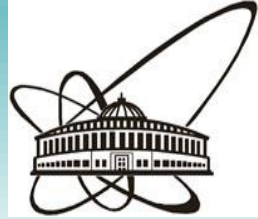




# Electromagnetic Properties of $\nu$

(effects of magnetic moments)



C.Giunti, A.Studenikin,

" $\nu$  electromagnetic interactions: A window to new physics", Rev.Mod.Phys, 2015

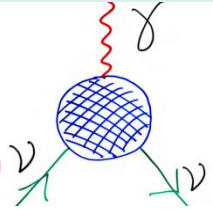
MSU

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Studenikin,

" $\nu$  electromagnetic interactions: A window to new physics - II", arXiv: 1801.18887



## 1 $\nu$ EP theory - $\nu$ vertex function

matrices in  $\nu$  mass eigenstates space

$$\Lambda_\mu^{if}(q) = f_Q^{if}(q^2)\gamma_\mu + f_M^{if}(q^2)i\sigma_{\mu\nu}q^\nu + f_E^{if}(q^2)\sigma_{\mu\nu}q^\nu\gamma_5 + f_A^{if}(q^2)(q^2\gamma_\mu - q_\mu\not{q})\gamma_5,$$

form factors  $f_X^{if}(q^2)$  at  $q^2 = 0$  static EP of  $\nu$

electric charge  
magnetic moment  
electric moment  
anapole moment

Dirac  $\nu$  Majorana

$q_{if}$	$q_{if} = 0$	} CPT + charge conservation
$\mu_{if}$	$\mu_{if}^{(i \neq f)}$	
$\epsilon_{if}$	$\epsilon_{if}^{(i \neq f)}$	
$a_{if}$	$a_{if}$	

Diagram of a neutrino mass matrix with a circled zero element.

Hermiticity and discrete symmetries of EM current put constraints on form factors

$$\langle \nu(p') | J_\mu^{EM} | \nu(p) \rangle = \bar{u}(p') \Lambda_\mu(q) u(p)$$

$$\mu_{jj}^D = \frac{3e_0 G_F m_j}{8\sqrt{2}\pi^2} \approx 3.2 \times 10^{-19} \mu_B \left( \frac{m_j}{1 \text{ eV}} \right)$$

- much greater values are Beyond Minimally Extended SM
- transition moments  $\mu_{i \neq f}, \epsilon_{i \neq f}$  are GIM suppressed reactor  $\nu$  scattering

## 3 $\nu$ EP experimental bounds

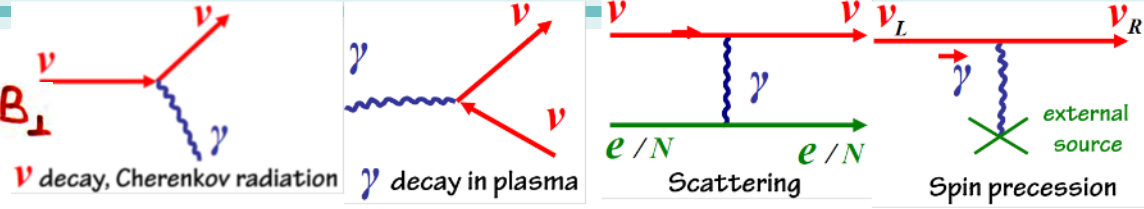
$\mu_\nu^{eff} < 2.9 \times 10^{-11} \mu_B$  GEMMA Coll. 2012  
 $\mu_\nu^{eff} < 2.8 \times 10^{-11} \mu_B$  Borexino Coll. 2017  
 $\sim 0.1 \mu_B$  Astrophysics, Raffelt ea 1988  
 Arcoa Dias ea 2015

$q_\nu < \begin{cases} \sim 10^{-12} \\ \sim 10^{-19} \\ \sim 10^{-21} \end{cases} e_0$  AS '14, Chen ea '14  
 AS '14 (astrophysics) neutrality of matter

# Effects of $\nu$ magnetic moment:

- spin precession and oscillations in  $B_{\perp}$

Cisneros, Okun, Voloshin, Vysotsky, Valle, Raffelt, Schechter, Petkov, Akhmedov, Lim, Marciano, Smirnov, Pulido, Dvornikov, Grigoriev, Lobanov, Lokhov, Kouzakov, Ternov, Studenikin et al



## New effects reported at Neutrino 2018

1 Electromagnetic interactions and oscillations of ultrahigh-energy cosmic  $\nu$  in interstellar space

Kouzakov & AS, poster # 174  
 PRD 96 (2017)  $L_B = \pi / \mu_{\nu} B$   
 $P_{\nu^L \rightarrow \nu^R}(x) = \sin^2 \left( \frac{\pi x}{L_B} \right)$

$$P_{\nu_e^L \rightarrow \nu_{\mu}^L}(x) = [1 - P_{\nu^L \rightarrow \nu^R}(x)] \sin^2 2\theta \sin^2 \left( \frac{\pi x}{L_{\text{vac}}} \right)$$

amplitude of flavour oscillations is modulated by  $\mu_{\nu} B$  frequency

2  $\nu$  flavour, spin and spin-flavour oscillations and consistent account for a constant magnetic field

Popov & AS, poster # 140 arXiv: 1803.05766  
 probability of spin oscillations depends on  $\Delta m^2$

$$P_{\nu_e^L \rightarrow \nu_e^R} = \left\{ \sin(\mu_+ B_{\perp} t) \cos(\mu_- B_{\perp} t) + \cos 2\theta \sin(\mu_- B_{\perp} t) \cos(\mu_+ B_{\perp} t) \right\}^2 - \sin^2 2\theta \sin(\mu_1 B_{\perp} t) \sin(\mu_2 B_{\perp} t) \sin^2 \frac{\Delta m^2}{4p} t$$

3  $\nu$  spin and spin-flavour oscillations engendered by transversal matter current

Pustoshny & AS, poster # 139 arXiv: 1801.08911  
 Studenikin 2004, 2017

• transversal matter currents  $j_{\perp}$  do change  $\nu$  helicity !

4 Spin-light of  $\nu$  in Gamma-Ray Bursts

new mechanism of EM radiation by  $\nu$   
 JCAP 1711 (2017) no. 11, 024

Grigoriev, Lokhov, Studenikin, Ternov, poster # 167