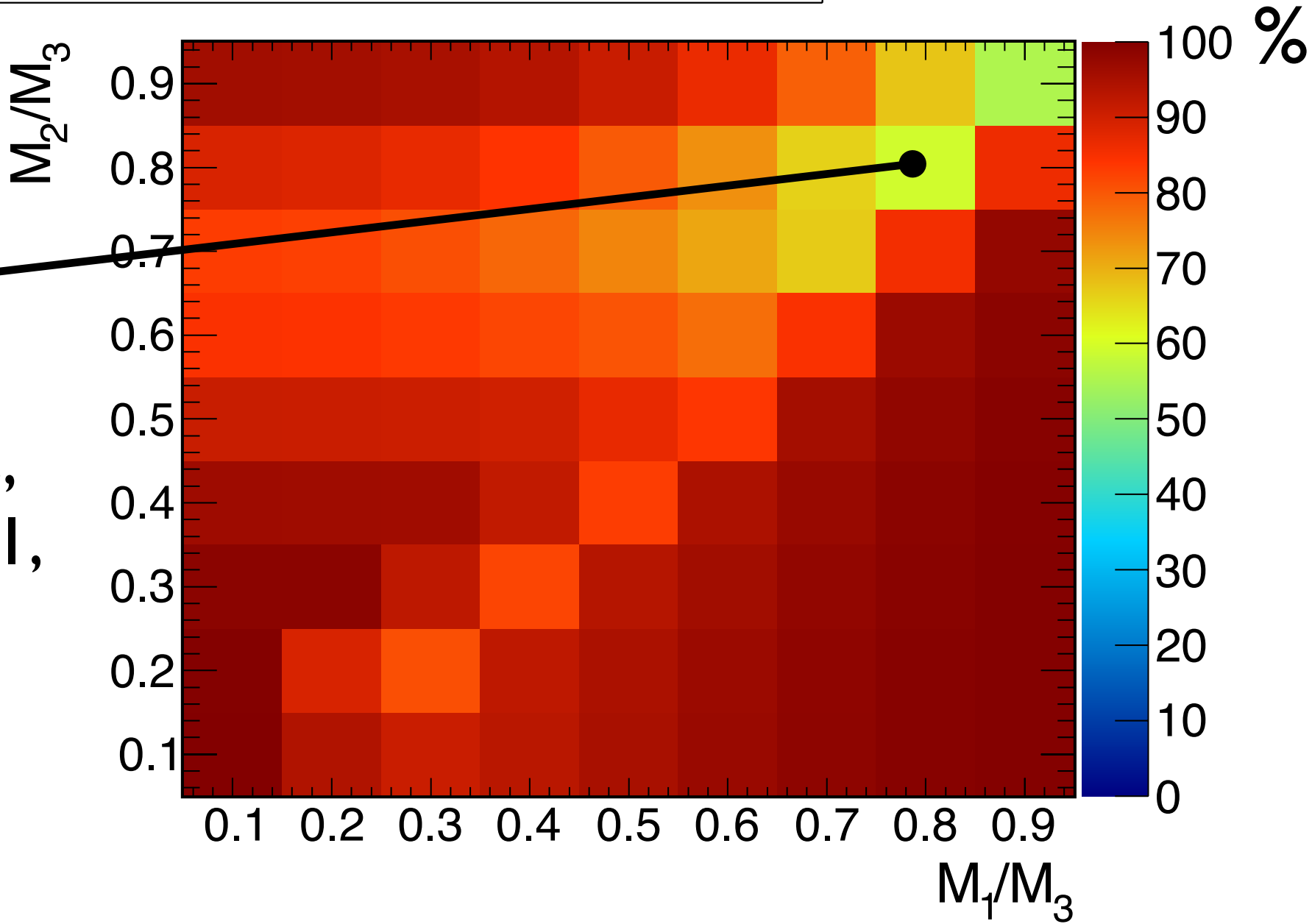
natural SUSY 8TeV: GN1+Q3N1+GT1N1+GB1N1 mG=1TeV



N2 is required

asymmetric process,
GtTltNI_GbBibNI,
required (4D)

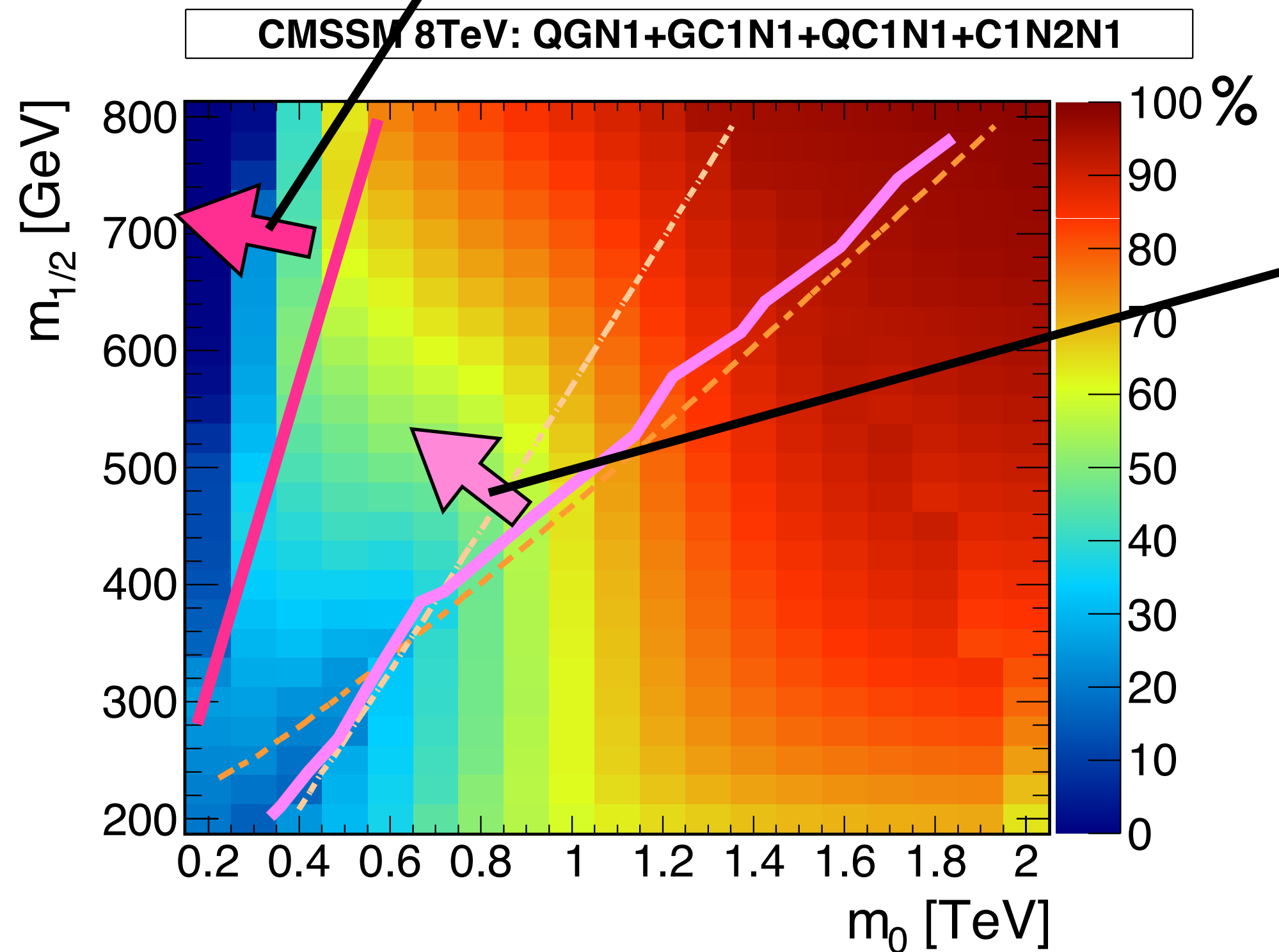
spread SUSY 8TeV: GC1N1 + C1N2N1 mG=1TeV



CMSSM coverage

$$\begin{aligned} Q &= \tilde{q} \\ G &= \tilde{g} \\ C1 &= \tilde{\chi}_1^\pm \\ N2 &= \tilde{\chi}_2^0 \\ N1 &= \tilde{\chi}_1^0 \end{aligned}$$

$$\begin{aligned} \tilde{\chi}_1^+ &\rightarrow \tilde{\tau}^+ \nu \rightarrow \tau^+ \nu \tilde{\chi}_1^0 \text{ open} \\ \tilde{\chi}_1^+ &\rightarrow W^+ \tilde{\chi}_1^0 \text{ suppressed} \end{aligned}$$



$$\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{b}_1 b \text{ open}$$

$$\tilde{t}_1 \rightarrow b \tilde{\chi}_2^+ \rightarrow \dots$$

$$\tilde{b}_1 \rightarrow b \tilde{\chi}_3^0 \rightarrow \dots$$

higgsinos make decay chains longer
BR is large due to the top Yukawa