



Probing Low *x* Gluons at STAR with Forward Asymmetry Measurements

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For the STAR Collaboration

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Probing Low *x* Gluons at STAR with Forward Asymmetry Measurements



- Current Understanding of $\Delta g(x)$
- STAR Detector and Datasets
- Pushing to Low *x* with Forward π^0 's
 - In the Endcap
 - In the Forward Calorimeter
- Prospects for Very Low *x* with Dijets and a Calorimeter Upgrade













- Integral of $\Delta g(x)$ in range 0.05 < x < 1.0increases substantially, now significantly above zero.
- Uncertainty shrinks substantially from DSSV* to new DSSV fit
- First firm evidence of non-zero gluon polarization!

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- Published results from 2006 longitudinal dataset at 200 GeV
 - $-\sim 7 \text{ pb}^{-1}$ recorded
- Work in progress from
 - 2009 200 GeV long. ~25 pb^-1
 - $-2012\ 510\ GeV\ long. \sim 80\ pb^{-1}$
 - 2013 510 GeV long. ~300 pb⁻¹
 recorded
- Caveats
 - Other datasets, publications, for other measurements
 - 2013 >450 pb⁻¹ delivered
 - Prescales can affect individual measurements



A. Gibson, Valparaiso; STAR Low x g; PANIC



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- STAR has measured $\pi^0 A_{LL}$ in three different pseudorapidity ranges
 - Different kinematics, different systematics
 - Here with data from 2006

•qg scattering dominates at high η with high x quarks and low x gluons

•No large asymmetries seen



STAR's Endcap Electromagnetic Calorimeter

mount-

ing

rina

SS back

nlate

post-

shower

Pb/SS

plastic

scint





- Scintillating strip SMD
 - $-\phi$ segmented into 12 sectors
 - Two active planes
 - 288 strips per plane
- Resolution of a few mm



- Nucl. Instrum. Meth. A 499 (2003) 740.
- Lead/scintillator sampling EM calorimeter
 - Covers $1.09 < \eta < 2.00$ over full 2π azimuth
 - 720 optically isolated projective towers (~22 X_0)
 - 2 pre-shower, 1 post-shower layers, and an additional shower maximum detector (SMD)
- Photon trigger places thresholds on maximum tower energy and the 3x3 patch of surrounding towers

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π^0 Background and Cross-Section Computation 0.8 < η < 2.0 with 2006 Dataset



- Inclusive π^0 mass distribution fit to MC templates, in bins of $\pi^0 p_T$
 - Signal
 - Conversion BG (π^0 candidate is from gamma \rightarrow e+ e-)
 - All other BG (extra or missing photons, π^0 candidate is gamma and e-, etc.)
 - Shapes from MC, relative fraction (and thus signal fraction) extracted from fit to data

STAR π^0 's at low and high p_T , for sqrt(s) 200 GeV (PYTHIA, unpolarized CTEQ 5L)







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 - Shapes from MC, relative fraction (and thus signal fraction) extracted from fit to data
- Lowest analyzed bin is 5-6 GeV $\pi^0 p_T$
 - Data-MC agreement unsatisfactory below this
 - Large amount of passive material, not well modeled
- Unfolded cross section calculated with a "smearing matrix"
 - Dominant systematic is EEMC energy scale
 - Consistent with NLO pQCD (thanks M. Stratmann)
 - B. Jaeger et al., Phys. Rev. D 67, 054005 (2003) CTEQ 6.5, DSS FF







- Raw longitudinal asymmetry corrected for
 - Luminosity asymmetries (small)
 - Beam polarizations
 - Background asymmetries
 - Estimated from mass sidebands, and consistent with zero (with uncertainty ~ 0.01)
- Statistical error (bars) dominate
- Systematic error (boxes)
 - Signal fraction uncertainties from template fits
 - Uncertainty on background asymmetry
- Integrated across p_T probably constrains GRSV Δg -max?
- PRD 89, 012001 (2014)







- Transverse asymmetries as well!
- Plotted in bins of $\pi^0 p_T$ (integrated over $0.06 < x_F < 0.27$), and in bins of x_F
- Statistical error (bars) dominates over systematic error (boxes)
- Twist-3 prediction
 - K. Kanazawa and Y. Koike,
 - Phys. Rev. D 83, 114024 (2011)
 STAR at sqrt(s) 200 GeV





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$\pi^0 A_{LL}$ Prospects in 2012 Dataset



FMS

Pb Glass EM Calorimeter pseudo-rapidity 2.7<η<4.0 Small cells: 3.81x3.81 cm Outer cells: 5.81 x 5.81 cm

MULLI **FPD EM Calorimeter** Small cells only Two 7x7 arrays AR 20 Forward EM Calorimetry In STAR.

- Pushing even further forward, with the FPD and FMS
- Work underway with large 2012 and 2013 datasets at 510 GeV
 - Prescales lead to effectively ~50 pb⁻¹ in 2012, focused at high p_T
 - And ~10 pb⁻¹ in 2013, focused at lower p_T
 - An older preliminary result also exists

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Dijet Measurements

- Inclusive measurements have been the workhorse of STAR ΔG program to date
- Broad *x* range sampled in each p_T bin
- Dijet or other correlation measurements which reconstruct the full final state are sensitive to initial kinematics at leading order

•Prospect of mapping out the shape of $\Delta g(x)$

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Jet Reconstruction

MC Jets **Jet Levels** Jet direction Detector GEANT Particle e, v, γ PYTHIA π, p, etc Parton q,g

STAR Detector has:

- Full azimuthal coverage
- Charged particle tracking from TPC for $|\eta| < 1.3$
- E/BEMC provide electromagnetic energy reconstruction for $-1 < \eta < 2.0$ STAR well suited for jet measurements

Anti-K_T Jet Algorithm:

- Radius = 0.6
- Used in both data and simulation

2009 Dijet Cross Section Results

STAR Di-jet Cross Section 10⁻¹ Data STAR PRELIMINARY Data Systematic Error d²ơ/dMd|ŋ| [µb/(GeV/c²) 0 0 0 NLO pQCD CTEQ6M (de Florian) + Had + UE pp @√s = 200 GeV Anti-k_τ, R = 0.6, |η| < 0.8 L dt = 17.1 $pb^{1} \pm 7.7\%$ 10⁻⁵ B. Page DNP 2013 30 20 50 60 70 80 90 100 40 Invariant Mass [GeV/c²] p. 24

Thickness of vertical black hashing represents size of statistical error on the measurement

Green hatched box is symmetric about data point and is the quadrature sum of all systematic errors

Thickness of blue box represents error on theory determined by changing factorization and renormalization scales by factor of 0.5 and 2

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2009 Dijet Cross Section at 500 GeV

- Also a preliminary 2006 dijet cross section at 200 GeV
 - T. Sakuma, M. Walker, Journal of Physics: Conference Series 295, 012068 (2011).

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Probing *very* low *x* gluons with Forward Calorimeter Upgrade: 2020

ECal:

Tungsten-Powder-Scintillating-fiber 2.3 cm Moliere Radius, Tower-size: 2.5x2.5x17 cm³ 23 X_o HCal:

Lead and Scintillator tiles, Tower size of 10x10x81 cm³ 4 interaction length

https://drupal.star.bnl.gov/STAR/starnotes/public/sn0605

STAR polarized p+p and p+A LoI

Dijet Projections with the Forward Calorimeter Upgrade

R_{cone}=0.7 L=1fb⁻¹ $-1 < \eta < 2$ $\sqrt{s} = 500 \ GeV$ E_{T1}>8 GeV P=60% 2.8<*n*<3.7 E_{T2}>5 GeV https://drupal.star.bnl.gov/STAR/starnotes/public/sn0605 B. Surrow PoS(DIS2014) 241 STAR polarized p+p and p+A LoI 0.2E-2 0.15 10⁸ تے 10 $0 < \eta_{3(4)} < 0.8 / 2.8 < \eta_{4(3)} < 3.7$ -0.8 < $\eta_{3\,(4)}$ < 0 / 2.8 < $\eta_{4\,(3)}$ < 3.7 do/dx, (do/dx₂) (pb) da/dx, (da/dx₂) (pb) (EAST / FCS) (WEST / FCS) -2 0.15 x 10 107 10 -2 0.1 (WEST / FCS) (EAST / FCS) x 10 10⁶ 10⁶ 0.1 0.05 10⁵ 10 5 0.05 10 4 10 4 -0 0 10³ 10 10² -0.05 -0.05 10 GRSV STD 10 10 -0.1 **GRSV STD** -0.1 DSSV 1 10⁻⁵ 10⁻⁵ -0.15 10^{-4} 10^{-3} 10^{-2} 10^{-1} 1 10^{-4} 10^{-3} 10^{-2} 10^{-1} 1 30 40 45 50 55 20 30 35 40 45 50 55 60 M (GeV) X, (X.) X1 (X2) M (GeV) A_{LL} 0.2 0.2 ALL 10 do/dx, (do/dx₂) (pb) $1.2 < \eta_{3(4)} < 1.8 / 2.8 < \eta_{4(3)} < 3.7$ 10 $\textbf{2.8} < \eta_{3~(4)} < \textbf{3.7} \ / \ \textbf{2.8} < \eta_{4~(3)} < \textbf{3.7}$ do/dx1 (do/dx2) (pb) (EEMC / FCS) (FCS / FCS) 10 10 0.15 0.15 (EEMC / FCS) (FCS / FCS) x 10⁻² -2 10⁶ 10⁶ x 10 0.1 0.1 10 10 0.05 0.05 10 ' 10 0 10 10 ³ 0 10² -0.05 10 -0.05 GRSV STD 10 10 GRSV STD -0.1 -0.1 1 1 [']10 ⁻⁵ 10 ⁻⁴ 10 ⁻³ 10 ⁻² 10 ⁻¹ 1 10 ⁻⁵ -0.15 10⁻¹ 10 1 10 15 20 25 35 10 11 12 13 14 15 16 17 18 19 20 **10**-3 X1 (X2) M (GeV) $x_{1}(x_{2})$ M (GeV) Probe gluons to $x \sim 10^{-3}$ An attractive probe at rather low *x* before the EIC era

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STAR

eRHIC and eSTAR (>2025) will offer unprecedented reach in Q^2 and x

STAR

- Δg coming into focus, but still poorly constrained at low x
- Can **push to lower** *x* **with forward measurements**, and higher sqrt(s)
- Proof of principle result exists with $\pi^0 A_{LL}$ in endcap ($0.8 < \eta < 2.0$)
 - PRD 89, 012001 (2014)
 - Work underway with large 2012 dataset at 510 GeV
- Preliminary $\pi^0 A_{LL}$ with forward calorimeters (~3 < η < ~4)
 - Work underway with large 2012 and 2013 datasets at 510 GeV
- Preliminary **dijet cross sections**
 - Exciting prospects for very low x gluons with upgraded calorimeter at forward η
- Large datasets on hand, analyses underway
 - 2012, 2013
- Detector upgrades continue
 - Forward calorimetry: FPS+FMS 2015, FCS 2020
- Continuing data taking planned
 - 2015, 2016, and beyond
- Stay tuned!

Backup

• DSSV++ arXiv:nucl-ex/1304.0079

• DSSV 14 PRL **113**, 012001 (2014)

FIG. 1 (color online). Gluon helicity distribution at $Q^2 = 10 \text{ GeV}^2$ for the new fit, the original DSSV analysis of [3], and for an updated analysis without using the new 2009 RHIC data sets (DSSV*, see text). The dotted lines present the gluon densities for alternative fits that are within the 90% C.L. limit. The *x* range primarily probed by the RHIC data is indicated by the two vertical dashed lines.

Strip Index

- EM Particle Reconstruction Procedure
 - Identify clusters in the u and v strips
 - Determine which u and v clusters to associate with incident particles
 - Compute energy of incident particles (e.g. photons) from the towers
 - Compute momentum from the vertex and SMD cluster positions
- SMD response (right) in π^0 candidate event from data
 - Blue histograms show energy response per strip
 - Red triangles represent clusters drawn at mean strip position, and 10% of the cluster energy
- SMD clusters are found by
 - Smoothing the histogram using the method of J. Tukey
 - Identifying clusters as a strip above an energy threshold, with +-3 adjacent strips with monotonically decreasing energy
 - Setting cluster position to energy-weighted mean position of strips
- EM particle candidates built from pairs of u-v clusters
 - Clusters matched by energy of u and v strips
 - Required to have associated tower energy above threshold
 - Often have e.g. two photons from one π^0 deposited in one tower
- Reconstruction difficulties include
 - Upstream passive material: π^0 opening angle on the same order as photon conversions
 - Single particles sometimes look like two particles, and vice versa

[GeV]

Energy [

0.002

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STAR 0

- day showing a glimpse at only a tiny fraction of the 2009 datasets
- Larger jet trigger patch allows events with more jet background, softer π^0
- Background somewhat higher than for photon triggers
 - But can probe to considerably lower π^0 pT
 - Very reasonable π^0 peak
 - Possibility to extend reach to lower x?
- Work remains on MC validation, understanding of π^0 's in "jettier" environment, etc.

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Relativistic Heavy Ion Collider as a Spin Collider

Concert of Facilities

• OPPIS \rightarrow LINAC \rightarrow AGS \rightarrow RHIC

Polarized-proton Collider

- Mitigate effects of depolarization resonances with "Siberian Snakes"
- Polarization measured with CNI polarimeter
- Spin rotators provide choice of spin orientation *independent of experiment*

RHIC Beam Characteristics

- Clockwise beam: "blue"; counter-clockwise beam: "yellow"
- Spin direction varies bucket-to-bucket (9.4 MHz)
- Spin pattern varies fill-to-fill

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STAR's Endcap Electromagnetic Calorimeter

SS hub

back

post

showe

Pb/SS

plastic

scint

.086

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 - Covers $1.09 < \eta < 2.00$ over full 2π azimuth
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- Scintillating strip SMD
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 - Two active planes
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- Resolution of a few mm

STAR Detector in 2014

HFT: Heavy Flavor Tracker, MTD: Muon Telescope Detector

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STAR Detector in 2015-2016

FMS: Forward Meson Spectrometer, **FPS**: Forward Preshower, **RP II***: Roman Pot Phase II*

STAR Detector in 2018-2019

iTPC: inner TPC, EPD: Event Plane and Centrality Detector, ETOF: End-cap TOF, Fixed Target

STAR Detector in 2021-2022

FCS/FTS: Forward Calrimeter/Tracking System, RP II: Full Roman Pot Phase II

STAR Detector in 2025+

CEMC: Central EM Calorimeter, **eTRK:** electron Tracker, **TRD:** Transition Radiation Detector ΔΔ

eSTAR Detector in 2025+

The very successful STAR detector will evolve into an EIC detector 61

Proton Helicity Structure

