

Where we do not want to go: avoiding charge- or  
color-breaking vacua

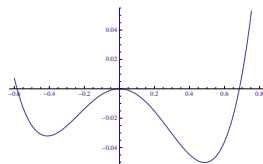
Ben O'Leary  
in collaboration with  
José Eliel Camargo Molina, Werner Porod, and Florian Staub

Julius-Maximilians-Universität Würzburg

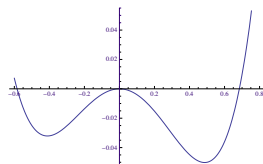
LHC Run1 Aftermath  
Bad Honnef,  
October 2nd, 2013



Even tree-level potentials for single scalars have in general multiple minima:

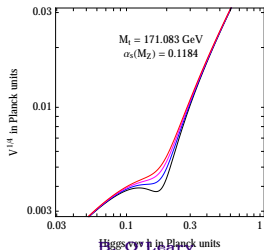


Even tree-level potentials for single scalars have in general multiple minima:

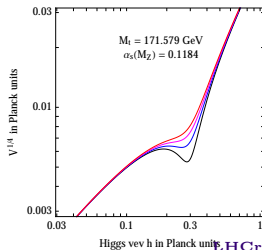


Top loops create extra minima for high Higgs field values in SM:

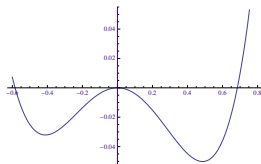
SM Higgs potential,  $M_h = 125$  GeV



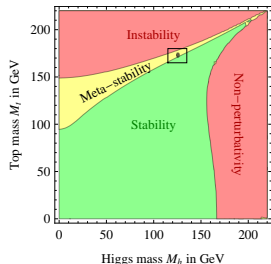
SM Higgs potential,  $M_h = 126$  GeV



Even tree-level potentials for single scalars have in general multiple minima:

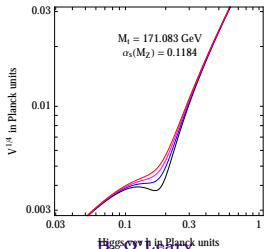


SM is probably metastable!

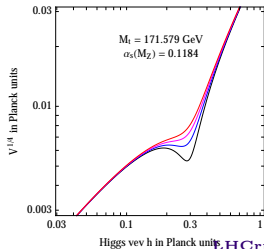


Top loops create extra minima for high Higgs field values in SM:

SM Higgs potential,  $M_h = 125$  GeV



SM Higgs potential,  $M_h = 126$  GeV

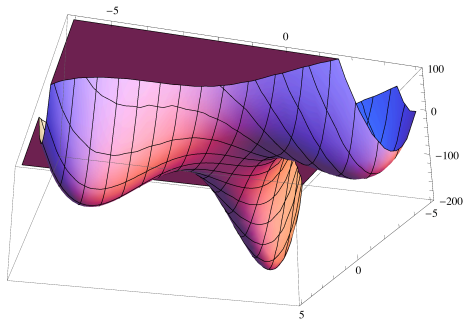


Figs. 2, 3, 4:  
Degrassi *et al.*, JHEP  
1208 (2012)

More scalars  $\Rightarrow$  more minima in general

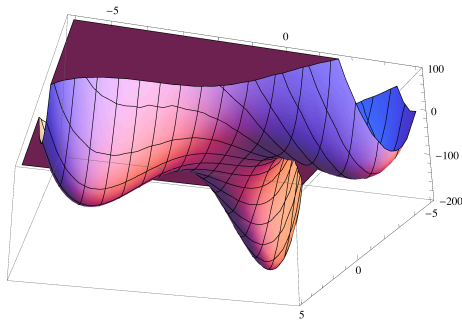
More scalars  $\Rightarrow$  more minima in general

Multiple scalars in general yield many vacua:



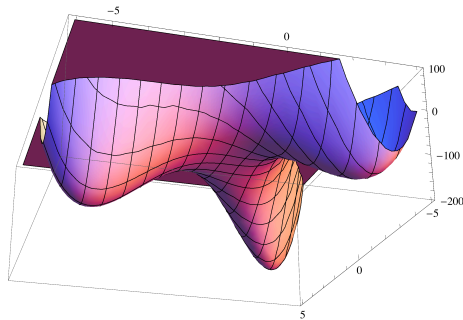
More scalars  $\Rightarrow$  more minima in general

Multiple scalars in general yield many vacua:



Finding the  
global  
minimum is  
not trivial!

Multiple scalars in general yield many vacua:

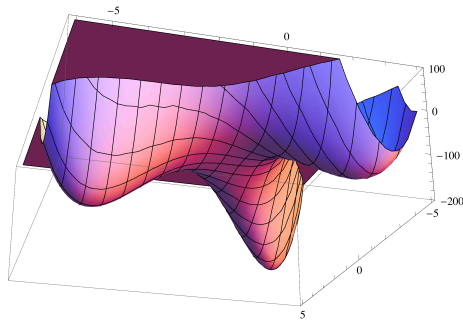


Finding the  
global  
minimum is  
not trivial!

- ▶ Charge- and/or color-breaking (CCB) minima (VEVs for charged or colored scalars)?



Multiple scalars in general yield many vacua:



Finding the  
global  
minimum is  
not trivial!

- ▶ Charge- and/or color-breaking (CCB) minima (VEVs for charged or colored scalars)?
- ▶ Desired VEV combination may not be global minimum (even non-CCB if there are enough VEVs required)

Charged or colored scalars?

### Charged or colored scalars?

- ▶ Supersymmetry: scalar partners for every SM fermion (plus lots of other stuff, not worth worrying about for this talk)

### Charged or colored scalars?

- ▶ Supersymmetry: scalar partners for every SM fermion (plus lots of other stuff, not worth worrying about for this talk)
- ▶ Staus: charged scalars, partners of  $\tau$  leptons
- ▶ Stops: charged and colored scalars, partners of top quarks
- ▶ Smuons, sneutrinos, sbottoms, *etc.*

## Charged or colored scalars?

- ▶ Supersymmetry: scalar partners for every SM fermion (plus lots of other stuff, not worth worrying about for this talk)
- ▶ Staus: charged scalars, partners of  $\tau$  leptons
- ▶ Stops: charged and colored scalars, partners of top quarks
- ▶ Smuons, sneutrinos, sbottoms, *etc.*
- ▶ Large Yukawa couplings and trilinear terms for (s)tops and, for large  $\tan \beta$ , (s)taus  $\Rightarrow$  focus on stau and stop VEVs for today

## Charged or colored scalars?

- ▶ Supersymmetry: scalar partners for every SM fermion (plus lots of other stuff, not worth worrying about for this talk)
- ▶ Staus: charged scalars, partners of  $\tau$  leptons
- ▶ Stops: charged and colored scalars, partners of top quarks
- ▶ Smuons, sneutrinos, sbottoms, *etc.*
- ▶ Large Yukawa couplings and trilinear terms for (s)tops and, for large  $\tan \beta$ , (s)taus  $\Rightarrow$  focus on stau and stop VEVs for today

Also, today we'll stick to the CMSSM:

## Charged or colored scalars?

- ▶ Supersymmetry: scalar partners for every SM fermion (plus lots of other stuff, not worth worrying about for this talk)
- ▶ Staus: charged scalars, partners of  $\tau$  leptons
- ▶ Stops: charged and colored scalars, partners of top quarks
- ▶ Smuons, sneutrinos, sbottoms, *etc.*
- ▶ Large Yukawa couplings and trilinear terms for (s)tops and, for large  $\tan \beta$ , (s)taus  $\Rightarrow$  focus on stau and stop VEVs for today

Also, today we'll stick to the CMSSM:

- ▶  $m_{\text{scalar}}^2(Q_{\text{GUT}}) = M_0^2$
- ▶  $m_{\text{gaugino}}(Q_{\text{GUT}}) = M_{1/2}$
- ▶ [scalar-scalar-scalar factor]( $Q_{\text{GUT}}$ ) =  $A_0$

Tree-level potential for non-zero stau VEVs:



Tree-level potential for non-zero stau VEVs:

$$\begin{aligned}
 V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
 = \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u \\
 + \frac{1}{2} (|\mu|^2(v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2) + \\
 \frac{1}{4} (Y_\tau^2(v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots
 \end{aligned}$$

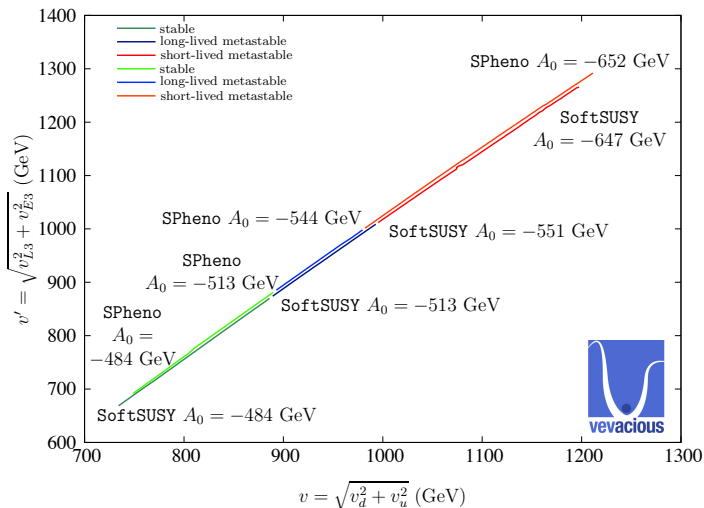
Tree-level potential for non-zero stau VEVs:

$$\begin{aligned}
 V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
 = \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u \\
 + \frac{1}{2} (|\mu|^2(v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2) + \\
 \frac{1}{4} (Y_\tau^2(v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots
 \end{aligned}$$

Minima could develop where  $v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u)$  gets more negative than “ $m^2 v^2 + \lambda v^4$ ” is positive



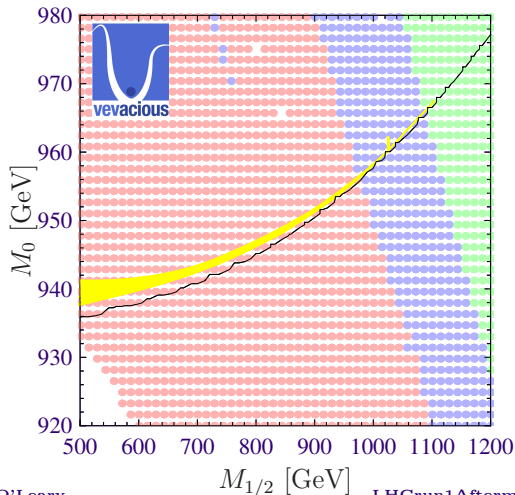
Camargo-Molina, BO'L, Porod, Staub, arXiv:1309.7212



$$m_0 = 400 \text{ GeV}, M_{1/2} = 300 \text{ GeV}, \tan \beta = 50, \mu > 0$$

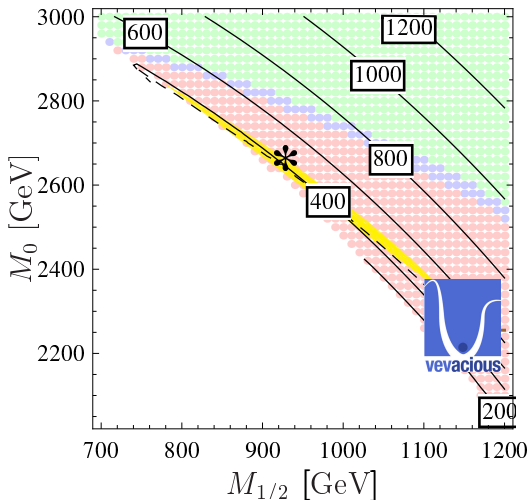


$A_0 = +3$  TeV,  $\tan\beta = 40$ ,  $\mu > 0$ ; (arXiv:1309.7212)  
 red/blue: metastable ( $\tau_{\text{tunnel}} < / > 3$  Gy); green: stable  
 yellow region: correct relic density; black:  $m_{\tilde{\tau}_1} = m_{\tilde{\chi}_1^0}$





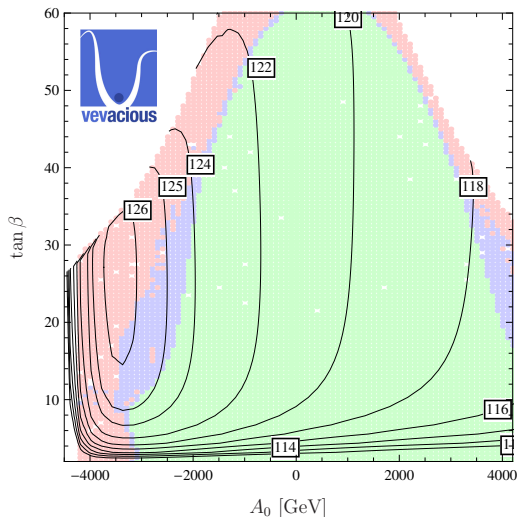
$A_0 = -6.444$  TeV,  $\tan \beta = 8.52$ ,  $\mu < 0$ ;  $m_{\tilde{t}_1}$  (GeV) contours colors as before, but dashed black for  $m_{\tilde{\tau}_1} = m_{\tilde{\chi}_1^0}$  (1309.7212)







$M_0 = M_{1/2} = 1$  TeV,  $\mu > 0$ ;  $m_h$  (GeV) contours  
 colors as before (1309.7212)



$$\begin{aligned}
& V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
&= \frac{1}{32} \left( g_1^2 (v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2 (v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2 \right) - B_\mu v_d v_u \\
&+ \frac{1}{2} \left( |\mu|^2 (v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2 \right) + \\
&\frac{1}{4} \left( Y_\tau^2 (v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots \right)
\end{aligned}$$

$$\begin{aligned}
& V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
&= \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u \\
&+ \frac{1}{2} (|\mu|^2(v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2) + \\
&\frac{1}{4} (Y_\tau^2(v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots
\end{aligned}$$

Some conditions in the literature have often been (mis-)used:

$$\begin{aligned}
& V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
&= \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u \\
&+ \frac{1}{2} (|\mu|^2(v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2) + \\
&\frac{1}{4} (Y_\tau^2(v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots
\end{aligned}$$

Some conditions in the literature have often been (mis-)used:

- ▶  $(A_0 - 0.5M_{1/2})^2 < 9M_0^2 + 2.67M_{1/2}^2$  [“GUT”, (if  $Y_f \ll 1$ )]

$$\begin{aligned}
& V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
&= \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u \\
&+ \frac{1}{2} (|\mu|^2(v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2) + \\
&\frac{1}{4} (Y_\tau^2(v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots
\end{aligned}$$

Some conditions in the literature have often been (mis-)used:

- ▶  $(A_0 - 0.5M_{1/2})^2 < 9M_0^2 + 2.67M_{1/2}^2$  [“GUT”, (if  $Y_f \ll 1$ )]
- ▶  $A_\tau^2 < 3(m_{H_d}^2 + |\mu|^2 + m_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2)$  [“ $A_\tau$ ”]

$$\begin{aligned}
V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
= \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u \\
+ \frac{1}{2} (|\mu|^2(v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2) + \\
\frac{1}{4} (Y_\tau^2(v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots
\end{aligned}$$

Some conditions in the literature have often been (mis-)used:

- ▶  $(A_0 - 0.5M_{1/2})^2 < 9M_0^2 + 2.67M_{1/2}^2$  [“GUT”, (if  $Y_f \ll 1$ )]
- ▶  $A_\tau^2 < 3(m_{H_d}^2 + |\mu|^2 + m_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2)$  [“ $A_\tau$ ”]
- ▶  $A_t^2 < 3(m_{H_u}^2 + |\mu|^2 + m_{\tilde{t}_L}^2 + m_{\tilde{t}_R}^2)$  [“ $A_t$ ”]

$$\begin{aligned}
V^{\text{tree}}(H_d = v_d/\sqrt{2}, H_u = v_u/\sqrt{2}, \tilde{\tau}_L = v_{\tilde{\tau}_L}/\sqrt{2}, \tilde{\tau}_R = v_{\tilde{\tau}_R}/\sqrt{2}) \\
= \frac{1}{32} (g_1^2(v_d^2 - v_u^2 + v_{\tilde{\tau}_L}^2 - 2v_{\tilde{\tau}_R}^2)^2 + g_2^2(v_d^2 - v_u^2 - v_{\tilde{\tau}_L}^2)^2) - B_\mu v_d v_u \\
+ \frac{1}{2} \left( |\mu|^2(v_d^2 + v_u^2) + m_{H_d}^2 v_d^2 + m_{H_u}^2 v_u^2 + m_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2 v_{\tilde{\tau}_R}^2 \right) + \\
\frac{1}{4} (Y_\tau^2(v_d^2 v_{\tilde{\tau}_L}^2 + v_d^2 v_{\tilde{\tau}_R}^2 + v_{\tilde{\tau}_L}^2 v_{\tilde{\tau}_R}^2) + \frac{Y_\tau}{\sqrt{2}} v_{\tilde{\tau}_L} v_{\tilde{\tau}_R} (A_\tau v_d - \mu v_u) + \dots
\end{aligned}$$

Some conditions in the literature have often been (mis-)used:

- ▶  $(A_0 - 0.5M_{1/2})^2 < 9M_0^2 + 2.67M_{1/2}^2$  [“GUT”, (if  $Y_f \ll 1$ )]
- ▶  $A_\tau^2 < 3(m_{H_d}^2 + |\mu|^2 + m_{\tilde{\tau}_L}^2 + m_{\tilde{\tau}_R}^2)$  [“ $A_\tau$ ”]
- ▶  $A_t^2 < 3(m_{H_u}^2 + |\mu|^2 + m_{\tilde{t}_L}^2 + m_{\tilde{t}_R}^2)$  [“ $A_t$ ”]
- ▶  $|(Y_\tau v_u \mu)/(\sqrt{2}m_\tau)| < 56.9\sqrt{m_{\tilde{\tau}_L} m_{\tilde{\tau}_R}} + 57.1(m_{\tilde{\tau}_L} + 1.03m_{\tilde{\tau}_R}) - 1.28 \times 10^4 \text{GeV} + \frac{1.67 \times 10^6 \text{GeV}^2}{m_{\tilde{\tau}_L} + m_{\tilde{\tau}_R}} - 6.41 \times 10^6 \text{GeV}^3 \left( \frac{1}{m_{\tilde{\tau}_L}^2} + \frac{0.983}{m_{\tilde{\tau}_R}^2} \right)$   
[“numeric”]

“GUT”: Ellwanger, Rausch de Traubenberg, Savoy, Nucl. Phys. B**492**

“ $A_\tau$ ”, “ $A_t$ ”: Alvarez-Gaumé, Polchinski, Wise, Nucl. Phys. B**221**;

“numeric”: Kitahara, Yoshinaga, arXiv:1303.0461, JHEP)



SPS4 ( $M_0 = 400\text{GeV}$ ,  $M_{1/2} = 300\text{GeV}$ ,  $\tan\beta = 50$ ,  $|\mu| > 0$ ,  
 $A_0 = 0\text{GeV}$ ) but with  $A_0 \rightarrow \dots$

SPS4 ( $M_0 = 400\text{GeV}$ ,  $M_{1/2} = 300\text{GeV}$ ,  $\tan\beta = 50$ ,  $|\mu| > 0$ ,  $A_0 = 0\text{GeV}$ ) but with  $A_0 \rightarrow \dots$

| $A_0$ | generator | $v_d$ | $v_u$ | $v_{\tilde{\tau}_L}$ | $v_{\tilde{\tau}_R}$ |
|-------|-----------|-------|-------|----------------------|----------------------|
| -484  | SPheno    | 184   | 726   | 409                  | 558                  |
| -484  | SoftSUSY  | 181   | 712   | 394                  | 540                  |
| -513  | SPheno    | 269   | 851   | 540                  | 701                  |
| -513  | SoftSUSY  | 274   | 846   | 532                  | 694                  |
| -544  | SPheno    | 326   | 927   | 620                  | 787                  |
| -551  | SoftSUSY  | 342   | 935   | 626                  | 795                  |
| -652  | SPheno    | 485   | 1110  | 819                  | 999                  |
| -647  | SoftSUSY  | 481   | 1097  | 800                  | 981                  |

$v_d : v_{\tilde{\tau}_L} : v_{\tilde{\tau}_R} \neq 1 : 1 : 1$  at CCB minimum

SPS4 ( $M_0 = 400\text{GeV}$ ,  $M_{1/2} = 300\text{GeV}$ ,  $\tan\beta = 50$ ,  $|\mu| > 0$ ,  $A_0 = 0\text{GeV}$ ) but with  $A_0 \rightarrow \dots$

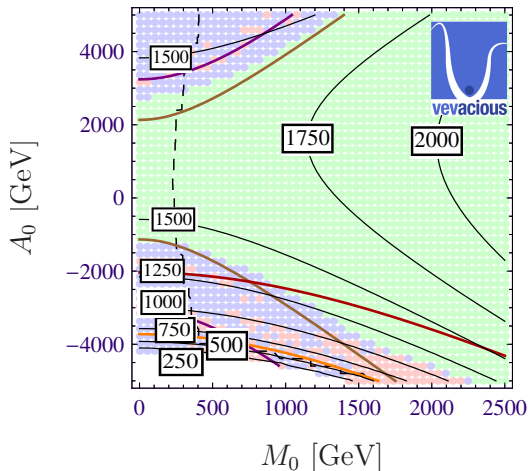
| $A_0$ | generator | $v_d$ | $v_u$ | $v_{\tilde{\tau}_L}$ | $v_{\tilde{\tau}_R}$ |
|-------|-----------|-------|-------|----------------------|----------------------|
| -484  | SPheno    | 184   | 726   | 409                  | 558                  |
| -484  | SoftSUSY  | 181   | 712   | 394                  | 540                  |
| -513  | SPheno    | 269   | 851   | 540                  | 701                  |
| -513  | SoftSUSY  | 274   | 846   | 532                  | 694                  |
| -544  | SPheno    | 326   | 927   | 620                  | 787                  |
| -551  | SoftSUSY  | 342   | 935   | 626                  | 795                  |
| -652  | SPheno    | 485   | 1110  | 819                  | 999                  |
| -647  | SoftSUSY  | 481   | 1097  | 800                  | 981                  |

$v_d : v_{\tilde{\tau}_L} : v_{\tilde{\tau}_R} \neq 1 : 1 : 1$  at CCB minimum  
 $\Rightarrow$  not on line of “ $A_\tau$ ”!

Sometimes it looks like analytic conditions do well

Sometimes it looks like analytic conditions do well

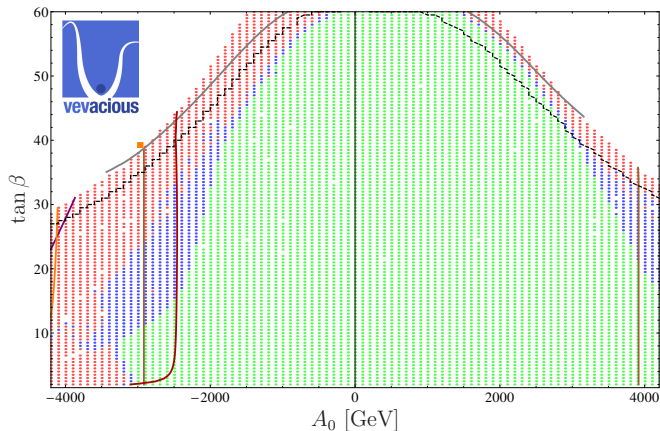
$M_{1/2} = 1$  TeV,  $\tan \beta = 10$ ,  $\mu > 0$ ;  $m_{\tilde{\tau}_1}$  (GeV) contours (1309.7212)



Brown: “GUT”; Purple: “ $A_\tau$ ”; Orange: “ $A_t$ ”  
 Dark red: improved “ $A_t$ ” (Casas, Lleyda, Munoz, Nucl. Phys. B471)  
 Dashed black:  $m_{\tilde{\tau}_1} = m_{\tilde{\chi}_1^0}$



$$M_{1/2} = 1000 \text{ GeV}, m_0 = 1000 \text{ GeV}, \mu > 0 \text{ (1309.7212)}$$

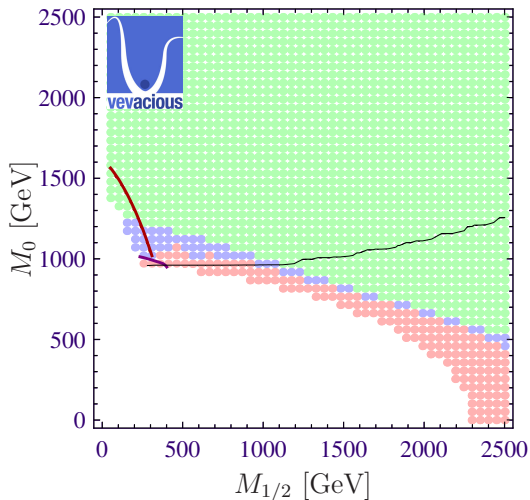


Brown: “GUT”;      Purple: “ $A_\tau$ ”;      Orange: “ $A_t$ ”  
 Grey: “numeric”;      Dark red: improved “ $A_t$ ”  
 Dashed black:  $m_{\tilde{\tau}_1} = m_{\tilde{\chi}_1^0}$

# Analytic conditions can completely fail



$$A_0 = +3 \text{ TeV}, \tan \beta = 40, \mu > 0 \text{ (1309.7212)}$$



Purple: “ $A_\tau$ ”; Dark red: improved “ $A_t$ ”; Black:  $m_{\tilde{\tau}_1} = m_{\tilde{\chi}_1^0}$

Vevacious: a tool to find global minima of multiscalar potentials!



v e v a c i o u s

Vevacious: a tool to find global minima of multiscalar potentials!

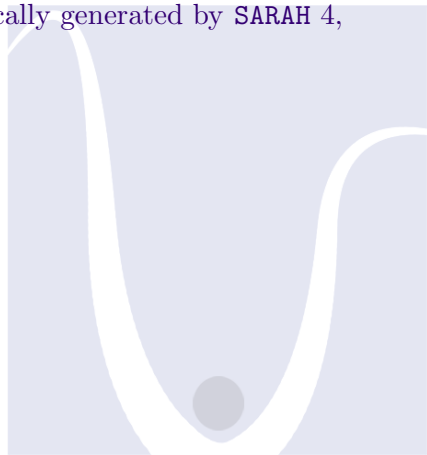
Vevacious is a new, publicly-available code, that:



v e v a c i o u s

Vevacious is a new, publicly-available code, that:

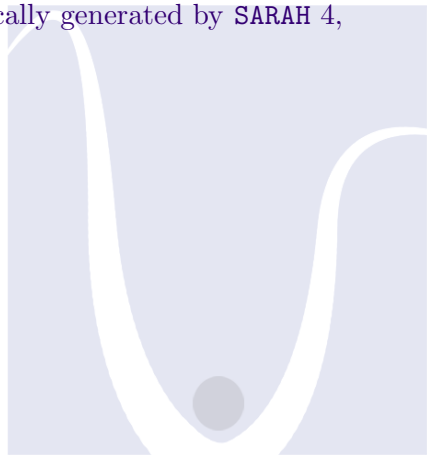
- ▶ takes a model file (automatically generated by SARAH 4, F. Staub [arXiv:1309.7223](https://arxiv.org/abs/1309.7223))



v e v a c i o u s

Vevacious is a new, publicly-available code, that:

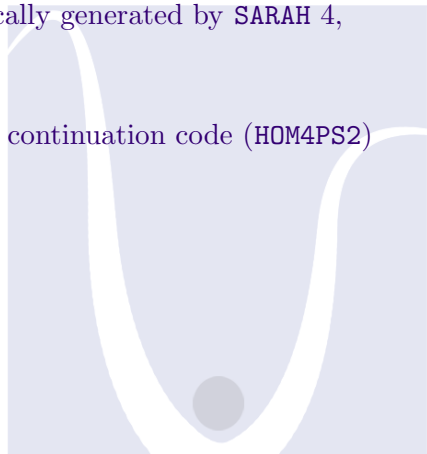
- ▶ takes a model file (automatically generated by SARAH 4, F. Staub [arXiv:1309.7223](https://arxiv.org/abs/1309.7223))
- ▶ takes an SLHA file



v e v a c i o u s

Vevacious is a new, publicly-available code, that:

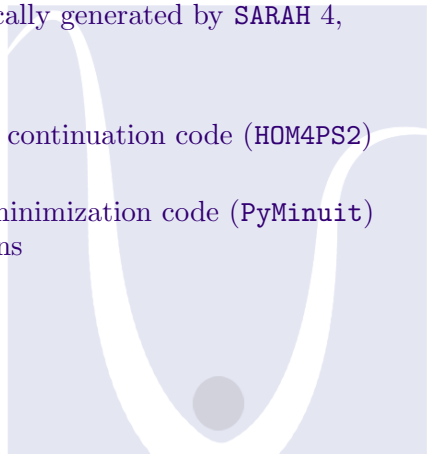
- ▶ takes a model file (automatically generated by SARAH 4, F. Staub [arXiv:1309.7223](https://arxiv.org/abs/1309.7223))
- ▶ takes an SLHA file
- ▶ prepares and runs homotopy continuation code (HOM4PS2) to find *all* tree-level extrema



v e v a c i o u s

Vevacious is a new, publicly-available code, that:

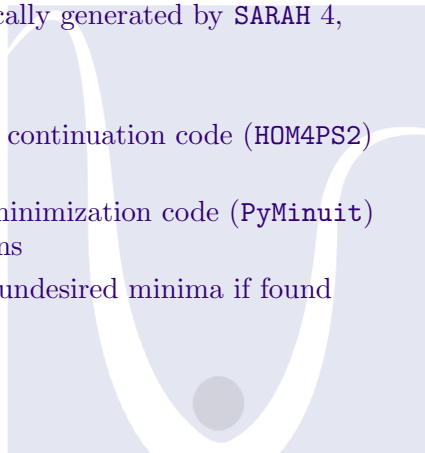
- ▶ takes a model file (automatically generated by SARAH 4, F. Staub [arXiv:1309.7223](https://arxiv.org/abs/1309.7223))
- ▶ takes an SLHA file
- ▶ prepares and runs homotopy continuation code (HOM4PS2) to find *all* tree-level extrema
- ▶ prepares and runs gradient minimization code (PyMinuit) to account for loop corrections



v e v a c i o u s

Vevacious is a new, publicly-available code, that:

- ▶ takes a model file (automatically generated by SARAH 4, F. Staub [arXiv:1309.7223](https://arxiv.org/abs/1309.7223))
- ▶ takes an SLHA file
- ▶ prepares and runs homotopy continuation code (HOM4PS2) to find *all* tree-level extrema
- ▶ prepares and runs gradient minimization code (PyMinuit) to account for loop corrections
- ▶ calculates tunneling time to undesired minima if found (CosmoTransitions)



v e v a c i o u s



Vevacious is a new, publicly-available code, that:

- ▶ takes a model file (automatically generated by SARAH 4, F. Staub arXiv:1309.7223)
- ▶ takes an SLHA file
- ▶ prepares and runs homotopy continuation code (HOM4PS2) to find *all* tree-level extrema
- ▶ prepares and runs gradient minimization code (PyMinuit) to account for loop corrections
- ▶ calculates tunneling time to undesired minima if found (CosmoTransitions)

<http://vevacious.hepforge.org/>

v e v a c i o u s



Coupled cubic equations are pretty damn hard!

Coupled cubic equations are pretty damn hard!  
Gröbner bases:

Coupled cubic equations are pretty damn hard!

Gröbner bases:

- ▶ Decomposition of system using fancy algebra
- ▶ Has been used to investigate NMSSM  
(Maniatis, von Manteuffel, Nachtmann, arXiv:hep-ph/0608314, EJPC)
- ▶ Computationally expensive, especially in terms of RAM

Coupled cubic equations are pretty damn hard!

Gröbner bases:

- ▶ Decomposition of system using fancy algebra
- ▶ Has been used to investigate NMSSM  
(Maniatis, von Manteuffel, Nachtmann, arXiv:hep-ph/0608314, EJPC)
- ▶ Computationally expensive, especially in terms of RAM

Homotopy continuation:

Coupled cubic equations are pretty damn hard!

Gröbner bases:

- ▶ Decomposition of system using fancy algebra
- ▶ Has been used to investigate NMSSM  
(Maniatis, von Manteuffel, Nachtmann, arXiv:hep-ph/0608314, EJPC)
- ▶ Computationally expensive, especially in terms of RAM

Homotopy continuation:

- ▶ Gradual deformation of simple system of equations into target system
- ▶ Has been used to investigate SM with up to 5 extra scalars  
(Maniatis, Mehta, arXiv:1203.0409, EPJ+)
- ▶  $\exists$  public codes and programs: PHCpack, Bertini, HOM4PS2

Coupled cubic equations are pretty damn hard!

Gröbner bases:

- ▶ Decomposition of system using fancy algebra
- ▶ Has been used to investigate NMSSM  
(Maniatis, von Manteuffel, Nachtmann, arXiv:hep-ph/0608314, EJPC)
- ▶ Computationally expensive, especially in terms of RAM

Homotopy continuation:

- ▶ Gradual deformation of simple system of equations into target system
- ▶ Has been used to investigate SM with up to 5 extra scalars  
(Maniatis, Mehta, arXiv:1203.0409, EPJ+)
- ▶  $\exists$  public codes and programs: `PHCpack`, `Bertini`, `HOM4PS2`

Before **Vevacious**: only implemented on a model-by-model basis, at tree-level!



Vevacious is fast enough for scans, can be adapted to new models easily



v e v a c i o u s

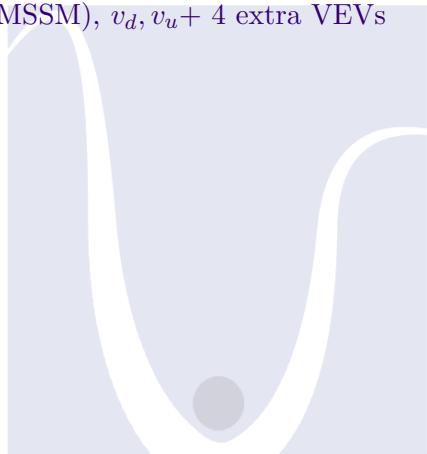
- ▶ Evaluating stability of a parameter point depends on model



v e v a c i o u s

- ▶ Evaluating stability of a parameter point depends on model

For example, MSSM (not just CMSSM),  $v_d, v_u + 4$  extra VEVs for  $\tilde{\tau}_{L,R}, \tilde{t}_{L,R}$ , on my laptop:

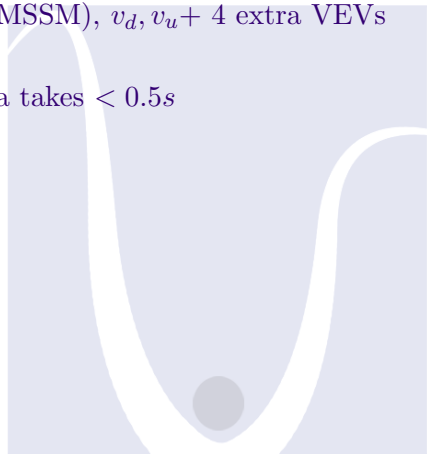


v e v a c i o u s

- ▶ Evaluating stability of a parameter point depends on model

For example, MSSM (not just CMSSM),  $v_d, v_u + 4$  extra VEVs for  $\tilde{\tau}_{L,R}, \tilde{t}_{L,R}$ , on my laptop:

- ▶ Finding *all* tree-level extrema takes  $< 0.5s$

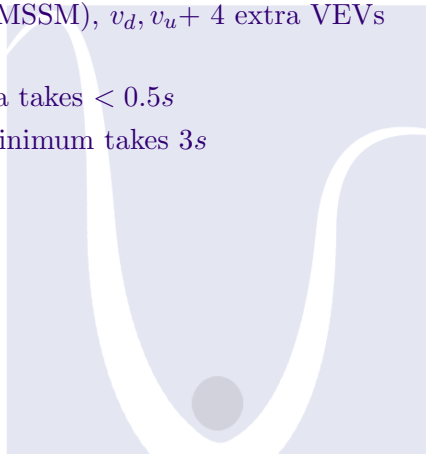


v e v a c i o u s

- ▶ Evaluating stability of a parameter point depends on model

For example, MSSM (not just CMSSM),  $v_d, v_u + 4$  extra VEVs for  $\tilde{\tau}_{L,R}, \tilde{t}_{L,R}$ , on my laptop:

- ▶ Finding *all* tree-level extrema takes  $< 0.5s$
- ▶ Determining 1-loop global minimum takes  $3s$



v e v a c i o u s

- ▶ Evaluating stability of a parameter point depends on model

For example, MSSM (not just CMSSM),  $v_d, v_u + 4$  extra VEVs for  $\tilde{\tau}_{L,R}, \tilde{t}_{L,R}$ , on my laptop:

- ▶ Finding *all* tree-level extrema takes  $< 0.5s$
- ▶ Determining 1-loop global minimum takes  $3s$
- ▶ Estimating tunneling time strongly depends on relative depth and location of global minimum compared to input minimum:  $15s$  typical,  $500s$  for borderline cases

v e v a c i o u s

- ▶ Evaluating stability of a parameter point depends on model

For example, MSSM (not just CMSSM),  $v_d, v_u + 4$  extra VEVs for  $\tilde{\tau}_{L,R}, \tilde{t}_{L,R}$ , on my laptop:

- ▶ Finding *all* tree-level extrema takes  $< 0.5s$
- ▶ Determining 1-loop global minimum takes  $3s$
- ▶ Estimating tunneling time strongly depends on relative depth and location of global minimum compared to input minimum:  $15s$  typical,  $500s$  for borderline cases

Making new model files with SARAH 4:

- ▶ Small modifications to normal SARAH model file:
  - ▶ specify VEVs for whichever scalars
  - ▶ modify mass eigenstates to accomodate more mixing

v e v a c i o u s

- ▶ Evaluating stability of a parameter point depends on model

For example, MSSM (not just CMSSM),  $v_d, v_u + 4$  extra VEVs for  $\tilde{\tau}_{L,R}, \tilde{t}_{L,R}$ , on my laptop:

- ▶ Finding *all* tree-level extrema takes  $< 0.5s$
- ▶ Determining 1-loop global minimum takes  $3s$
- ▶ Estimating tunneling time strongly depends on relative depth and location of global minimum compared to input minimum:  $15s$  typical,  $500s$  for borderline cases

Making new model files with SARAH 4:

- ▶ Small modifications to normal SARAH model file:
  - ▶ specify VEVs for whichever scalars
  - ▶ modify mass eigenstates to accomodate more mixing
- ▶ Creating model file with SARAH takes minutes:

MakeVevacious[]



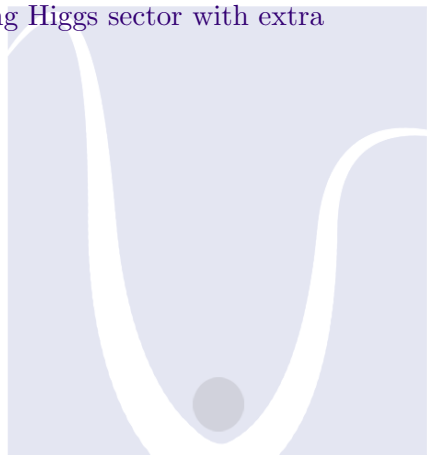
Minimizing potentials not trivial:



v e v a c i o u s

Minimizing potentials not trivial:

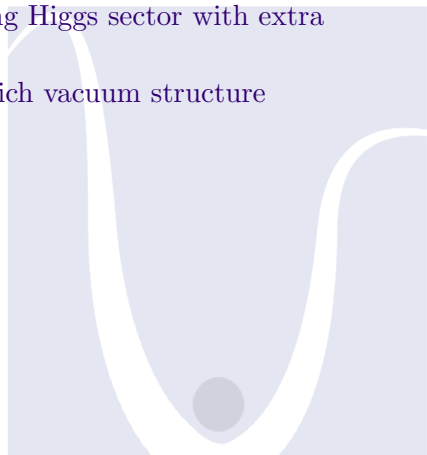
- ▶ Vitrally important if extending Higgs sector with extra scalars



v e v a c i o u s

Minimizing potentials not trivial:

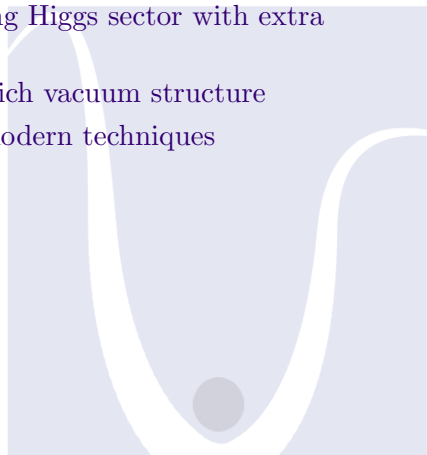
- ▶ Vitally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure



v e v a c i o u s

Minimizing potentials not trivial:

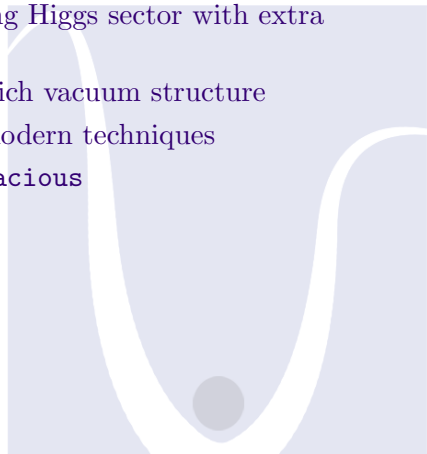
- ▶ Vitrally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure
- ▶ Difficult, but feasible with modern techniques



v e v a c i o u s

Minimizing potentials not trivial:

- ▶ Vitrally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure
- ▶ Difficult, but feasible with modern techniques
- ▶ Automated by **SARAH** + **Vevacious**

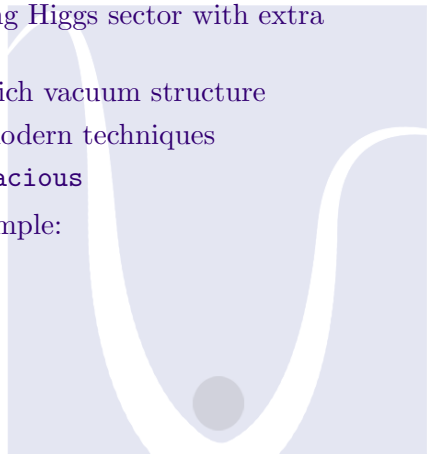


v e v a c i o u s

Minimizing potentials not trivial:

- ▶ Vitrally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure
- ▶ Difficult, but feasible with modern techniques
- ▶ Automated by **SARAH** + **Vevacious**

The CMSSM is an excellent example:



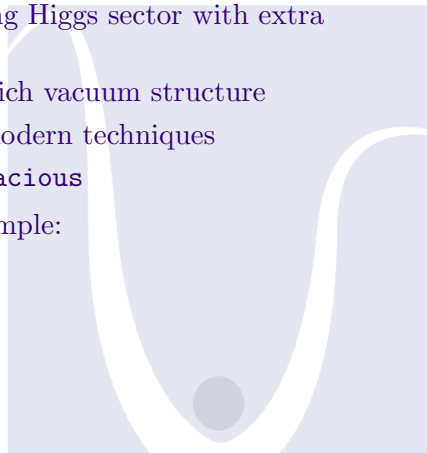
v e v a c i o u s

Minimizing potentials not trivial:

- ▶ Vitally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure
- ▶ Difficult, but feasible with modern techniques
- ▶ Automated by **SARAH + Vevacious**

The CMSSM is an excellent example:

- ▶ Non-trivial VEV structure



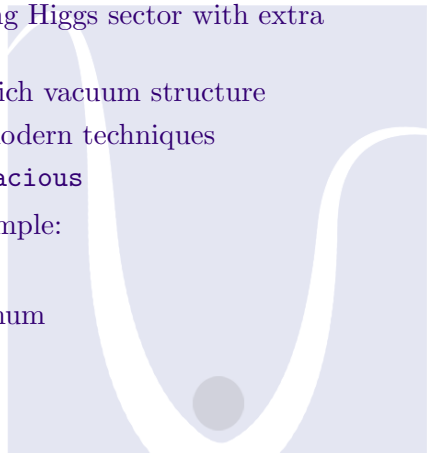
v e v a c i o u s

Minimizing potentials not trivial:

- ▶ Vitrally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure
- ▶ Difficult, but feasible with modern techniques
- ▶ Automated by **SARAH + Vevacious**

The CMSSM is an excellent example:

- ▶ Non-trivial VEV structure
- ▶ Often has CCB global minimum



v e v a c i o u s



Minimizing potentials not trivial:

- ▶ Vitrally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure
- ▶ Difficult, but feasible with modern techniques
- ▶ Automated by **SARAH + Vevacious**

The CMSSM is an excellent example:

- ▶ Non-trivial VEV structure
- ▶ Often has CCB global minimum

Avoiding short tunneling times to CCB global minima should be a concern in choosing benchmarks and parameter fits!

v e v a c i o u s

Minimizing potentials not trivial:

- ▶ Vitrally important if extending Higgs sector with extra scalars
- ▶ Multiple VEV-ing fields  $\rightarrow$  rich vacuum structure
- ▶ Difficult, but feasible with modern techniques
- ▶ Automated by **SARAH** + **Vevacious**

The CMSSM is an excellent example:

- ▶ Non-trivial VEV structure
- ▶ Often has CCB global minimum

Avoiding short tunneling times to CCB global minima should be a concern in choosing benchmarks and parameter fits!

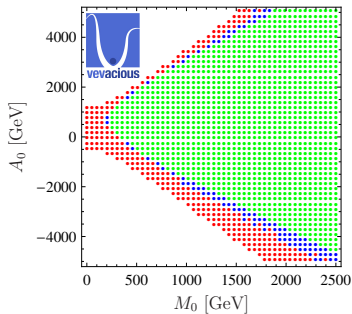
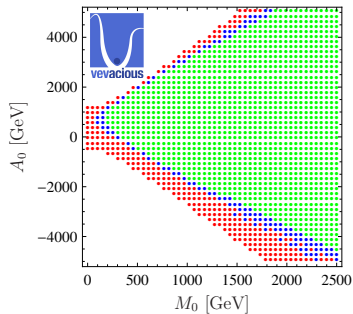
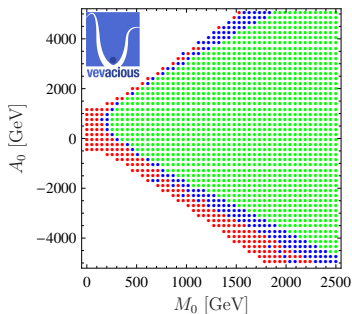
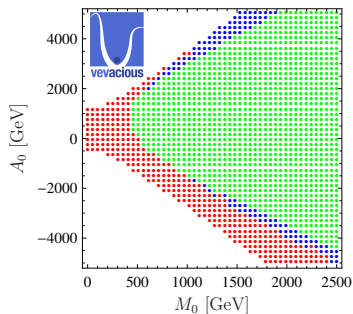
Thank you for your attention!

v e v a c i o u s

Backup slides

- ▶  $\Gamma / \text{volume} = Ae^{-B/\hbar}(1 + \mathcal{O}(\hbar))$
- ▶  $A$  is solitonic solution, should be  $\sim$  energy scale of potential
- ▶  $B \sim ([\text{surface tension}]/[\text{energy density difference}])^3$
- ▶ typically TeV-scale energy barriers, energy depth differences  $\Rightarrow$  roughly tunneling times of (factors of  $16\pi^2$  etc.)/TeV  $\ll$  age of Universe

# Scale and loop order dependence: halving $Q$



# Scale and loop order dependence: doubling $Q$

