



ECHo Experiment



Loredana Gastaldo
for the ECHo collaboration

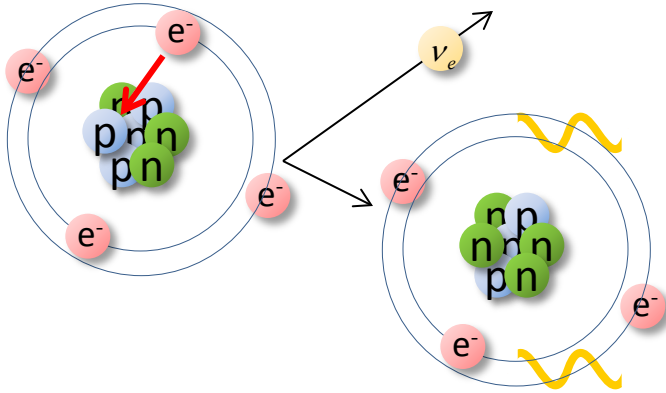


Contents

- Electron capture process: The case of ^{163}Ho
- Metallic Magnetic Calorimeters
- Recent results
- ECHo experiment

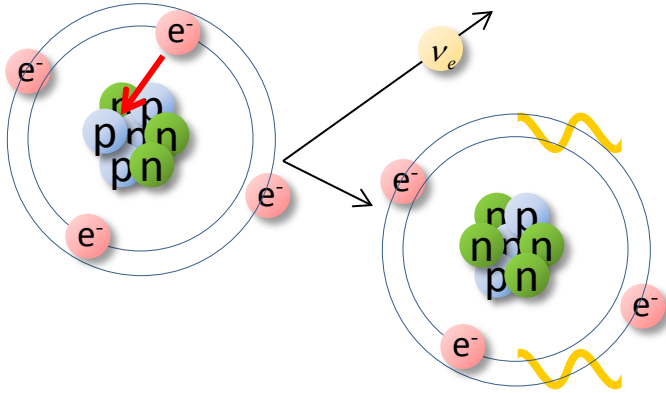


Electron Capture



A non- zero neutrino mass affects
the **de-excitation energy spectrum**

Electron Capture

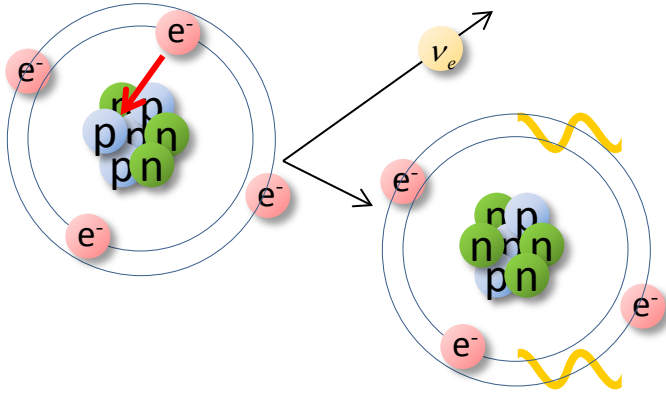


A non- zero neutrino mass affects the **de-excitation energy spectrum**

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Electron Capture



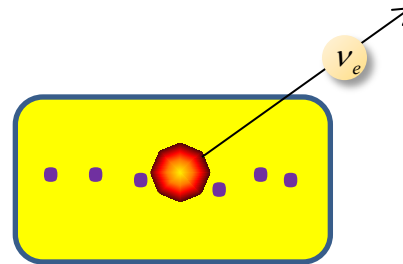
A non- zero neutrino mass affects the **de-excitation energy spectrum**

Atomic de-excitation:

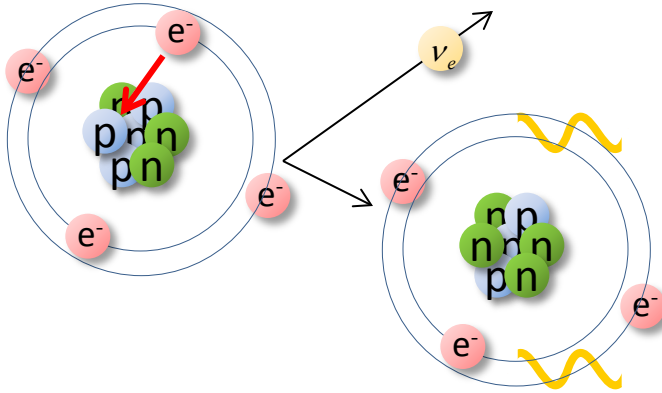
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- Coster-Kronig transitions



Calorimetric measurement



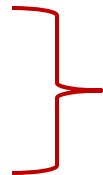
Electron Capture



A non- zero neutrino mass affects the **de-excitation energy spectrum**

Atomic de-excitation:

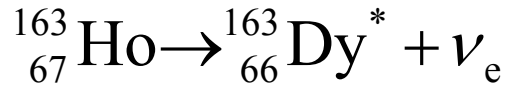
- X-ray emission
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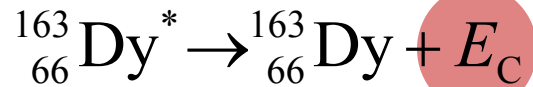
Calorimetric measurement

$$\frac{dW}{dE_C} = A(Q_{EC} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{EC} - E_C)^2}} \sum_H B_H \phi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$

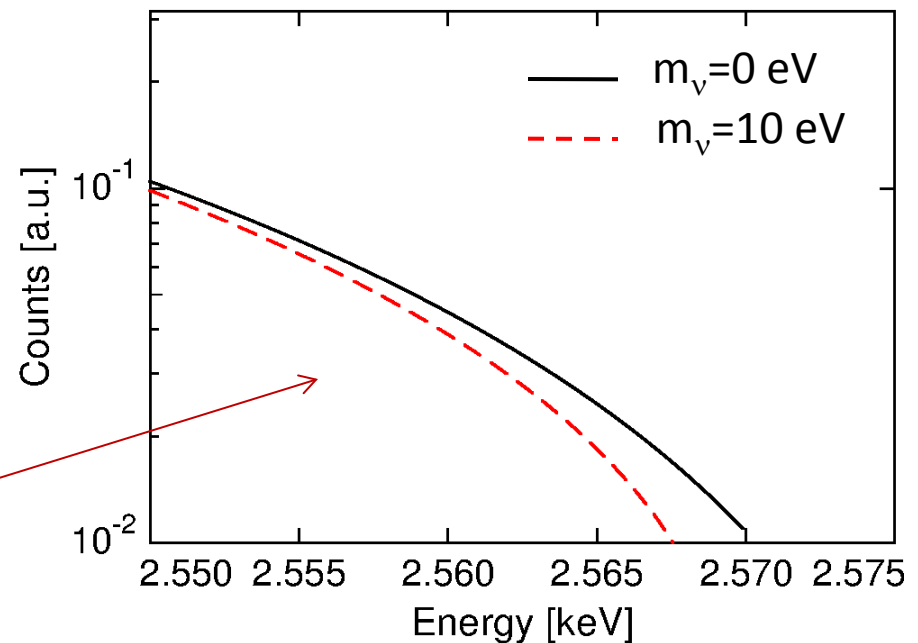
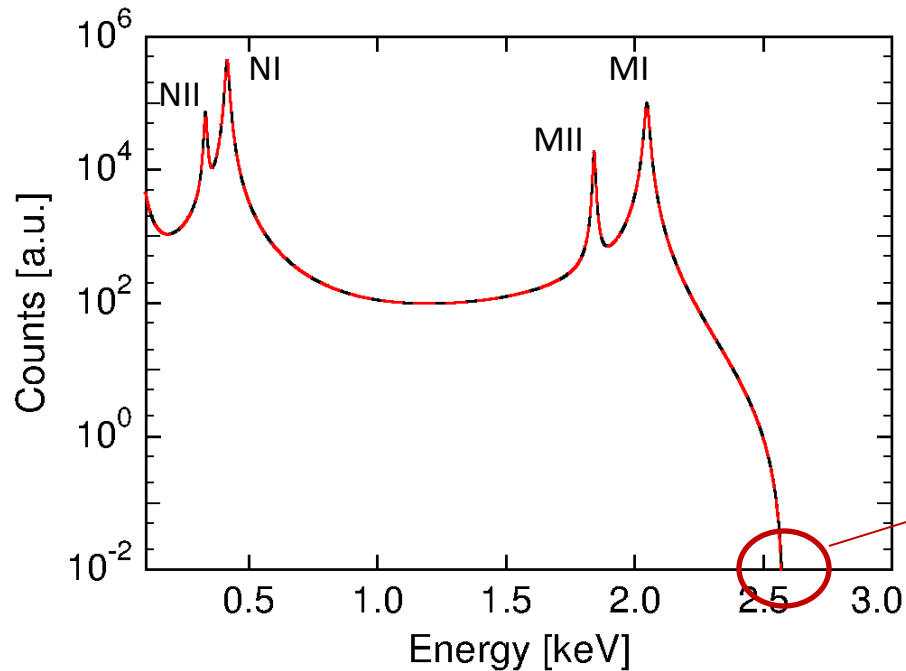
The case of ^{163}Ho



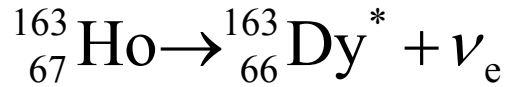
- $Q_{\text{EC}} \cong 2.5 \text{ keV}$



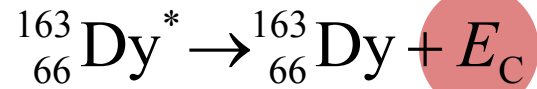
- $\tau_{1/2} \cong 4570 \text{ years}$



The case of ^{163}Ho



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9 December 1982

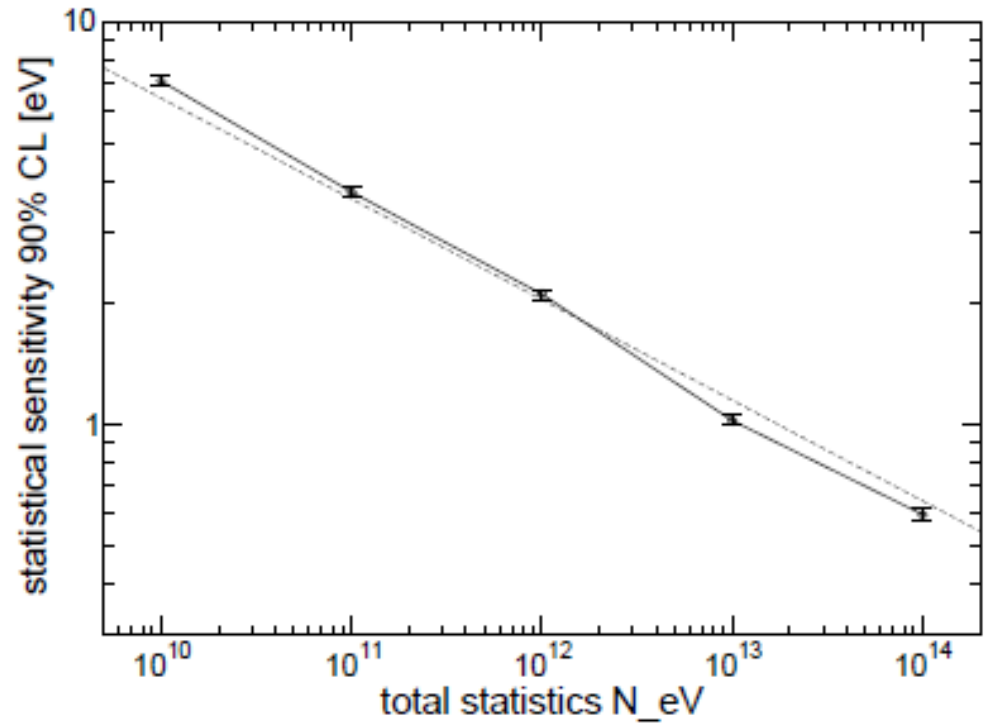
CALORIMETRIC MEASUREMENTS OF ^{163}Ho DECAY AS TOOLS TO DETERMINE THE ELECTRON NEUTRINO MASS

A. DE RÚJULA and M. LUSIGNOLI ¹

CERN, Geneva, Switzerland

The case of ^{163}Ho : Sub-eV sensitivity

Statistics

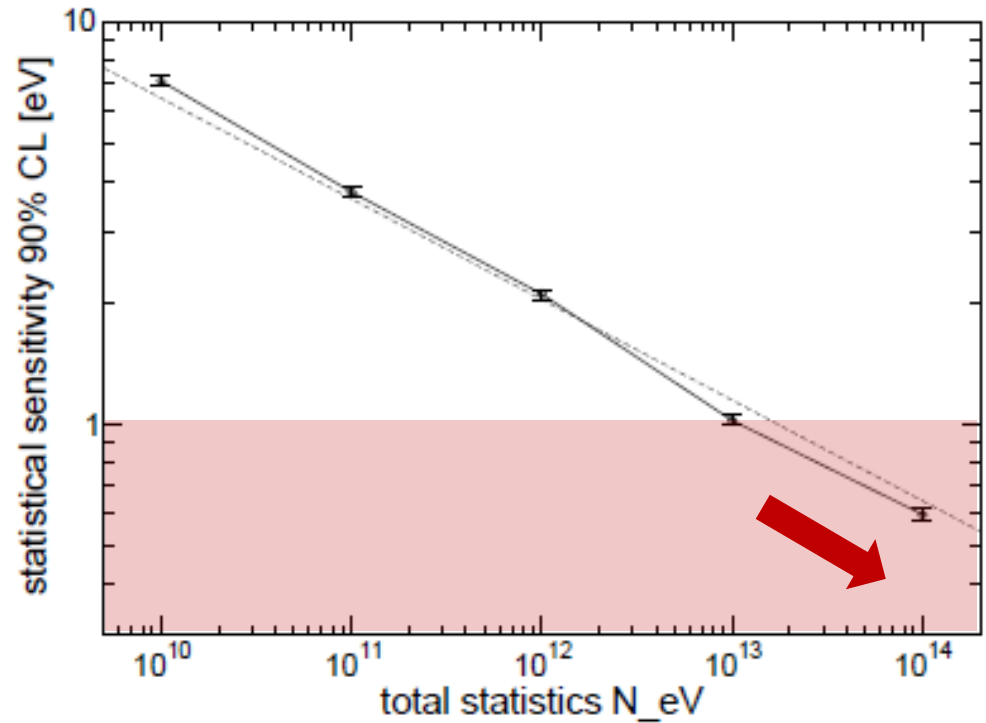


$$\Delta E_{\text{FWHM}} = 1 \text{ eV}, f_{\text{pp}} = 10^{-5}, Q_{\text{EC}} = 2600 \text{ eV}$$

The case of ^{163}Ho : Sub-eV sensitivity

$$N_{\text{ev}} > 10^{14}$$

Statistics

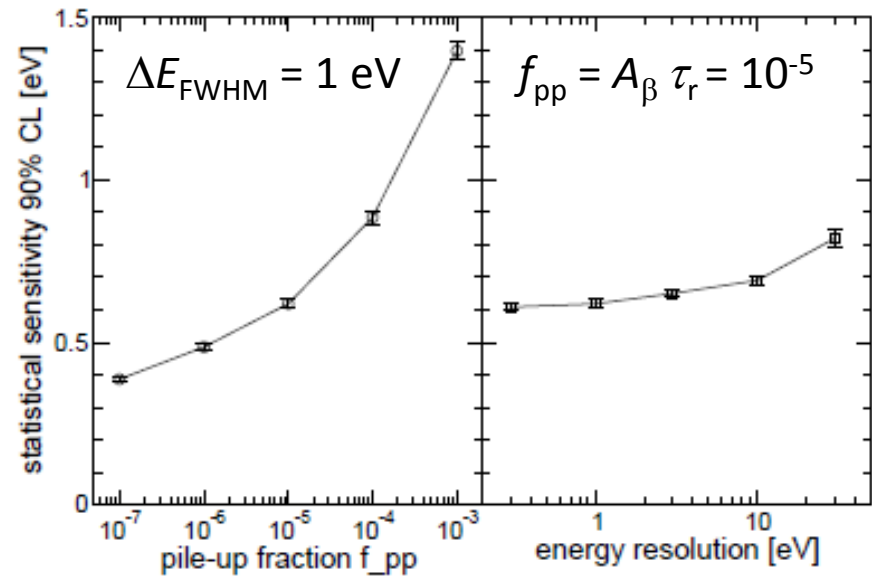


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The case of ^{163}Ho : Sub-eV sensitivity

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Detector performance



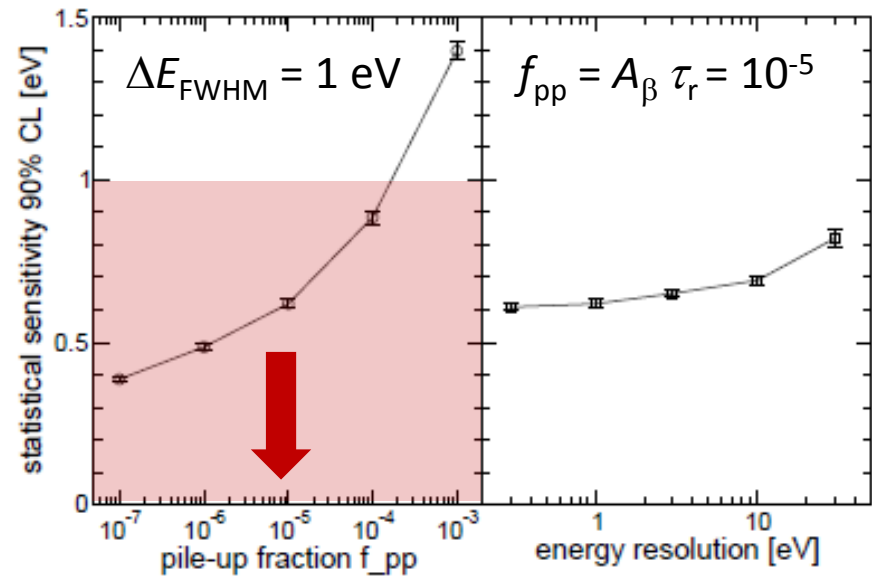
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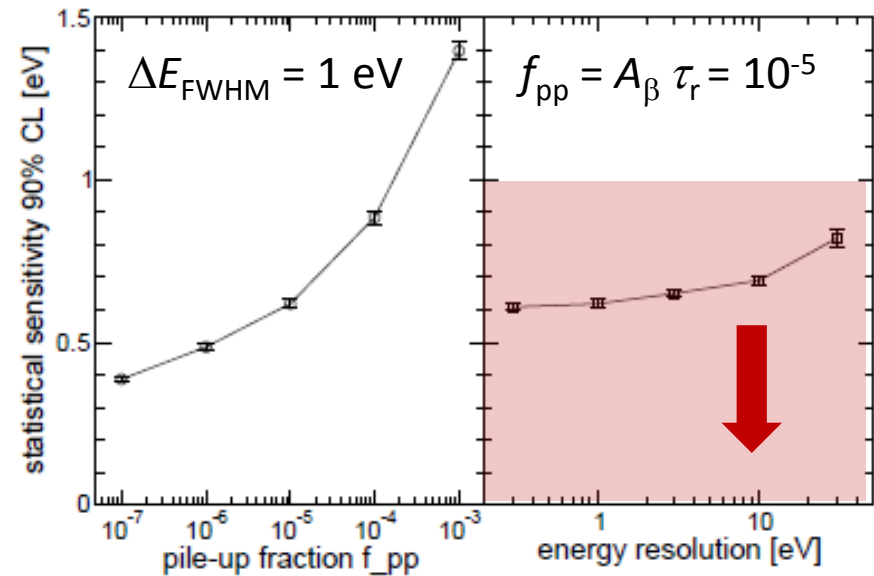
The case of ^{163}Ho : Sub-eV sensitivity

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$$f_{\text{pp}} < 10^{-5}$$

$$\Delta E_{\text{FWHM}} < 10 \text{ eV}$$

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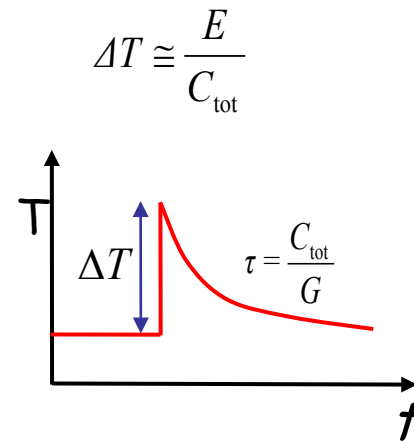
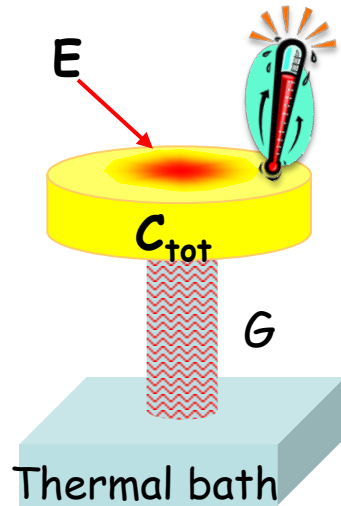
$$\begin{array}{l} N_{\text{ev}} > 10^{14} \\ f_{\text{pp}} < 10^{-5} \\ \Delta E_{\text{FWHM}} < 10 \text{ eV} \end{array} \longrightarrow \begin{array}{l} \tau_r \sim 0.1 \mu\text{s} \\ A_{\beta} \approx 10 \text{ s}^{-1} \end{array} \longrightarrow \geq 10^5 \text{ detectors}$$

The case of ^{163}Ho : Sub-eV sensitivity

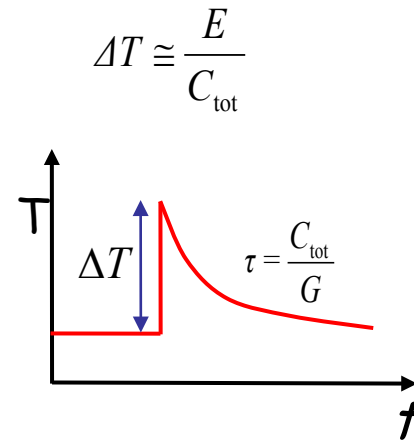
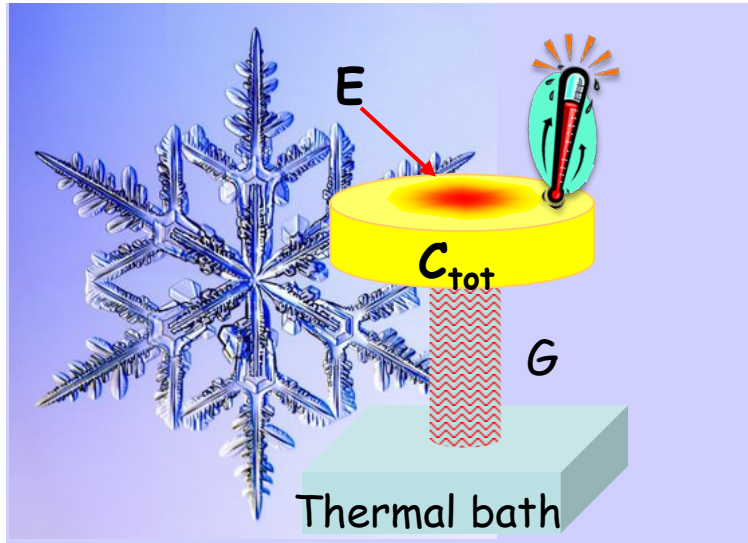
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Low temperature
Metallic Magnetic Calorimeter

MMCs: Concept

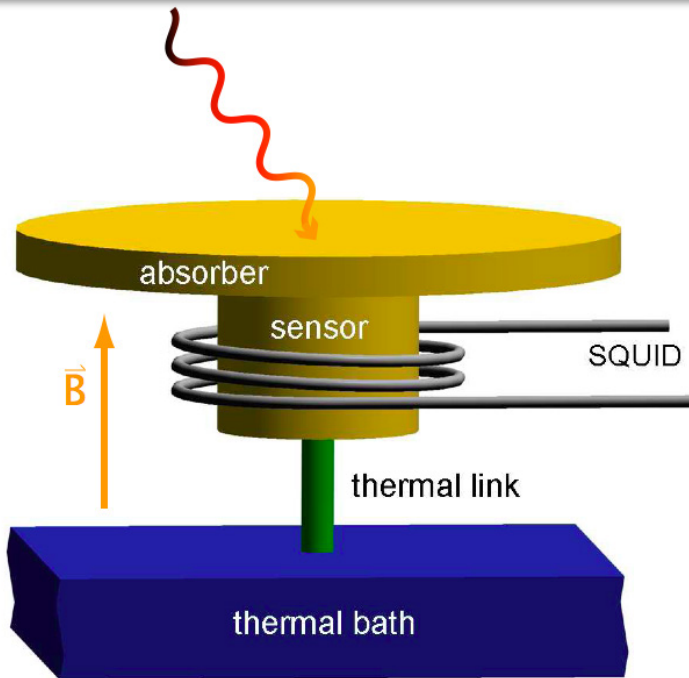


MMCs: Concept

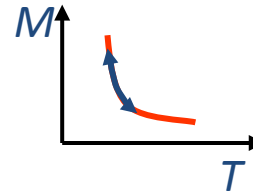


- Working temperature below 100 mK
small specific heat
large temperature change
small thermal noise
- Very sensitive temperature sensor

MMCs: Concept



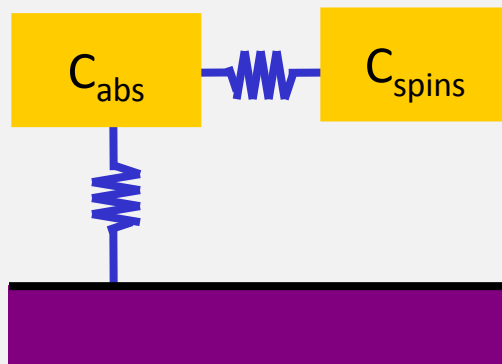
Paramagnetic sensor: **Au:Er_{500ppm}**



Signal size:

$$\delta M = \frac{\partial M}{\partial T} \delta T = \frac{\partial M}{\partial T} \frac{E_\gamma}{C_{\text{tot}}}$$

Energy resolution --- Why ,micro'-calorimeter



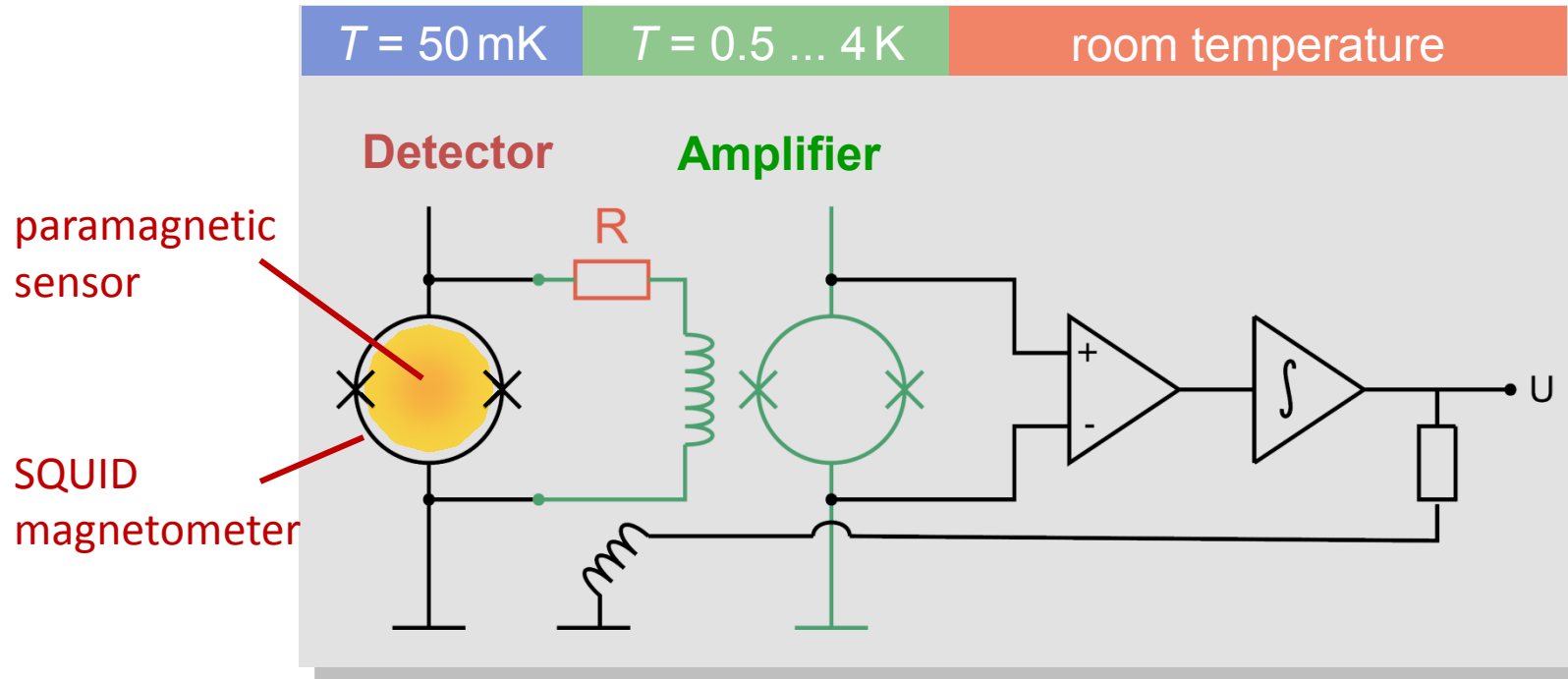
Thermal fluctuations of energy

between absorber, thermometer and bath lead to

$$\Delta E_{\text{FWHM}} \simeq 2,36 \sqrt{4k_B C_{\text{Abs}} T^2} \sqrt{2} \left(\frac{\tau_0}{\tau_1} \right)^{1/4}$$

e.g. **1eV** for $C = 1 \text{ pJ/K}$ at $T = 50 \text{ mK}$

MMCs: Readout

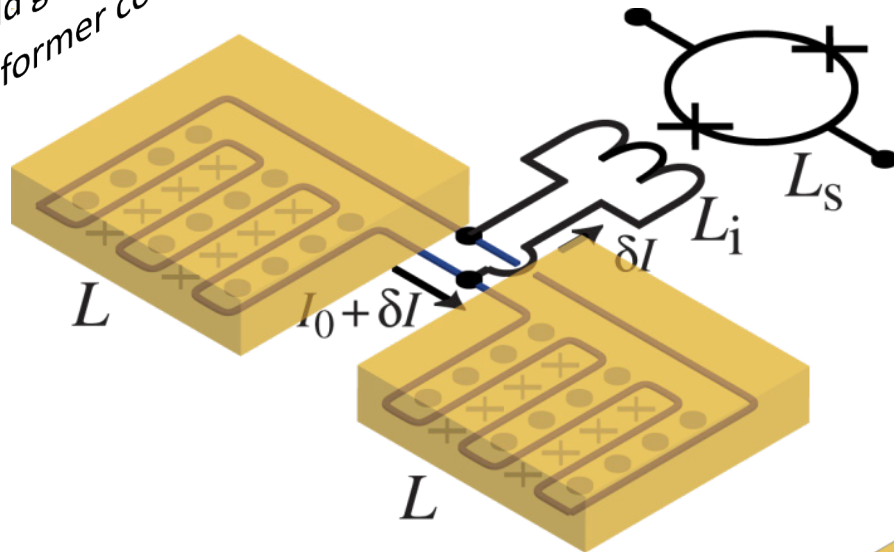


Two-stage SQUID setup with flux locked loop to linearize the first stage SQUID allows for:

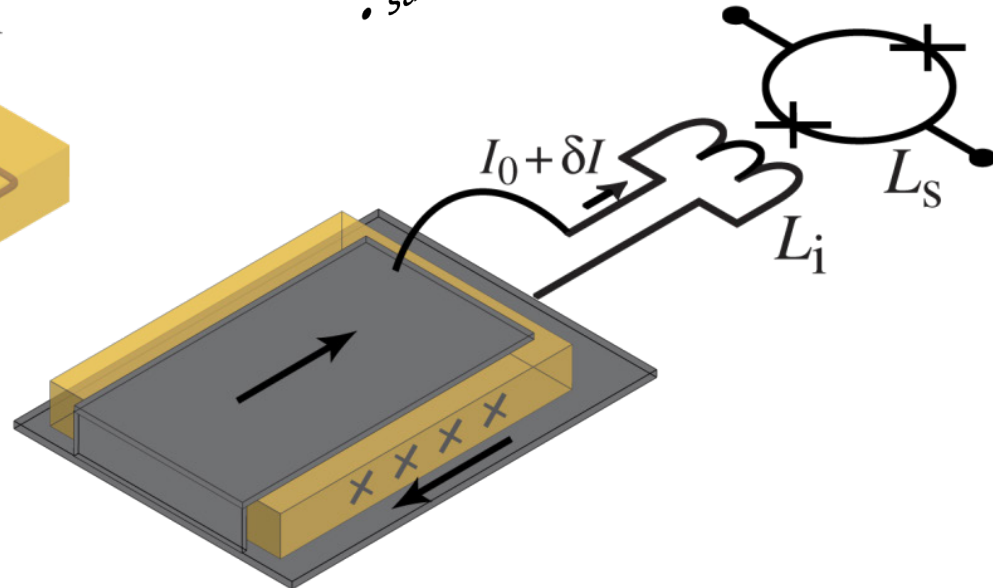
- low noise
- large bandwidth / slewrate
- small power dissipation on detector SQUID chip (voltage bias)

MMCs: Geometries

- Well established:**
- planar T -sensor
 - superconducting meander shaped pickup loop
 - B-field generated by persistent current
 - transformer coupled to SQUID

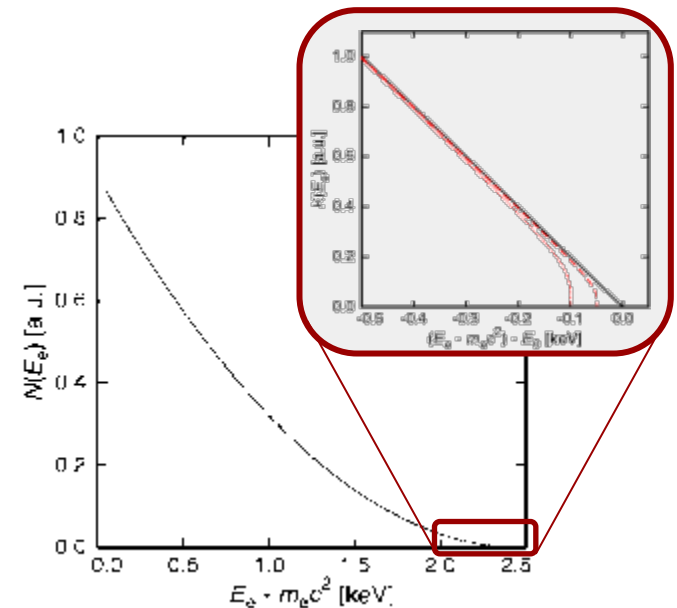
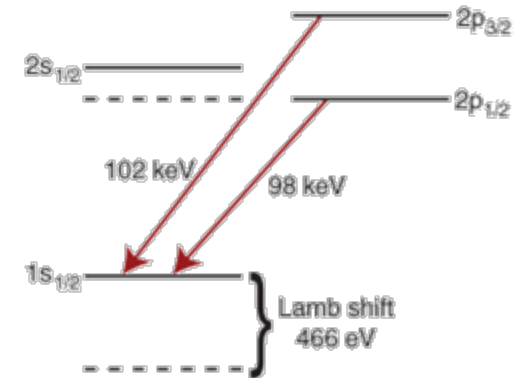
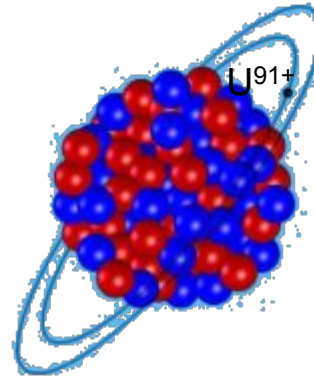


- New sandwich geometry:**
- best magn. flux coupling,
 - planar sensor
 - sandwiched between stripline

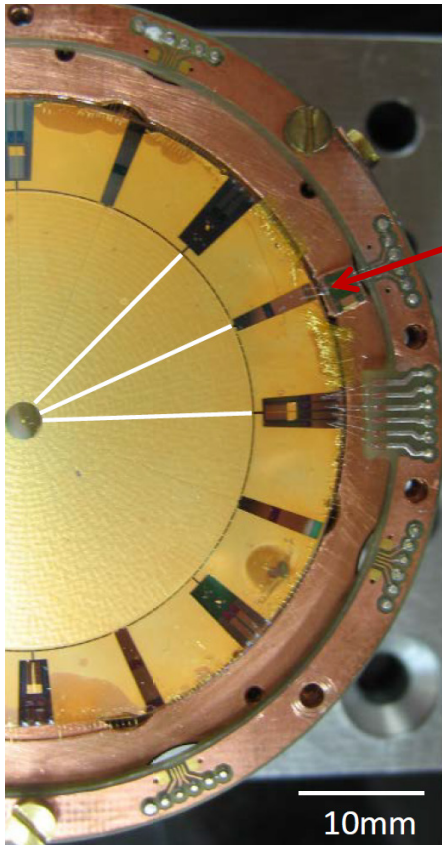


MMCs: Applications

- X-ray spectroscopy
 - atomic physics
 - astronomy
- X-ray imaging
 - large MMC arrays
 - microwave SQUID multiplexing
- Detection of molecular fragments
- Radiation standards for metrology
- Experiments for neutrino physics
 - β decay of ^{187}Re (MARE)
 - EC of ^{163}Ho (ECHO)
 - $0\nu 2\beta$ (AMoRE & LUMINEU)



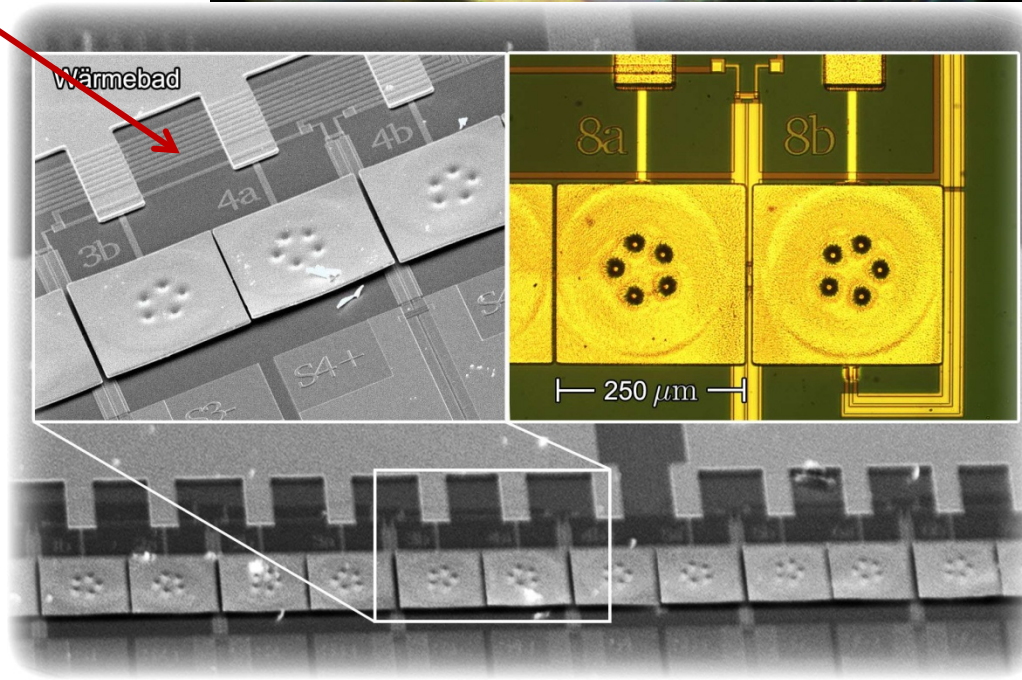
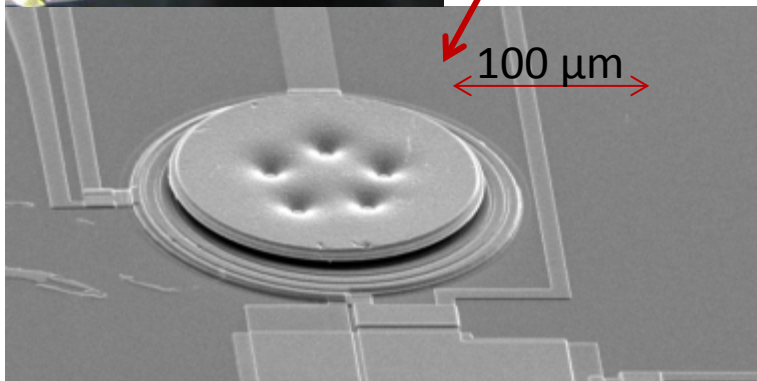
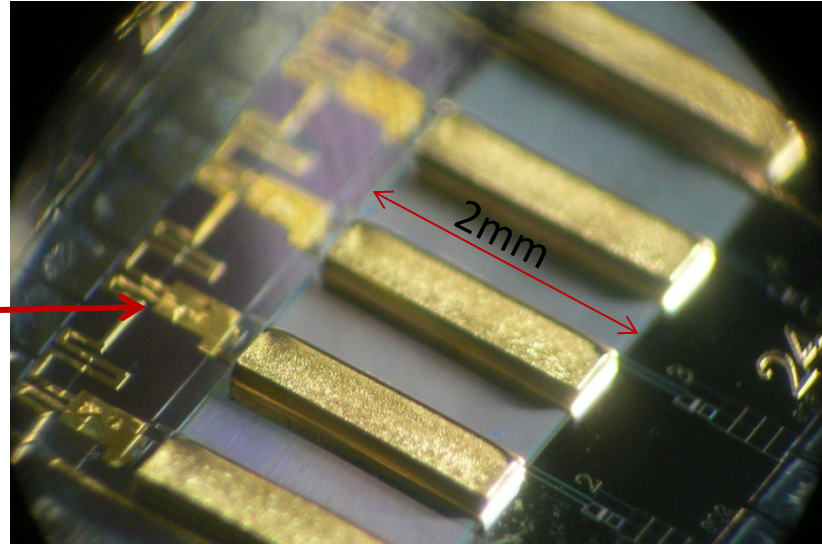
MMCs: Examples



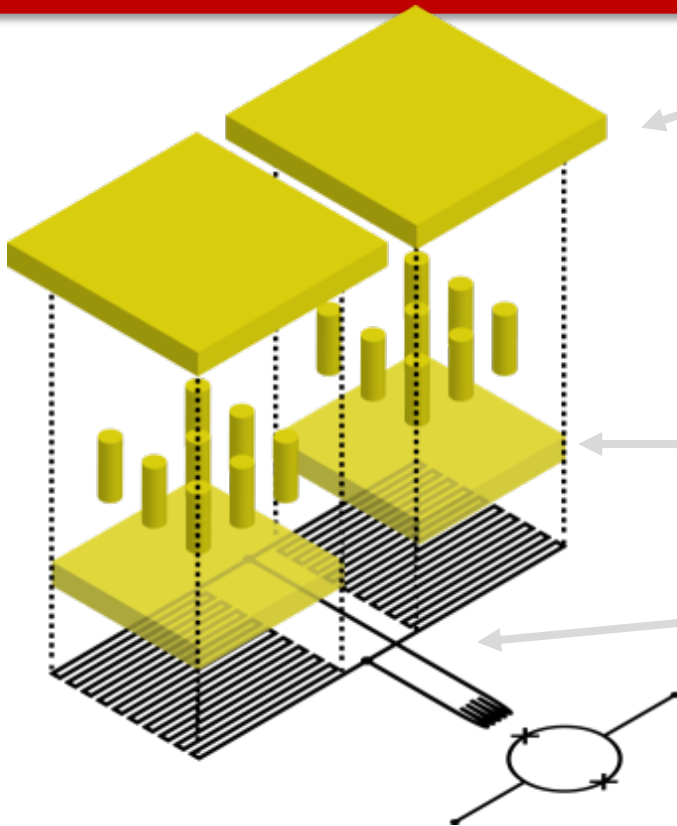
Pizza detector
for molecular fragments

Hard x-rays detectors

Soft x-rays
detectors



maXs: 1d-array for soft x-rays



- **1×8 x-ray absorbers**

- 250 μm ×250 μm gold, 5 μm thick
- >98% Qu.-Eff. @ 6 keV
- electroplated into photoresist mold (RRR>15)
- mech/therm contact to sensor by stems to prevent loss of initial hot phonons

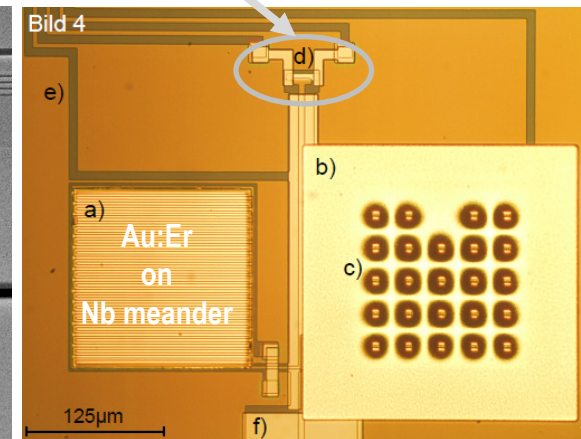
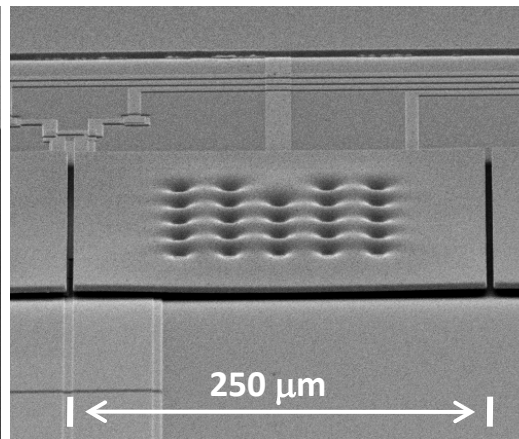
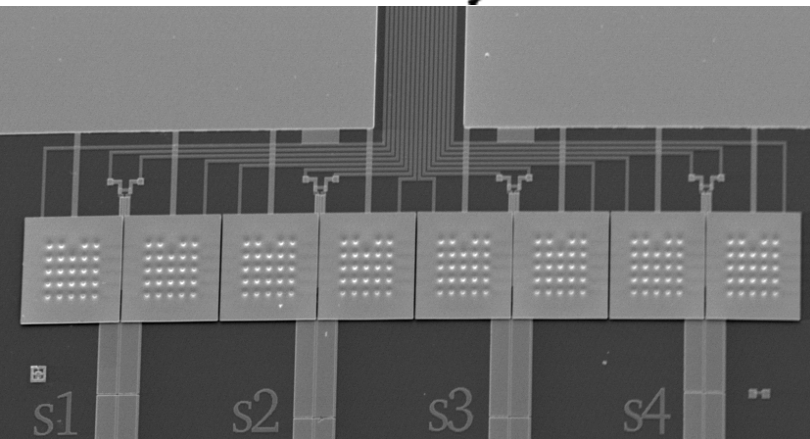
- **Au: $^{166}\text{Er}_{300\text{ppm}}$ temperature sensors**

- co-sputtered from pure Au and high conc. AuEr target

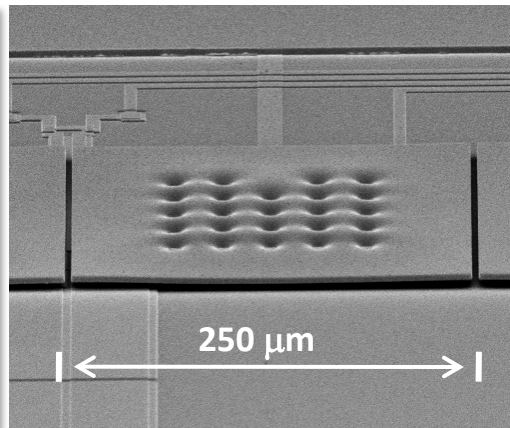
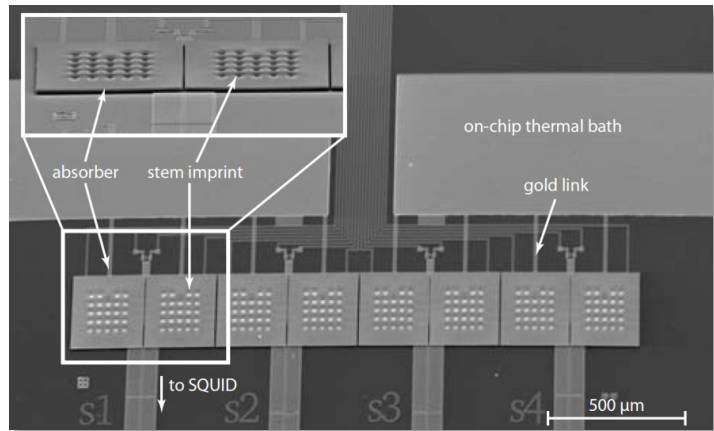
- **Meander shaped pickup coils**

- 2.5 μm wide Nb lines, ~400 nm thick
- $I_c \approx 100\text{mA}$

- **On-chip persistent current switch (AuPd)**

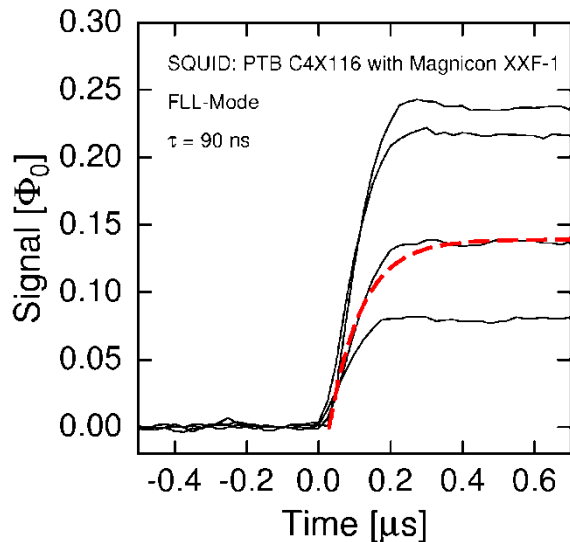


maXs: 1d-array for soft x-rays ($T=20$ mK)

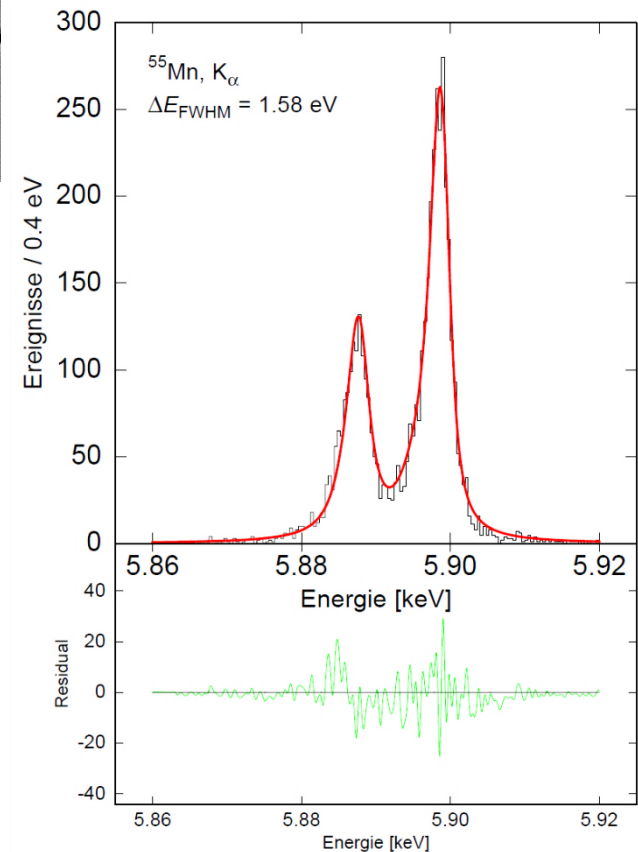
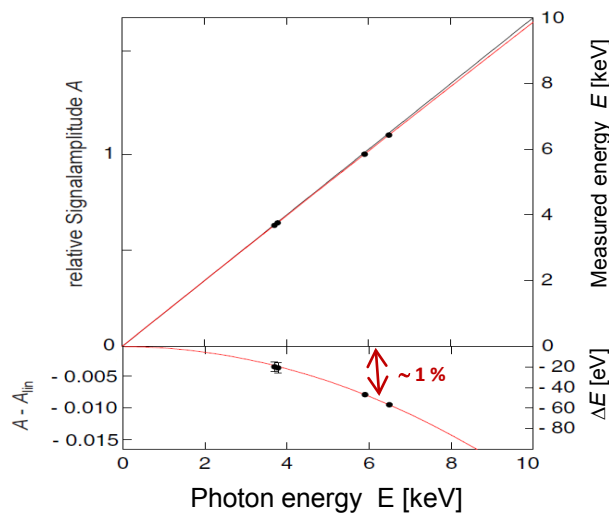


$$\Delta E_{\text{FWHM}} = 1.6 \text{ eV @ } 6 \text{ keV}$$

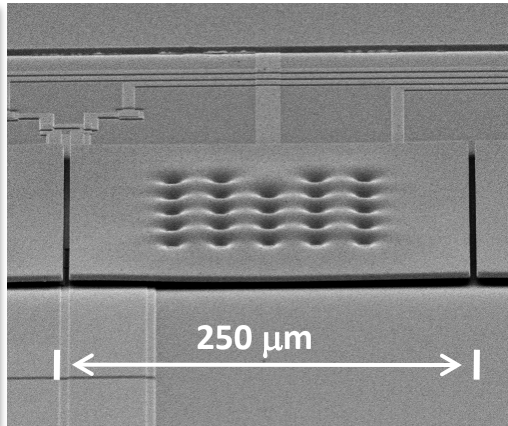
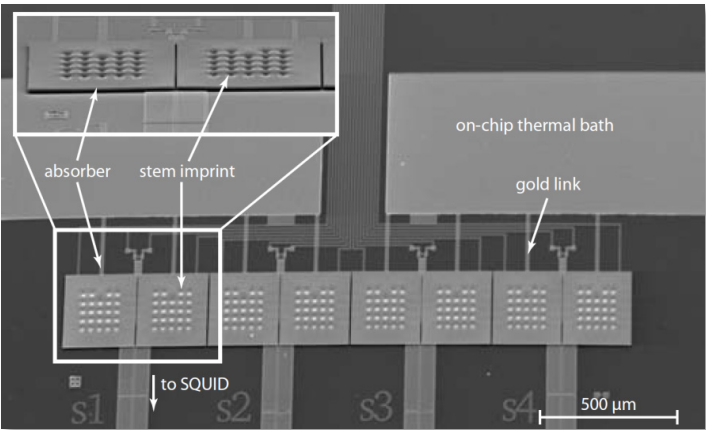
Rise Time: 90 ns



Non-Linearity < 1% @6keV

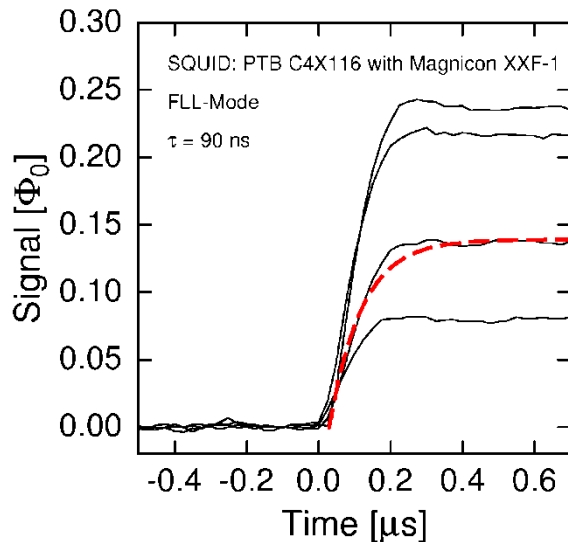


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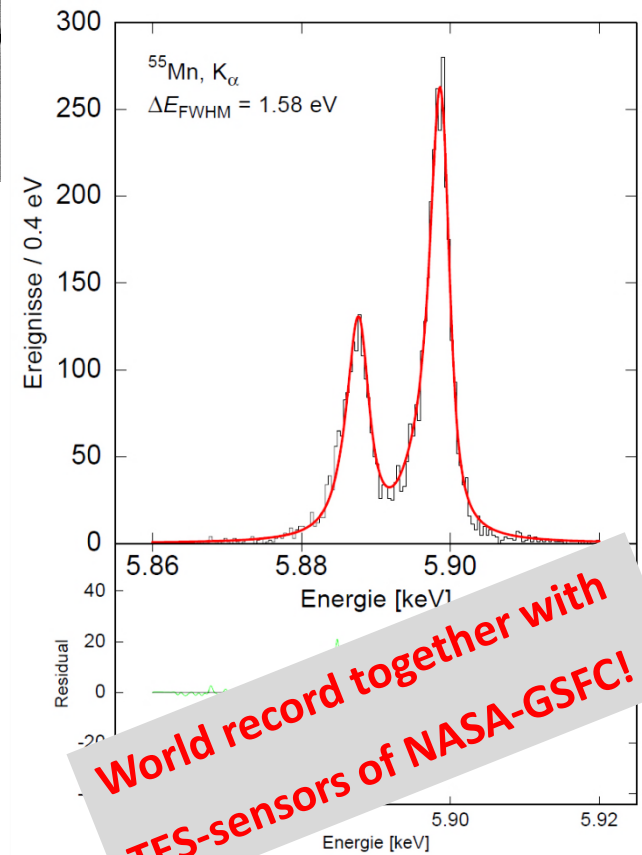
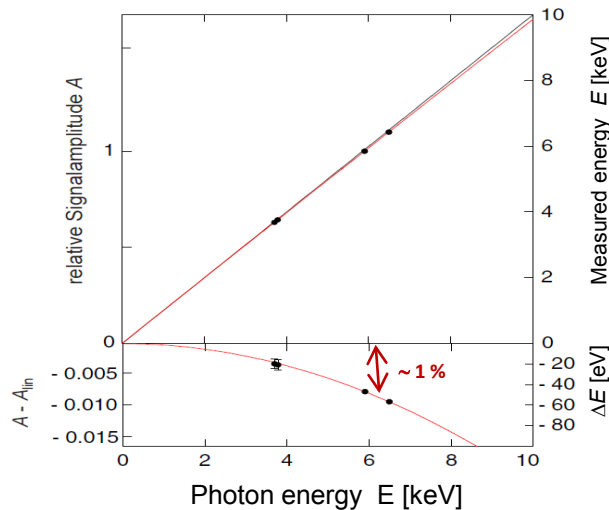


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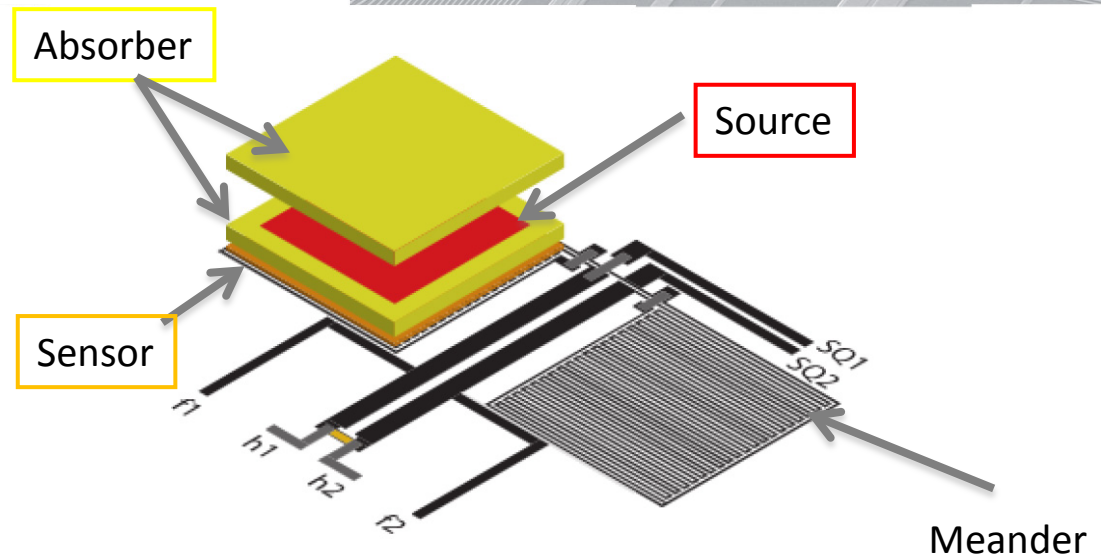
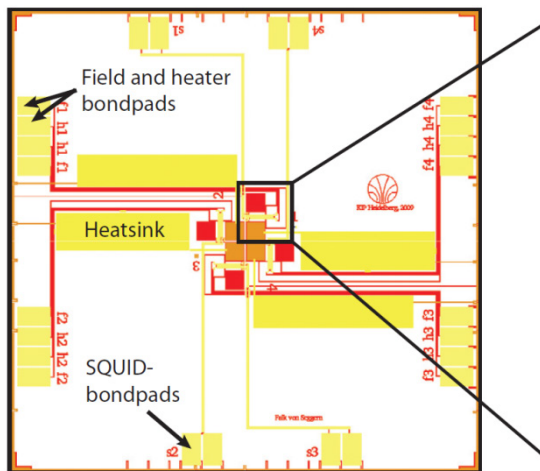
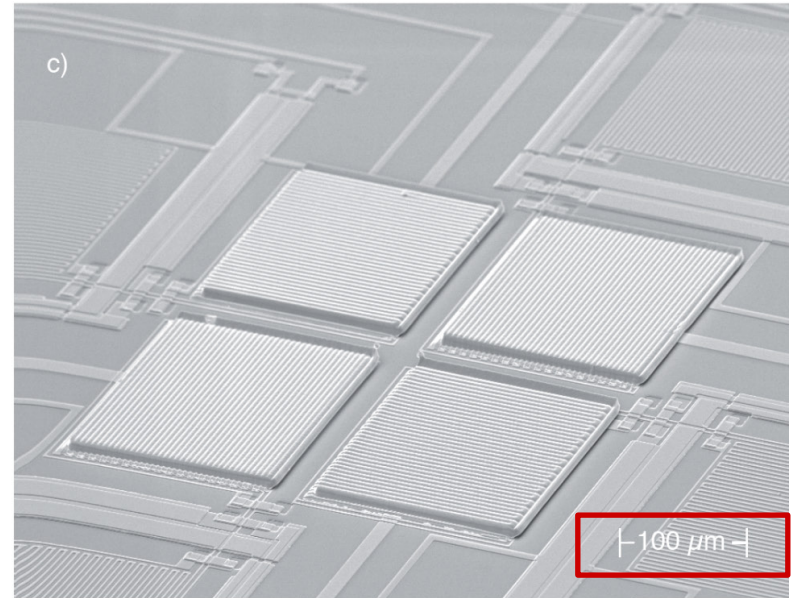
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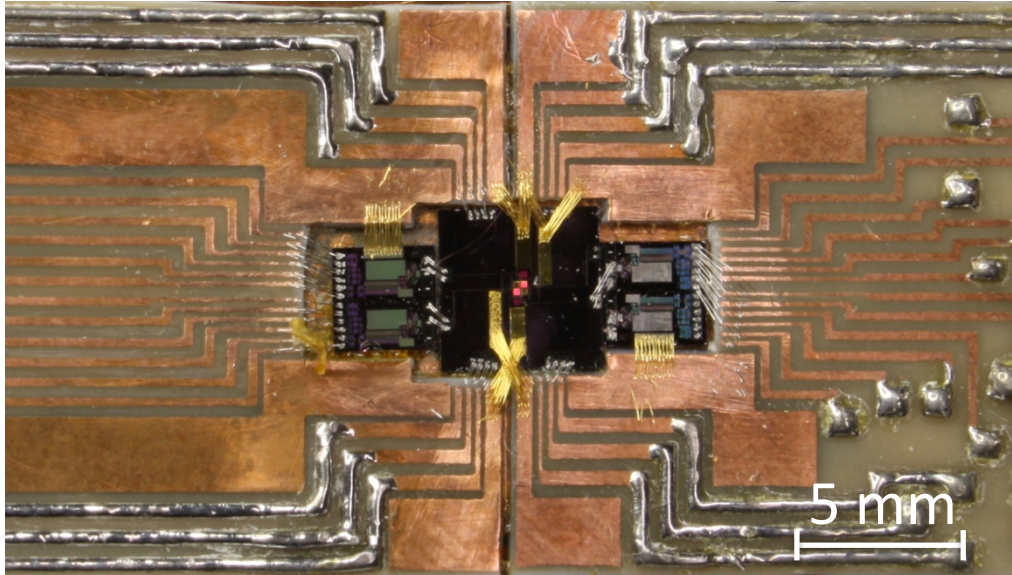
World record together with
TES-sensors of NASA-GSFC!

^{163}Ho Detector: First prototype

- Absorber for calorimetric measurement → ion implantation @ ISOLDE-CERN
- Two pixels have been simultaneously measured
- ^{55}Fe calibration source was collimated only on one pixel

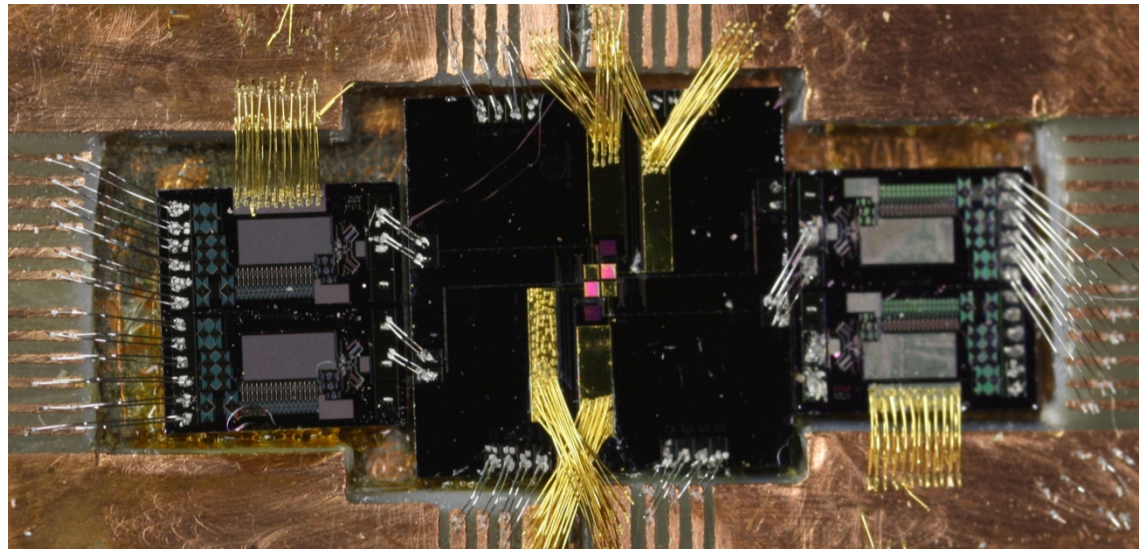


^{163}Ho Detector: First prototype

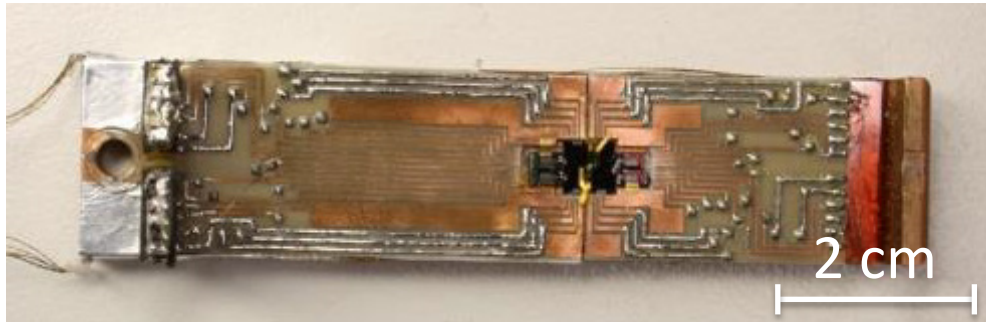


- Newest generation C6X114W SQUIDs (PTB Berlin)
→ best noise performance

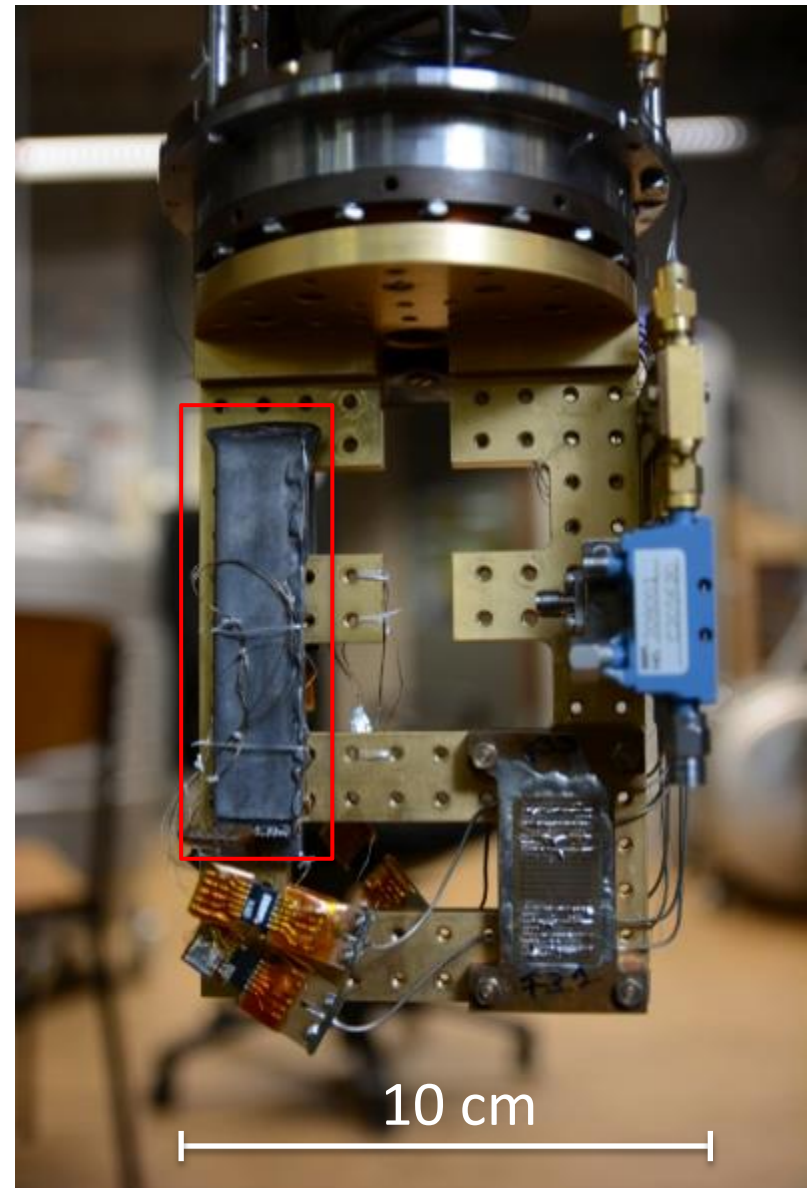
- Al bonding wires for electrical connections
- Au bonding wires for thermal connections



^{163}Ho Detector: Experimental environment



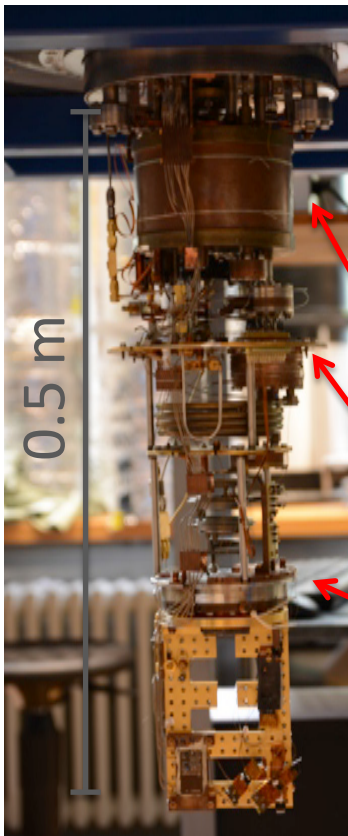
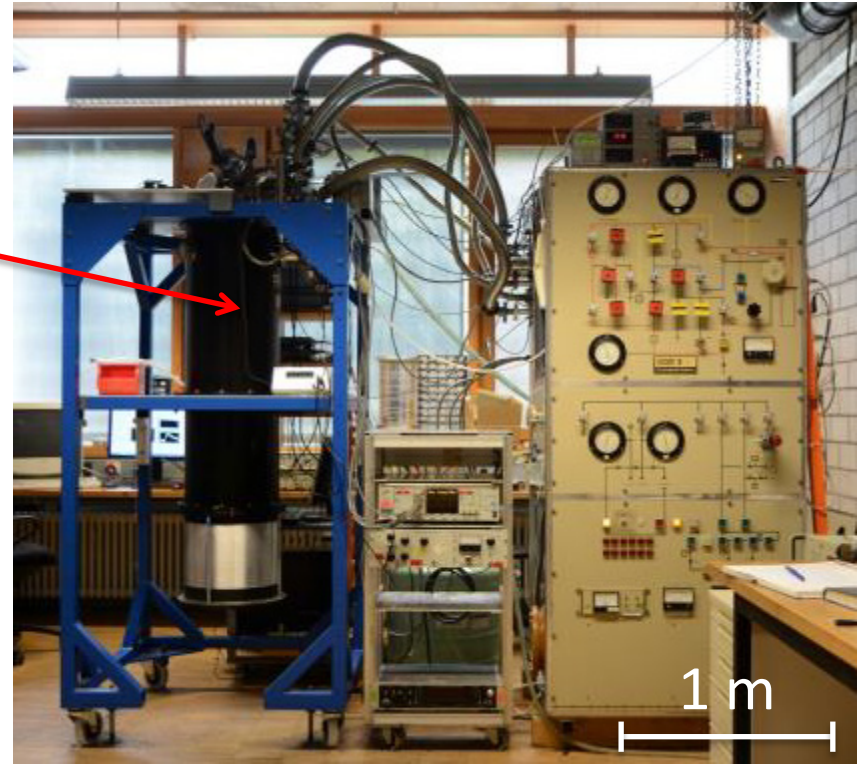
- Cu holder
- CuFlon circuit board
- Pb shielding
- $^3\text{He}/^4\text{He}$ dilution refrigerator (Oxford instruments)
- Cu experimental platform (Au plated)
- $T \approx 20 \text{ mK}$



^{163}Ho Detector: Experimental environment

Layered precooling:

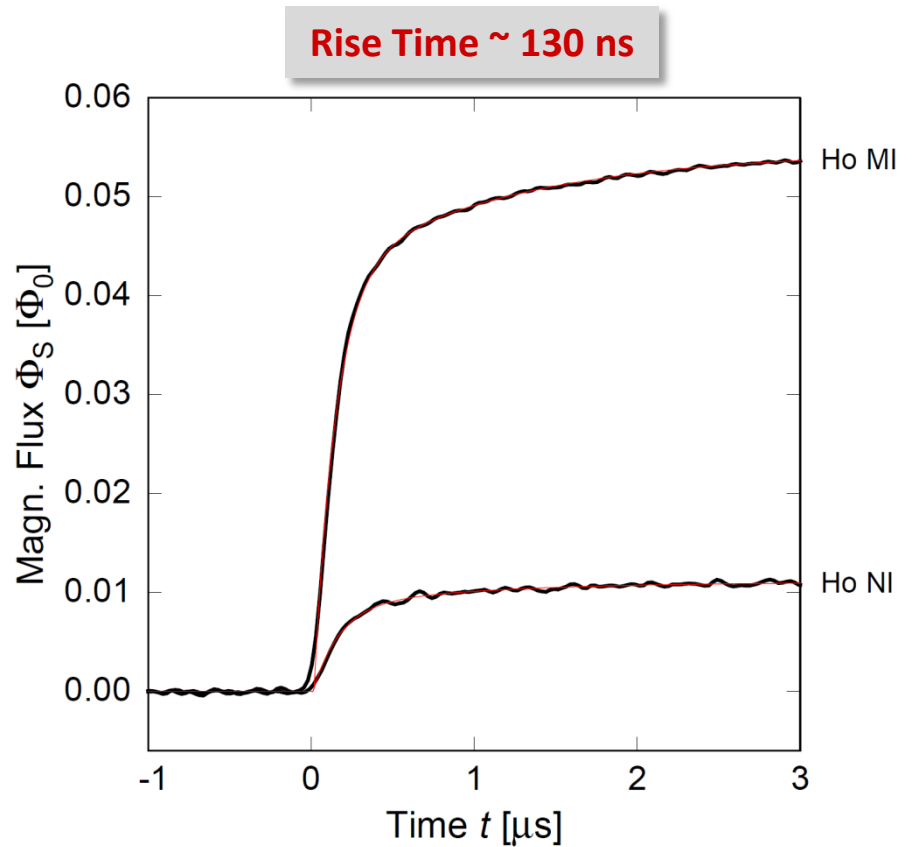
- Liquid $\text{N}_2 \sim 77 \text{ K}$
- Liquid $^4\text{He} \sim 4.2 \text{ K}$



Dilution unit

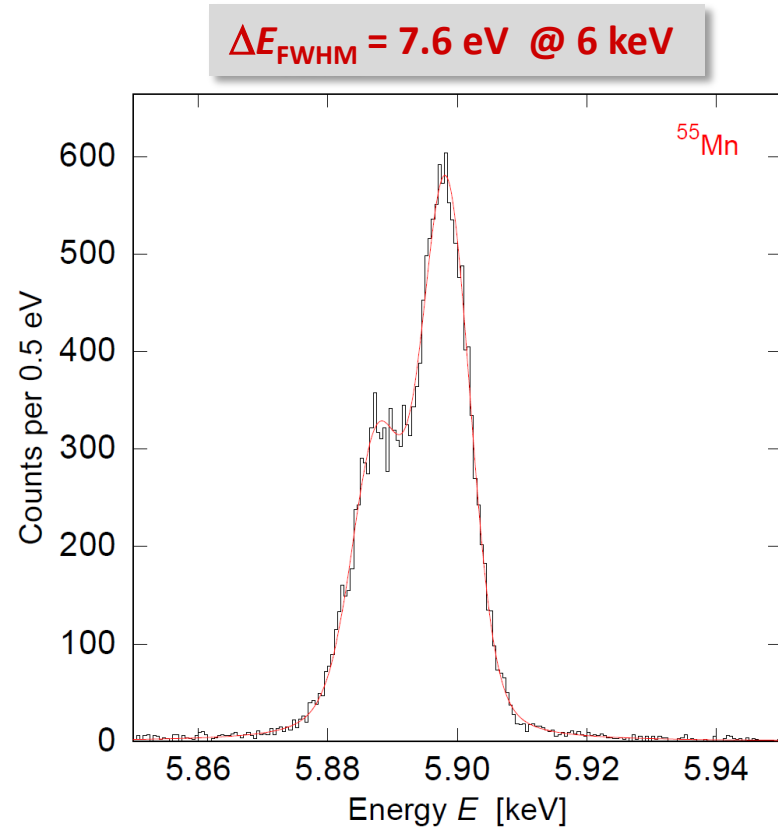
- 1st stage $\sim 1.5 \text{ K}$
- 2nd stage $\sim 0.6 \text{ K}$
- Experimental platform $\sim 20 \text{ mK}$

^{163}Ho Detector: Calorimetric spectrum



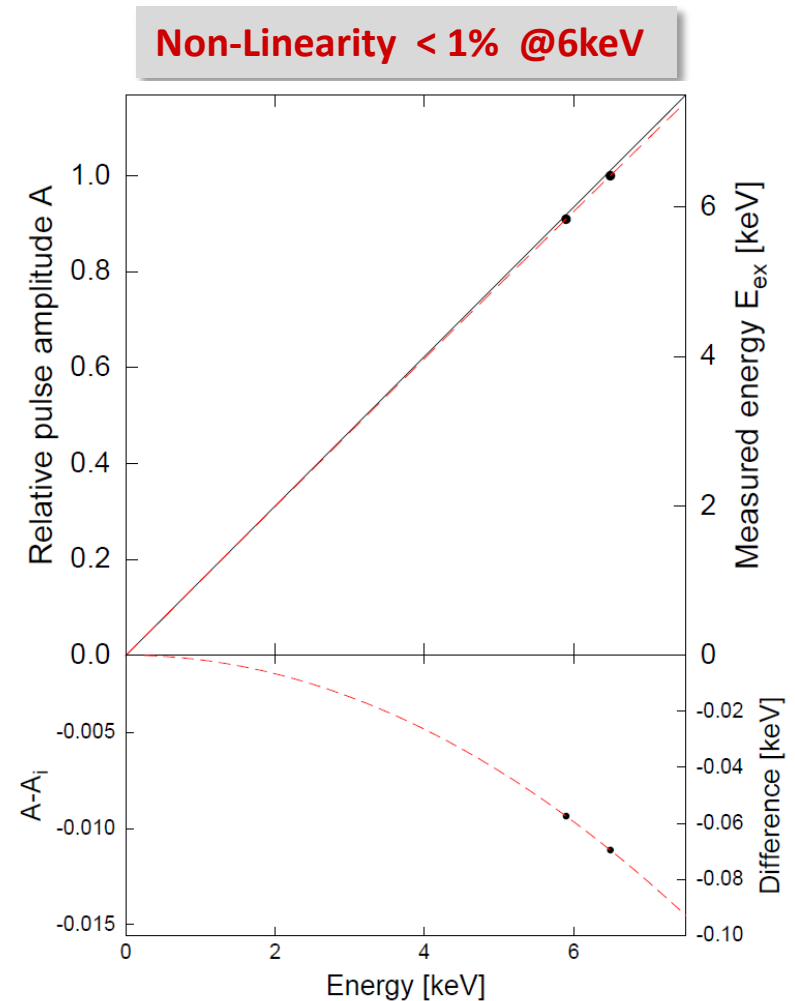
^{163}Ho Detector: Calorimetric spectrum

- Rise Time ~ 130 ns



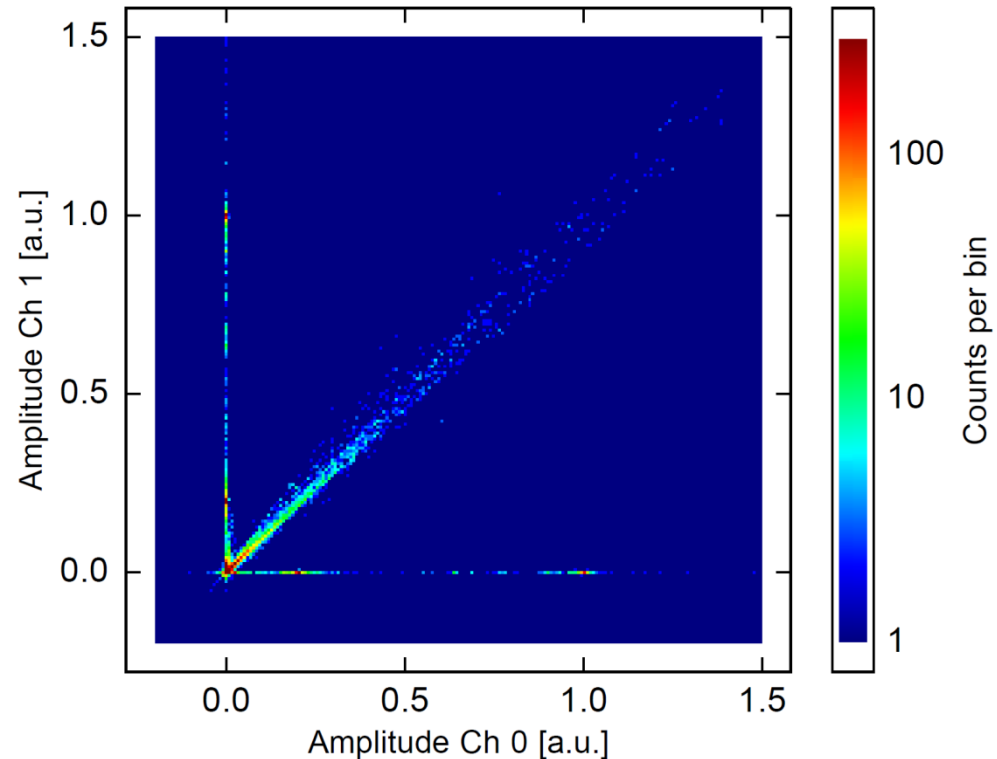
^{163}Ho Detector: Calorimetric spectrum

- Rise Time ~ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6$ eV @ 6 keV



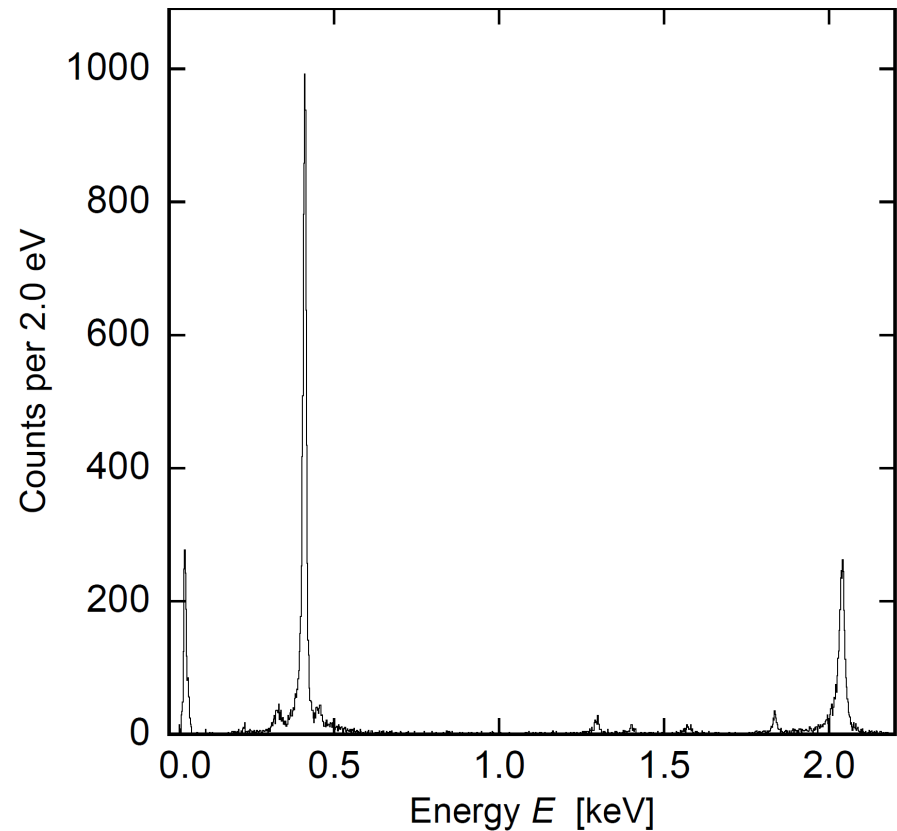
^{163}Ho Detector: Calorimetric spectrum

- Rise Time ~ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6$ eV @ 6 keV
- Non-Linearity $< 1\%$ @6keV



^{163}Ho Detector: Calorimetric spectrum

- Rise Time ~ 130 ns
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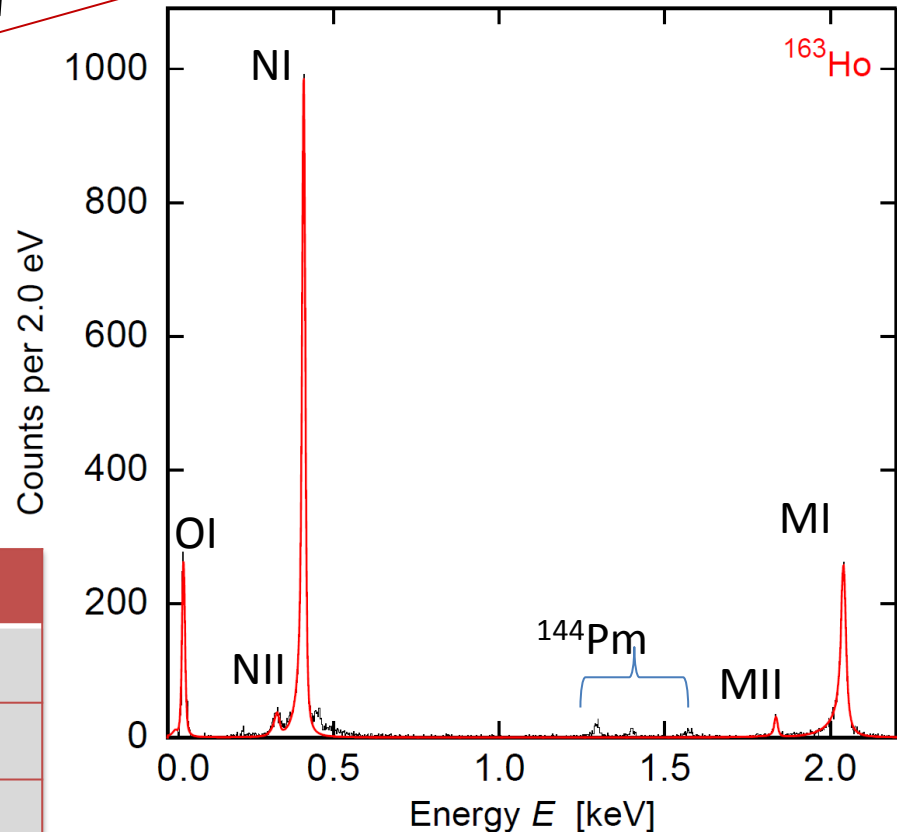


^{163}Ho Detector: Calorimetric spectrum

- Rise Time ~ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6$ eV @ 6 keV
- Non-Linearity $< 1\%$ @6keV
- Most precise ^{163}Ho spectrum

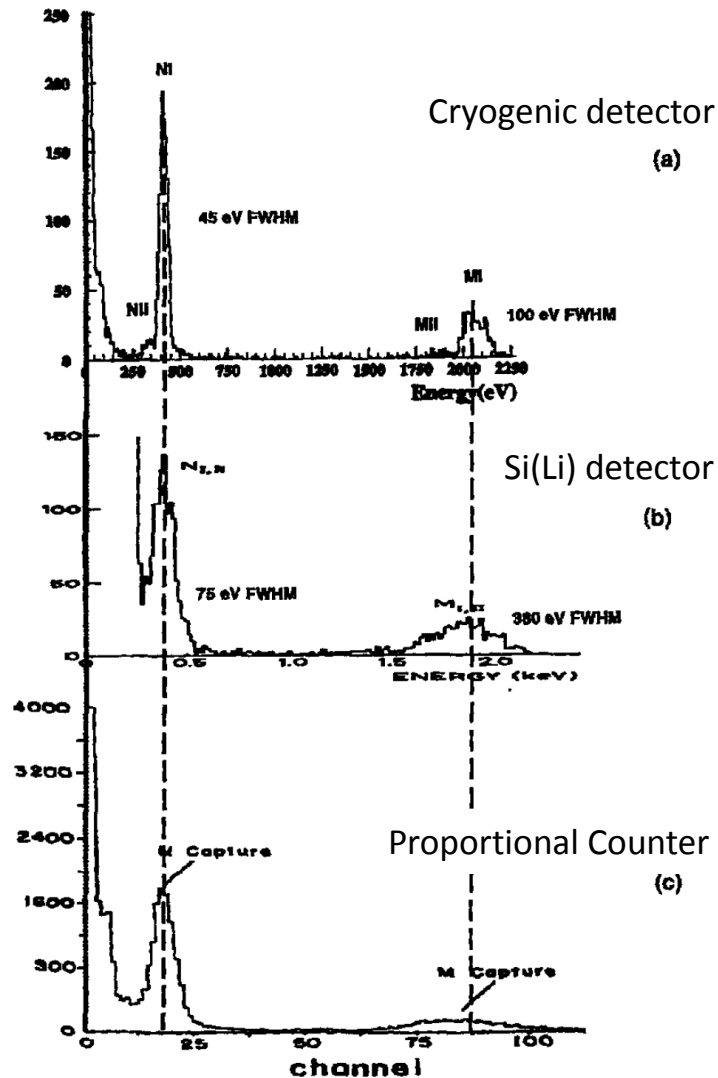
Preliminary analysis

	E_{H} lit.	E_{H} exp.	Γ_{H} lit.	Γ_{H} exp
MI	2.047	2.040	13.2	13.7
MII	1.845	1.836	6.0	7.2
NI	0.420	0.411	5.4	5.3
NII	0.340	0.333	5.3	8.0
OI	0.050	0.048	5.0	4.3

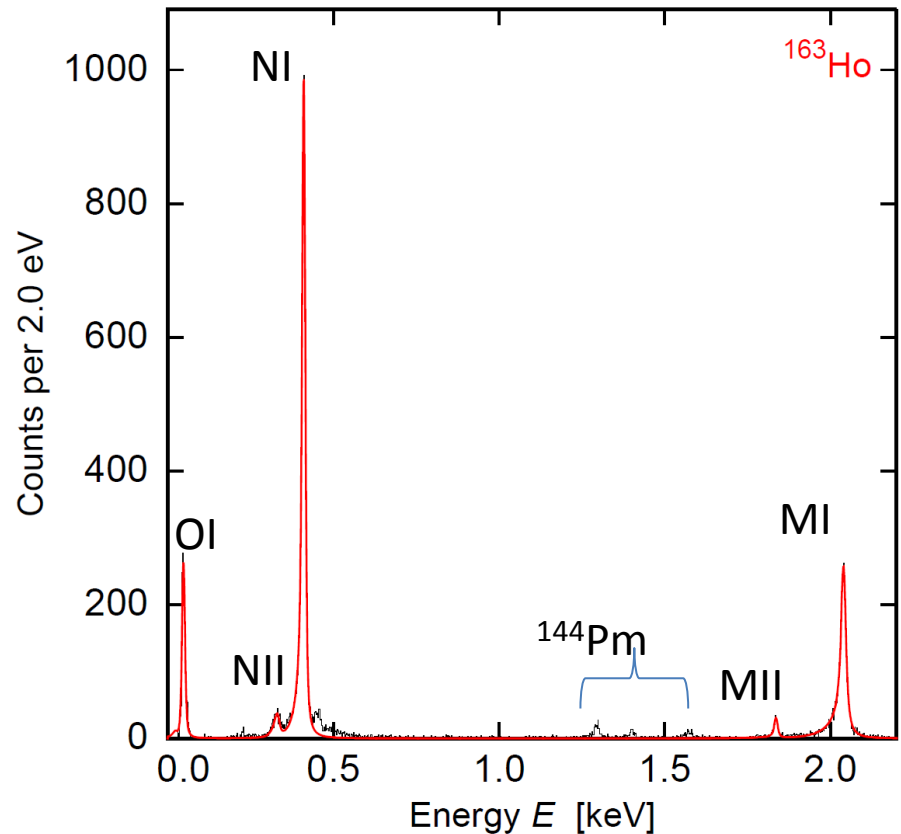


$$Q_{\text{EC}} = (2.80 \pm 0.08) \text{ keV}$$

^{163}Ho Detector: Calorimetric spectrum



F. Gatti et al., Physics Letters B 398 (1997) 415-419

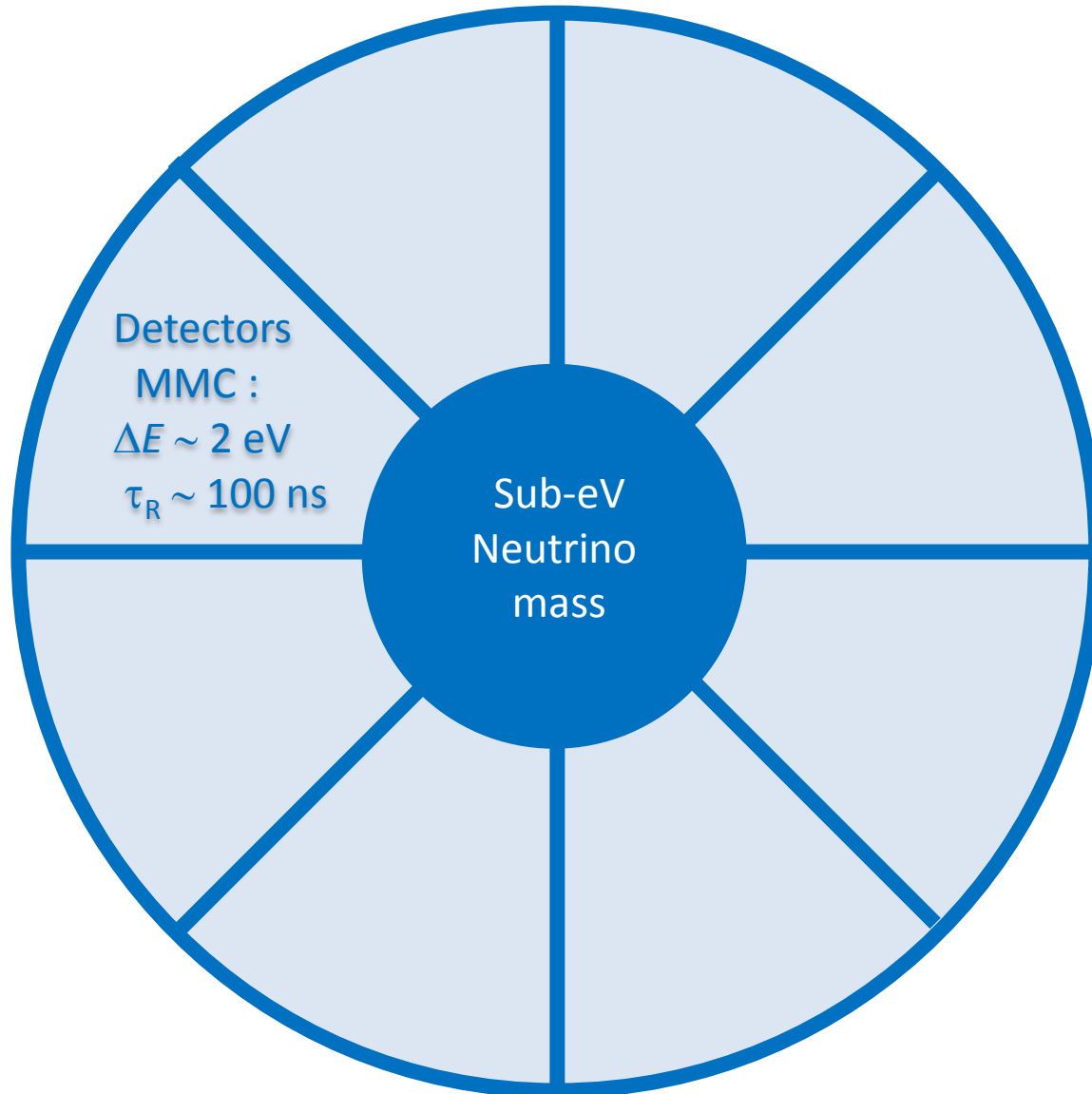


(a) F. Gatti et al., Physics Letters B 398 (1997) 415-419

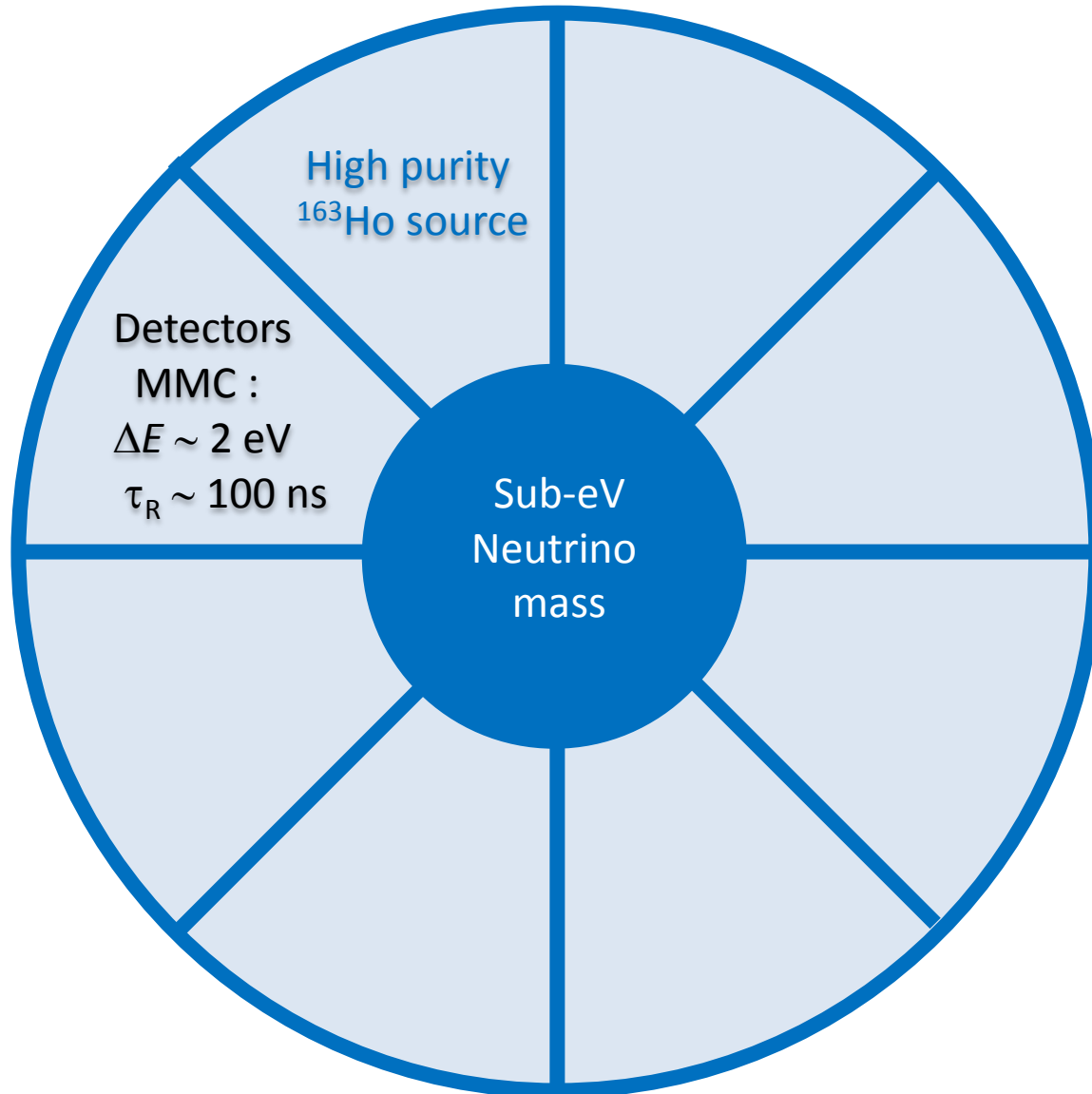
(b) E. Laesgaard et al., Proceeding of 7th International Conference on Atomic Masses and Fundamental Constants (AMCO-7), (1984).

(c) F.X. Hartmann and R.A. Naumann, Nucl. Instr. Meth. A 313 (1992) 237.

ECHo experiment



ECHo experiment



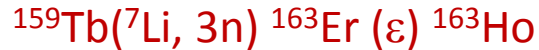
ECHo experiment: ^{163}Ho source

- Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17} \text{ atoms}$

EChO experiment: ^{163}Ho source

- Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

➤ ^{163}Ho can be produced by charged particle activation through direct or indirect way

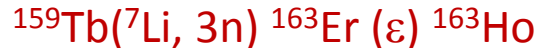


Er155 5.3 m 7/2- C, α	Er156 19.5 m 0+ EC	Er157 18.65 m 3/2- EC *	Er158 2.29 h 0+ EC	Er159 36 m 3/2- EC	Er160 28.58 h 0+ EC	Er161 3.21 h 3/2- EC	Er162 0+ 0.14	Er163 75.0 m 5/2- EC	Er164 0+ 1.61	Er165 10.36 h 5/2- EC	Er166 0+ 33.6
Ho154 11.76 m (2)- C, α *	Ho155 48 m 5/2+ EC	Ho156 56 m (4+) EC *	Ho157 12.6 m 7/2- EC	Ho158 11.3 m 5+ EC *	Ho159 33.05 m 7/2- EC *	Ho160 25.6 m 5+ EC *	Ho161 2.48 h 7/2- EC *	Ho162 15.0 m 1+ EC *	Ho163 4570 y 7/2- EC *	Ho164 29 m 1+ EC, β^- *	Ho165 7/2- 100
Dy153 6.4 h 7/2(-) α	Dy154 3.0E+6 y 0+ α	Dy155 9.9 h 3/2- EC	Dy156 0+ 0.06	Dy157 8.14 h 3/2- EC *	Dy158 0+ 0.10	Dy159 144.4 d 3/2- EC	Dy160 0+ 2.34	Dy161 5/2+ 18.9	Dy162 0+ 25.5	Dy163 5/2- 24.9	Dy164 0+ 28.2
Tb152 17.5 h 2- C, α *	Tb153 2.34 d 5/2+ EC	Tb154 21.5 h 0 EC, β^- *	Tb155 5.32 d 3/2+ EC	Tb156 5.35 d 3- EC, β^- *	Tb157 71 y 3/2+ EC	Tb158 180 y 3- EC, β^- *	Tb159 3/2+ 100	Tb160 72.3 d 3- β^-	Tb161 6.88 d 3/2+ β^-	Tb162 7.60 m 1- β^-	Tb163 19.5 m 3/2+ β^-

EChO experiment: ^{163}Ho source

- Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

- ^{163}Ho can be produced by charged particle activation through direct or indirect way



- ^{163}Ho can be produced by via (n, γ) -reaction on ^{162}Er

Er155 5.3 m 7/2- C,α	Er156 19.5 m 0+ EC	Er157 18.65 m 3/2- EC *	Er158 2.29 h 0+ EC	Er159 36 m 3/2- EC	Er160 28.58 h 0+ EC	Er161 3.21 h 3/2- EC	Er162 0+ 0.14	Er163 75.0 m 5/2- EC	Er164 0+ 1.61	Er165 10.36 h 5/2- EC	Er166 0+ 33.6
Ho154 11.76 m (2)- C,α *	Ho155 48 m 5/2+ EC	Ho156 56 m (4+) EC *	Ho157 12.6 m 7/2- EC	Ho158 11.3 m 5+ EC *	Ho159 33.05 m 7/2- EC *	Ho160 25.6 m 5+ EC *	Ho161 2.48 h 7/2- EC *	Ho162 15.0 m 1+ EC *	Ho163 4570 y 7/2- EC *	Ho164 29 m 1+ EC,β- *	Ho165 7/2- 100
Dy153 6.4 h 7/2(-) α	Dy154 3.0E+6 y 0+ α	Dy155 9.9 h 3/2- EC	Dy156 0+ 0.06	Dy157 8.14 h 3/2- EC *	Dy158 0+ 0.10	Dy159 144.4 d 3/2- EC	Dy160 0+ 2.34	Dy161 5/2+ 18.9	Dy162 0+ 25.5	Dy163 5/2- 24.9	Dy164 0+ 28.2
Tb152 17.5 h 2- C,α *	Tb153 2.34 d 5/2+ EC	Tb154 21.5 h 0 EC,β- *	Tb155 5.32 d 3/2+ EC	Tb156 5.35 d 3- EC,β- *	Tb157 71 y 3/2+ EC	Tb158 180 y 3- EC,β- *	Tb159 3/2+ 100	Tb160 72.3 d 3- β-	Tb161 6.88 d 3/2+ β-	Tb162 7.60 m 1- β-	Tb163 19.5 m 3/2+ β-

ECHo experiment: ^{163}Ho source

- Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17} \text{ atoms}$
- ^{163}Ho can be produced by charged particle activation through direct or indirect way
 - $^{\text{nat}}\text{Dy}(p, xn) ^{163}\text{Ho}$
 - $^{\text{nat}}\text{Dy}(\alpha, xn) ^{163}\text{Er} (\varepsilon) ^{163}\text{Ho}$
 - $^{159}\text{Tb}(^7\text{Li}, 3n) ^{163}\text{Er} (\varepsilon) ^{163}\text{Ho}$
- ^{163}Ho can be produced by via (n, γ) -reaction on ^{162}Er

Two sources already produced

- ✓ Helmholtz Zentrum Berlin
- ✓ Institut Laue-Langevin in Grenoble

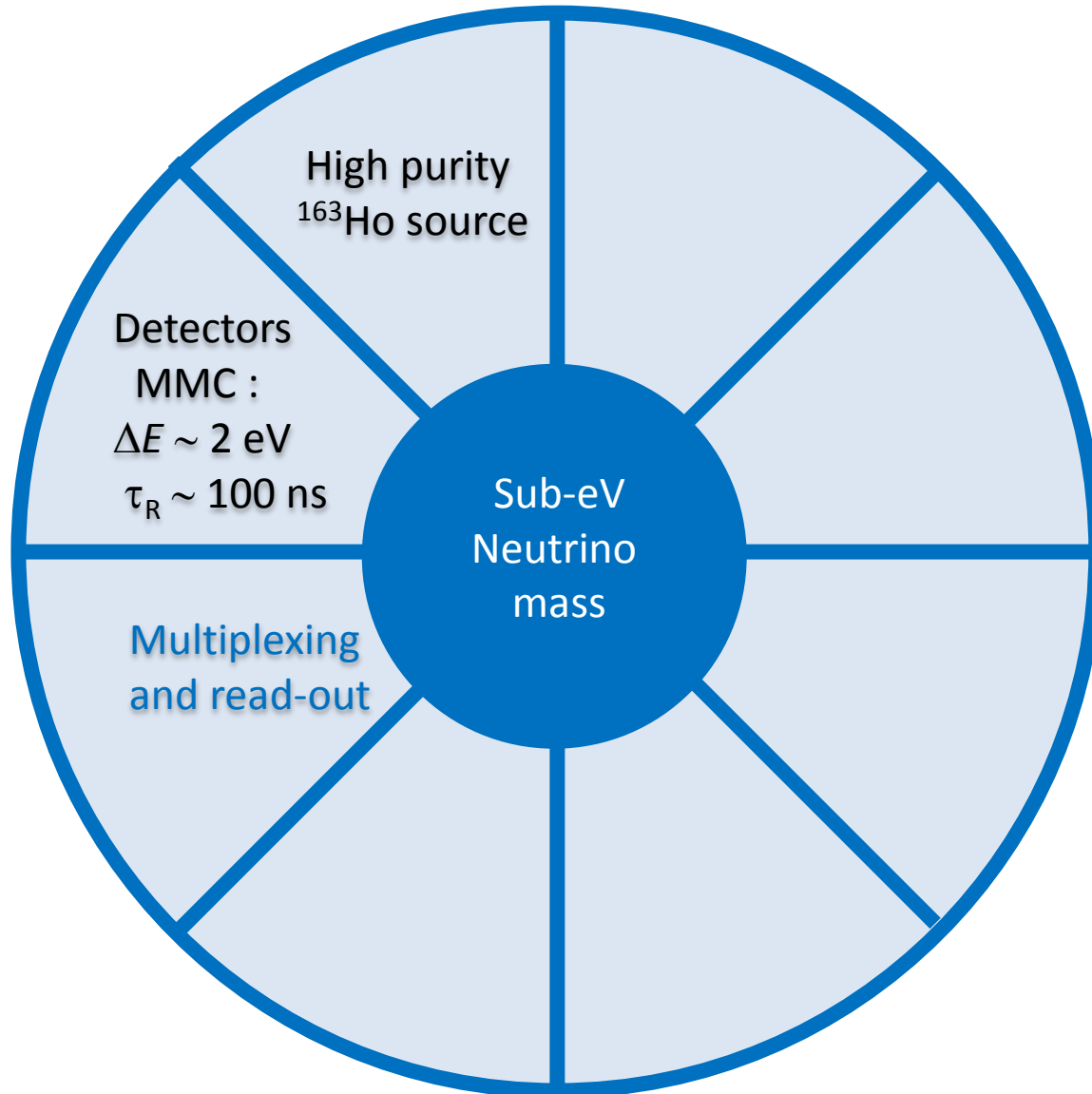
EChO experiment: ^{163}Ho source

- Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17} \text{ atoms}$
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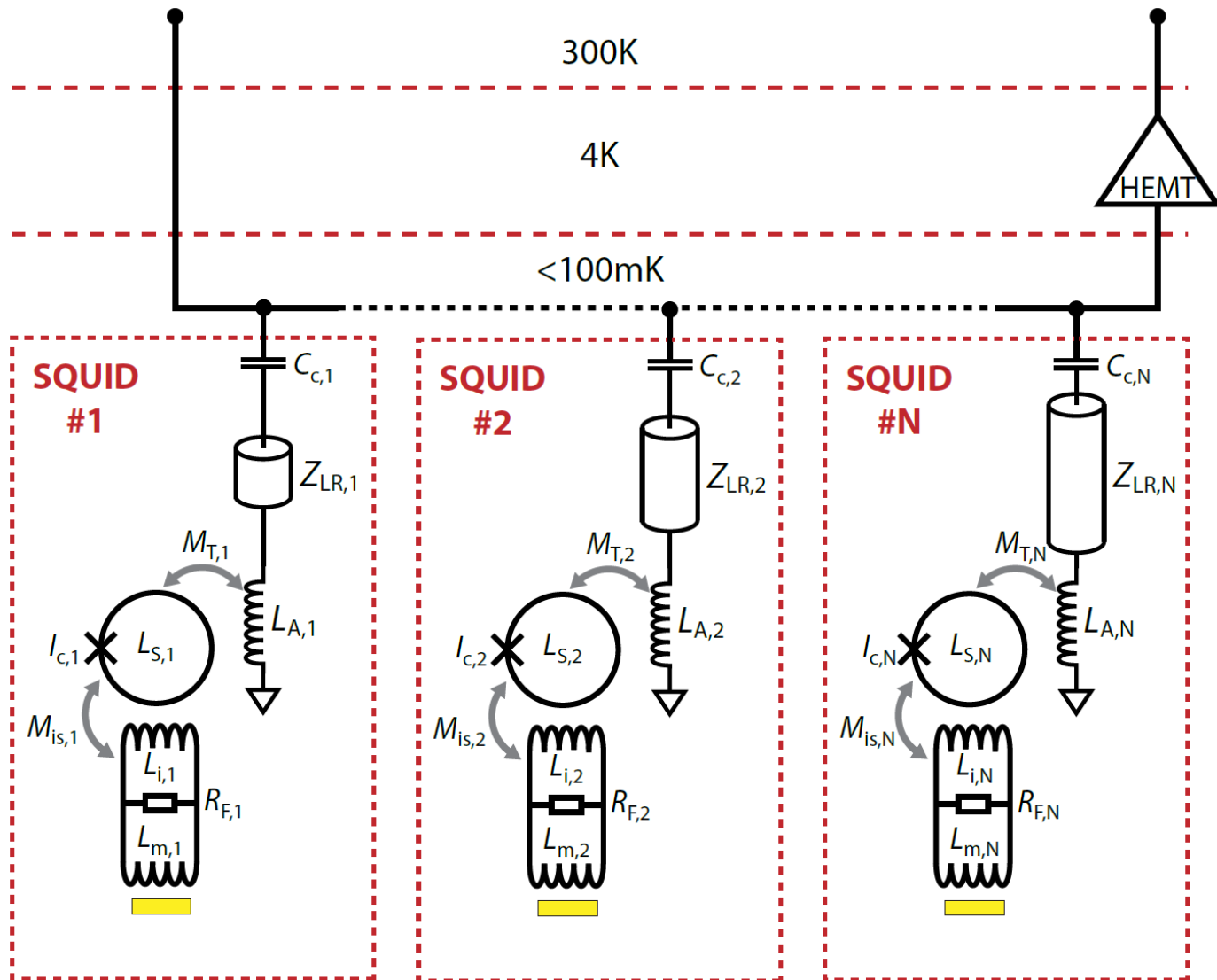
Two sources already produced

 - ✓ Helmholtz Zentrum Berlin
 - ✓ Institut Laue-Langevin in Grenoble
- Purity: No radioactive contaminants and removed target material
- High efficiency purification methods
- Chemical form: depends on the absorber preparation (ion implantation, dilute alloys)

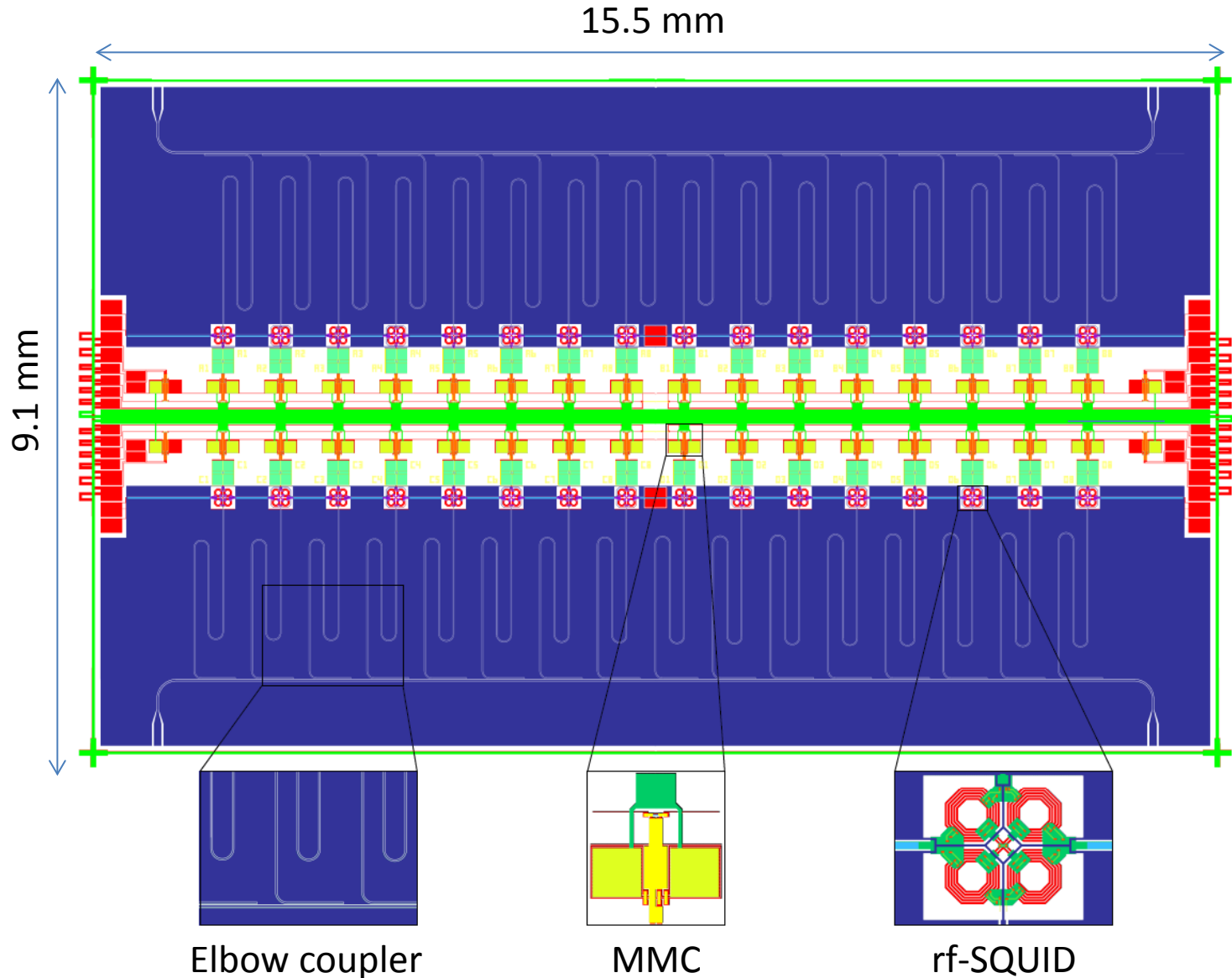
ECHo experiment



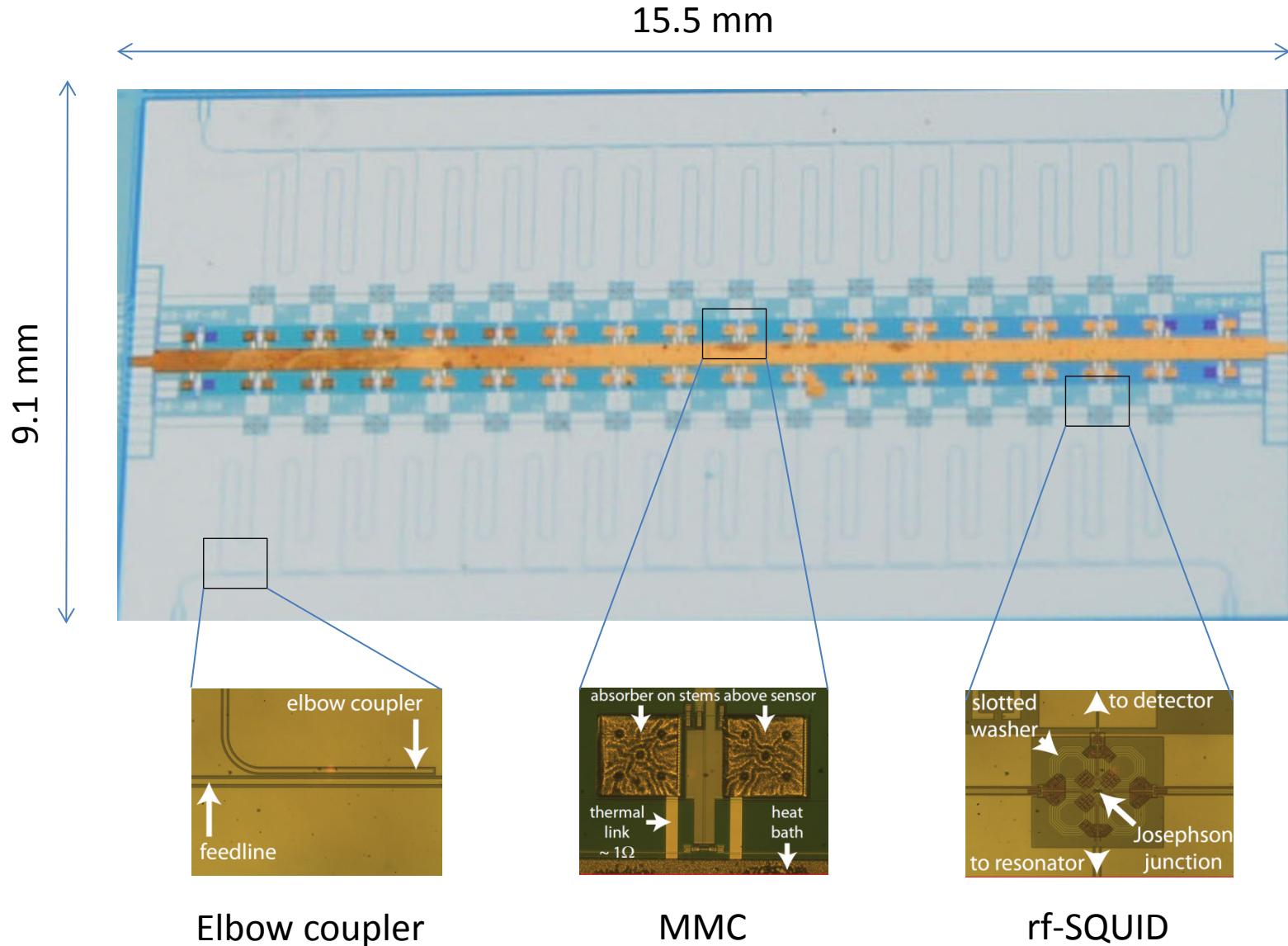
ECHo experiment: μ wave SQUID multiplexing



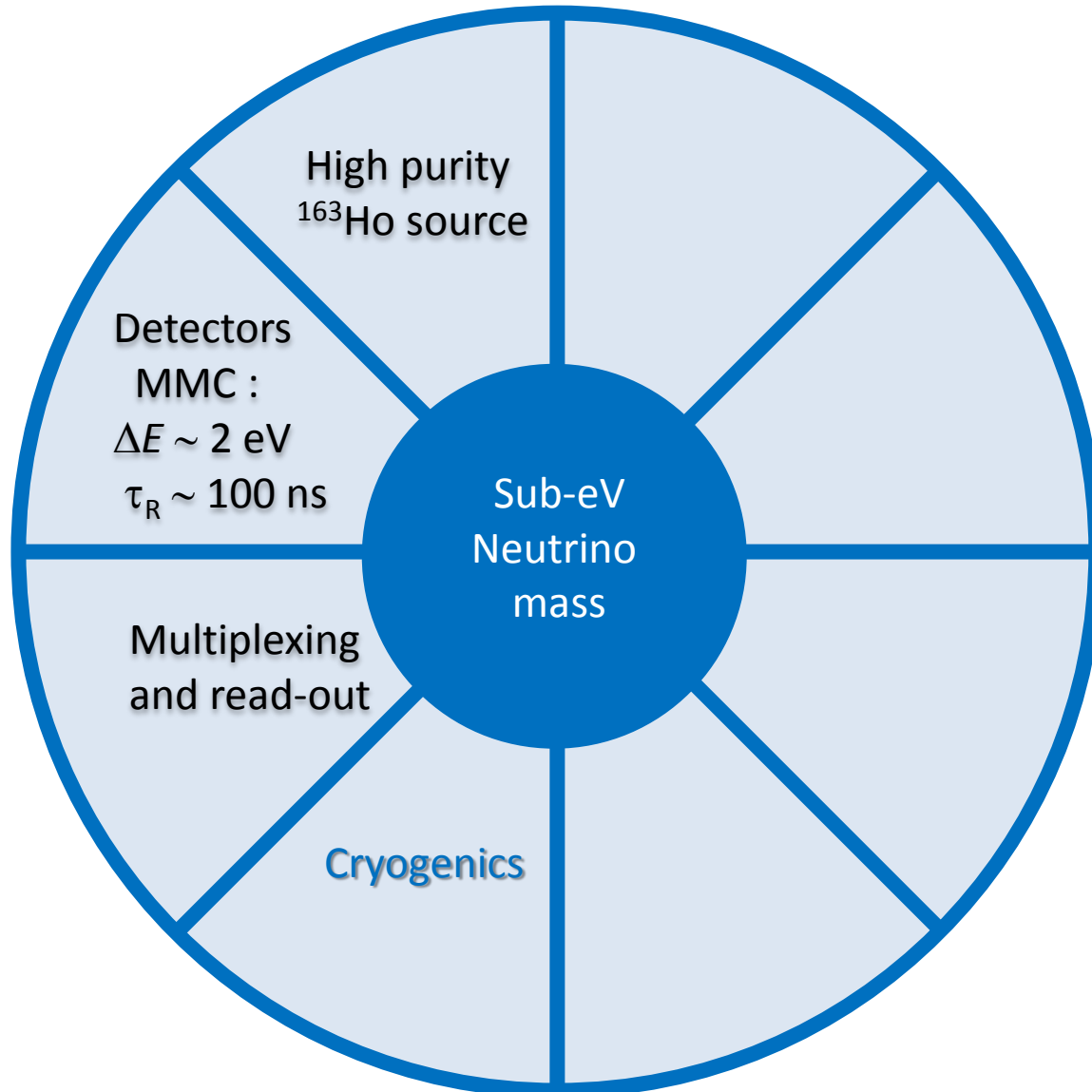
ECHo experiment: 64-pixel chip



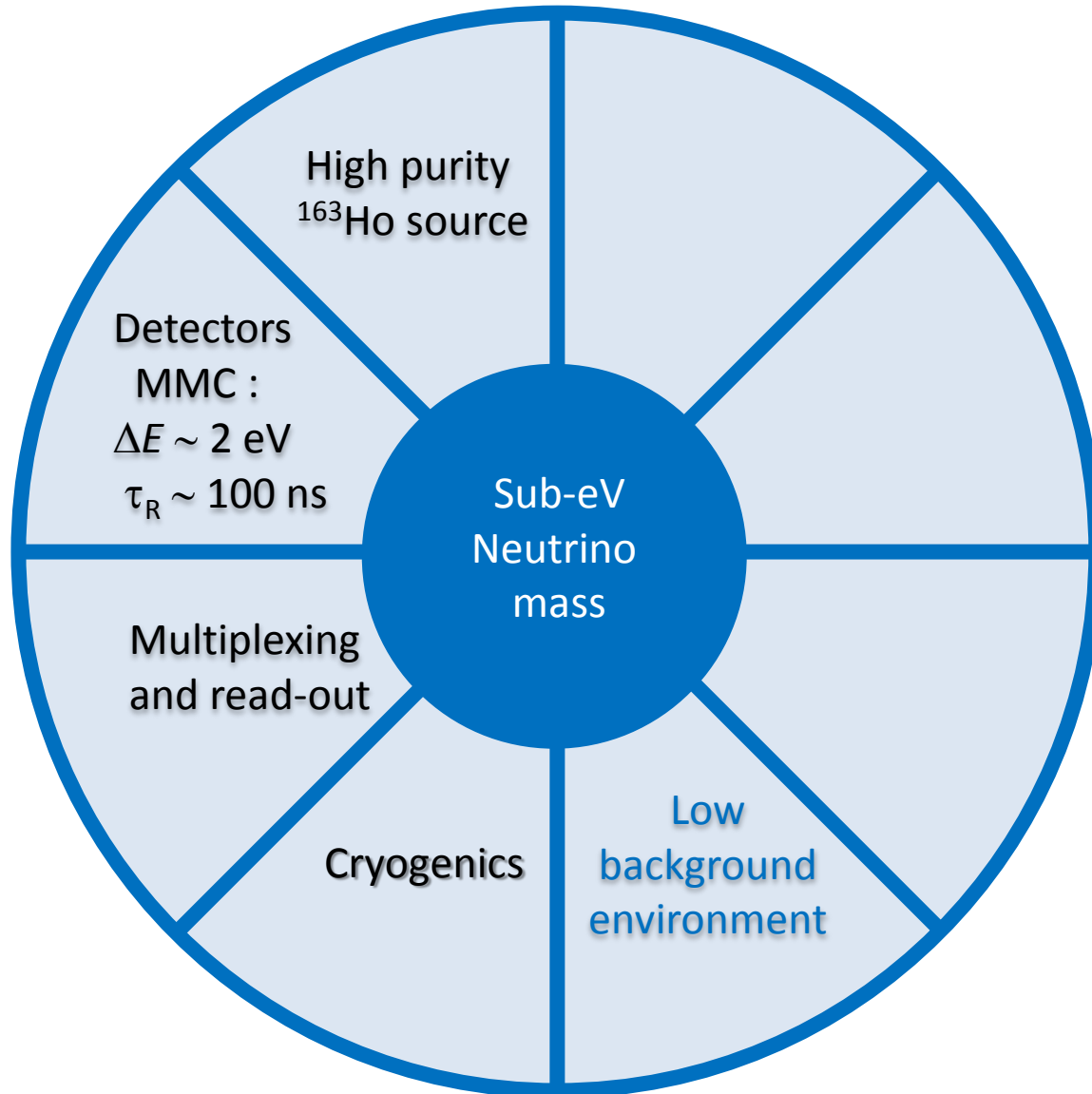
ECHo experiment: 64-pixel chip



ECHo experiment



ECHo experiment



ECHo experiment: Low background

Background sources:

- Contamination of the source → $^{166\text{m}}\text{Ho}$
- Environmental radioactivity
- Induced secondary radiation by cosmic rays

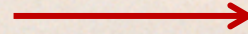


ECHo experiment: Low background

Background sources:

- Contamination of the source → $^{166\text{m}}\text{Ho}$
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Monte Carlo simulations
Dedicated experiments



ECHo experiment: Low background

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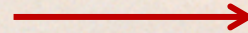
Monte Carlo simulations
Dedicated experiments

First measurements
underground in Modane
winter 2013

ECHo experiment: Low background

Background sources:

- Contamination of the source → ^{166m}Ho
- Environmental radioactivity
- Induced secondary radiation by cosmic rays

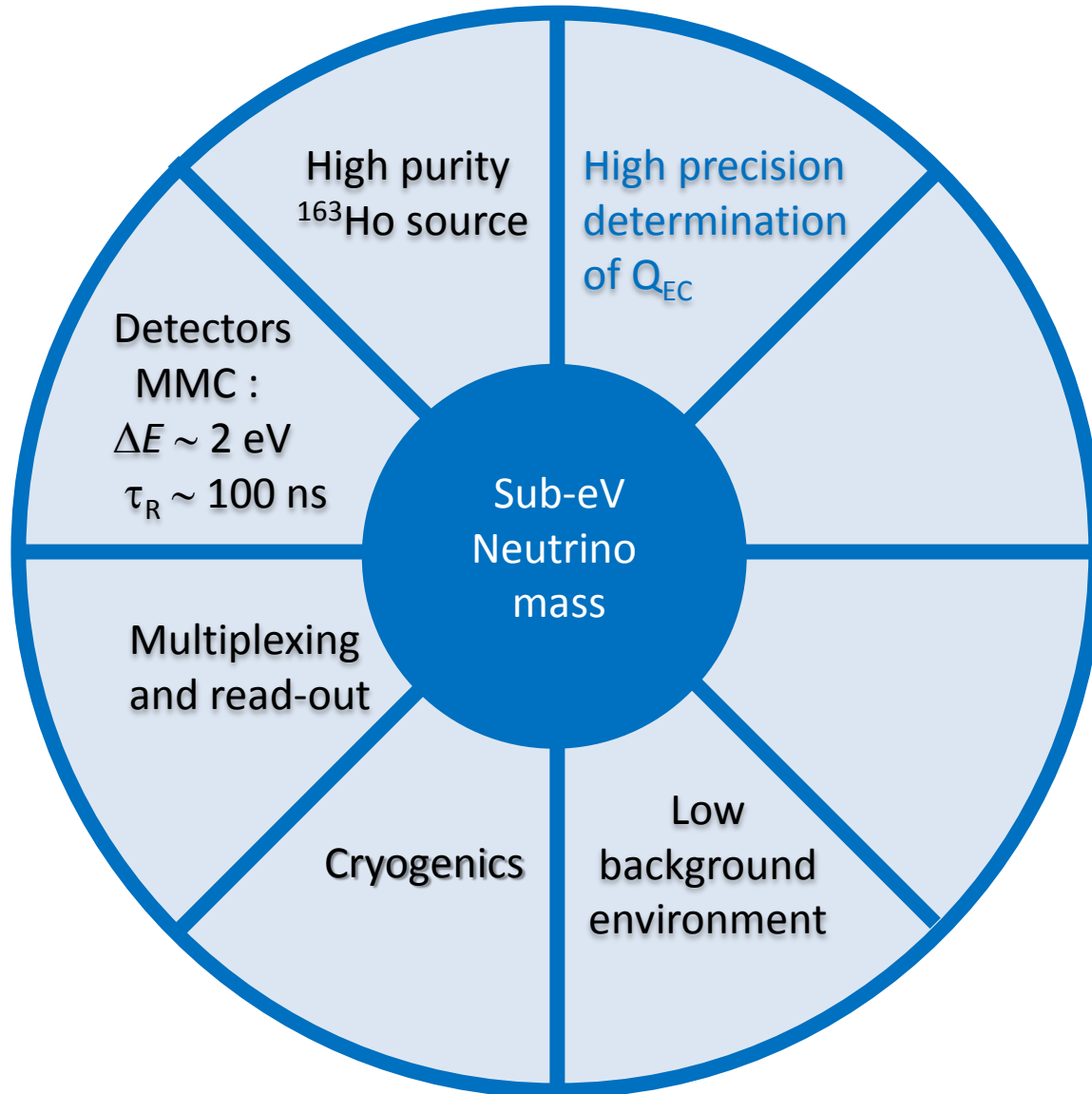


Monte Carlo simulations
Dedicated experiments

First measurements
underground in Modane
winter 2013

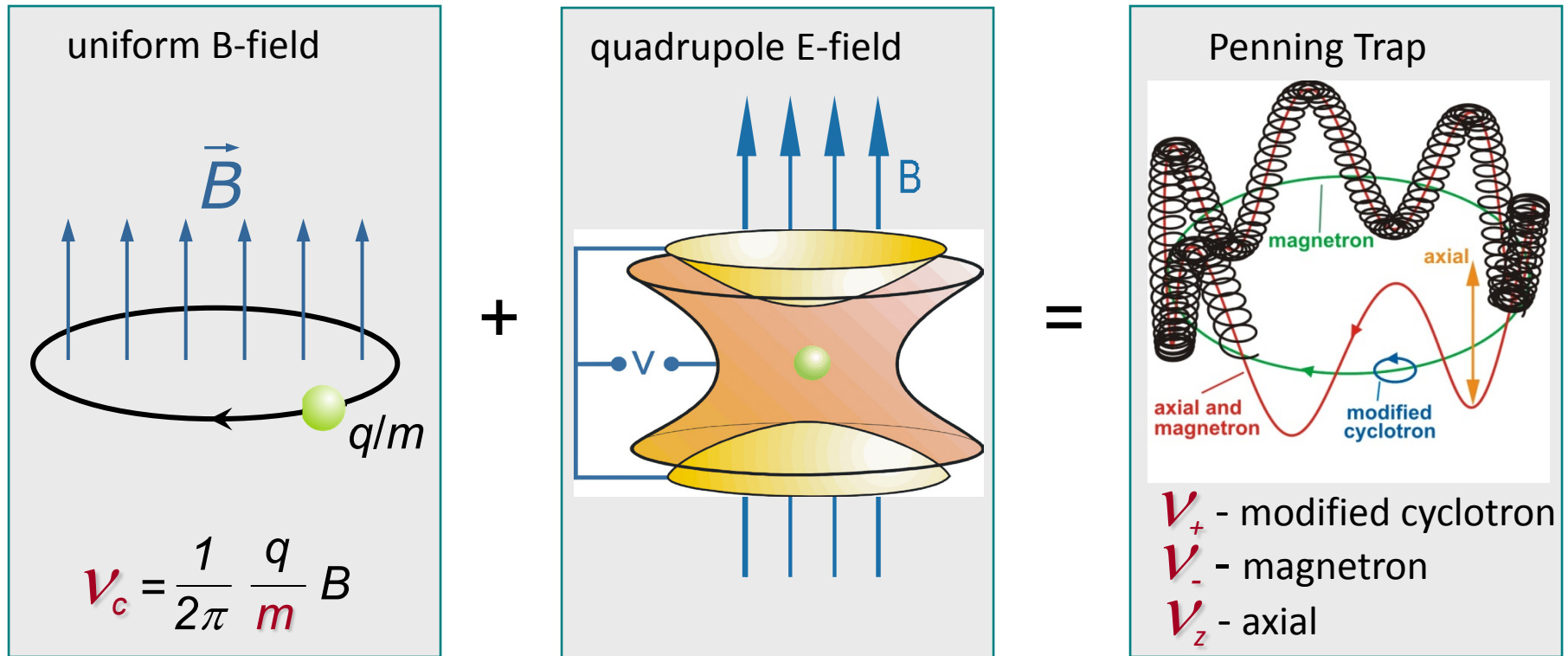
- ✓ Selection of materials for the experimental set-up
- ✓ Definition of detector design → veto
- ✓ Designs for the shielding

ECHo experiment



EChO experiment: Q_{EC} determination

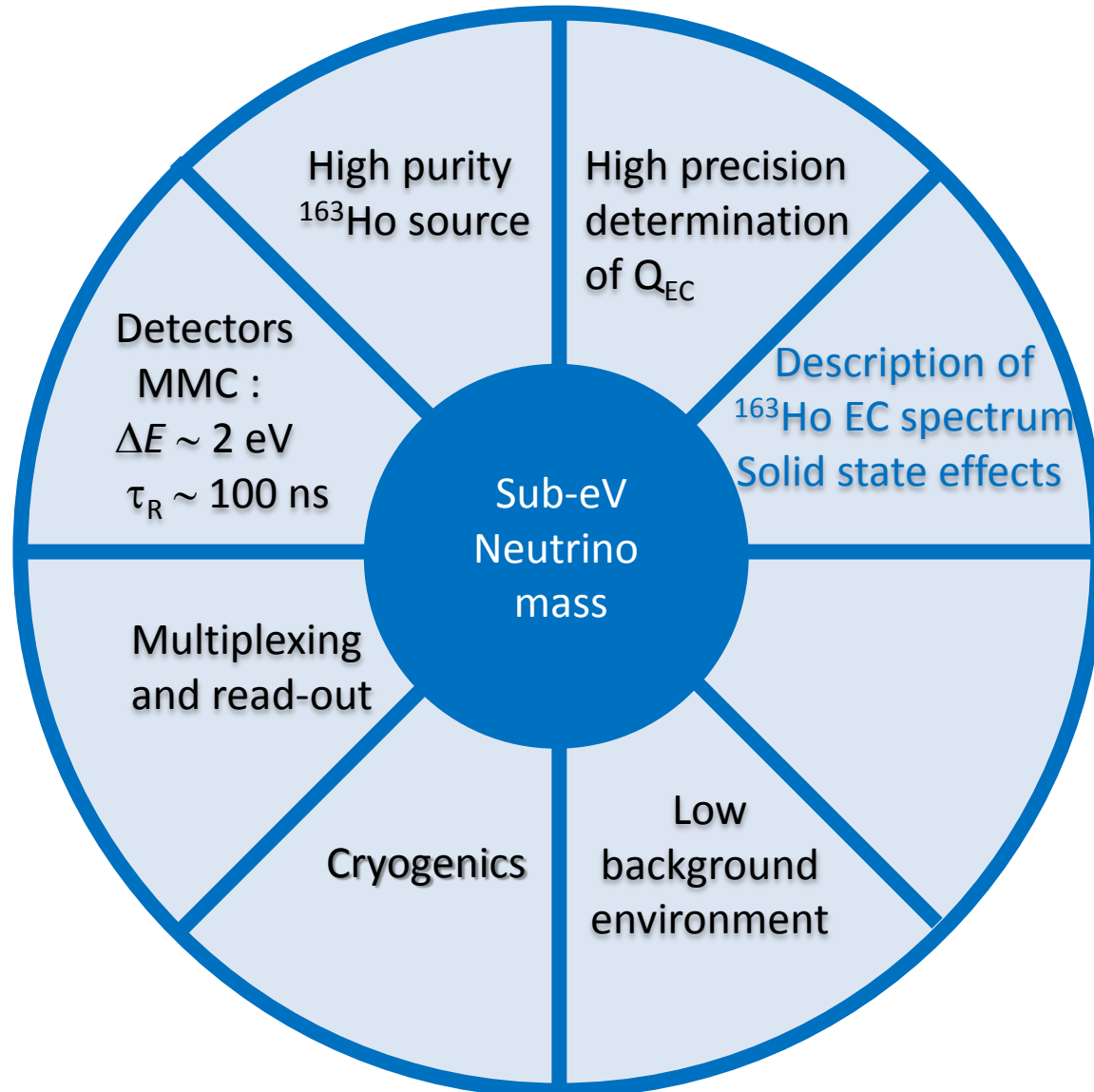
Penning Trap mass spectroscopy



Next future : **SHIPTRAP (GSI)** $\rightarrow Q_{EC}$ determination within **100 eV**

In few years: **PENTATRAP (MPI-K HD)** $\rightarrow Q_{EC}$ determination within **1 eV**

ECHo experiment



ECHo experiment: spectral shape

How precise do we know the calorimetric spectrum of ^{163}Ho ?

$$\frac{dW}{dE_C} = A (Q_{\text{EC}} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{\text{EC}} - E_C)^2}} \sum_H B_H \phi_H^2(0) \frac{\Gamma_H}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$

Density Functional Theory (DFT) & Quasiparticle Random Phase Approximation (QRPA)

SOLID STATE EFFECTS

Experimental investigations:

^{163}Ho implanted in different materials

Simple ^{163}Ho molecules implanted in Au

Core Level binding energy Shift (CLS):

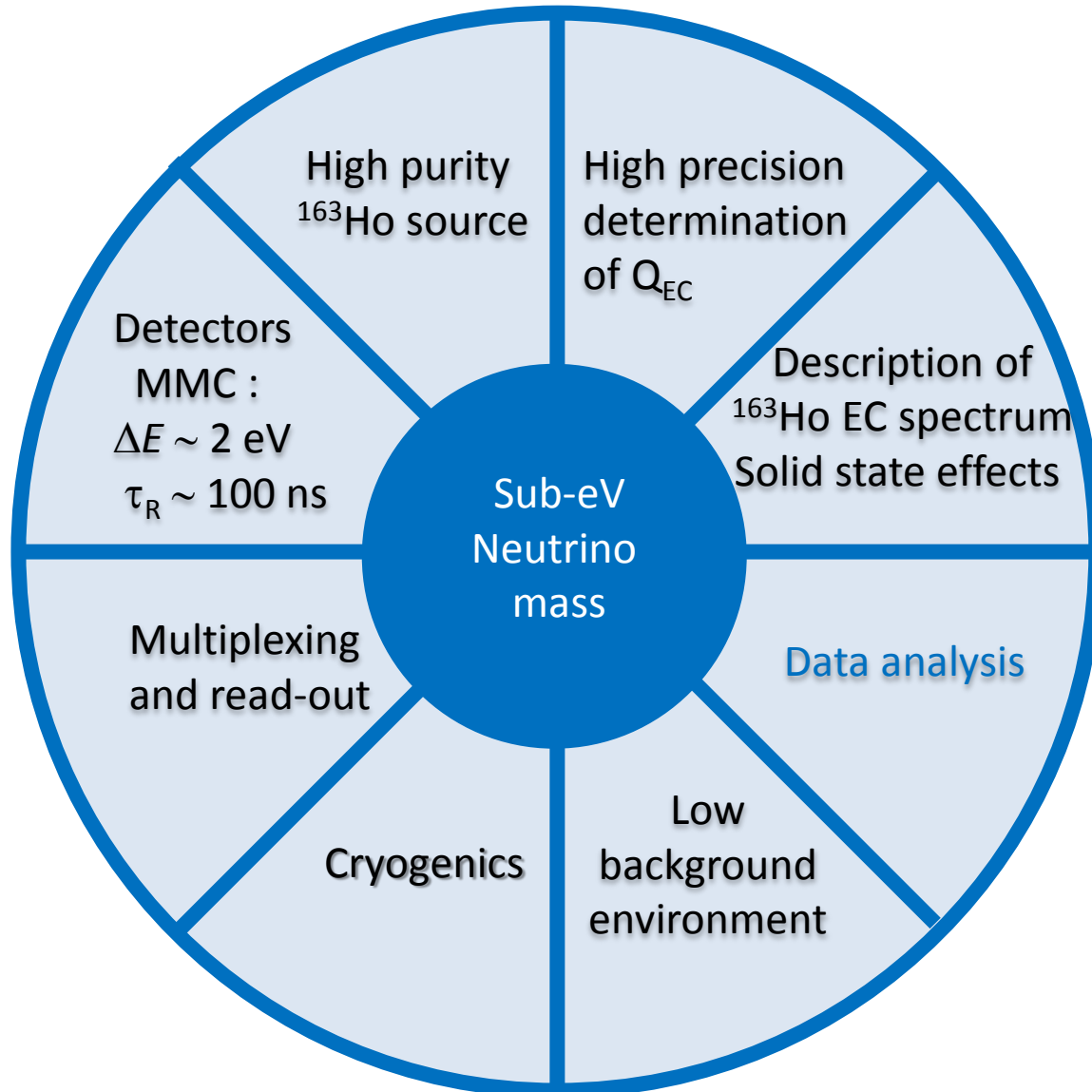
calculated using the complete screening approximation of DFT

Sensitivity to:

keV sterile neutrino

Cosmic neutrino background

ECHo experiment



^{163}Ho experiments



- ◆ Started R&D in 2011
- ◆ Small scale experiment with ~ 100 pixels within the next three years
- ◆ Large scale experiment to reach sub-eV sensitivity to neutrino mass

<http://www.kip.uni-heidelberg.de/echo/>

^{163}Ho experiments



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HOLMES

- ◆ Established in 2013 (ERC Advanced Grants for Prof. S. Ragazzi)
- ◆ Some R&D done already within the MARE experiment

<http://artico.mib.infn.it/nucrimib/general-infos/holmes-approved>

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OTHERS

- ◆ **LANL + NIST** (last two years)
 - investigation for source production
 - detector development for calorimetric measurements

<http://conference.ipac.caltech.edu/ltd-15> (Kunde, Schmidt, Croce, Fowler)

Conclusion

4

A. De Rujula

arXiv:1305.4857v1 [hep-ph] 21 May 2013

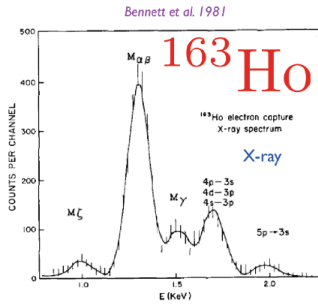


FIG. 5: IBEC spectrum in ^{163}Ho decay [22], showing prominent X-ray lines.

Some early measurements with a ^{163}Ho source [22, 23] were based on IBEC (Internal Bremsstrahlung in Electron Capture), the first-principle theory of which is fiendishly complex both above [24] and – more so – below [4] the energies coinciding with X-ray resonances. One example is shown in Fig. 5. Other measurements were calorimetric [25], see Fig. 6. The most stringent of the early mass limits, from [23] and [26] were, respectively:

$$\begin{aligned} m_\nu &< 225 \text{ eV at } 95\% \text{ CL,} \\ m_\nu &< 490 \text{ eV at } 68\% \text{ CL.} \end{aligned} \quad (8)$$

The recent progress may be illustrated by comparing Fig. 6 [25] with the preliminary results shown in Fig. 7, from the incipient experiment ECHO [27], which employs MMCs (Magnetic Metallic Calorimeters). The unlabeled peaks in Fig. 7 are due to ^{144}Pm , an impurity accompanying ^{163}Ho at the implantation stage at ISOLDE-CERN, an early test of source-preparation techniques.

One cannot resist the temptation of showing a scheme and a picture of the set of four MMCs in the ^{163}Ho detector prototype of ECHO [27]: Figs. 8 and 9. There is satisfaction associated with the possibility of measuring a tiny quantity – the neutrino mass – with nano-scale detectors. Even with the associated cryogenics and electronics, the apparatuses are still table-top.

V. THE THEORY OF EC IN ^{163}Ho

The EC process, all by itself, does not yield any information on the neutrino mass, or on anything else, for that matter. The mere information that “it happened” is

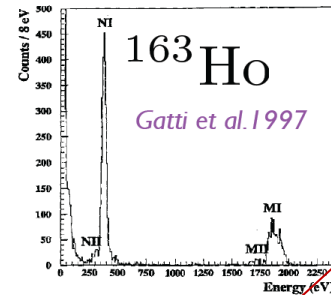


FIG. 6: Results of an early ^{163}Ho calorimetric spectrum [25].

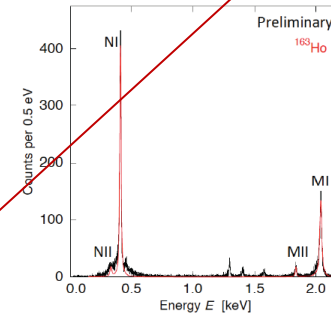


FIG. 7: Test results of ECHO [27] for the calorimetric spectrum of ^{163}Ho decay. The unlabeled impurities are ^{144}Pm . The continuous (red line) theory [5] is based on Eq. (9).

provided by the fact that the daughter atom, and sometimes its nucleus, are unstable. The hole in an atomic shell, for instance, results in observable X-rays, as the outer electrons cascade inwards, see Fig. 5.

The measured $Q = M(^{163}\text{Ho}) - M(^{163}\text{Dy})$ is so small that EC is only energetically allowed from ^{163}Ho orbitals with principal quantum number $n > 2$. The emission of X-rays from holes in such external shells is negligible compared to that of atomic de-excitations involving electron emission (in the classical parlance, the “fluorescence yields” are tiny). The electron-emitting transitions have

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Thank you!

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