

selected topics



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the underlying

physics

neutrino



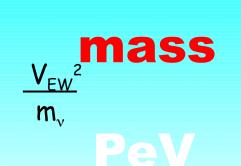
Additional materials

$\mathbf{3}v$ - paradigm

All well established/confirmed three neutrinos results fit well a framework with interactions described by the standard model with masses and mixing It is widely believed that Connection orana ature exist Zere conserved charges Ma Where is V_{R} 9

Scales

of new physics



LHC

eV-

GUT - Planck

Electroweak -

28 orders of magnitude High scale seesaw Quark- lepton symmetry /analogy GUT

Low scale seesaw, radiative mechanisms, RPV, high dimensional operators

Scale of neutrino masses themselves Relation to dark energy, MAVAN?



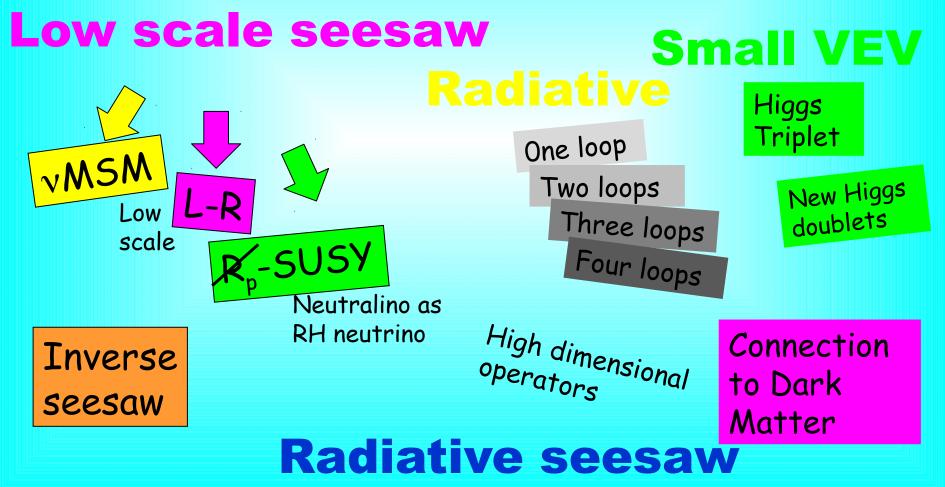
Neutrino mass itself is the fundamental scale of new physics

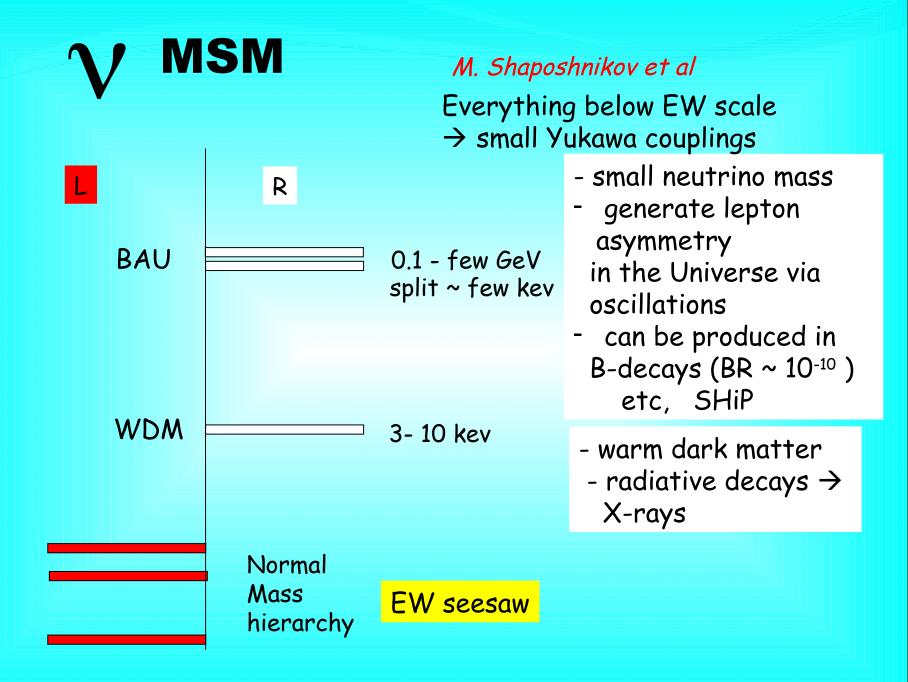
Spurious scale?

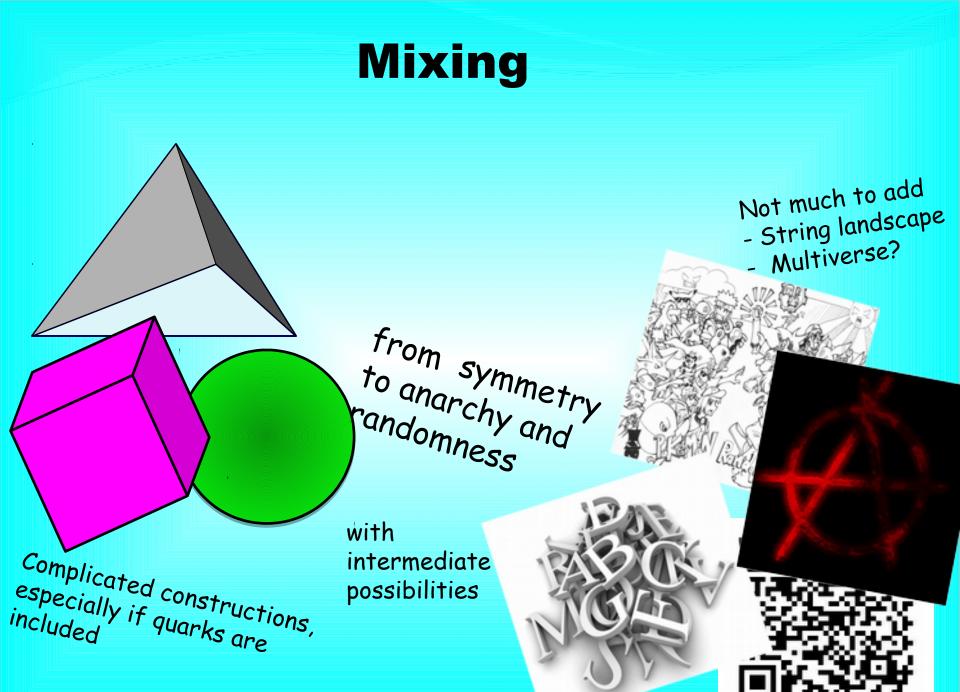
m

EW – LHC scale

- No hierarchy problem (even without SUSY)
- testable at LHC, new particles at 0.1 few TeV scale
- LNV decays







Trends and implications

No new physics at LHC in particular new physics which could be associated to low scale mechanisms of neutrino mass generation

- right Handed neutrinos, new heavy leptons
- right handed gauge bosons of the L-R symmetric models
- double charged scalars (of seesaw type II), etc
- new fermions and scalars which can participate in the radiative mechanism of neutrino mass generation

 \rightarrow Bounds on masses / couplings of these new particles

No Lepton number violation, MEG,

Nothing yet at well motivated TeV-scale. The next motivated scales are intermediate and then GUT



observations

and hints

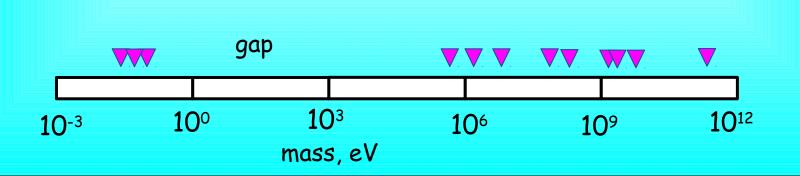
Analysing data

Smallness of mass?

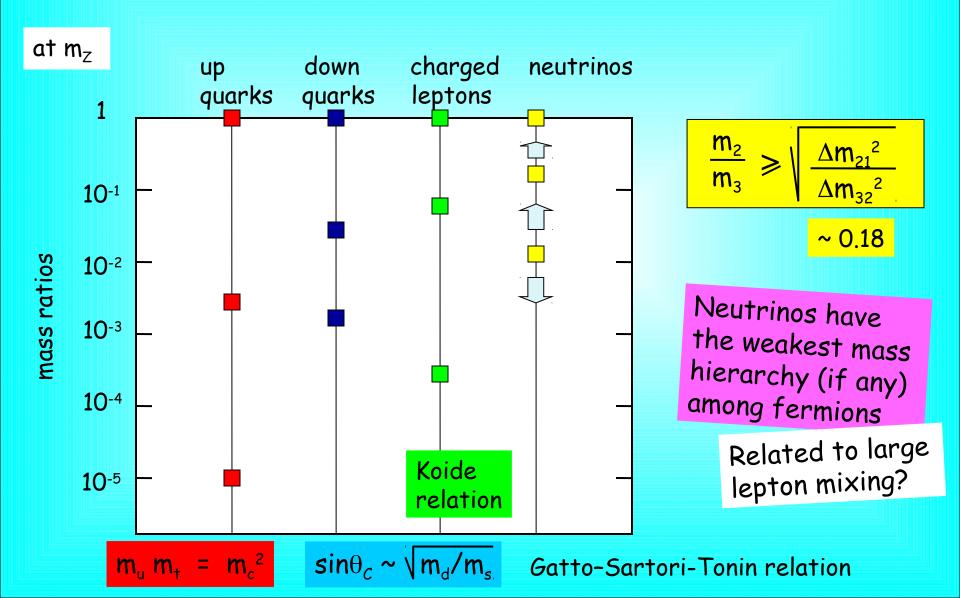
 \mathbf{m}_{t}

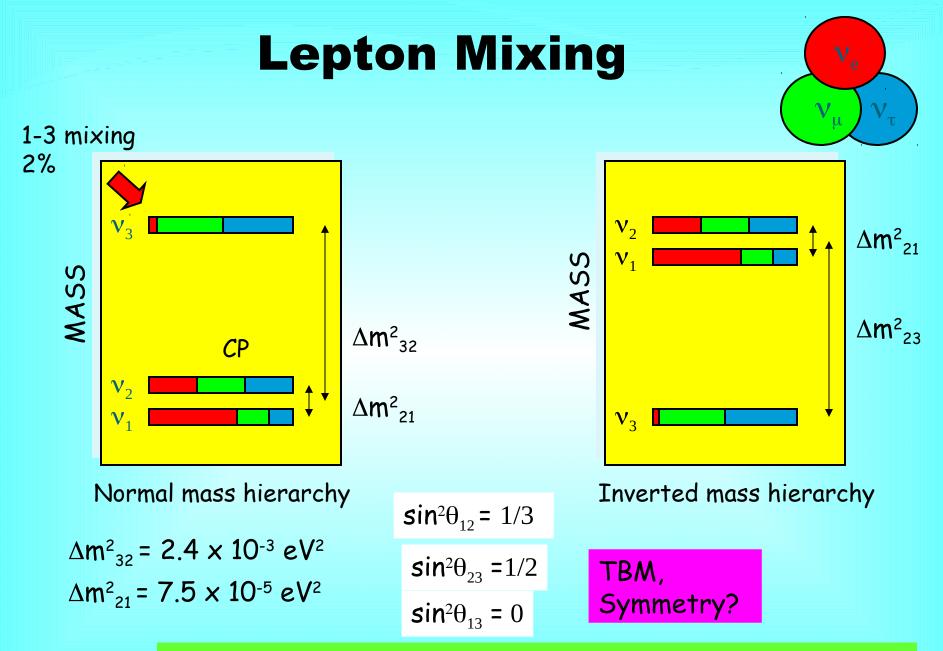
m_e

Special	comparing within generation:
$\frac{m_3}{m_{\tau}} \sim 3.10$	Similar for other generations if spectrum is hierarchical
Normal?	Neutrinos: no clear generation structure and correspondence light flavor – light mass, especially if the mass hierarchy is inverted or spectrum is quasi-degenerate
$\frac{m_3}{2} \sim 3$	10^{-6} $\frac{m_e}{m}$ ~ 3 10^{-6}



Mass hierarchies





For antineutrinos spectra are different (distribution of the v_{μ} and v_{τ} - flavors in v_1 and v_2) due to possible CP-violation

TBM Mixing pattern

$$U_{tbm} = \begin{pmatrix} \sqrt{2/3} & \sqrt{1/3} & 0 \\ -\sqrt{1/6} & \sqrt{1/3} & -\sqrt{1/2} \\ -\sqrt{1/6} & \sqrt{1/3} & \sqrt{1/2} \\ 0.62 \\ 0.78 \end{pmatrix}$$

$$U_{tbm} = U_{23}(\pi/4) U_{12}$$

sin² θ_{12} = 1/3 0.30- 0.31

Accidental, numerology, useful for bookkeeping

> Accidental symmetry (still useful)

Not accidental

Lowest order approximation which corresponds to weakly broken (flavor) symmetry of the Lagrangian with some other physics and structures associated

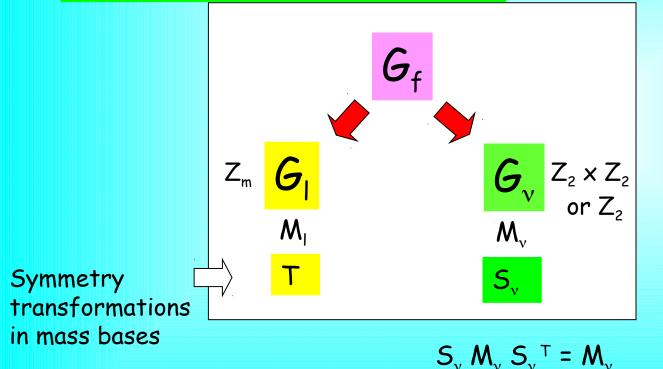
There is no relation of mixing with masses (mass ratios)

Parameters look like C-G coefficients

Residual symmetries approach ^{E. Ma,} C. S.

Mixing appears as a result of different ways of the flavor symmetry breaking in the neutrino and charged lepton (Yukawa) sectors.

No connection of masses and mixing



 $\begin{array}{|c|c|c|c|} A_4 & S_4 & T_7 \\ \hline & T_7 & T_7 \end{array}$

Residual symmetries of the mass matrices

Lam

Generic symmetries which do not depend on values of masses

Discrete finite groups Flavons to break symmetries avons

Realizations: too complicated

Intrinsic symmetries

Realized for arbitrary values of neutrino and charged lepton mass

Assumptions

Majorana neutrinos
symmetry can be embedded in SU(3)

D. Hernandez, A.S.

in the mass basis

$$m = diag(m_1, m_2, m_3)$$

$$G_{v}$$

$$S_1 = diag(1, -1, -1)$$

 $S_i^2 = I \qquad Z_2 \times Z_2$

The Klein group

$$m_{I} = diag(m_{e}, m_{\mu}, m_{\tau})$$

$$G_{I}$$

$$T = diag(e^{i\phi_{e}}, e^{i\phi_{\mu}}, e^{i\phi_{\tau}})$$

$$\phi_{\alpha} = 2\pi k_{\alpha}/m$$

$$T^{m} = I \qquad Z_{m}$$

$$\Sigma \phi_{\alpha} = 0$$

the simplest discrete symmetries which left the matrix diagonal

in flavor basis $S_i^{f} = U_{PMNS} S_i U_{PMNS}^{+}$

nsic symmetries = residual symmetri

If intrinsic symmetries are residual symmetries of the unique symmetry group (follow from breaking of unique group) → bounds on elements of mixing matrix 1204.0445

P -integer Symmetry group

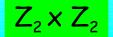
Equation gives 2 relations between mixing parameters condition

 $(U_{PMNS}S_{i}U_{PMNS}^{+}T)^{p} = I$

Two such equations for i = 1,2 fix the mixing matrix completely → TBM

Equivalent to

$$\begin{array}{|c|c|c|} Tr \left(U_{PMNS} S_{i} U_{PMNS}^{+} T \right) = a \\ a = \sum_{j} \lambda_{j} \\ \lambda_{j}^{p} = 1 \quad j = 1,2,3 \end{array} \end{array} \begin{array}{|c|} Tr \left(W_{iU} \right) = a \\ Tr \left(W_{iU} \right) = a \\ \lambda_{j} - three \ eigenvalues \\ of \ W_{iU} \end{array}$$



Example

$$G_{v} = Z_{2}$$

Two relations

D. Hernandez, A Y S. 1304.7738 [hep-ph]

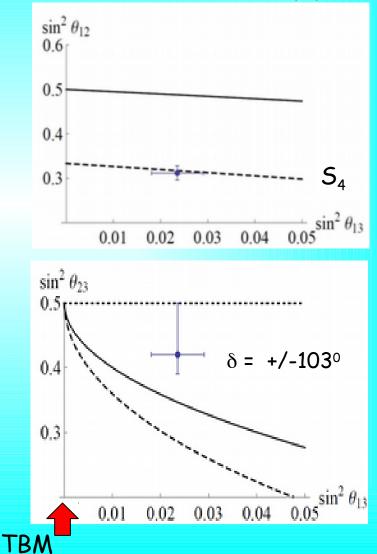
In the residual symmetries approach

for column of the mixing matrix:

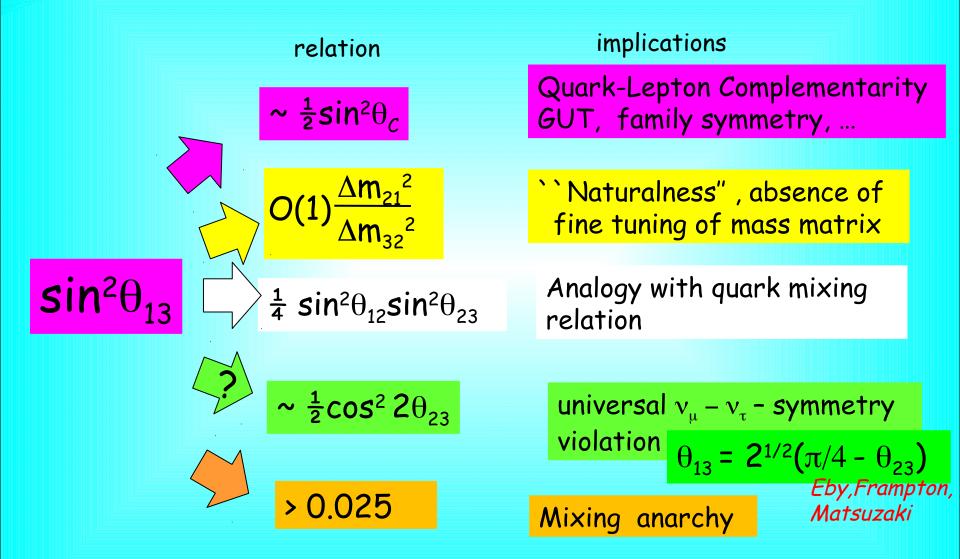
 $|U_{\beta i}|^2 = |U_{\gamma i}|^2$ $|U_{\alpha i}|^2 = \frac{1+a}{4 \sin^2(\pi k/m)}$

k, m, p integers which determine symmetry group

 $k_{\alpha} = 0$



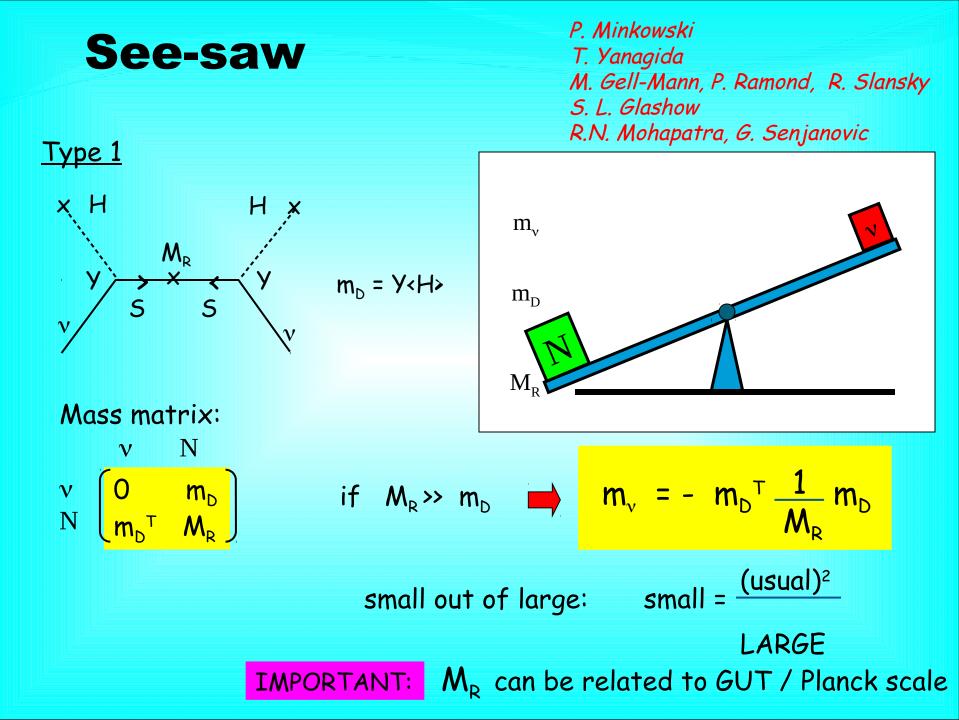
olation of TBM: 1-3 mixing ent implications



One can use $G_v = Z_2$ to relax symmetry restrictions and accommodate non-zero 1-3 mixing

Smallness of

Neutrino mass and seesaw



$\nu\text{-}$ mass and Higgs physics

Correction to λ - 4 point

 V_{R}

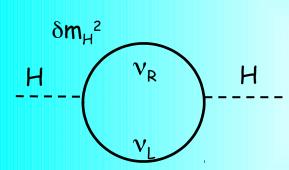
coupling - vacuum stability

H

H

bottom -up

Correction to Higgs mass



Upper bound on mass M_R < 10⁷ GeV

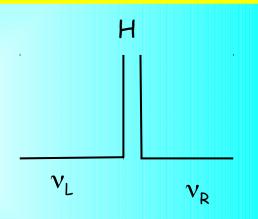
- → leptogenesis ?
- cancellation (a kind of SUSY)
- F. Vissani ... J Elias-Miro et al, R Volkas, et al, M. Fabbrichesi ...

Other contributions from particles associated to neutrino mass generation, e.g. Higgs triplets

H

C. Bonila et al, 1506.04031

Higgs as composite state of neutrinos



New strong int. Generate 4 fermionic coupling

Recent: J. Krog, C. T. Hill 1506.02843

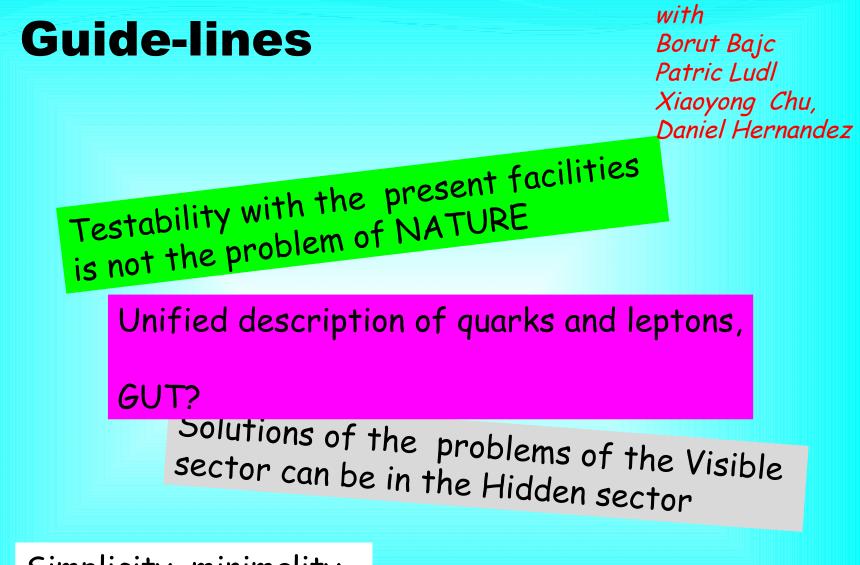
Neutrinos and

Hidden world

Theoretical

perspective

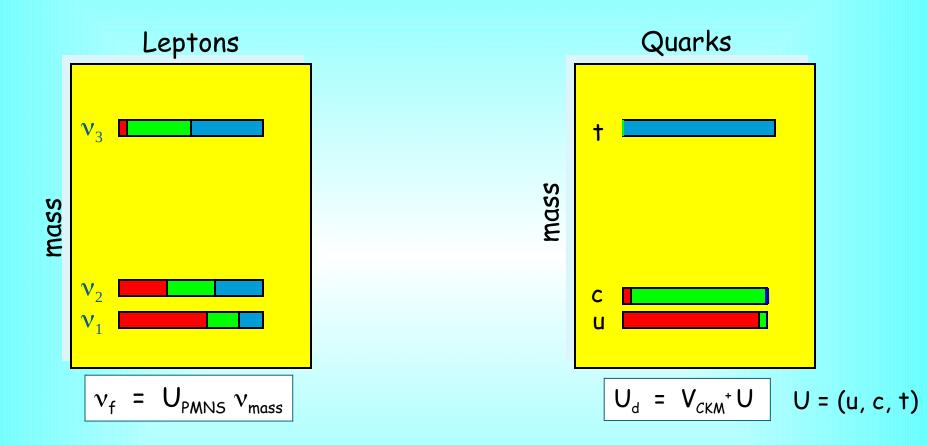




Simplicity, minimality

Old does not mean wrong

Leptons versus quarks



Mixings of quarks and leptons are strongly different but still related

Observation:

$$\theta_{12}^{|} + \theta_{12}^{|} \sim \pi/4$$

 $\theta_{23}^{|} + \theta_{23}^{|} \sim \pi/4$

Sum up to maximal mixing angle kind of complementarity

In general

C. Giunti, M. Tanimoto H. Minakata, A Y S Z - Z. Xing J Harada S Antusch , S. F. King Y Farzan, A Y S M Picariello ,etc.

$$U_{PMNS} = V_{CKM} U_X$$

Ux special matrix close to bi-maximal or TBM matrix:

$$U_{X} = \begin{cases} U_{BM} = U_{23}(\pi/4) U_{12}(\pi/4) \\ U_{TBM} = U_{23}(\pi/4) U_{12}(\theta_{TBM}) \end{cases}$$

Prediction:if $U_x = U_{23}(\pi/4) U_{12}$ $\theta_{13}^x \sim 0$

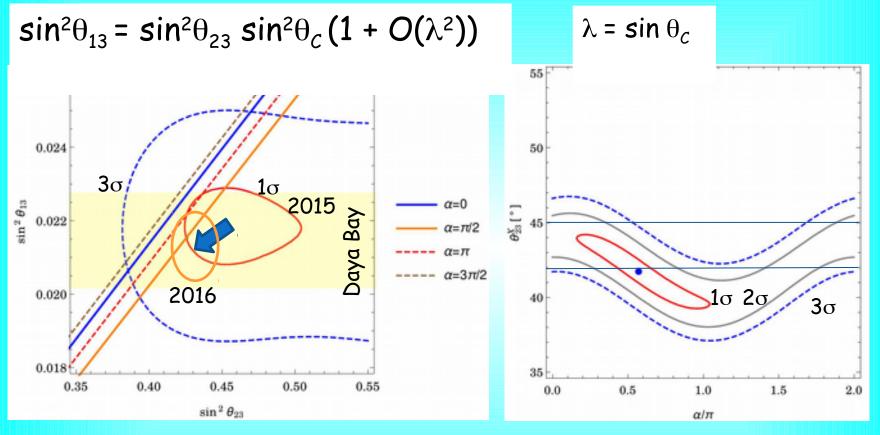
permutation - to reduce the lepton mixing matrix to the standard form gives

$$\sin \theta_{13} \sim \sqrt{\frac{1}{2}} \sin \theta_c$$

Now more than 3σ off

General relation

Normal mass ordering



Dependence of 1-3 mixing on 2-3 mixing for different values of the phase α . Allowed regions from the global fit NuFIT 2015

Allowed values of parameters of U_X Best fit value: $\theta_{23}^{\times} = 42^{\circ}$

RGE effect from maximal mixing value at high scale

What does this mean?

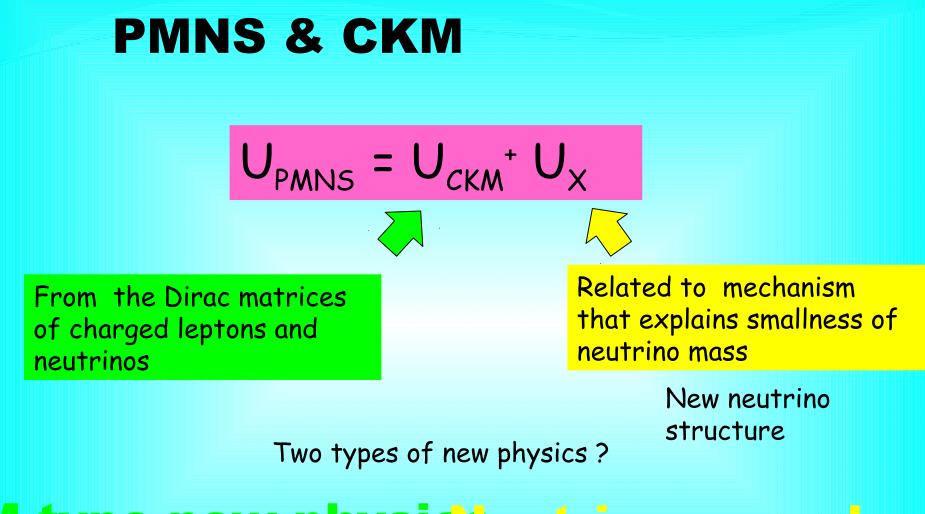


Quarks and leptons know about each other, Q L unification, GUT or/and Common flavor symmetries



Some additional physics is involved in the lepton sector which explains smallness of neutrino mass and difference of the quark and lepton mixing patterns





M type new physicheutrino new physic

Can be naturally realized in the seesaw type I which after all is the most appealing mechanism of explanation of smallness of neutrino mass

More than usual see-saw?

ale of see-sa
$$\mathbf{W} = -\mathbf{m}_{D}^{T} \frac{1}{\mathbf{m}_{v}} \mathbf{m}_{D}$$

 $q - l similarity: \mathbf{m}_{D} \sim \mathbf{m}_{q} \sim \mathbf{m}_{l}$
for one third generations $\mathbf{M}_{R} \sim 2 \ 10^{14} \ \text{GeV}$
new scale ?

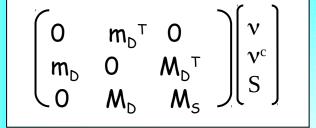
avor structureDifficult to reproduce

Can be explained in the framework of double seesaw

Double Seesaw

R.N. Mohapatra J. Valle

Three additional singlets S which couple with RH neutrinos



 $M_s >> M_D$ M_s - scale of B-L violation

RH neutrinos get mass via see-saw

$$\mathbf{M}_{\mathsf{R}} = \mathbf{M}_{\mathsf{D}}^{\mathsf{T}} \mathbf{M}_{\mathsf{S}}^{-1} \mathbf{M}_{\mathsf{D}}$$

This explains

1. strong mass hierarchy $M_{\rm D} \sim m_{\rm D}$ and $M_{\rm S}$ has no strong hierarchy

 T_{D}

- 2. intermediate scale of masses if $M_{s} \sim M_{Pl}$, $M_{D} \sim M_{GU}$
- 3. Flavor structure:

$$m_{v} = m_{D}^{T} M_{D}^{-1T} M_{S} M_{D}^{-1} m_{s}$$

$$f m_{D} = A M_{D} \implies m_{v} \sim M_{S}$$

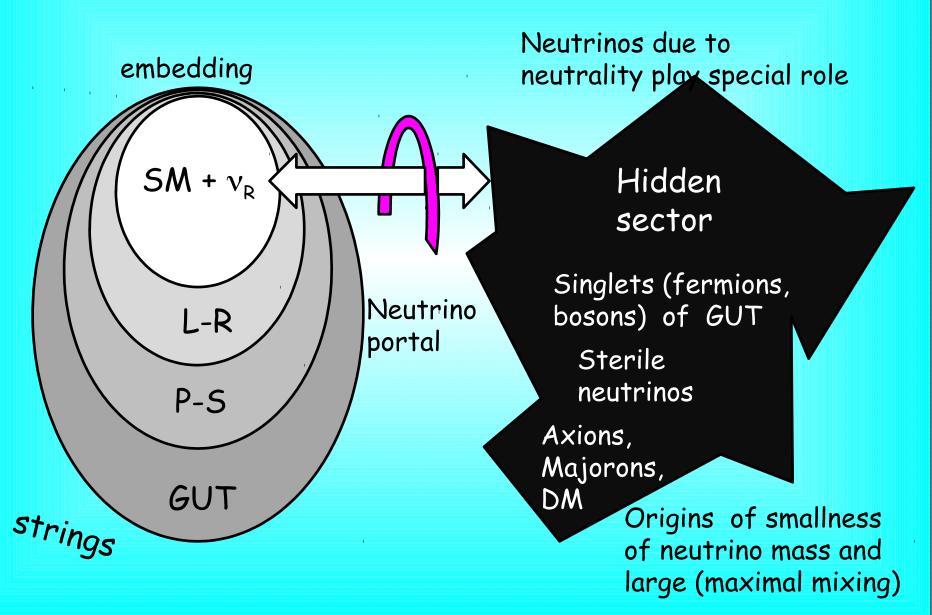
Screening of the Dirac structure

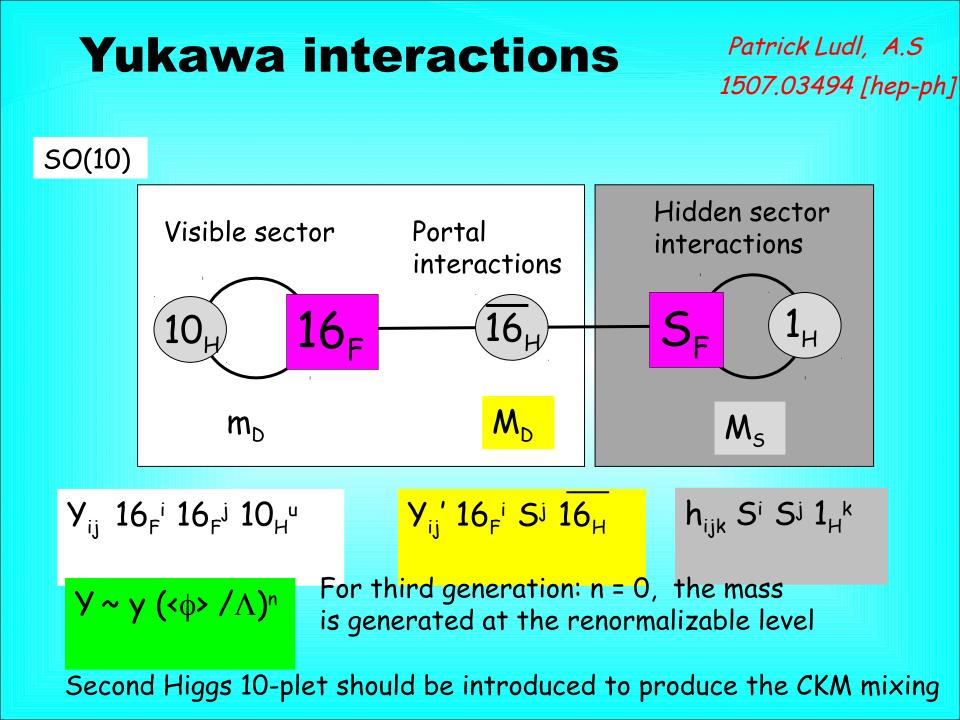
may have certain M. Lindner, symmetries

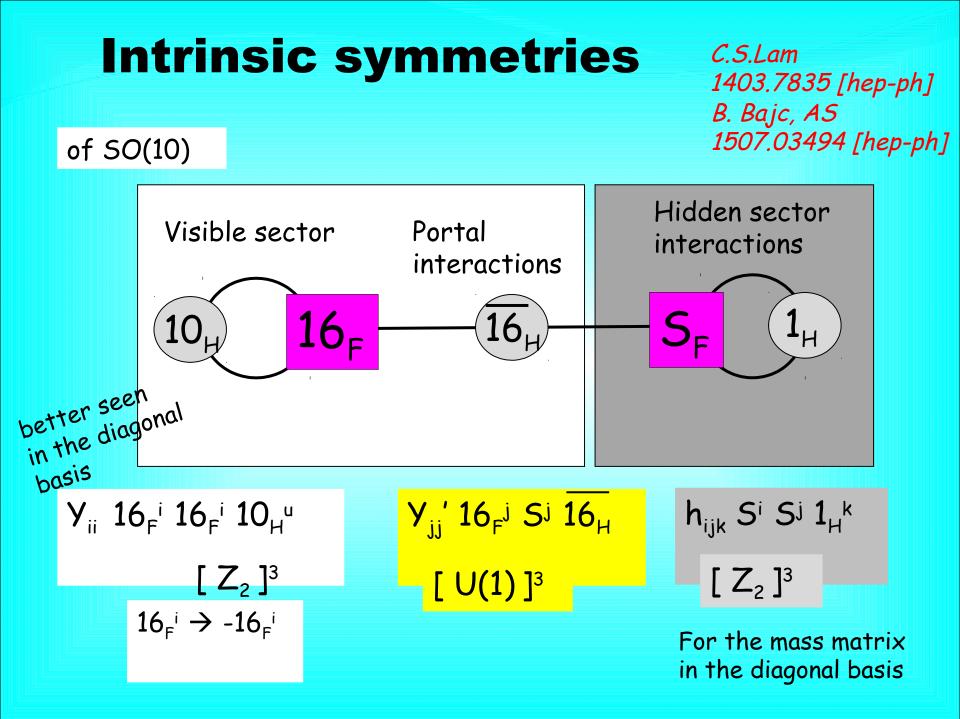
A.Y.5 M.A. Schmidt A.Y.S

Setup

S - from the Hidden sector?

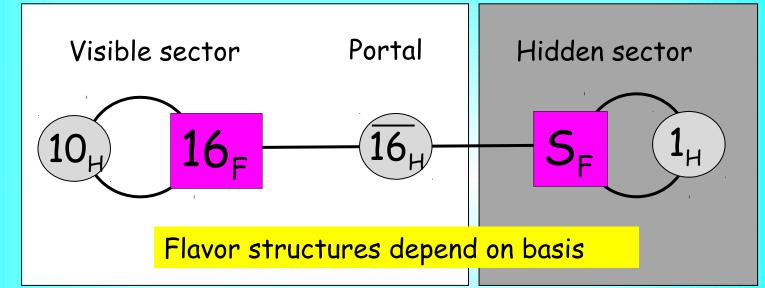






munication to the Hidden work of the field of the state o

Information about flavor structures in the hidden sector should be communicated (transferred) to the observable sector



Minimal way – to communicate info about bases we use in the Hidden (S) and visible sectors (16)

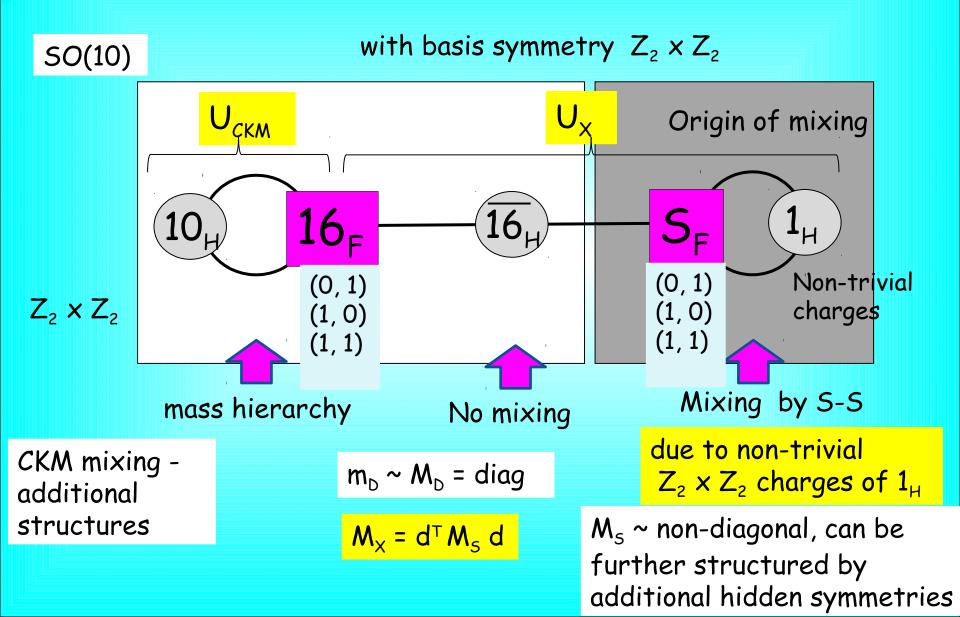
Introduce the same basis symmetry G_{basis} for F and S and prescribe certain charges for them

 $G_{\text{basis}} = G_{\text{intrinsic}} = Z_2 \times Z_2$

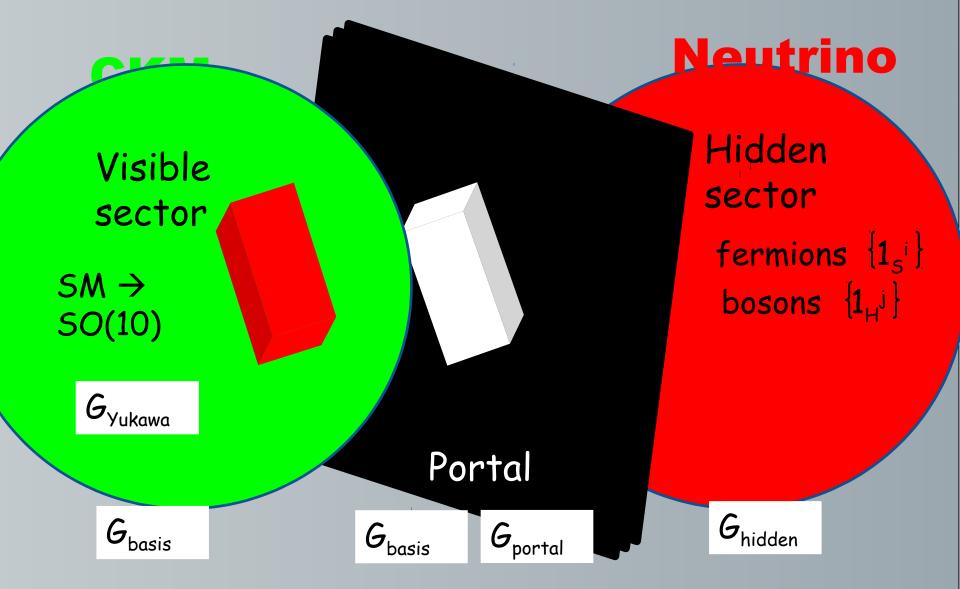
(subgroup)

Scheme

Patrick Ludl, A.S arXiv:1507.03494 [hep-ph]

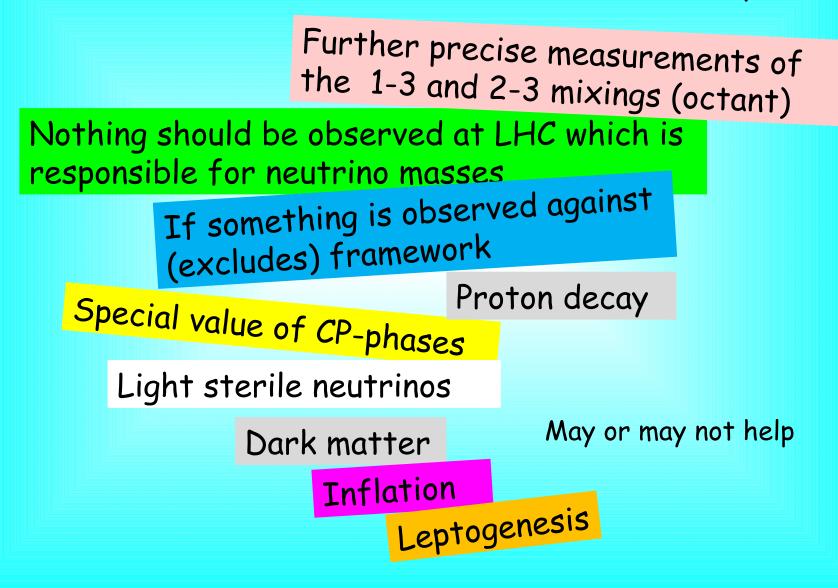


Symmetries and set-up Patrick Ludl A.5



How to test?

Hopeless?



Framework

In general

Take

1) whole V_{CKM} with small elements V_{td} , V_{cb} , etc. \rightarrow this will give also corrections to 2-3 mixing 2) non-maximal rotation $U_{23}(\theta_{23}^{*})$:

 $U_{x} = \Gamma(\alpha) U_{23}(\theta_{23}^{*}) U_{12}(\theta_{12}^{*}) \qquad \Gamma(\alpha) = \text{diag}(1, 1, e^{i\alpha})$

$$\sin^{2}\theta_{13} = \sin^{2}\theta_{23} \sin^{2}\theta_{c} \left\{ -2\cot^{2}\theta_{23} \cos(\alpha - \phi_{td}) |V_{td}| / |V_{cd}| \right\}$$

where $\phi_{td} = \operatorname{Arg} V_{td}$

In Wolfenstein parametrization

$$\sin^2 \theta_{13} = \sin^2 \theta_{23}^x \lambda^2 \left\{ - 2A\lambda^2 [(1 - \rho)^2 + \eta^2]^{1/2} \cot^2 \theta_{23}^x \cos(\alpha - \phi_{td}) \right\}$$

tan² θ₂₃ = tan²θ[×]₂₃ (1 -
$$\lambda^2$$
) $\frac{1}{2}$ - 2A λ^2 sin⁻¹ 2θ[×]₂₃ cos α

$$\tan^2 \theta_{23} = \tan^2 \theta_{23}^{\times} \kappa (\alpha) \qquad \kappa = (1 - \lambda^2) \left[1 - 4A\lambda^2 \cos \alpha \right]$$

Excluding θ_{23}^{x} sin² θ_{13} = f(θ_{23} , α)

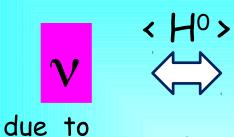
Hidde	n sector	singlet of the SM symmetry group, also invisible
Populated by new	fermions – sterile neutrin scalar bosons – axions, m Gauge bosons – dark phot	ajorons, flavons
Fill in The Dark universe	Responsible for Anomalies	explains Neutrino masses
DM DE BAU	LSND/MiniBo Reactor Gallum 2 5 koV/line2	
	3.5 keV line?	

It looks like, The Problems are in the Visible sector Solutions - in the Hidden one

Neutrino portal

mix

Neutrinos are special providing a portal to the Hidden sector





Mixing due to interaction with Higgs field

Fermions from Hidden sector

Through this mixing

neutrality

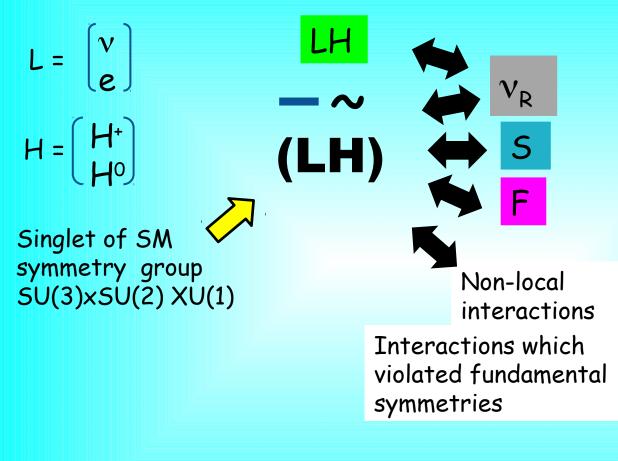
Neutrinos can get mass

Observed pattern of lepton mixing, difference from The quark mixing can be produced

Hidden sector can be probed provide window

Particles fron the Hidden sector can decay/ transform to particles of the visible sector

Neutrino Portal

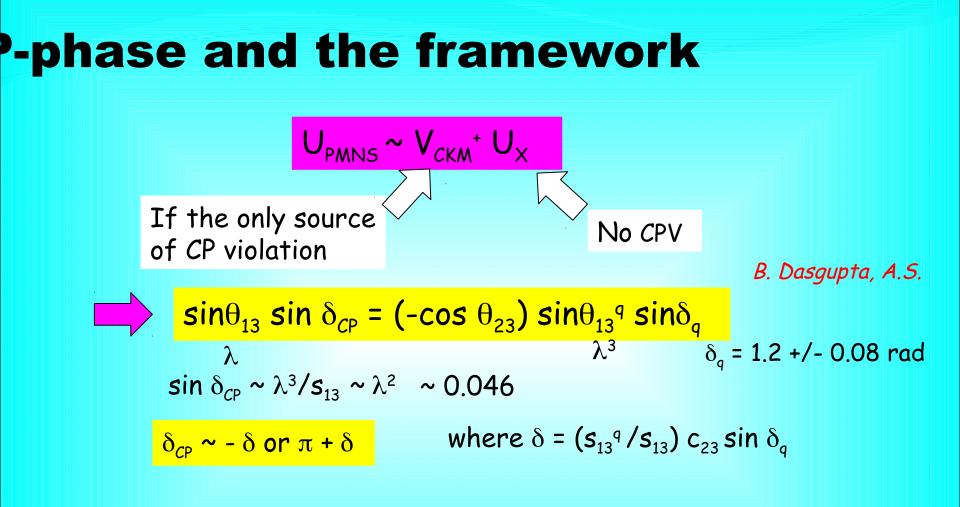


$$\frac{1}{\Lambda^{n(F)-3/2}} LHF$$

F is composite fermionic operator

singlet of symmetry group of hidden sector

In GUT $(\overline{5}_{F} \ 5_{H})$ in SU(5) $(16_{F} \ \overline{16}_{H})$ in SO(10)



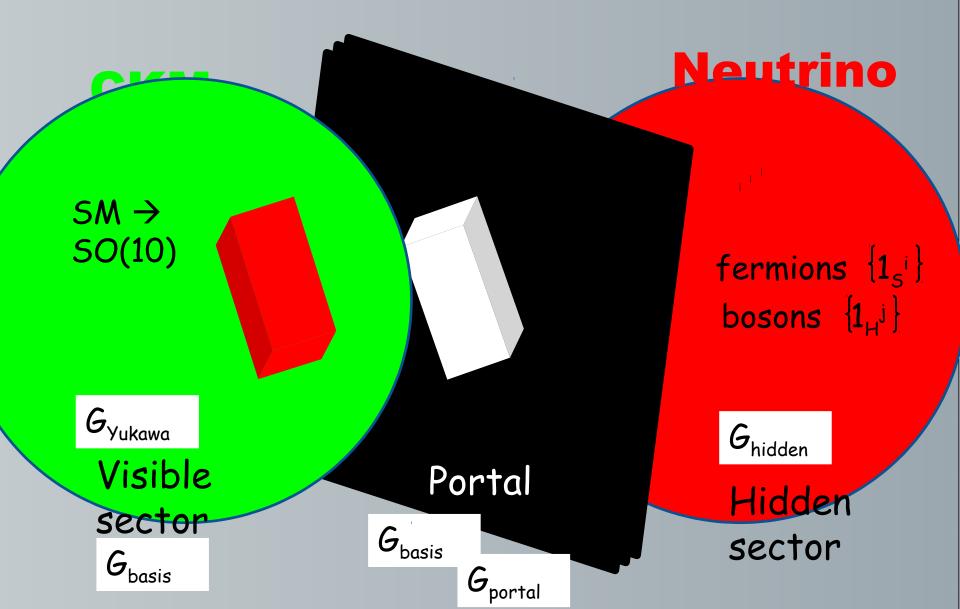
Implications

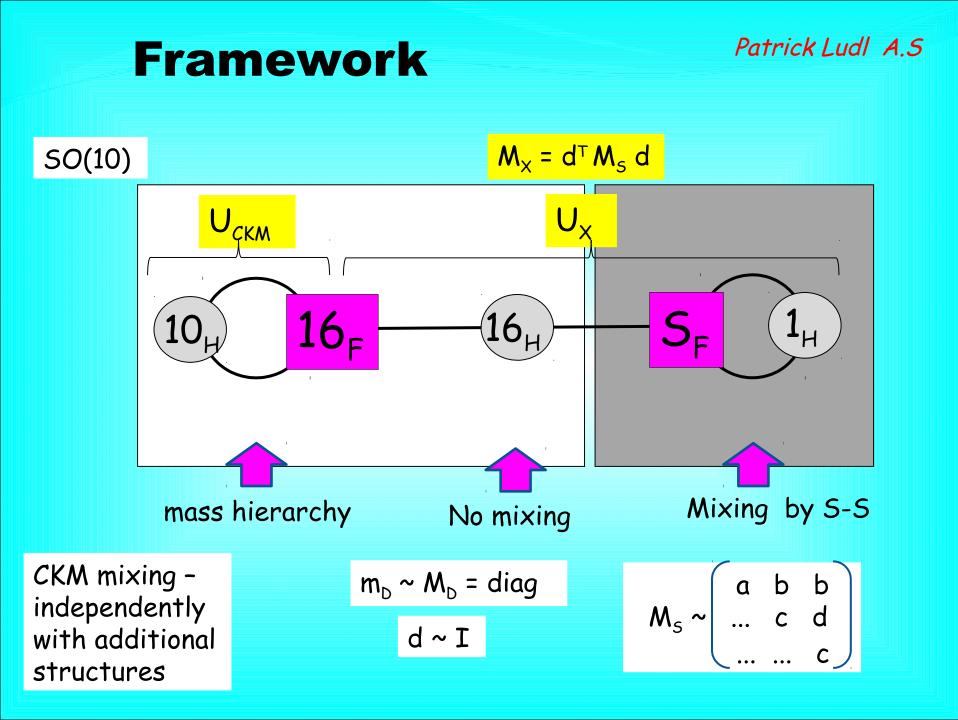
If the phase δ_{CP} deviates substantially from 0 or π , new sources of CPV beyond CKM

New sources may have specific symmetries which lead to particular values of δ_{CP} e.g. $-\pi$ /2

Set-up

Patrick Ludl A.S





Yukawa interactions

 $\mathbf{m}_{\mathrm{D}}^{\mathrm{I}} = \mathbf{m}_{\mathrm{D}}$

 \mathbf{m}_{D}

MD

 M_{S}

Visible sector

F-5 Portal interactions Y_{ij} 16_F 1_S 16_H

5 Hidden sector interactions ${\sf Y}_{ij} \ {\sf 16}_{\sf F}{}^i \ {\sf 16}_{\sf F}{}^j \ {\sf 10}_{\sf H}{}^u$ ${\sf Y}_{ij} \ {\sf 16}_{\sf F}{}^i \ {\sf 1}_{S}{}^j \ {\sf 16}_{\sf H}$

h_{ijk} 1_sⁱ 1_s^j 1_H^k

Υ~ y (<φ> /Λ)ⁿ

In general

For third generation: n = 0, i.e. the mass is generated at the renormalizable level

No mixing is generated by F-F at this level Allow to separate CKM and Neutrino new physics

Basis-fixing symmetry

Although mixing is not generated, masses are generated by the F-F term

Information about states with definite masses should be communicated to the Hidden sector

- For this symmetry which
 - distinguishes three 16-plets, and
 - makes the F-F interactions diagonal

The smallest group is

with charge assignment

matrix of charges of F-F couplings $G_{\text{basis}} = Z_2 \times Z_2$ (1, 0), (0,1), (0,0) powers of $e^{i\pi}$ (0, 0), (1,1), (1,0)
(1, 1), (0,0), (0,1)
(1, 0), (0.1), (0.0)

Both distinguishes states and makes couplings diagonal

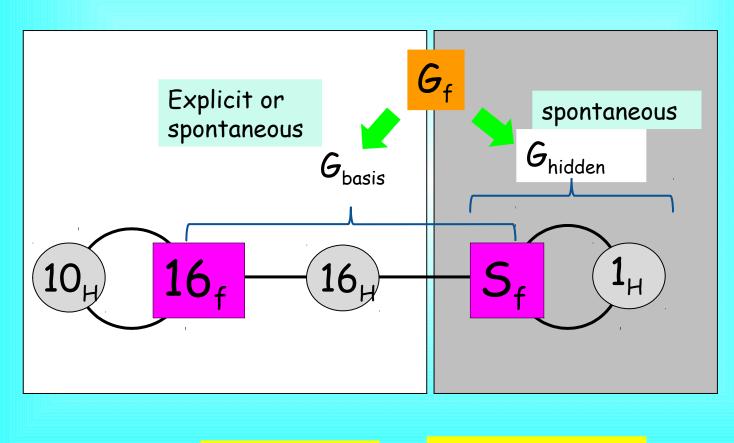
assignment

(1, 0), (0,1), (1,1) with 3 nontrivial charges fixes number of generations



Non-Abelian hidden symmetry



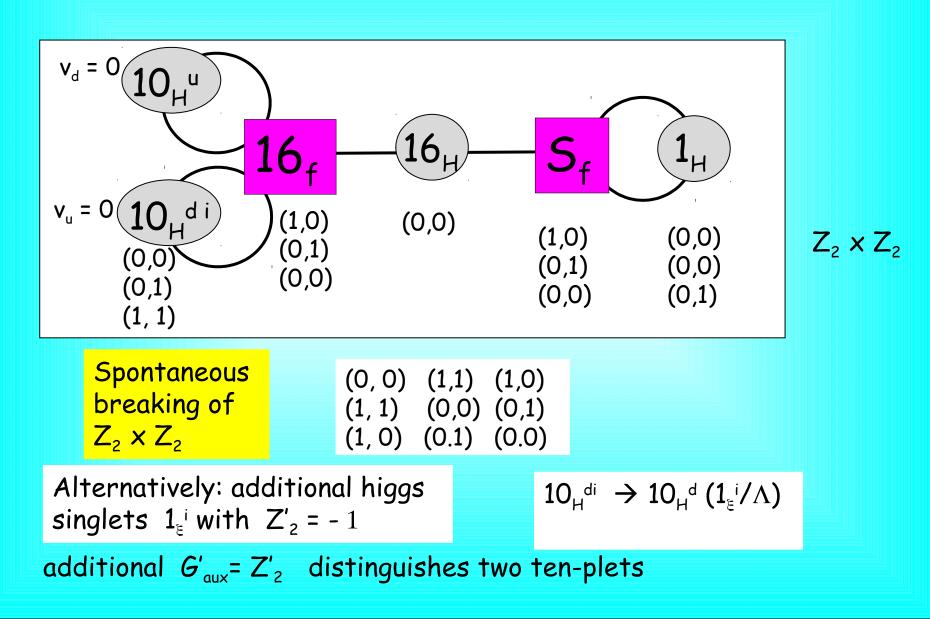


 $G_{\rm f}$ = $A_4 \times Z_2$

 $G_{\text{basis}} = Z_2 \times Z_2$

CKM mixing

Patrick Ludl A.S



CKM physics

- hierarchy of Yukawas Y_i, Y'_i
- CKM mixing
 - difference of masses of upper and down quarks
 - difference of masses of down quarks and charged leptons

Related to breaking of SO(10)

Framework allows to disentangle the CKM physics and neutrino physics

Hierarchy of masses

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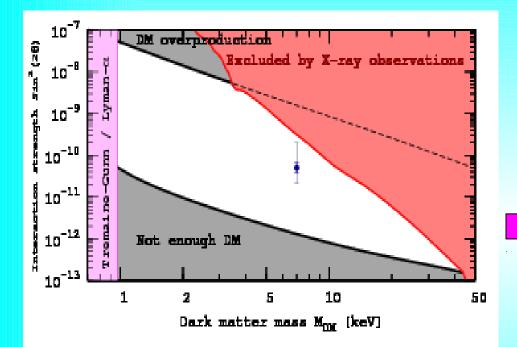
- due to hierarchy of couplings: $Y_1 \leftrightarrow Y_2 \leftrightarrow Y_3$, $Y'_1 \leftrightarrow Y'_2 \leftrightarrow Y'_3$
 - → in turn, due to operators of different order:

3rd generation renormalizable coupling with , $Y_3 = 1$, $Y'_3 = 1$

2nd generation:

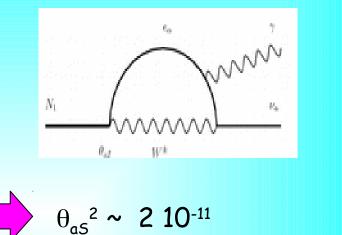
1st generation:

WDM from Hidden sector



The blue point: the best-fit value from M31 (Andromeda galaxy). Thick error bars are $\pm 1\sigma$ limits on the flux. Thin error bars correspond to the uncertainty in the DM distribution in the center of M31.

A Boyarsky et al, 1402.4119



 $\delta m \sim \theta_{aS}^2 m_s \sim (1 - 2) \, 10^{-7} \, eV$

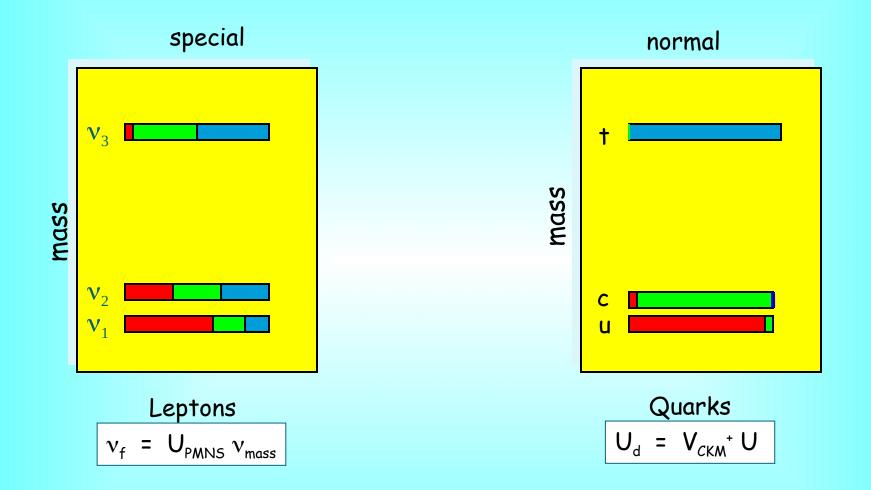
below any relevant scale in the neutrino mass matrix ~ 10⁻³ eV

 \rightarrow does not participate in neutrino mass generation

 \rightarrow is not RH, but some singlet from HS beside 3 v_{R}

 \rightarrow high mass scales are involved

Leptons versus quarks



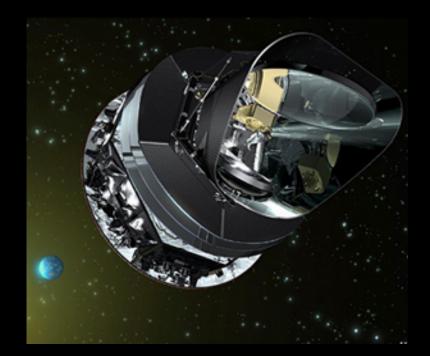
U = (u, c, t)

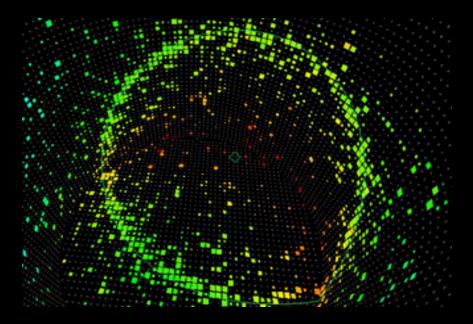
combination of down-quarks produced with a given up quark

Three neutrino

paradigm

Cosmic neutrinos and Supernova





High scale line: the problem

Simplest seesaw implies new physical scale

$$M_{R} \sim m_{D}^{2}/m_{v} \sim 10^{14} \text{ GeV}$$
 << M_{P}

(Another indication: unification of gauge couplings)



Н V_L *F. Vissani hep-phl9709409* <u>H</u> *J Elia 1112.3*

J Elias-Miro et al, 1112.3022 [hep-ph]

$$\delta m_{H^{2}} \sim \frac{Y^{2}}{(2 \pi)^{2}} M_{R^{2}} \log (q / M_{R})$$

~
$$\frac{M_{R}^{3} m_{v}}{(2 \pi v)^{2}} \log (q / M_{R})$$

New physics below Planck scale M_R < 10⁷ GeV

Small Yukawas, Leptogenesis ? M. Fabbrichesi Cancellation?

"Partial" SUSY?

1-3 mixing: Normal?

``Naturalness" : absence of fine tuning of mass matrix Connecting solar and atmospheric neutrino sectors

$$\sin^2\theta_{13} = O(1) \frac{\Delta m_{21}^2}{\Delta m_{32}^2}$$

E. K. Akhmedov, G.C. Branco, M.N. Rebelo Phys.Rev.Lett. 84 (2000) 3535, [hep-ph/9912205]

Very small 1-3 mixing would be something special (symmetry)

Yet another "normal" relation: almost the same relation in the quark and lepton sectors $sin^2\theta_{13} = C sin^2\theta_{12}sin^2\theta_{23}$ K. Patel, A. Y. S. $C_q = 0.380 + - 0.020$ $C_1 = 0.407 + - 0.033$ Similar structure of the mass matrices Abelian symmetry

Neutrinos - Dark matter

Is the (hot) component of the DM

Mechanism of generation of small neutrino masses is related to DM

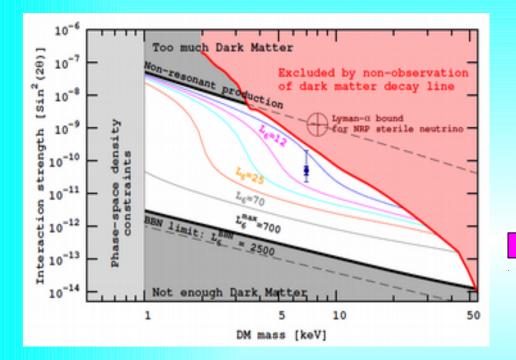


Neutrino portal connects DM and neutrinos

DM particles participate (appear in loops) in generation of neutrino mass

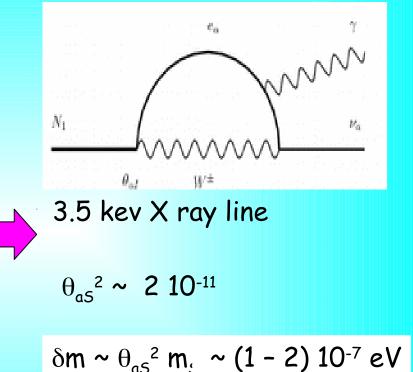
The same symmetry is responsible for smallness of neutrino mass and stability of the DM

WDM sterile neutrino ?



The blue point: the best-fit value from M31 (Andromeda galaxy). Thick error bars are $\pm 1\sigma$ limits on the flux. Thin error bars correspond to the uncertainty in the DM distribution in the center of M31.

A Boyarsky et al, 1402.4119

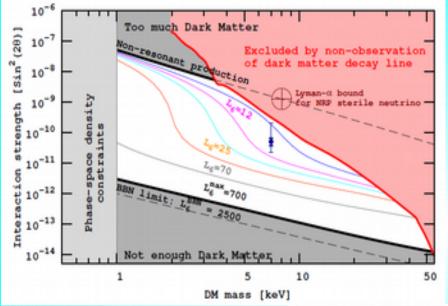


 \rightarrow does not participate in

neutrino mass generation

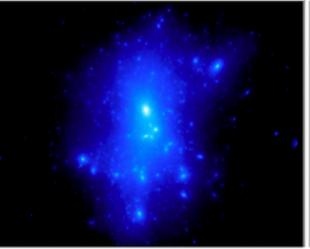
 \rightarrow is not RH, but some singlet from HS beside 3 v_{R}

WDM sterile neutrino ? A Boyarsky et al, 1402.4119

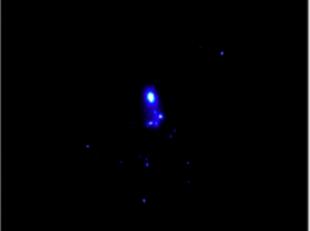


3.5 kev X ray line $\theta_{as}^2 \sim 2 \ 10^{-11}$ m_s ~ 7 keV Contribution to active neutrino masses $\delta m \sim \theta_{as}^2 m_s \sim (1 - 2) \ 10^{-7} eV$

Too small v_s is not RH neutrino Draco dwarf Galaxy: observations analysis



DM decay signal from a galaxy

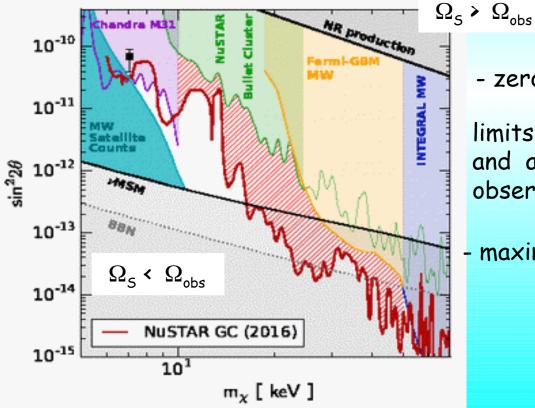


DM annihilation signal from a galaxy

Sterile Neutrinos as Dark matter

(Almost) Closing the Sterile Neutrino Dark Matter Window with NuSTAR

K. Perez, et al. arXiv:1609.00667 [astro-ph.HE]



Nuclear Spectroscopic Telescope Array, Galactic Center

- zero lepton asymmetry

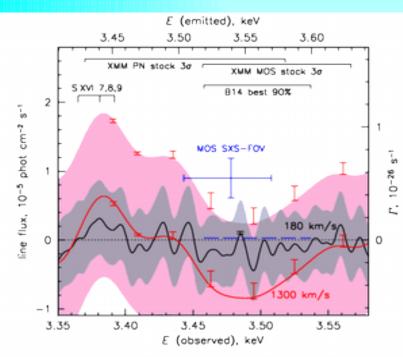
limits from structure formation and astrophysical X-ray observations the colored, regions.

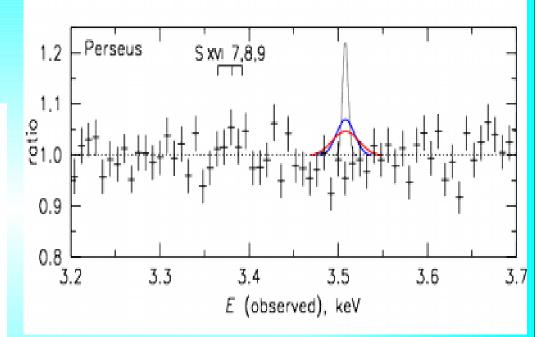
maximal lepton asymmetry

Hitomi constraints

on the 3.5 keV line in the Perseus galaxy cluster

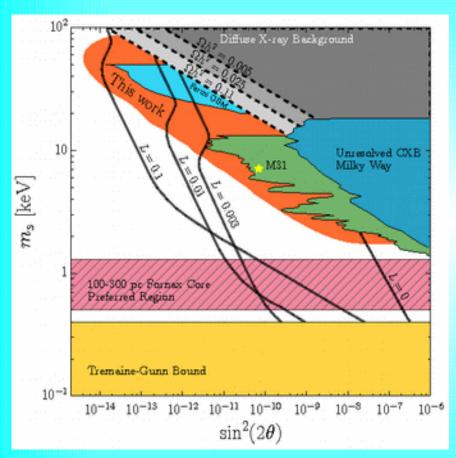
Hitomi Collaboration (F. Aharonian, et al.) arXiv:1607.07420 [astro-ph.HE]





Ratio of data to best-fit model . A line at 3.57 keV (rest-frame) with a flux derived in the SXS FOV is shown with curves of different colors, which denote different l.o.s. velocity dispersions (gray: 180 km/s, blue: 800 km/s, red: 1300 km/s. Position of S feature is marked.

Bounds from Supernova



Supernova bounds on the sterile neutrino parameters from cooling effect $E_s < E^{tot}$

C. A. Argüelles, et al. arXiv:1605.00654 [hep-ph]

Similar bound from SN1987A from the condition that the flux of antinu_e is not suppressed stronger than by factor 10, otherwise no signal could be observe

S.P. Mikheev, A.Yu. S. JETP Lett. 46 (1987) 10-14

dification of the production mechani

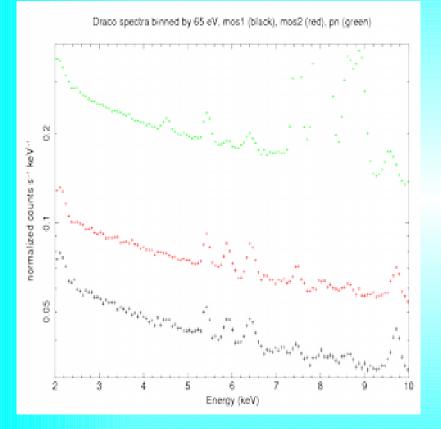
decreasing mixing

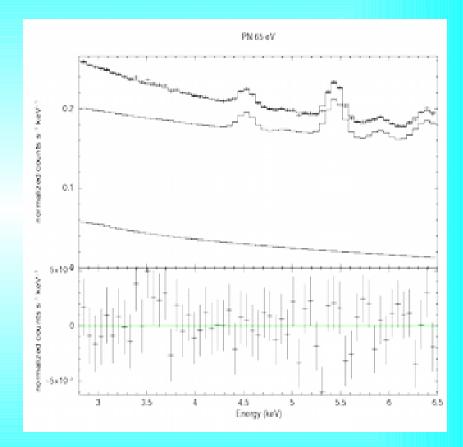
A Berlin, D. Hooper, 1610.03849 "axion assisted ...

Large mixing in
earlier epoch during
production of SSmall mixing
later, to suppress
X ray production"axion assisted ... $g_a \frac{a}{f_a} S LH + \frac{1}{2} m_s SS$ $\sum in^2 2\theta \sim \frac{g_a^2 V^2 \rho_a}{m_s}$ a - axion
 $f_a \sim 10^9 GeV - scale of
PQ symmetry breaking<math>g_a \sim 10^{-9}$

3.5 kev y-line

O. Ruchayskiy et al, 1512.07217 [astro-ph.HE]



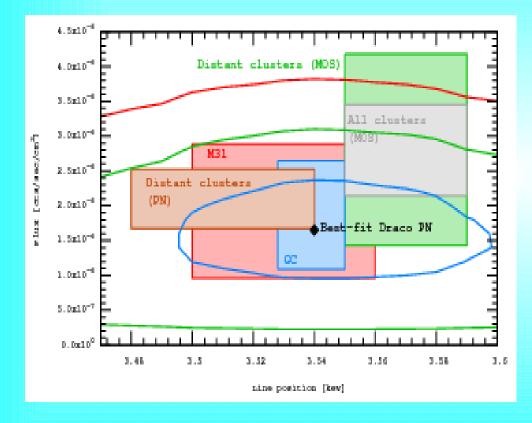


Spectra of Draco dwarf spheroidal seen by MOS1 (black), MOS2 (red) and PN (green) cameras

PN spectrum with unmodeled feature at 3.54 keV.

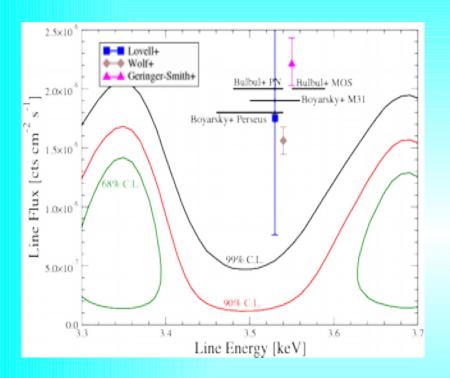
3.5 kev y-line

O. Ruchayskiy et al, 1512.07217 [astro-ph.HE]

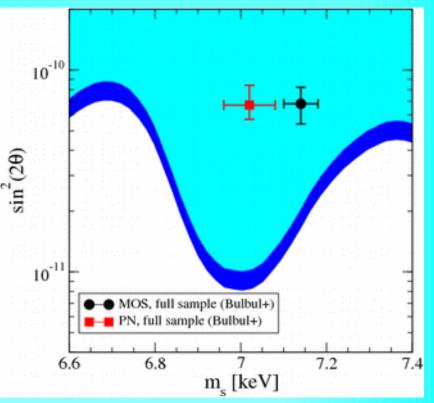


3.5 kev y-line

T. Jeltema and S Profumo, 1512.01239 [astro-ph.HE]



Limits on the flux of a line from MOS observations of the Draco dSph, and predictions for the flux of a 3.5 keV line assuming a dark matter decay origin for the line detected at that energy from stacked clusters of galaxies and from the Milky Way center.



Constraints on mass of sterile neutrino and its mixing angle with active neutrinos from Draco MOS observations. The most Milky Way dark matter density profile from Malyshev et al. (2014)

High scale seesaw, unification

$$m_v = -m_D^T \frac{1}{M_R} m_D$$

q - 1 similarity: $m_D \sim m_q \sim m_l$ Lepton number violation

$$M_{R} \sim \begin{cases} M_{GUT} \sim 10^{16} \text{ GeV} \\ 10^{8} - 10^{14} \text{ GeV} \\ 10^{16} - 10^{17} \text{ GeV} \end{cases}$$

for the heaviest in the presence of mixing $\frac{M_{GUT}^2}{M_{Pl}}$ double seesaw many heavy singlets (RH neutrinos)string theory N ~ 10²

In favor Gauge coupling unification

Leptogenesis

Seesaw sector is responsible for inflation (scalar which breaks B-L and gives masses of RH neutrinos), dark matter

eV - sub eV scale physics

Very light sector which may include

- new scalar bosons, majorons, axions,
- new fermions (sterile neutrinos, baryonic nu),
- new gauge bosons (e.g. Dark photons)

M. Pospelov

Maybe related to Dark energy, MAVAN

Generate finite neutrino masses, usual Dirac masses can be suppressed by seesaw with $M_R = M_{Pl}$

eV scale Seesaw with RH neutrinos A. De Gouvea for sterile anomalies LSND/ MiniBooNE

Tests:	5 th force searches experiments	
	Modification of dynamics of neutrino oscillations	
	Checks of standard oscillation formulas, searches for deviations	

utrino mass scale - fundamental sca

Neutrino mass scale is not a spurious scale but the fundamental new physics

This new physics contribution becomes dominant because usual contributions are strongly suppressed

by seesaw with $M_R = M_{PI}$

By multi-singlet mechanisms

It could be whole invisible low scale sector with

- new scalar bosons, majorons, axions,
- new fermions (sterile neutrinos, baryonic nu),
- new gauge bosons (e.g. Dark photons)

Data confirm prediction

 $\theta_{13} \sim \sqrt{\frac{1}{2}} \theta_{c}$ $\sin^2\theta_{13} \sim \frac{1}{2}\sin^2\theta_c$



Phenomenological level

C. Giunti, M. Tanimoto



From QLC (Quark-Lepton Complementarity)

H. Minakata, A Y S



From TBM-Cabibbo scheme

S. F. King et al

Now accuracy of measurements permits detailed comparison

eesaw and PMNS-CKM relation

Right handed components of neutrinos N Neutrinos get usual Dirac mass terms m_D = Y<H> N have large Majorana masses M_R >> m_D

 $\mathbf{v} \quad \mathbf{N} \\
 \mathbf{v} \quad \left[\begin{array}{c} \mathbf{v} & \mathbf{N} \\ \mathbf{0} & \mathbf{m}_{\mathrm{D}} \\ \mathbf{m}_{\mathrm{D}}^{\mathrm{T}} & \mathbf{M}_{\mathrm{R}} \end{array} \right]$

diagonalization $m_v = -m_D (M_R)^{-1} m_D^{T}$ P. Minkowski T. Yanagida M. Gell-Mann, P. Ramond, R. Slansky S. L. Glashow R.N. Mohapatra, G. Senjanovic

 $m_{D} = U_{L} (m_{D}^{diag}) U_{R}^{+}$ If $U_{L} = V_{CKM}^{*}$ -- realized in the simplest SO(10) $m_{v} = -U_{L} (m_{D}^{diag}) U_{R}^{+} (M_{R})^{-1} U_{R}^{*} (m_{D}^{diag}) U_{L}^{\top}$ produces U_{x}^{\vee} - we realize the connection so U_{x} should diagonalize that is $M_{x} = -m_{D}^{diag} U_{R}^{+} (M_{R})^{-1} U_{R}^{*} m_{D}^{diag}$ $M_{x} = U_{x} M_{x}^{diag} U_{x}^{\top} = U_{x} m_{v}^{diag} U_{x}^{\top} \sim m_{TBM}$

How to reconcile all this?

Special symmetry in lepton sector no symmetry in quark sector

Smallness of neutrino mass

Complementarity relation