

# Amplitude Method – summary results Veta GHENESCU, Titi PREDA

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# Outline

- ✓ Test Beam set-up
- ✓ Amplitude Method
  - Amplitude algorithm
  - Signal candidates
  - MPV uniformity in space&time
  - Signal to noise ratio
- ✓ Edge effects:
  - Track reconstruction
  - Induced signal in neighboring pads
  - Results
- ✓ Shower analysis
  - Longitudinal & transversal evolution of showers
- ✓ Conclusions





- DESY II Synchrotron provide electrons with up to 1000 particles per cm<sup>2</sup> and second, energies from 1 to 6 GeV;
- Test Beam took place in beam line 22 of DESY II ring in Hamburg, from 04.11.2011 to 22.11.2011;
- Used 2GeV or 4 GeV electron beams;





 $t = [2X_0 \div 14X_0];$ 

step =  $2X_0$ 



- ZEUS telescope planes (1, 2, 3):
  - Si planes: 300  $\mu m$  thick
  - Active area: 32 x 32mm<sup>2</sup>
  - Double perpendicular layers,
  - 640 strip channels (50µm)
- > Trigger scintillators (4,5) :
  - Trigger window: 7 x 7mm<sup>2</sup>
- BeamCal Sensor (6)
  - GaAs:Cr sensor



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## Amplitude method (MAX)

#### 1. Data

At each trigger the ADC counts stored in a matrix  $A = ||A_{i,j}||, i \in [0, 31], j \in [0, 31]$ 



2. Amplitude method finds for every pad (channel) the maximum of ADC counts for  $j \in [20, 31]$ 

 $Amax_i = MAX(A_{i,20}, ..., A_{i,31}) \longrightarrow Amax_i = A_{i,smax}$ , where smax represents the sample corresponding to  $Amax_i$ 

**3.** Pedestals are computed for every run and every pad (channel) in the sample interval  $j \in [0,19]$ Pedestal values are the mean values,  $P_i$ , of the  $f_i(A_{i,0}, ..., A_{i,19})$  distributions which have the  $RMS_i$  values

$$f_i(A_{i,0}, \dots, A_{i,19}) = f_{i,0}(A_{i,0}) \times \dots \times f_{i,19}(A_{i,19})$$
, where  $i \in [0, 31]$ 

ISS

### Amplitude method (MAX)





1. The maximum count has to satisfy:

$$Amax_i \ge P_i + coef \times RMS_i$$
, we used  $coef = 3$  (1)

2. At least one of the nearest samples has the count:

$$A_{i,smax-1} \text{ or } A_{i,smax+1} \ge P(i) + coef \times RMS_i$$
 (2)

The Signal Amplitude is:

$$S_i = Amax_i - P_i$$

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#### Track reconstruction method

- DigXs and DigYs coordinates have been took for all telescope planes;
- Hits number/plane =  $1 \rightarrow$  one EM shower/event

#### where:

• (X<sub>im</sub>,Y<sub>im</sub>) = measured coordinates or given coordinates by TelAna, *i* € [1, 3];

• 
$$(X_{im}, Y_{im}) = (DigXs, DigYs);$$

•  $(X_{ip}, Y_{ip})$  = predicted coordinates given by line intersection with each telescope plane

$$Min (d^{2}) = Min\left(\sum_{i=1}^{3} \left( \left( x_{ip} - x_{im} \right)^{2} + \left( y_{ip} - y_{im} \right)^{2} \right) \right)$$



#### Track reconstruction method

#### **Spatial resolution:**

ollaboration recision design

- Sigma from fits are smaller than about  $30\mu m$
- The Si chamber alignment was make with a maximum 100 µm shift







ollaboration recision design







### The energy deposited dependence by tungsten radiation lengths for experimental data

## Shower analysis







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#### Conclusions

## In the paper could be included:

- Amplitude method
  - Algorithm
  - Results (uniformity, S/N)
- Edge effect
  - Track reconstruction
  - Results

