Why does black hole describe the deconfinement phase?

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M.H.-Hyakutake-Nishimura-Takeuchi, PRL (2009) M.H.-Hyakutake-Ishiki-Nishimura, Science (2014) M.H.-Maltz-Susskind, hep-th (2014) + work in progress.



 $\lambda = \infty$, $N = \infty$ corresponds to supergravity.

assumed to be correct without proof, and applied to QGP etc

SYM

STRING

large-N, strong coupling

large-N, finite coupling

finite-N, finite coupling SUGRA

tree-level string (SUGRA+α')



Quantum string (g_{string}>0)

SYMdifficult

large-N, strong coupling

large-N, finite coupling

finite-N, finite coupling

STRING

SUGRA easier

tree-level string (SUGRA+α') more difficult Quantum string (g_{string}>0) very difficult

SYMdifficult

STRING

large-N, strong coupling

large-N, finite coupling

finite-N, finite coupling SUGRA easier

tree-level string (SUGRA+α') more difficult Quantum string (g_{string}>0) very difficult

STRING **SYM**_{difficult} **SUGRA** large-N, strong coupling easier tree-level string large-N, (SUGRA+ α ') finite coupling more difficult Quantum string finite-N, (g_{string}>0) finite coupling very difficult The opposite (actually the <u>original</u>)

direction of the dictionary is useful!



.... and string should tell us more about real SU(3) QGP.

(1) Is it correct?

(1-1)Is it correct only at large-N, strong coupling?

(supergravity, or Einstein gravity)

(1-2) Or correct including $1/\lambda$ and 1/N corrections? (superstring theory)

(2) If correct, why? Can we understand it intuitively?

Important problems for particle & nuclear physics, because

 We want to understand quantum gravity, which is an important piece of physics beyond standard model.

— We want to understand thermalization of QGP.





(Maldacena 1997; Itzhaki-Sonnenschein-Maldacena-Yankielowicz 1998)

Quantitative test is possible by studying SYM numerically.

Is it correct?



M.H.-Hyakutake-Nishimura-Takeuchi, PRL 2009

 $(\lambda^{-1/3}T : dimensionless effective temperature)$

Maldacena conjecture is correct at finite coupling & temperature!



M.H.-Hyakutake-Nishimura-Takeuchi, PRL 2009



Dual gravity prediction (Y. Hyakutake 2013)

$$E/N^{2} = 7.41T^{2.8} - 5.58T^{4.6} + \dots + (1/N^{2})(-5.77T^{0.4} + aT^{2.2} + \dots) + (1/N^{4})(bT^{-2.6} + cT^{-2.0} + \dots) + \dots + \dots$$

Can it be reproduced from YM?



M.H.-Hyakutake-Ishiki-Nishimura, Science 2014



Simulations back up theory that Universe is a hologram

A ten-dimensional theory of gravity makes the same predictions as standard quantum physics in fewer dimensions.

'most-read news in 2013' @ nature.com

1,200,000 hits in less than one week

Gauge/gravity duality in tabloids



- Holographic principle claims gravity comes from thin, vibrating strings
- These strings are holograms of events that take place in a flatter cosmos
- According to this theory, everything we experience can be described as
 events that take place in this flatter location
- This is the first time the validity of the model has been mathematically tested

By ELLIE ZOLFAGHARIFARD

PUBLISHED: 08:03 EST, 12 December 2013 | UPDATED: 09:43 EST, 12 December 2013

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The universe is a hologram and everything you can see - including this article and the device you are reading it on - is a mere projection.

This is according to a controversial model proposed in 1997 by theoretical physicist Juan Maldacena.

Until now the bizarre theory had never been tested, but recent mathematical models suggest that the mind-boggling principle could be true.





shared 34 timesmore than Justin Bieber's marijuana scandal.



1600 times more than Lady Gaga

O(1,000,000) shares/tweets/comments

not counted as citations?

Other Simulations

- Two-point functions (M.H.-Nishimura-Sekino-Yoneya, PRL 2010, JHEP 2011)
- Polyakov loop (Anagnostopoulos-M.H.-Nishimura-Takeuchi, PRL 2008; M.H.-Miwa-Nishimura-Takeuchi, PRL 2009)
- 2d SYM vs black string (Catterall-Joseph-Wiseman, JHEP 2010)
- 4d SYM vs AdS₅ (Honda et al 2011–, Catterall et al 2012–)

Why is it correct?

Can we understand it intuitively?

microscopic descriptions of the black hole (black brane)

(1) D-branes + open strings Polchinski, ...

(2) condensation of closed strings

Susskind, Horowitz-Polchinski, ...



Black hole from closed string

(e.g. Susskind 1993)



Consider a long, winding string with length L. energy = tension \times L entropy \sim L

when L >> 1, huge energy and entropy are packed in a small region $\rightarrow \underline{black \ hole}$



On D-dim square lattice,

of possible shapes \sim (2D-1)^L entropy \sim L×log(2D-1)

How are they related?

long, winding strings = black brane + open strings



The meaning of N (# of D-branes) becomes clear later.

Gauge theory description

confining phase: 't Hooft, 1974 deconfining phase: M.H.-Maltz-Susskind, 2014 Lattice gauge theory description at strong coupling

Understand it by using the Hamiltonian formulation of lattice gauge theory (Kogut-Susskind, 1974)

$$H = \frac{\lambda N}{2} \sum_{\vec{x}} \sum_{\mu} \sum_{\alpha=1}^{N^2} \left(E^{\alpha}_{\mu,\vec{x}} \right)^2 + \frac{N}{\lambda} \sum_{\vec{x}} \sum_{\mu < \nu} \left(N - \operatorname{Tr}(U_{\mu,\vec{x}}U_{\nu,\vec{x}+\hat{\mu}}U^{\dagger}_{\mu,\vec{x}+\hat{\nu}}U^{\dagger}_{\nu,\vec{x}}) \right)$$
$$[E^{\alpha}_{\mu,\vec{x}}, U_{\nu,\vec{y}}] = \delta_{\mu\nu} \delta_{\vec{x}\vec{y}} \cdot \tau^{\alpha} U_{\nu,\vec{y}}$$

$$\sum_{\alpha=1}^{N^2} \tau_{ij}^{\alpha} \tau_{kl}^{\alpha} = \frac{\delta_{il} \delta_{jk}}{N^2}$$
Hilbert space is expressed by
Wilson loops.
(closed string)





splitting ~ 1/N joining ~ 1/N 1/N² for each loop of closed strings



Strings out of YM: <u>deconfining</u> phase

M.H.-Maltz-Susskind, 2014

Hilbert space is always the same. Why don't we express the deconfining phase by using Wilson loops?

• interaction (joining/splitting) is 1/N-suppressed

"large-N limit is the theory of free string"

- It is true when L is $O(N^0)$. (\rightarrow confining phase)
- In deconfinement phase, total length (energy) of the strings is O(N²) → number of intersections is O(N²) → interaction is **not** negligible

large-N limit is still very dynamical!

confining phase = gas of short strings

long and winding string, which is interpreted as BH, appears





as the density of strings increase, interaction between strings becomes important,and...

Why L \sim N²?

• Tr(UU'U''....) length \gtrsim N² \longrightarrow factorizes to shorter traces

> N² is the upper bound. Beyond there, the counting changes; <u>not much gain for the entropy.</u>



long, winding QCD-strings = black brane + open QCD-strings



open strings = Wilson lines, which have N color d.o.f at endpoints \rightarrow black brane is made from N Dp-branes

D-dim square lattice at strong coupling

deconfinement temperature



Real-time study of BH thermalization

(in progress)

matrix models for black holes









 $U_1 \bigcirc \bigcirc \bigcirc$ $U_2 \bigcirc \bigcirc \bigcirc$ $U_1, U_2 \bigcirc \bigcirc \bigcirc$





it 'thermalizes' soon; better thermalization with more links.





Maldacena's conjecture is correct at finite temperature, including 1/λ and 1/N corrections, at least to the next-leading order.

so, lattice/nuclear theorists can study quantum gravity, by studying field theory. You can do something string theorists cannot do.



RHIC/LHC are machines for quantum gravity!

conclusion(2)

deconfinement _____ phase



Strong coupling limit contains the essence. Time evolution can be studied.

Future directions

 Lattice gauge theory/nuclear physics + Black hole physics + Information science should provide us with new research field.

Complete solution of Hawking's paradox

 Cosmology from string theory by lattice gauge theory/nuclear physics methods

New applications of lattice/Nuclear techniques!!

• Gravity calculation beyond supergravity (full string theory) will tell us about finite-N, finite-coupling gauge theory, especially the real-time dynamics.





M.H.-Hyakutake-Ishiki-Nishimura, Science 2014



M.H.-Hyakutake-Ishiki-Nishimura, Science 2014

E/N² - (7.41T^{2.8}-5.77T^{0.4}/N²) vs. I/N⁴

