



Basic Considerations on Spatial Resolution in Patterned Detectors

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 Warmup: Resolution with binary readout – optimal signal width

- Error of Center-of-Gravity: When do we need a fit?
- Influence of noise on spatial resolution
 - Higher Moments
 - Correlated Noise
 - 2D structures
 - Wide signals
- Error when doing 'Eta-reconstruction'
 - Search for 'best' response function

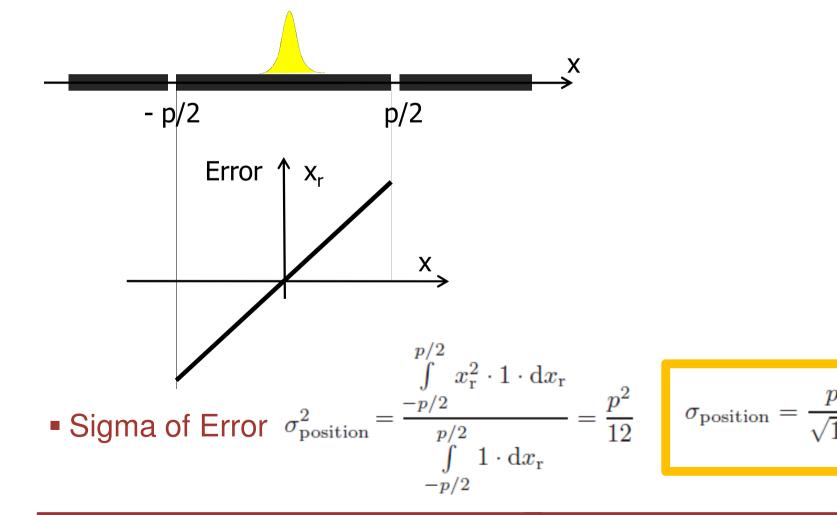
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BINARY READOUT OF BOX SIGNALS

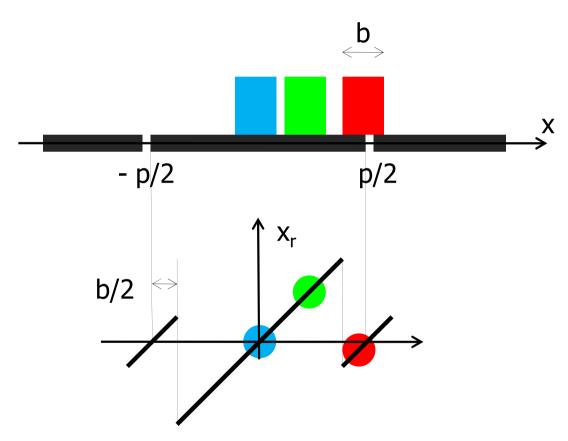
Spatial Resolution of Narrow Signals

- Consider very narrow signal
- Only one strip is hit
- Reconstructed position = strip center, error = offset in strip



Resolution with wider Signals (Binary Readout!)

- Consider 'Box' Signals for simplicity
- When 2 strips are hit \rightarrow reconstruct at edge \rightarrow small error



Minimum Error for b = p/2. Error becomes half: \$\sigma = \frac{1}{2} \end{p} \sigma 12\$
Note that we have 50% single / 50% double hits

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CENTER-OF-GRAVITY RECONSTRUCTION: WHEN IS IT SUFFICIENT - OR -WHEN DO WE NEED A FIT?

The question:

- A 1D signal with (spatial) shape f(x) falls onto a strip structure with pitch a
- We assume $\int f(x) dx = 1$ and f(x) symmetric.
- This generates (analogue) signals on several strips.
- We assume for now that **noise = 0**.
- Question:
 - What is the reconstruction error for GoG reconstruction ?
 - More precisely: Error for a single event? Average error? Sigma?

• We expect that the answer depends on

- signal shape
- Strip pitch a
- signal position (for single events)

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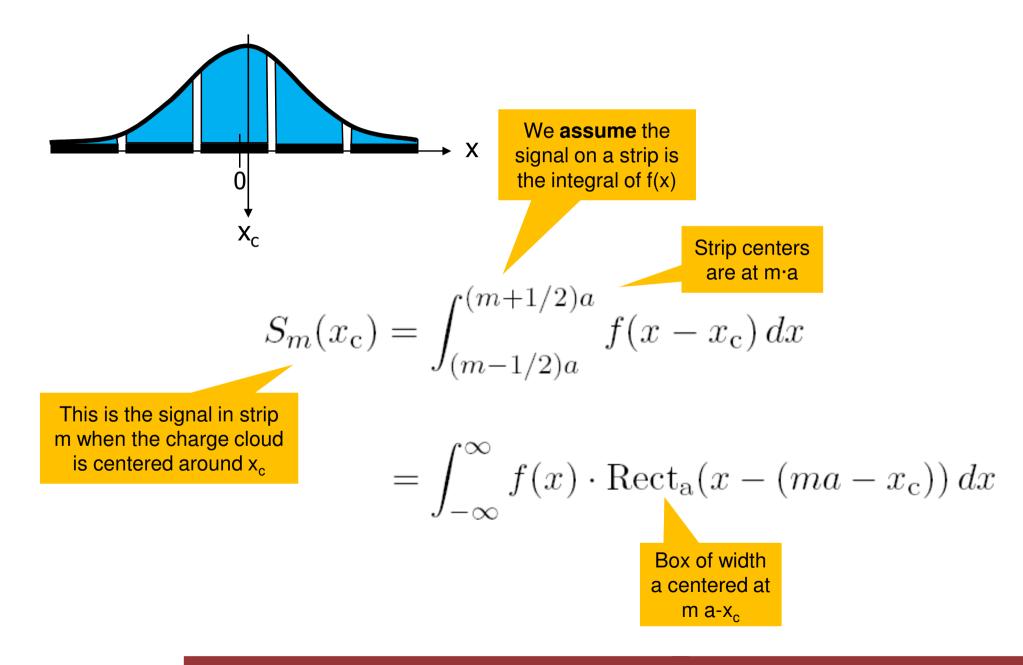


- The following calculation involves partial integrals over arbitrary function.
- Normally we must give up soon analytically (consider Gaussians..)
- But it turns out that we can...

 I could only show the result, but I like the fact that so many 'simple' aspects of basic Analysis are required...

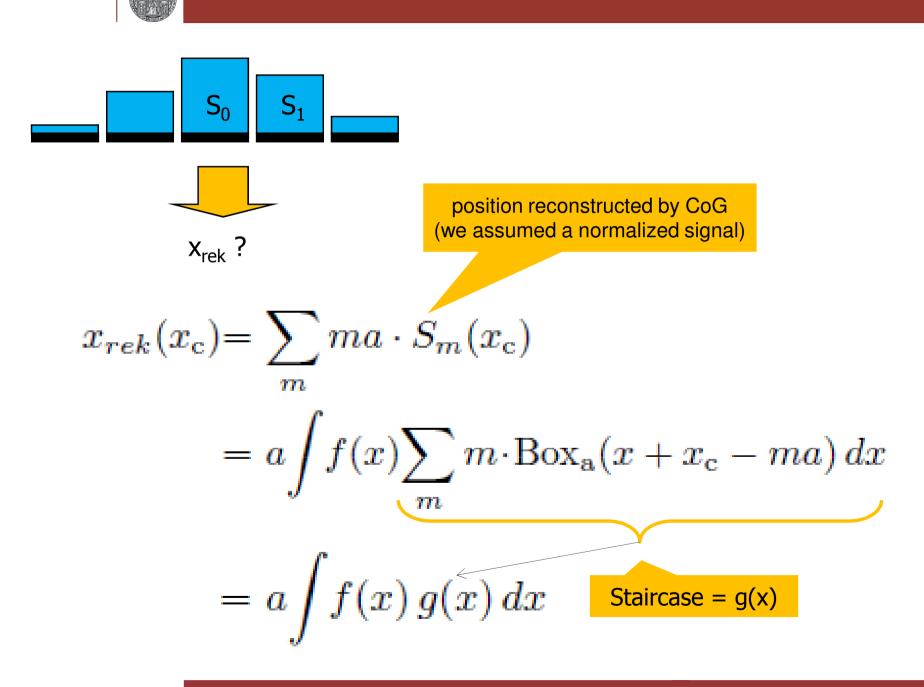


1. Signal on Strips



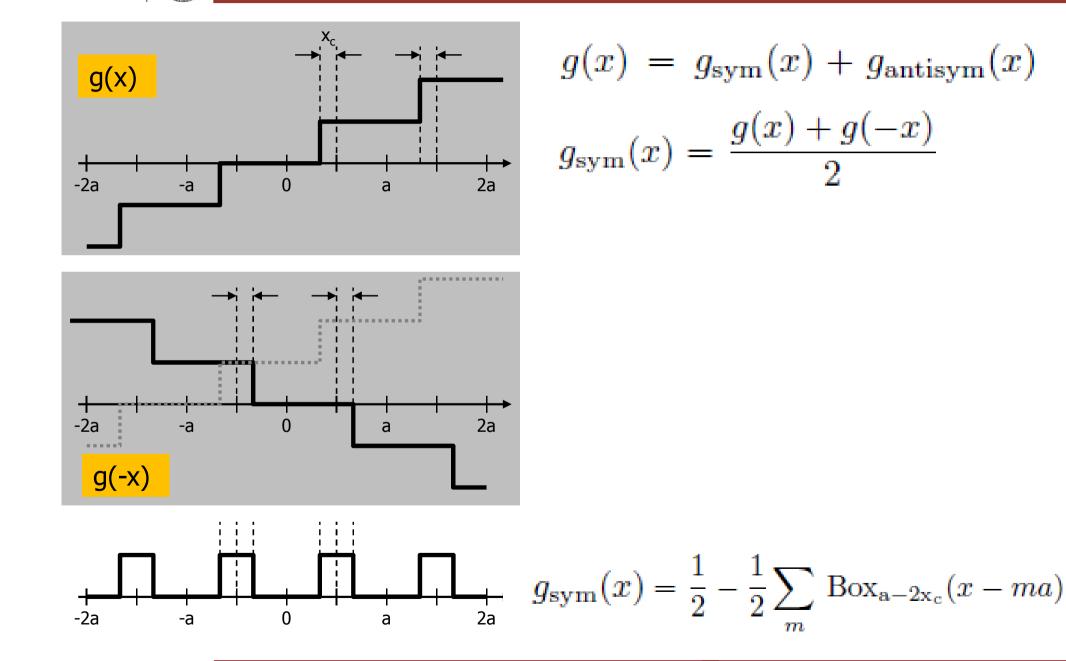


2. Reconstructed Position





3. Divide Staircase in sym. / antisym. parts



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4. Simplify the integrals, Move to Fourier Space

- Integral of g_{antisym}(x) is zero because f is symmetric
- We are left with

$$\begin{aligned} x_{rek}(x_{c}) &= a \int f(x) g_{sym}(x) dx \\ &= \frac{a}{2} - \frac{a}{2} \int f(x) \sum_{m} \operatorname{Box}_{a-2x_{c}}(x - ma) dx \\ &= \frac{a}{2} - \frac{a}{2} \int f(x) \left[\operatorname{Box}_{a-2x_{c}}(x) \star \operatorname{Comb}_{a}(x) \right] dx. \end{aligned}$$

Write Sum of Boxes as convolution of a single Box with Dirac Comb

To solve this, move to Fourier Space with

$$\widetilde{f}(k) := \int f(x) \, e^{-2\pi i k x} dx$$

• We can use $\widetilde{a \star b} = \widetilde{a} \cdot \widetilde{b}$ and $\int a(x) b(x) dx = \int \widetilde{a}(k) \widetilde{b}(k) dk$ (for symmetrical a,b)



5. Get rid of the Integral

$$x_{rek}(x_{c}) = \frac{a}{2} - \frac{a}{2} \int f(x) \left[\text{Box}_{a-2x_{c}}(x) \star \text{Comb}_{a}(x) \right] dx$$

$$= \frac{a}{2} - \frac{a}{2} \int \tilde{f}(k) \cdot \widetilde{\text{Box}}_{a-2x_{c}}(x) \cdot \widetilde{\text{Comb}}_{a}(x) dk$$
integral can be carried out. Sum is left
$$= \frac{a}{2} - \frac{a}{2} \sum_{m=-\infty}^{\infty} \tilde{f}\left(\frac{m}{a}\right) \frac{\sin\left(m\pi - \frac{2\pi mx_{c}}{a}\right)}{m\pi}$$



6. Use Symmetry, Simplify the Sin() function

$$\begin{aligned} x_{rek} &= \frac{a}{2} - \frac{a}{2} \sum_{m=-\infty}^{\infty} \tilde{f}\left(\frac{m}{a}\right) \frac{\sin\left(m\pi - \frac{2\pi mx_c}{a}\right)}{m\pi} & \text{Use symmetry} \\ &= x_c - \frac{a}{\pi} \sum_{m=1}^{\infty} \tilde{f}\left(\frac{m}{a}\right) \frac{\sin\left(m\pi - \frac{2\pi mx_c}{a}\right)}{m} & \text{Treat m=0:} \\ &\sim f(0) = 1, \\ &\text{Sin(m k)/m ->k} \end{aligned}$$
Center position x_c shows up ! $\sum_{\substack{n=1\\ \text{sin}(m\pi - x) = \sin(m\pi)\cos(x) - \cos(m\pi)\sin(x) \\ &= -(-1)^m \sin(x)} \\ &\swarrow \end{aligned}$

$$x_{err}(x_c) = \frac{a}{\pi} \sum_{m=1}^{\infty} \frac{(-1)^m}{m} \tilde{f}\left(\frac{m}{a}\right) \sin\left(\frac{2\pi mx_c}{a}\right)$$

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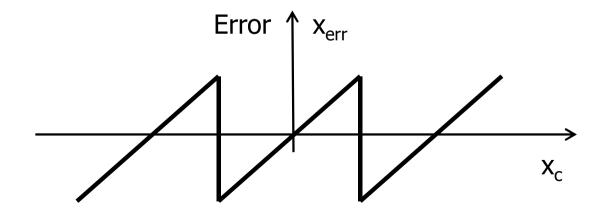


$$\boldsymbol{x_{err}}(\boldsymbol{x_{c}}) = \frac{a}{\pi} \sum_{m=1}^{\infty} \frac{(-1)^{m}}{m} \tilde{f}\left(\frac{m}{a}\right) \sin\left(\frac{2\pi m \boldsymbol{x_{c}}}{a}\right)$$

• For very narrow f(x), we have $\tilde{f}(k) \rightarrow 1$ so that

$$x_{err}(x_{c}) = \frac{a}{\pi} \sum_{m=1}^{\infty} \frac{(-1)^{m}}{m} \sin\left(m\pi \frac{2x_{c}}{a}\right)$$

This is the Fourier Series of a Saw-Tooth, as expected!

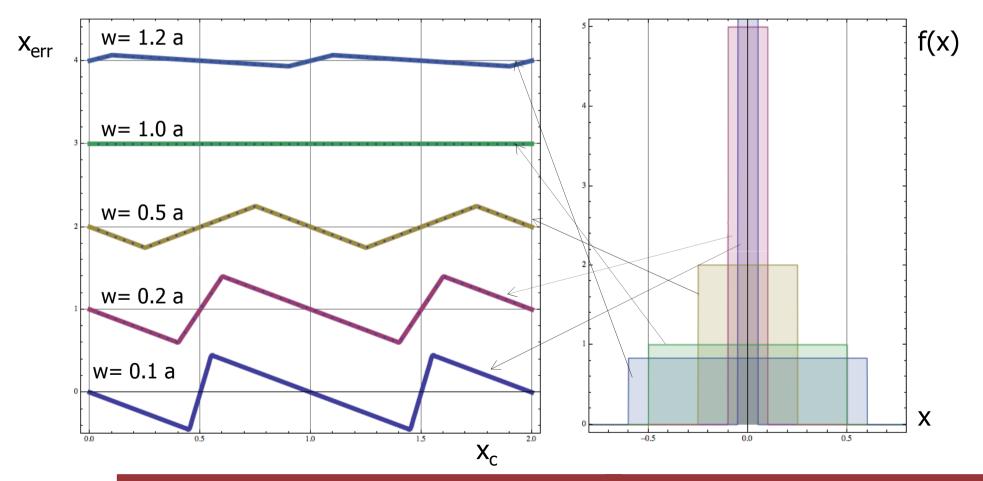




Check with Box f(x)

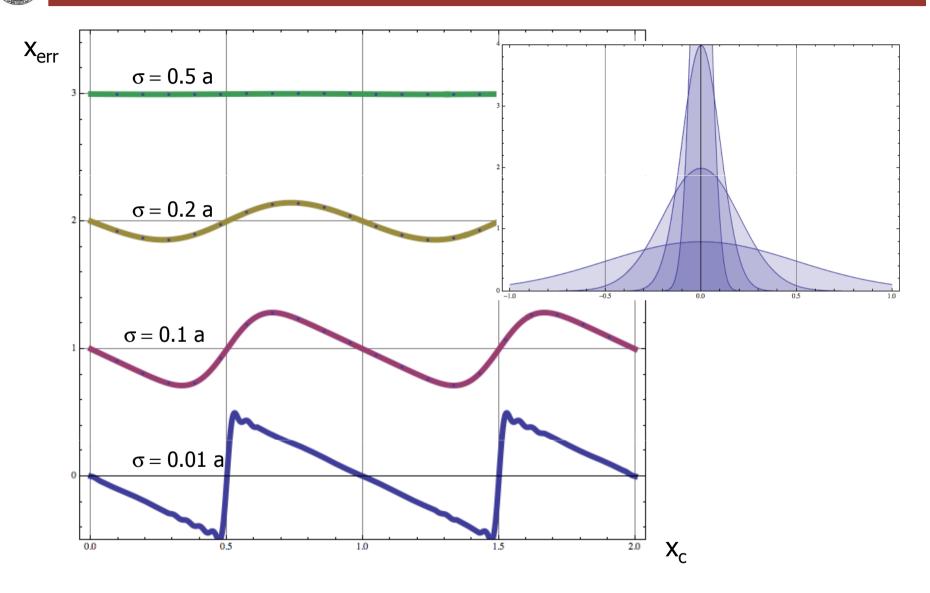
• For Box of width a, $\tilde{f}\left(\frac{m}{a}\right) = \sin(m\pi)/m\pi$ is zero for integer m. \rightarrow reconstruction is perfect..

$$\boldsymbol{x_{err}(x_c)} = \frac{a}{\pi} \sum_{m=1}^{\infty} \frac{(-1)^m}{m} \tilde{f}\left(\frac{m}{a}\right) \sin\left(\frac{2\pi m x_c}{a}\right)$$





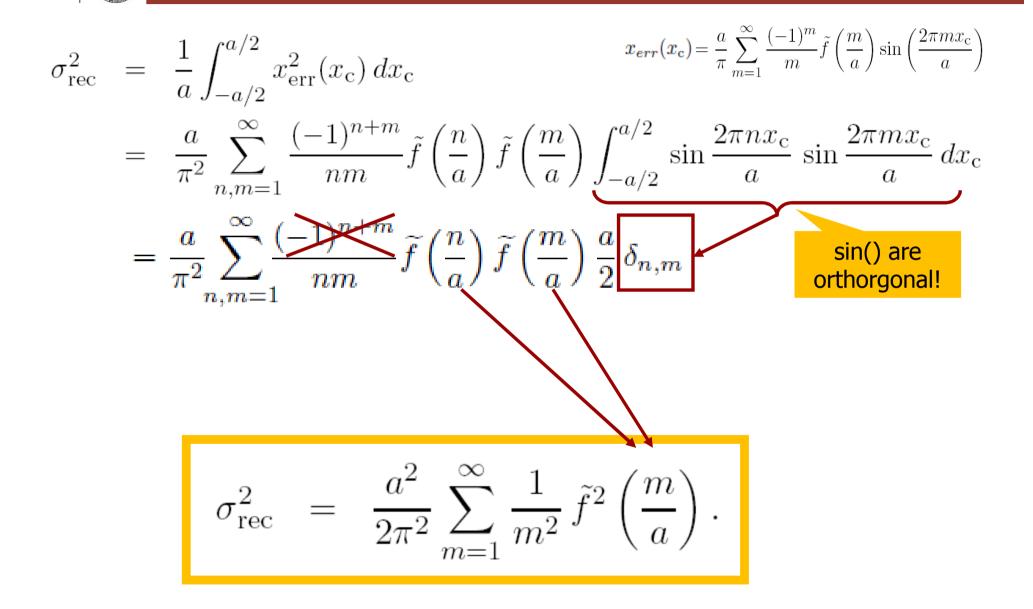
Check with Gaussians



• Error already very small for $\sigma = 0.5a$



Going Further: **Sigma** of x_{err}?





Another Check

$$\sigma_{\rm rec}^2 = \frac{a^2}{2\pi^2} \sum_{m=1}^{\infty} \frac{1}{m^2} \tilde{f}^2\left(\frac{m}{a}\right).$$

• For very narrow signals, we have again $\widetilde{f}(k) \rightarrow 1$ so that

$$\left(\frac{\sigma_{\text{err}}}{a}\right)^2 = \frac{1}{2\pi^2} \sum_{m=1}^{\infty} \frac{1}{m^2} = \frac{1}{12} \quad \text{as expected}...$$
$$\pi^2/6$$

• This is probably the most complicated way to get the 1/12...



f(x) = Gaussian / Box

 $\hfill \ensuremath{\,^{\circ}}$ For a Gaussian signal with width σ

$$G(x) = \frac{1}{\sqrt{2\pi\sigma_s}} \exp\left(-\frac{x^2}{2\sigma_s^2}\right) \quad \text{with} \quad \tilde{G}(k) = \exp\left(-2\pi^2 k^2 \sigma_s^2\right)$$
$$\text{we get} \quad \left(\frac{\sigma_{\text{err}}}{a}\right)^2 = \frac{1}{2\pi^2} \sum_{m=1}^{\infty} \frac{\exp\left(-\frac{4m^2 \pi^2 \sigma_s^2}{a^2}\right)}{m^2}$$

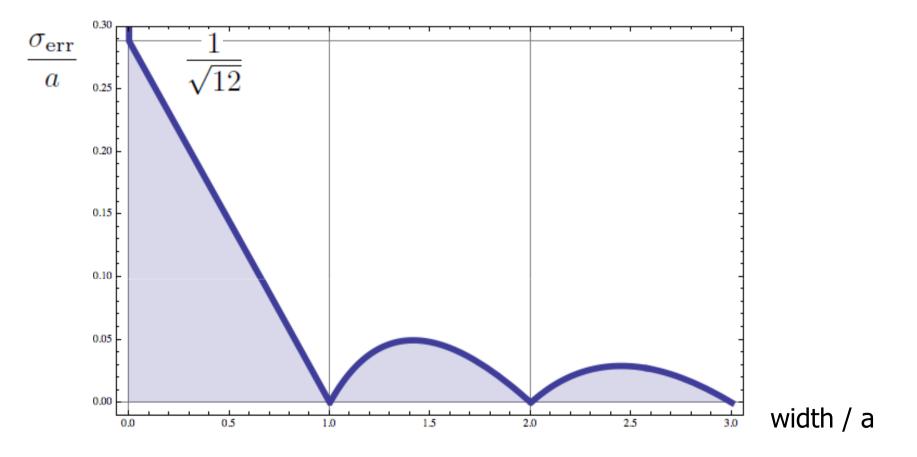
For a Box of width s a

$$\left(\frac{\sigma_{\rm err}}{a}\right)^2 = \frac{1}{360s^2} - \frac{1}{4\pi^4 s^2} \sum_{m=1}^{\infty} \frac{\cos(2\pi ms)}{m^4}$$

which is zero for integer s thanks to $\Sigma(1/n^4) = \pi^4/90..$

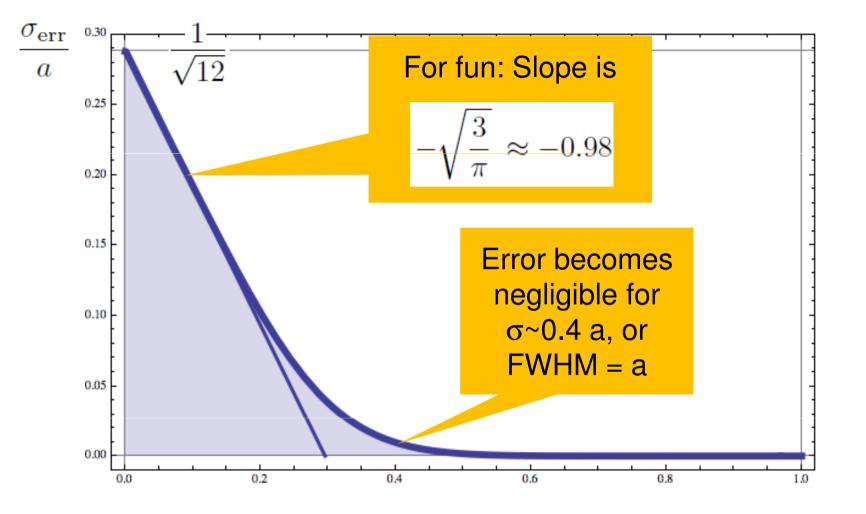
Plot this for f(x) = Box

- Error zero for integer box width.
- Behavior in between not trivial...





Plot this for f(x) = Gauss



The result 'Error \approx 0 for FWHM \approx a' can be found for many pulse shapes. We knew this... but now we know *for sure*...

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LIMIT OF SPATIAL RESOLUTION FROM NOISE



The Question

- How is spatial resolution degraded by noise?
- We all know $\sigma_{\rm err} = \kappa \cdot \sigma_{\rm n} = \frac{\kappa}{SNR}$

This states, that the resolution degrades with noise 'linearly to first order'.

- The proportionality κ is empirical. We want to calculate it
- We also want to check what happens with correlated noise
- We want to see what happens to higher order
 - What is this here? It is the distribution of the noise...
- We assume we can reconstruct with CoG (more later...)
- We restrict on a 1D treatment, but 2D is straight forward

Write down x_{rek} with noise

- A Signal at \vec{x} is distributed over N strips at positions \vec{x}_i
- Signal on i-th strip is $S_i(\vec{x})$
- The sum of all signals shall be normalized to 1:

$$\sum S_{i} = 1$$

Assume we can perfectly reconstruct position as center of gravity:

$$\vec{x} = \frac{\sum S_{i} \vec{x}_{i}}{\sum S_{i}} = \sum S_{i} \vec{x}_{i}$$

- Now assume noise $n_{
 m i}$ on all strips
- the reconstructed position is:

$$\vec{x}_{\rm rek}(\vec{x}) = \frac{\sum (S_{\rm i} + n_{\rm i})\vec{x}_{\rm i}}{\sum (S_{\rm i} + n_{\rm i})} = \frac{\vec{x} + \sum n_{\rm i}\vec{x}_{\rm i}}{1 + \sum n_{\rm i}}$$



Assume noise is small. Get the standard dev.

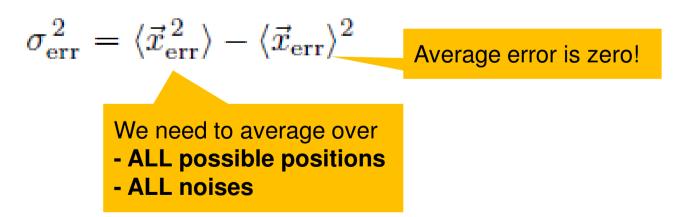
This becomes (Taylor Expansion of Denominator):

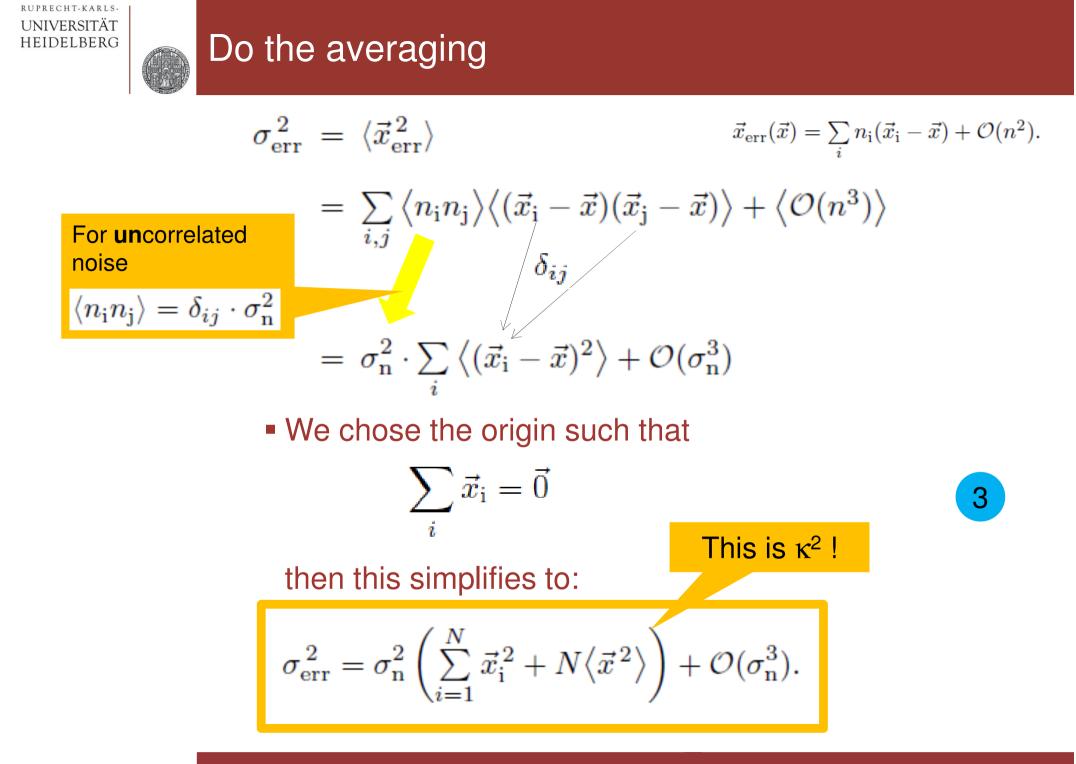
$$\vec{x}_{\rm rek}(\vec{x}) = \frac{\vec{x} + \sum n_{\rm i}\vec{x}_{\rm i}}{1 + \sum n_{\rm i}} = \left(\vec{x} + \sum n_{\rm i}\vec{x}_{\rm i}\right) \left(1 - \sum n_{\rm i} + \mathcal{O}(n^2)\right)$$

The reconstruction error is:

$$\vec{x}_{\mathrm{err}}(\vec{x}) = \sum_{i} n_{\mathrm{i}}(\vec{x}_{\mathrm{i}} - \vec{x}) + \mathcal{O}(n^2).$$

We need the standard deviation:





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Example: Strips

- Consider two strips at $x_1 = -a/2$ and $x_2 = +a/2$ (N = 2)
- Signals for a hit at x are

 $S_1(x) = (x_2 - x)/a$ and $S_2(x) = (x + x_2)/a$

• 1, 2 and 3 are fulfilled:

S₁ + S₂ = 1; X₁ S₁ + X₂ S₂ = X; X₁ + X₂ = 0 • We get $\left(\frac{\sigma_{\text{err}}}{\sigma_{\text{n}}}\right)^2 \approx x_1^2 + x_2^2 + \frac{2}{a} \int_{x_1}^{x_2} x^2 dx = \frac{2}{3} a^2$

• Or
$$\sigma_{\rm err} = 0.816 \cdot a \cdot \sigma_{\rm n}$$

For σ_n = 0.1 (Signal/Noise = 10), resolution = 8% · a
Resolution is better than optimal binary readout for S/N>5.6



Correlated Noise ?

• For FULLY correlated noise, $n_i = n_i$ and $\langle n_i n_j \rangle = \sigma_n^2$

• We get
$$\sigma_{
m err}^2 pprox \sigma_{
m n}^2 N^2 \langle ec{x}^2
angle$$

• For the strip example

$$\sigma_{err} = a \sigma_n / \sqrt{3} = 0.57 a \sigma_n$$

- Correlated noise is less harmful than 'normal' noise
- Note: For mixed noise, superimpose both components)
- Note: If the Amplitude of the signal is KNOWN (X-ray), noise becomes correlated and resolution improves!



Higher Orders (in noise)

- Noise can have different distributions.
- They have different higher moments:

$$\begin{array}{lll} \langle n_i^2 \rangle & = & \sigma_{\rm n}^2, \\ \langle n_i^4 \rangle & = & \beta \cdot \sigma_{\rm n}^4. \end{array}$$

 $\underline{2}$

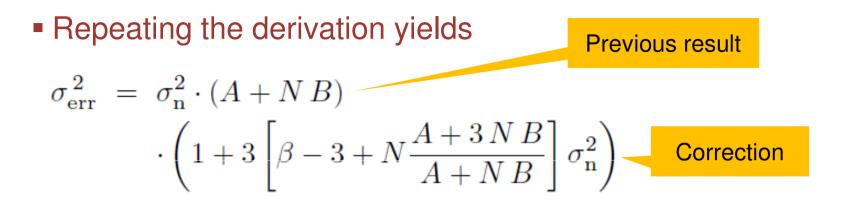
• They are
$$\beta := \frac{\int n^4 p(n) dn}{\left(\int n^2 p(n) dn\right)^2} = \begin{cases} 3 : \text{Gauss} \\ 9/5 : \text{Box} \\ 1 : \text{Peaks} \end{cases}$$

• We need then higher order correlations (not trivial..):

$$\begin{aligned} \langle n_{\rm i} \rangle &= 0\\ \langle n_{\rm i} n_{\rm j} \rangle &= \delta_{ij} \sigma_{\rm n}^2\\ \langle n_{\rm i} n_{\rm j} n_{\rm k} \rangle &= 0\\ \langle n_{\rm i} n_{\rm j} n_{\rm k} n_{\rm l} \rangle &= \delta_{ij} \delta_{jk} \delta_{kl} \left(\beta - 3\right) \sigma_{\rm n}^4 \quad + (\delta_{ij} \delta_{kl} + \delta_{ik} \delta_{jl} + \delta_{il} \delta_{jk}) \sigma_{\rm n}^4 \end{aligned}$$



Higher Orders

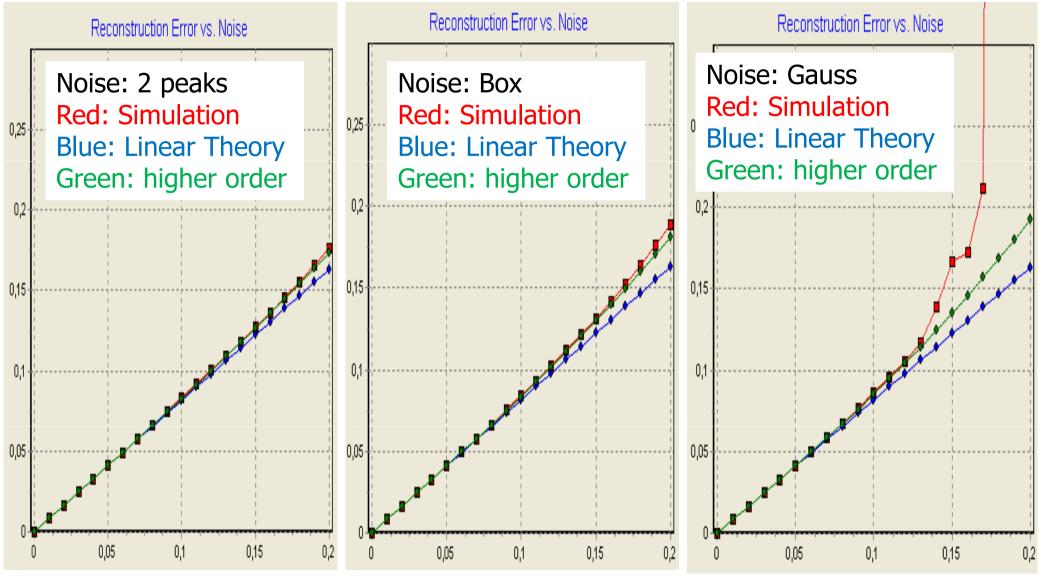


- Only the *correction* depends on the 'type' (shape) of noise.
- For small noise, there is *no need* to simulate Gaussian noise:

Randomly adding or subtracting $\pm \sigma_n$ has the same effect!

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Is this true? \rightarrow Small Monte Carlo



Reconstruction for Gauss noise fails in few cases

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- Can be treated similarly
- Observations:
 - Small number of electrodes is good
 - Well confined acceptance 'circle' is good

$$\sigma_{\rm err}^2 = \sigma_{\rm n}^2 \left(\sum_{i=1}^N \vec{x}_i^2 + N(\vec{x}^2) \right) + \mathcal{O}(\sigma_{\rm n}^3).$$

Geometry	σ_{1D} theory		numerical
	linear	correction	(A=p=1)
strips	$\sqrt{\frac{2}{3}} \cdot p$	$\sqrt{1+3\beta\sigma_{\rm n}^2}$	0.8165
square	$\frac{2}{\sqrt{3}} \cdot \sqrt{A}$	$\sqrt{1+3(3+\beta)\sigma_{\rm n}^2}$	1.1547
hexagon	$\frac{\sqrt{5} \cdot 3^{1/4}}{6} \cdot \sqrt{A}$	$\sqrt{1+3\left(\frac{6}{5}+\beta\right)\sigma_{\mathrm{n}}^2}$	0.4905

Hexagons are best.

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BACK TO GOG Now with Noise

Problems with Centroid

- Resolution for small σ is bad \rightarrow make f(x) wide
- BUT: The 'infinite' sum creates infinite noise (N high)
- Must chose N so that reconstruction is ok.
 - Obviously N ~ σ
- The choice is fairly arbitrary
- And:
 - In real system, there is often a **threshold** (hits below this are not read out)
 - The reconstructed amplitude is wrong (signals below threshold are lost)
 - Broken pixels need special treatment

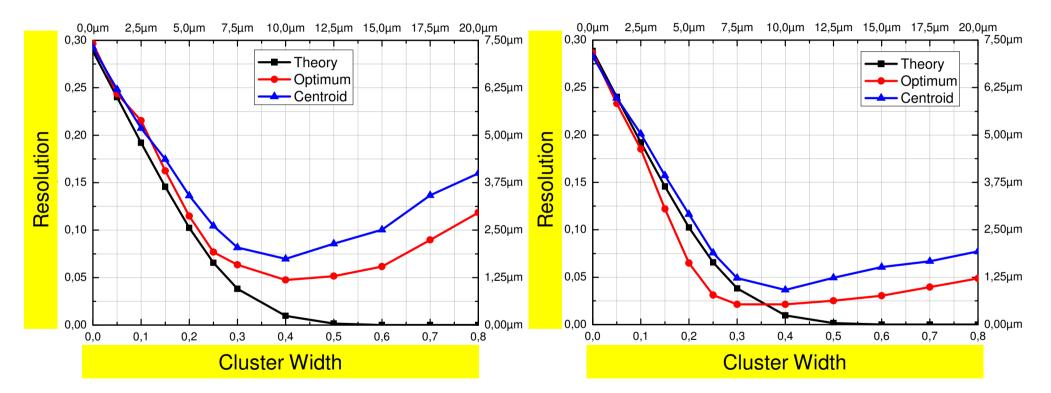


Monte Carlo Simulation

Many possibilities... I do not go in details

S/N = 40





The optimum signal width is still close to FWHM = a!

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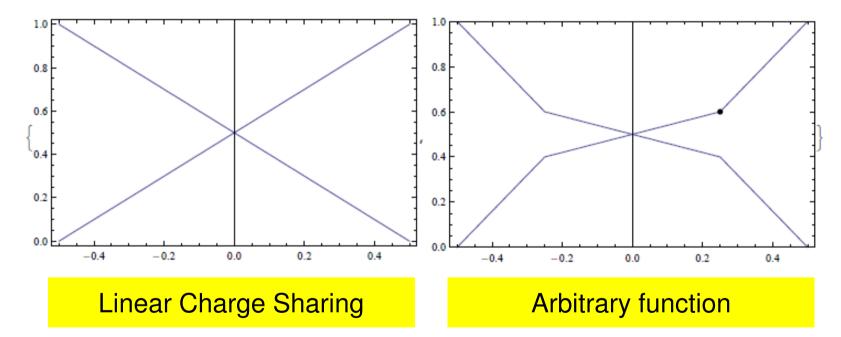


ETA FUNCTION



Motivation

- Often the Signals Distribution function (e.g. on 2 strips) is not linear.
- This is related to the 'famous' eta-function.



- The position then cannot be calculated by GoG, but by using the inverse function (or the 'eta'-lookup table)
- Question: How does resolution depend on f(x)?



1D Case: Reconstruction with Inverse Function

The signals on the two strips shall be $S_1(x) = Qf(x)$

$$S_2(x) = Q - S_1(x) = Q (1 - f(x))$$

(we assume no signal is lost, i.e. we require $S_1+S_2 = Q$

We require

- f(x) is strictly monotonic (obvious)

• f(x) shall be symmetric in x (may not always be the case)

Obviously

$$x_{rek} = f^{-1} \left[\frac{S_1}{S_1 + S_2} \right]$$



Adding Noise

With Noise on S₁ and S₂ we get

$$\begin{split} x_{rek} &= f^{-1} \begin{bmatrix} f(x) + n_1 \\ 1 + n_1 + n_2 \end{bmatrix} & \text{Add noise} \\ &\approx f^{-1} \left[(f(x) + n_1)(1 - n_1 - n_2) \right] & \text{Taylor (as before)} \\ &\approx f^{-1} \left[f(x) + n_1(1 - f(x)) - n_2 f(x) \right] \\ &\approx x + \frac{\mathrm{d}f^{-1}(s)}{\mathrm{d}s} \bigg|_{f(x)} \cdot \left[n_1(1 - f(x)) - n_2 f(x) \right] \\ &\qquad \mathrm{Taylor Series for f^1 around f(x)} \end{split}$$

$$x_{err} = \frac{n_1(1 - f(x)) - n_2 f(x)}{f'(x)}$$

A 'forgotten' math theorem: The derivative of the inverse function is the inverse of the derivative



Sigma – Averaging over Noise

To get

$$\sigma_{err}^2 = \langle x_{err}^2 \rangle - \langle x_{err} \rangle^2$$

$$x_{err} = \frac{n_1(1 - f(x)) - n_2 f(x)}{f'(x)}$$

• we average first over noise. We get

$$\sigma_{err}^2 = \sigma_n^2 \left\langle \frac{1 - 2f + 2f^2}{f'^2} \right\rangle + 2 \left\langle n_1 n_2 \right\rangle \left\langle \frac{f^2 - f}{f'^2} \right\rangle$$

- Coefficients depend on the signal shape
- They are small where the response function is steep (obvious..)

For uncorrelated noise, only the first term matters

$$\frac{\sigma_{err}^2}{\sigma_n^2} = \left\langle \frac{1-2f+2f^2}{f'^2} \right\rangle \quad \begin{array}{l} \text{Average over} \\ \text{position} \end{array}$$

Back to linear Interpolation

- What does this mean for linear interpolation, f(x) = x+0.5 ?
- Let us first look at the *position dependent* error

$$\frac{\sigma_{err}^2}{\sigma_n^2} = \frac{1 - 2f + 2f^2}{f'^2} = \frac{1}{2} + 2x^2$$

- This is NOT constant. It doubles at the edges !!!
- The average error is

$$\frac{\sigma_{err}^2}{\sigma_n^2} = \int_{-1/2}^{1/2} \left(\frac{1}{2} + 2x^2\right) = \frac{2}{3} \qquad \text{(s. page 28)}$$

as before.

- Very exciting: Can we find a f(x) such that the integral is better than with linear interpolation
 - Probably not. But let's see...
- Can we find a distribution function so that the error is independent of position?
- One line of Mathematica is enough:

$$f_{flat}(x) = \frac{1}{2} \left(1 - \operatorname{Sinh}[2x\operatorname{ArcSinh}(s)] \right)$$

• The average σ^2 is 0.643, which is (a bit) **better** than 2/3 !!

• We found a distribution which is better than lin. interpolation!

- Basic Algebra is fun.....
- CoG is 'perfect' when signal width = strip width
- Larger signals are worse when there is noise
- Ideal κ for strips is 0.816
- Analogue readout for S/N<6 is useless.</p>
- Noise shape does not matter for S/N > 10
- Correlated noise is less harmful
- Hexagons have better resolution
- Linear interpolation has more error at the edges.
- There is a better reconstruction function than linear
 - but the difference is negligible....





Thank you for your attention!

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